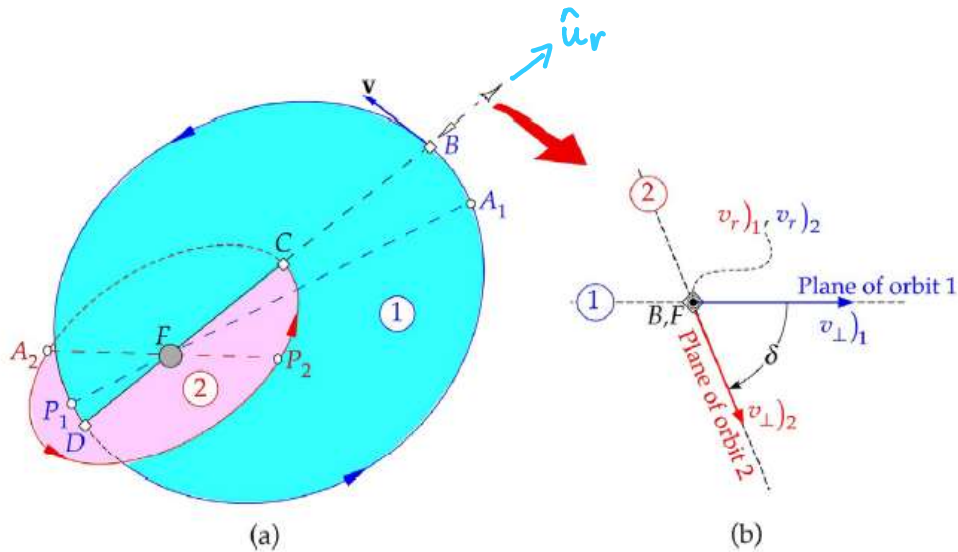
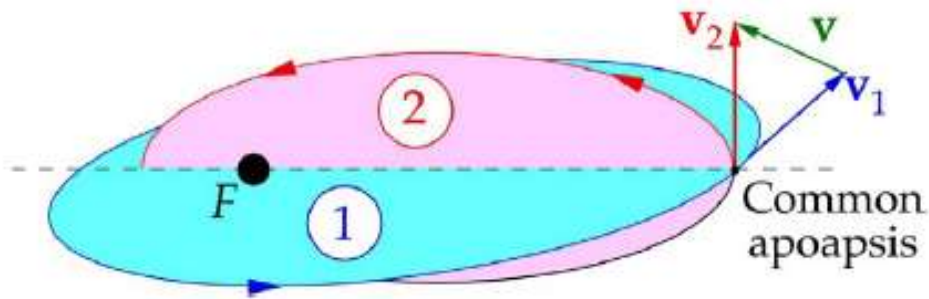


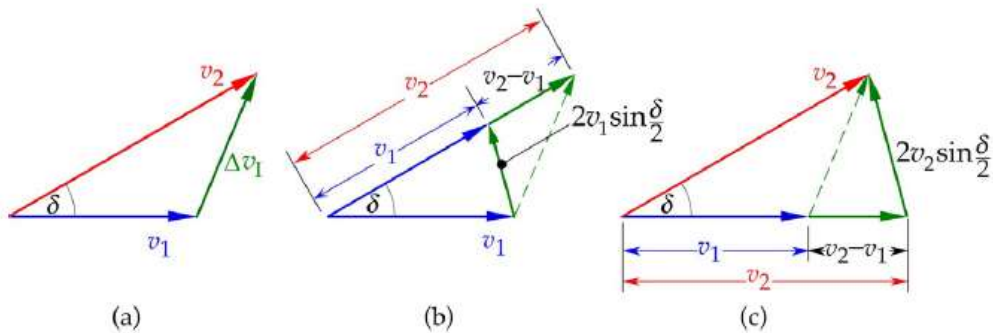
Plane Change Manoeuvres



- $\|V_1\| = v_{r1} \hat{u}_r + v_{\perp 1} \hat{u}_{\perp 1}$
 $\|V_2\| = v_{r2} \hat{u}_r + v_{\perp 2} \hat{u}_{\perp 2}$
- $\Delta V = V_2 - V_1 = (v_{r1} - v_{r2}) \hat{u}_r + v_{\perp 2} \hat{u}_{\perp 2} - v_{\perp 1} \hat{u}_{\perp 1}$
- $\|\Delta V\| = \sqrt{\Delta V \cdot \Delta V} = \sqrt{(v_{r2} - v_{r1})^2 + v_{\perp 1}^2 + v_{\perp 2}^2 - 2v_{\perp 1}v_{\perp 2}\cos\delta}$
- $v_{r1} = \|V_1\| \sin\gamma_1$, $v_{\perp 1} = \|V_1\| \cos\gamma_1$
 $v_{r2} = \|V_2\| \sin\gamma_2$, $v_{\perp 2} = \|V_2\| \cos\gamma_2$
- $\|\Delta V\| = \sqrt{\|V_1\|^2 + \|V_2\|^2 - 2\|V_1\|\|V_2\|[\cos\Delta\gamma - \cos\gamma_2\cos\gamma_1(1 - \cos\delta)]}$
- $\|\Delta V\| = \sqrt{\|V_1\|^2 + \|V_2\|^2 - 2\|V_1\|\|V_2\|\cos\Delta\gamma}$ ($\delta = 0^\circ$)



$$- \|\Delta V\| = \sqrt{v_1^2 + v_2^2 - 2v_1v_2\cos\delta} \quad (v_{r1}=v_{r2}=0, v_{\perp1}=v_1, v_{\perp2}=v_2)$$

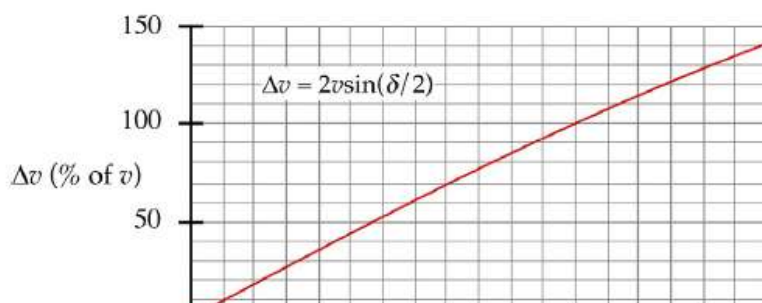


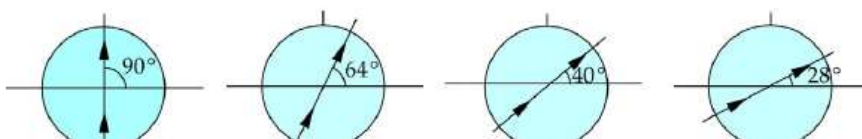
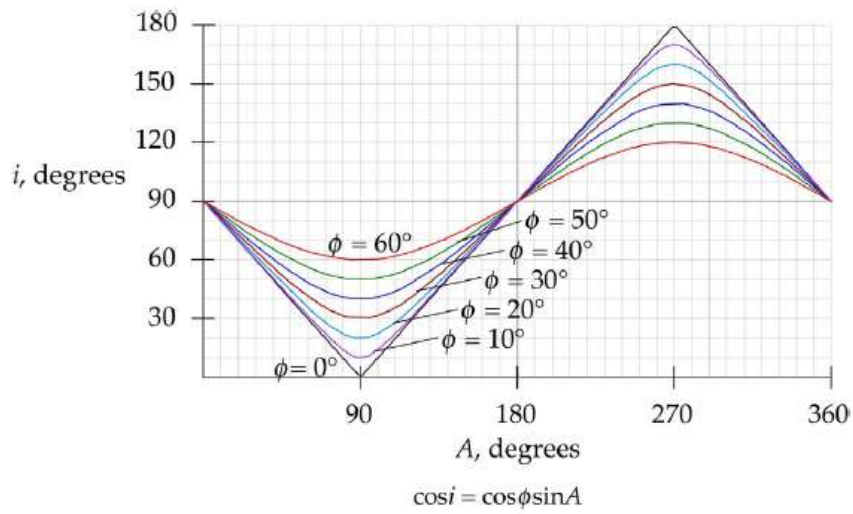
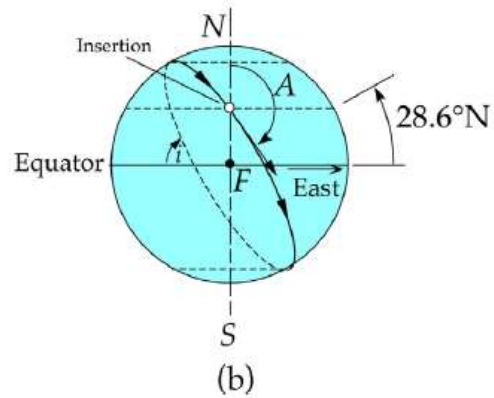
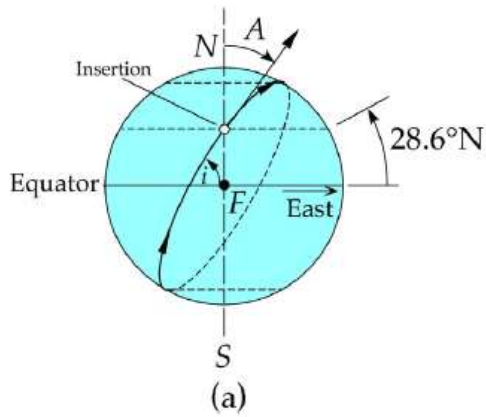
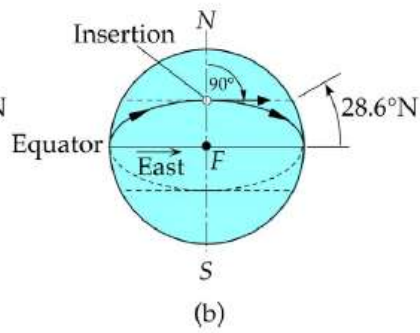
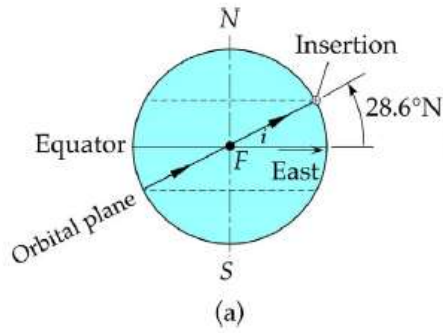
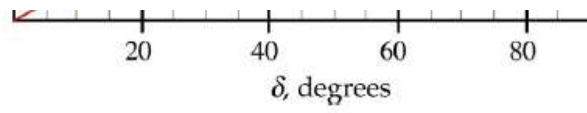
$$- \Delta V_I = \sqrt{(v_2 - v_1)^2 + 4v_1v_2\sin^2\delta/2}$$

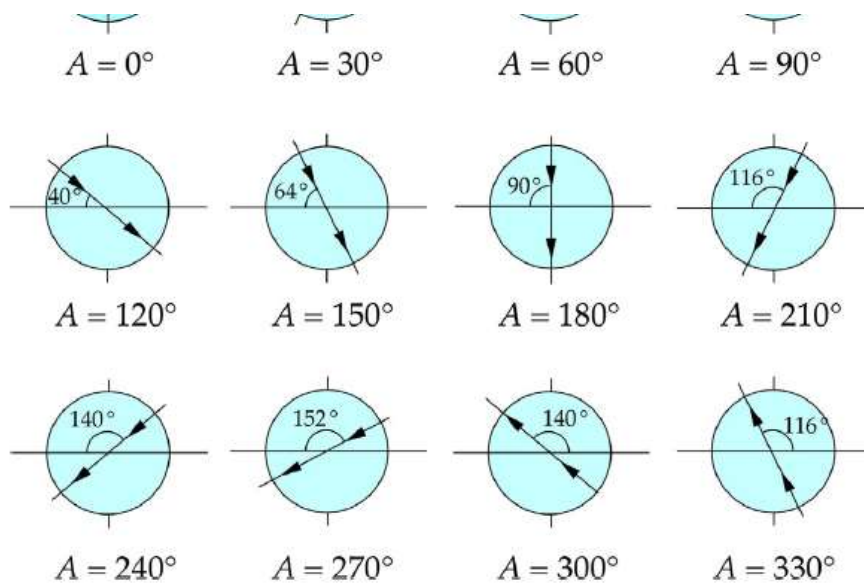
$$- \Delta V_{II} = 2v_1\sin\delta/2 + |v_2 - v_1|$$

$$- \Delta V_{III} = |v_2 - v_1| + 2v_2\sin\delta/2$$

$$- \Delta V_{II} > \Delta V_I, \Delta V_{III} > \Delta V_I$$

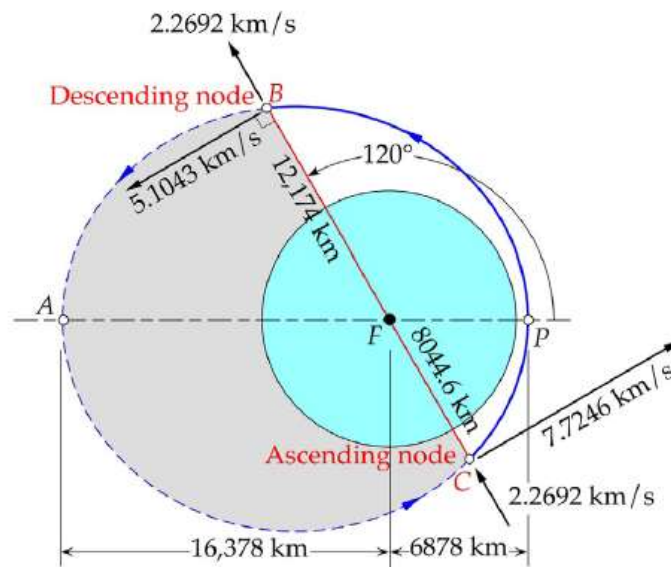






Example

A spacecraft is in a 500 km by 10,000 km altitude geocentric orbit that intersects the equatorial plane at a true anomaly of 120° (see Fig. 6.33). If the orbit's inclination to the equatorial plane is 15° , what is the minimum velocity increment required to make this an equatorial orbit?



Details

$$\|e\| = \frac{r_A - r_P}{r_A + r_P}, \quad \|h\| = \sqrt{2\mu} \sqrt{\frac{r_A r_P}{r_A + r_P}}$$

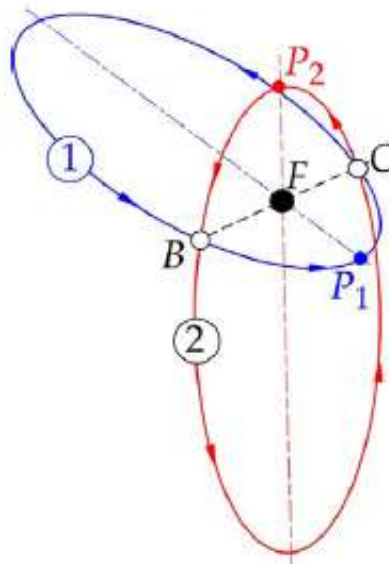
$$r_B = \frac{\|h\|^2}{\mu} \frac{1}{1 + \|e\| \cos \theta_B}, \quad V_{\perp B} = \frac{\|h\|}{r_B}, \quad V_{r_B} = \frac{\mu}{\|h\|} \sin \theta_B$$

$$r_C = \frac{\|h\|^2}{\mu} \frac{1}{1 + \|e\| \cos \theta_C}, \quad V_{\perp} = \frac{\|h\|}{r_C}, \quad V_{r_C} = \frac{\mu}{\|h\|} \sin \theta_C$$

$$\Delta V = 2V_{\perp} \sin \delta/2$$

Example

Orbit 1 in Fig. 6.34 has angular momentum h and eccentricity e . The direction of motion is shown. Calculate the Δv required to rotate the orbit 90° about its latus rectum BC without changing h and e . The required direction of motion in orbit 2 is shown.



Details

$$\theta_B)_1 = -90^\circ, \quad \theta_B)_2 = 90^\circ$$

$$r_B = \frac{h^2}{\mu}$$

$$V_{\perp B})_1 = \frac{h}{r_B} = \frac{\mu}{h}$$

$$V_{r_B})_1 = \frac{\mu}{h} e \sin \theta_B)_1 = -\frac{\mu e}{h}$$

$$V_{r_B})_2 = \frac{\mu}{h} e \sin \theta_B)_2 = \frac{\mu e}{h}$$

$$\Delta V_B = \dots$$