On Calibrating a Projector-Camera Rig to Enable Spatial Augmented Reality Use Cases

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# Motivation

Spatial augmented reality (SAR) is a branch of augmented reality whereby the augmentation of the scene is carried out using a projector. While pointing a projector at a scene and projecting an image is on its own not a challenge, doing so in a manner such that the projected content **appear to be** **naturally part of the scene** is more tricky. The intuition for the general problem we face can be drawn from using a projector to project an image to a flat wall: unless the projector faces the wall frontally, the bounds of a projected rectangular image will not appear rectangular, but more generally as a trapezoid (i.e., the image will appear distorted). The cause for this effect is that the geometry of the scene as it relates to the geometry of the projector plays a role in *how the pixels of the projector’s image plane map to points in the scene*.

The central **objective** we set for ourselves is to **enable correcting for distortion** caused by projecting to arbitrary scene geometry in an **automatic** fashion. This is to be done in a manner that lends itself to practical application in a real-world setting where the **scene** to be augmented—e.g., an **object** being gradually assembled and its **placement** relative to the projector—**does not remain static**. Moreover, the object in question may but may not in its entirety fall within the projector’s **field of view (FOV)**. Finally, the solution is ideally to rely on the existing **hardware setup**—i.e., projector with steerable mirror system and stereo camera—mounted in the Pilotfabrik[[1]](#footnote-1) of TU Vienna.

# Approach

The problem of correcting for distortions as outlined in Section 1 requires knowledge of the manner in which the respective rays through the pixels of the projector’s image plane fan out into the scene, and the geometry of the scene itself within at least the projector’s field of view (determined in part by the steerable mirror),[[2]](#footnote-2) as illustrated in Figure 1. This is because the scene point ‘illuminated’ by a pixel in the projector’s image plane is given by intersecting its corresponding ray with the geometry of the scene surface. To model this interaction calls for two components in particular: (i) one-time **projector calibration** and (ii) **recovery of** **scene geometry** whenever the scene undergoes change relative to the field of view of the projector.

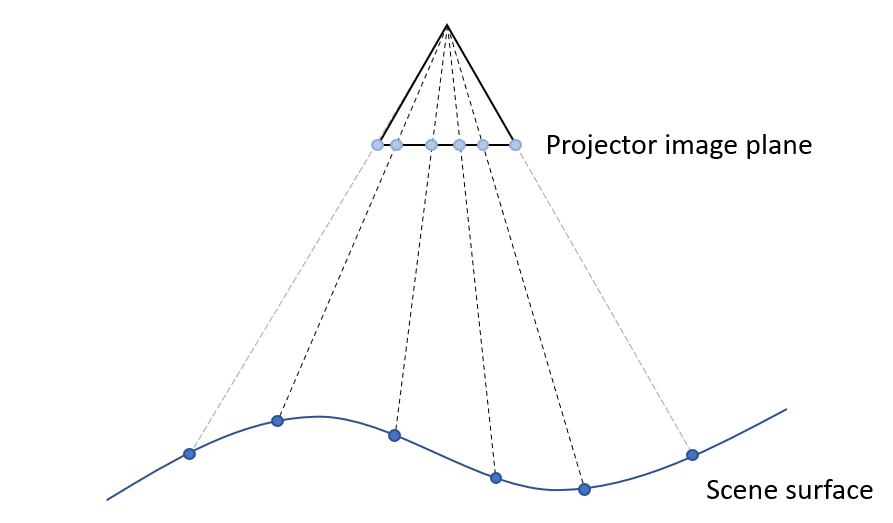


Figure 1: Projector-scene geometry. The way rays ‘fan out’ through the pixels of the projector’s image plane is given by calibrating the projector (note that the projector’s field of view is the portion of the scene in between the gray rays). The scene point ‘illuminated’ by a pixel in the projector’s image plane is given by intersecting its corresponding ray with the geometry of the scene surface. Projector calibration and recovery of scene geometry are accordingly prerequisites of the image warping that it is our ambition to carry out.

The task of finally *warping* imagery is left outside the scope of this document, as a subject to be addressed only after the topics of projector calibration and recovery of scene geometry have been nailed, since these two topics are prerequisites of such image warping. The remainder of this section is organized in accordance with these two components: one-time projector calibration, and recovery of scene geometry.

## One-time Projector Calibration

The ‘output’ of projector calibration[[3]](#footnote-3) is a mathematical model that determines, for each pixel **x**proj in the projector’s image plane *π*proj, the ray from the projector’s center of projection **C**proj through **x**proj, along which the pixel is projected *to the scene*. Inversely, the model also determines the projection **x**’proj of any point **X**’proj in the scene *to the projector’s image plane* *π*proj, obtained by intersecting *π*proj with the ray from **C**proj through **X**’proj. Such a model is said to function in accordance with the principle of **central projection**.

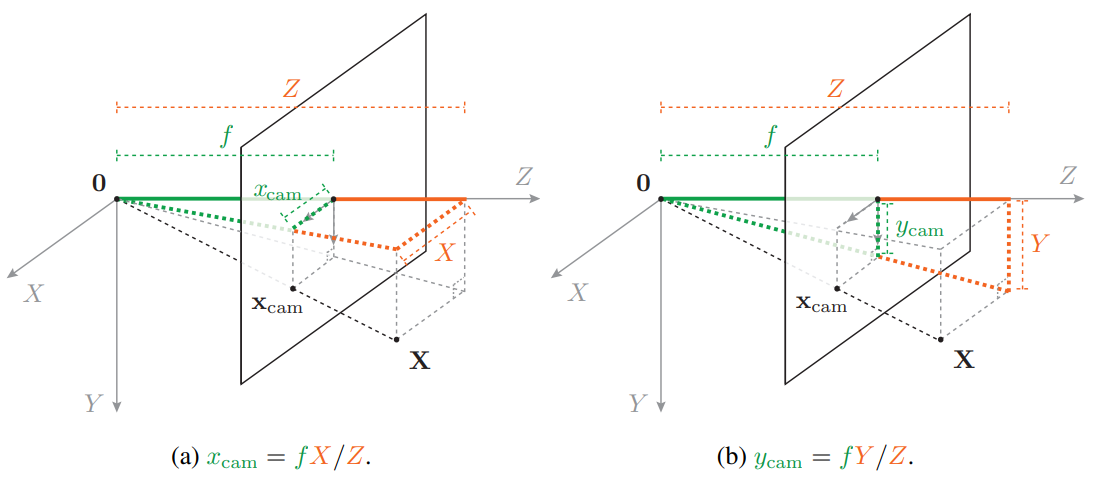


Figure 2: Pinhole camera model. The image plane π is perpendicular to the Z axis of the camera’s coordinate frame; its location is given by the focal length f, the distance of π along the Z axis. The camera’s center of projection **C** is located at the origin **0** of the camera coordinate frame. The projection **x** = (xcam, ycam) of a point **X** = (X, Y, Z) to π, by central projection, follows from the principle of similar triangles. Note that the origin of the coordinate system of the image plane—called the principal point—is situated at the intersection of the Z axis with π. Note that the Z component of a point expressed in the camera’s coordinate frame is referred to as the point’s depth relative to the camera.

Notably, central projection is the same principle that underlies the modelling of conventional cameras. A camera modeled in accordance with the principle of central projection is called a **pinhole camera**. A pinhole camera—as illustrated in Figure 2—can be expressed in terms of a 3 × 3 ‘camera calibration’ matrix **K**,[[4]](#footnote-4) derived from the camera’s focal length and principal point. In practice, a pinhole camera model works well in modelling how the respective rays through the pixels of the image plane fan out into the scene, on the condition that images to which the model is applied have been corrected for **lens** **distortions** caused by the camera’s lens system (cf. Figure 3 for an example). The ‘output’ of camera calibration is accordingly typically not only the matrix **K**, but the associated **lens distortions model coefficients** as well. The lens distortion model coefficients enable applying a lens distortion model to **undistort** any image acquired by the camera.

In practice, calibrating a *camera* relies on (i) establishing **2D-3D correspondences** between pixels in the camera’s image plane and corresponding points in the scene, and on (ii) using those correspondences as input to an **optimization procedure** that outputs the calibration matrix **K** and the associated distortion model coefficients. Calibration of a *projector* can be carried out in precisely the same manner insofar as step (ii) is concerned; *the major difference in projector calibration relative to camera calibration concerns the manner in which 2D-3D correspondences are identified*, i.e.,between pixels in the image plane of the projector and the corresponding points in the scene.

What remains of this section first, for context, reviews how 2D-3D correspondences are conventionally obtained between camera and scene. Next, we address how 2D-3D correspondences are conventionally obtained between projector and scene. Finally, we outline the manner in which 2D-3D correspondences are used to compute the calibration matrix **K** and associated lens distortion model coefficients. Note that these steps constitute fundamental tasks in computer vision and can be carried out largely using functionality available in OpenCV.

### Obtaining Camera-Scene 2D-3D Correspondences

The recovery of 2D-3D correspondences in support of camera calibration is carried out, conventionally, by relying on a planar **calibration surface** of known **scale** to automatically identify correspondences between the 3D points on the calibration surface and their 2D correspondences in the image plane. A widely used example of such a calibration object is a **chessboard pattern**. The 3D **corner points** of the chessboard are provided manually in a coordinate system defined in the plane of the chessboard.[[5]](#footnote-5) Next, using the findChessboardCorners() function in OpenCV[[6]](#footnote-6) the corresponding 2D pixels of the corner points in any single image of the chessboard are obtained as a list of lists (one list per input image); an example is given in Figure 3.

In practice, multiple images of the chessboard are acquired, from a set of **widely varying viewpoints**.[[7]](#footnote-7) The rationale behind doing so is to capture the varying effects of lens distortion across the full extent of the camera’s field of view; notably, images of the chessboard acquired only at approximately the middle of the image plane would fail to provide information about radial distortions present at the image plane’s boundaries.

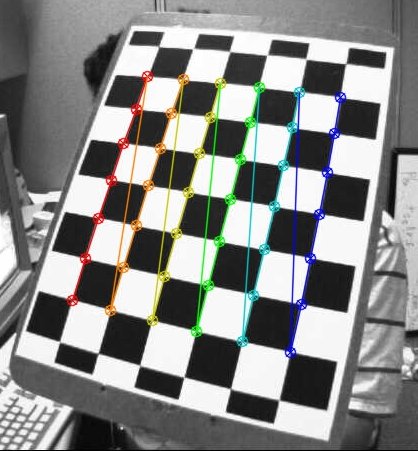


Figure 3: Visualization of chessboard corners recovered automatically using OpenCV using the findChessboardCorners() function, color coded using the drawChessboardCorners() function. Note the distortions caused by the camera’s lens system, clearly visible in the image: the edges of the chessboard should be straight lines, but instead appear to bend outwards.

### Obtaining Projector-Scene 2D-3D Correspondences

Todo: describe how we are now obtaining 3D points via intersection with ground plane of back-projections of detected circles…

### Computing the Calibration

Given a set of 2D-3D correspondences in the form of a list of lists[[8]](#footnote-8) of 2D pixel positions (‘image points’) and a corresponding list of 3D points (‘object points’), the OpenCV[[9]](#footnote-9) function calibrateCamera() can be used to carry out the actual calibration. The function returns the camera calibration matrix **K** and the lens distortion model coefficients.[[10]](#footnote-10) Additionally, the function returns what amounts to a list of 3D rigid body transformations whose length is the same as the number of lists of 2D-3D correspondences provided as input.

Any image subsequently acquired using a camera calibrated in this manner can be undistorted using the OpenCV function undistort(), whereupon the camera can be treated as a pinhole camera modelled using the matrix **K**. Note that in order to use the matrix **K** to model a projector as a pinhole ‘camera’, the *inverse* distortion would have to be applied to the *projecting image*.[[11]](#footnote-11)

## Recovery of Scene Geometry

Although useful for obtaining 2D-3D correspondences in support of projector calibration (cf. Section 2.1.2), we are in principle not restricted to structured light stereo for obtaining scene geometry in general. Structured light stereo, however, *is* the most attractive of the options that present themselves given the hardware setup in the Pilotfabrik, and is a **gold standard** approach for reconstruction of **indoor poorly textured surfaces** given its favorable quality-to-price tradeoff.

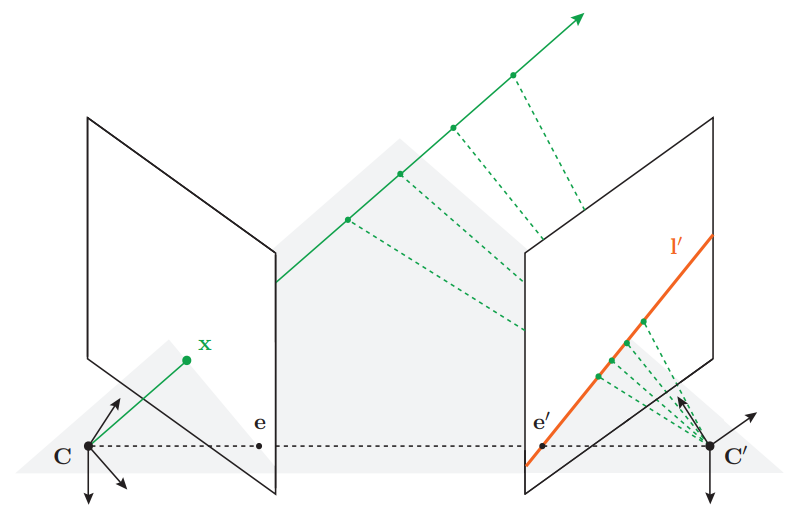


Figure 4: Epipolar geometry (i.e., the geometry of stereo). The 3D point **X** that projects to the pixel **x** in the left camera’s image plane must lie along the ray passing through **x** from the left camera’s center of projection **C**. Accordingly, the projection **x**’ of **X** to the image plane of the right camera must lie along the epipolar line **l**’ spanned by **x**, obtained by projecting that ray to the right camera’s image plane. This restricts the search for **x**’ to only the pixels along that line, a constraint common to all stereo approaches. Once **x**’ has been identified, the point **X** can be recovered by triangulation.

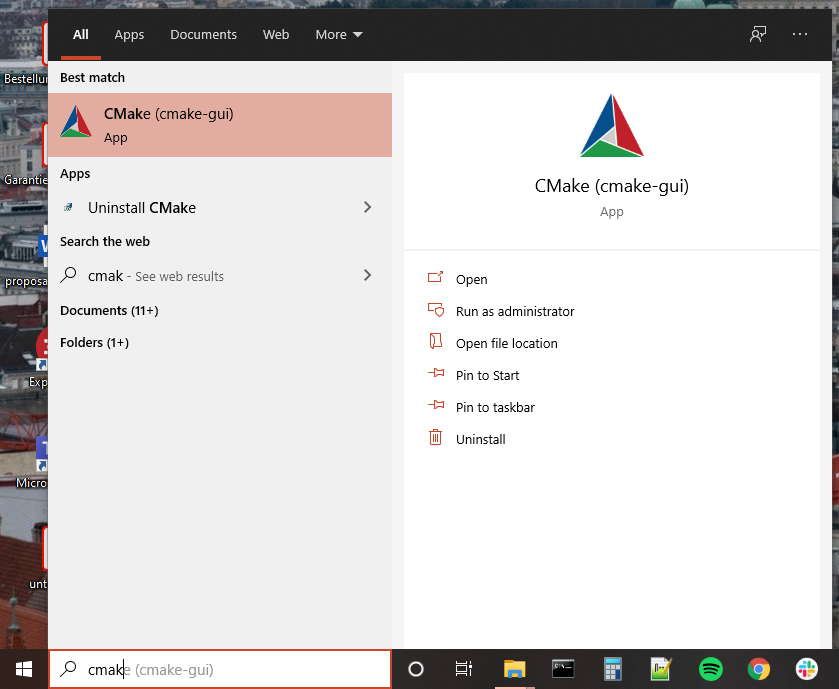
Fundamental to all variants of stereo is that, given a pixel **x** in the left image, the search for the corresponding pixel **x**’ in the right image is restricted to only pixels situated along the **epipolar line** spanned by **x**, as illustrated in Figure 4. *Where stereo approaches differ is primarily in how the ‘best’ match along an epipolar line is determined.* Classical **correlation-based** stereo approaches rely on minimizing a similarity score over a window about **x** in the left image and the corresponding windows about respective candidate matching pixel in the right image. However, a correlation-based approach can be expected to work well only in the presence of adequately **discriminative texture**.[[12]](#footnote-12) In structured light stereo, matching is carried out with respect to binary codes derived from projecting successive strip patterns, thereby in effect *introducing* discriminative texture into the scene and enabling the reconstruction of poorly textured surfaces. As it relies on illumination introduced to the scene, structured light stereo is termed an **active** sensing approach; in contrast, correlation-based stereo is considered a **passive** sensing approach.

Todo…

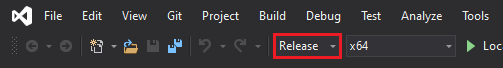
# Building the Code

Let CPSSPATIALAR\_BASEDIR point to the base directory of your local copy the cpsspatialar Git repository[[13]](#footnote-13) (such that src and docs are subdirectories of CPSSPATIALAR\_BASEDIR):

1. Install **Visual Studio Professional 2019**[[14]](#footnote-14)
2. Install **CMake**[[15]](#footnote-15)
3. Let EIGEN\_BASEDIR refer to CPSSPATIALAR\_BASEDIR\src\external\eigen; download version 3.4 of the source code of **Eigen**[[16]](#footnote-16) and extract to EIGEN\_BASEDIR (such that the CMakeLists.txt file is located in EIGEN\_BASEDIR)
4. Let OPENCV\_BASEDIR refer to CPSSPATIALAR\_BASEDIR\src\external\opencv; download version 4.5.1 of the source code of **OpenCV**[[17]](#footnote-17) and extract to OPENCV\_BASEDIR (such that the CMakeLists.txt file is located in OPENCV\_BASEDIR)
5. Run CMake (cmake-gui):



1. Enter OPENCV\_BASEDIR in the field to the right of ‘Where is the source code’ and OPENCV\_BASEDIR\build in the field to the right of ‘Where to build the binaries’ (note that backslashes are converted automatically to forward slashes)
2. Click ‘Configure’ and click ‘Yes’ if prompted to create the directory OPENCV\_BASEDIR
3. If prompted to specify the generator for this project and to optionally indicate the platform for the generator, provide ‘Visual Studio 16 2019’ and ‘x64’, respectively, and click ‘Finish’
4. Click ‘Open Project’ or, equivalently, open OPENCV\_BASEDIR\build\OpenCV.sln; this opens the Visual Studio solution OpenCV.sln in Visual Studio Professional 2019
5. Select ‘Release’ from the Solution Configuration pull-down menu and select ‘Build:Build Solution’ from the menu bar to finally build (i.e., compile) OpenCV in release mode (note that it will take some minutes for building to complete):



1. Let FREEGLUT\_BASEDIR refer to CPSSPATIALAR\_BASEDIR\src\external\ freeglut; download version 2.8.1 of **freeglut**[[18]](#footnote-18) pre-built binaries to FREEGLUT\_BASEDIR (such that the CMakeLists.txt file is located in FREEGLUT\_BASEDIR)
2. Let GLEW\_BASEDIR refer to CPSSPATIALAR\_BASEDIR\src\external\glew; download version 2.1.0 of the source code of **glew**[[19]](#footnote-19) and extract to GLEW\_BASEDIR (such that the Makefile file is located in GLEW\_BASEDIR)
3. Run CMake (cmake-gui)
4. Enter GLEW\_BASEDIR\build\cmake in the field to the right of ‘Where is the source code’ and GLEW\_BASEDIR in the field to the right of ‘Where to build the binaries’ (note that backslashes are converted automatically to forward slashes)
5. Click ‘Configure’
6. If prompted to specify the generator for this project and to optionally indicate the platform for the generator, provide ‘Visual Studio 16 2019’ and ‘x64’, respectively, and click ‘Finish’
7. Click ‘Open Project’ or, equivalently, open GLEW\_BASEDIR\glew.sln; this opens the Visual Studio solution glew.sln in Visual Studio Professional 2019
8. Select ‘Release’ from the Solution Configuration pull-down menu and select ‘Build:Build Solution’ from the menu bar to build glew in release mode
9. Let GLFW\_BASEDIR refer to CPSSPATIALAR\_BASEDIR\src\external\glfw; download version 3.3.4 of the source code of **glfw**[[20]](#footnote-20) and extract to GLFW\_BASEDIR (such that the Makefile file is located in GLFW\_BASEDIR)
10. Run CMake (cmake-gui)
11. Enter GLFW\_BASEDIR in the field to the right of ‘Where is the source code’ and GLFW\_BASEDIR\build in the field to the right of ‘Where to build the binaries’ (note that backslashes are converted automatically to forward slashes)
12. Click ‘Configure’
13. If prompted to specify the generator for this project and to optionally indicate the platform for the generator, provide ‘Visual Studio 16 2019’ and ‘x64’, respectively, and click ‘Finish’
14. Click ‘Open Project’ or, equivalently, open GLEW\_BASEDIR\glew.sln; this opens the Visual Studio solution glew.sln in Visual Studio Professional 2019
15. Select ‘Release’ from the Solution Configuration pull-down menu and select ‘Build:Build Solution’ from the menu bar to build glew in release mode
16. Run CMake (cmake-gui)
17. Enter CPSSPATIALAR\_BASEDIR\src in the field to the right of ‘Where is the source code’ and CPSSPATIALAR\_BASEDIR\build in the field to the right of ‘Where to build the binaries’ (note that backslashes are converted automatically to forward slashes)
18. Click ‘Configure’ and click ‘Yes’ if prompted to create the directory CPSSPATIALAR\_BASEDIR
19. If prompted to specify the generator for this project and to optionally indicate the platform for the generator, provide ‘Visual Studio 16 2019’ and ‘x64’, respectively, and click ‘Finish’
20. Click ‘Open Project’ or, equivalently, open CPSSPATIALAR\_BASEDIR\build\cpspatialar.sln; this opens the Visual Studio solution cpspatialar.sln in Visual Studio Professional 2019
21. Select ‘Release’ from the Solution Configuration pull-down menu and select ‘Build:Build Solution’ from the menu bar to build cpsspatialar in release mode

The **resulting executables** can be found at CPSSPATIALAR\_BASEDIR\build\bin\Release

1. <https://www.pilotfabrik.at/?lang=en> [↑](#footnote-ref-1)
2. Cf. Section 4.1.2 of the Spatial Augmented Reality textbook of Bimber and Raskar, available at <http://pages.cs.wisc.edu/~dyer/cs534/papers/SAR.pdf>. Note that the authors also mention **user location**; this is relevant for some—but not all—use cases, and we shall leave it too outside the scope of this document. [↑](#footnote-ref-2)
3. Cf. again Section 4.1.2 of Bimber and Raskar. [↑](#footnote-ref-3)
4. If you’ll forgive a brazen bit of self-indulgence, have a look at Chapter 2 of my dissertation at <https://publik.tuwien.ac.at/files/PubDat_243668.pdf> for a more detailed attempt at a self-contained but accessible introduction to the geometry pinhole cameras, the derivation of the camera calibration matrix **K**, and other topics of relevance to this document (including epipolar geometry). [↑](#footnote-ref-4)
5. E.g., (0,0,0), (1,0,0), (2,0,0), …, (6,5,0) for a chessboard with 7 × 6 corners, with each square of length and width of 1 unit, respectively. [↑](#footnote-ref-5)
6. See <https://docs.opencv.org/master/dc/dbb/tutorial_py_calibration.html> for a comprehensive tutorial. [↑](#footnote-ref-6)
7. Beware: for any given viewpoint, the *entirety* of the chessboard must be visible if relying on OpenCV’s findChessboardCorners(). Otherwise, the function will fail to find any corners in the image acquired from that viewpoint. [↑](#footnote-ref-7)
8. Each list contains a list corresponding to each respective image. [↑](#footnote-ref-8)
9. See again <https://docs.opencv.org/master/dc/dbb/tutorial_py_calibration.html> for a comprehensive tutorial. [↑](#footnote-ref-9)
10. See <https://www.learnopencv.com/understanding-lens-distortion/> for more details on how these particular coefficients are interpreted mathematically. [↑](#footnote-ref-10)
11. If this turns out to be necessary (it may well be the case that our projector already adequately corrects for lens distortions), we might have to implement our own function; cf. e.g. <https://stackoverflow.com/questions/21615298/opencv-distort-back>. [↑](#footnote-ref-11)
12. Objects like the white walls of a room *or typical manmade objects* are too poorly textured to enable reliable matching via correlation-based stereo. [↑](#footnote-ref-12)
13. <https://github.com/m-hornacek/cpsspatialar> [↑](#footnote-ref-13)
14. Available via TU Wien Campus Software at <https://www.it.tuwien.ac.at/en/services/software/software-provision/campus-software/>. [↑](#footnote-ref-14)
15. Via Windows win64-x64 Installer at <https://cmake.org/download/>. [↑](#footnote-ref-15)
16. From <https://gitlab.com/libeigen/eigen/-/archive/3.4/eigen-3.4.zip>. [↑](#footnote-ref-16)
17. From <https://github.com/opencv/opencv/archive/4.5.1.zip>. [↑](#footnote-ref-17)
18. From <https://www.transmissionzero.co.uk/files/software/development/GLUT/older/freeglut-MSVC-2.8.1-1.mp.zip>. [↑](#footnote-ref-18)
19. From <https://sourceforge.net/projects/glew/files/glew/2.1.0/glew-2.1.0.zip/download>. [↑](#footnote-ref-19)
20. From <https://github.com/glfw/glfw/releases/download/3.3.4/glfw-3.3.4.zip>. [↑](#footnote-ref-20)