

# ME 104 Lec 9

Recall: Two equivalent versions of the work-energy theorem:

$$\textcircled{1} \quad W_{AB}^{\text{Tot}} = \sum_{\text{All forces}} \int_{\text{Path } \underline{r}_A \rightarrow \underline{r}_B} \underline{F}^i \cdot d\underline{r} = \frac{1}{2}mv_B^2 - \frac{1}{2}mv_A^2.$$

$$\textcircled{2} \quad W_{AB}^{\text{nc}} = \sum_{\text{Non-conservative forces}} \int_{\text{Path } \underline{r}_A \rightarrow \underline{r}_B} \underline{F}^{\text{nc}} \cdot d\underline{r} = E_B - E_A$$

where  $E = \frac{1}{2}mv^2 + \sum_c U^c$  → Total potential energy of all conservative forces.  
( $\underline{F}^c = -\nabla U^c$ )

Benefits of using work-energy:

Can determine certain features of the motion without having to

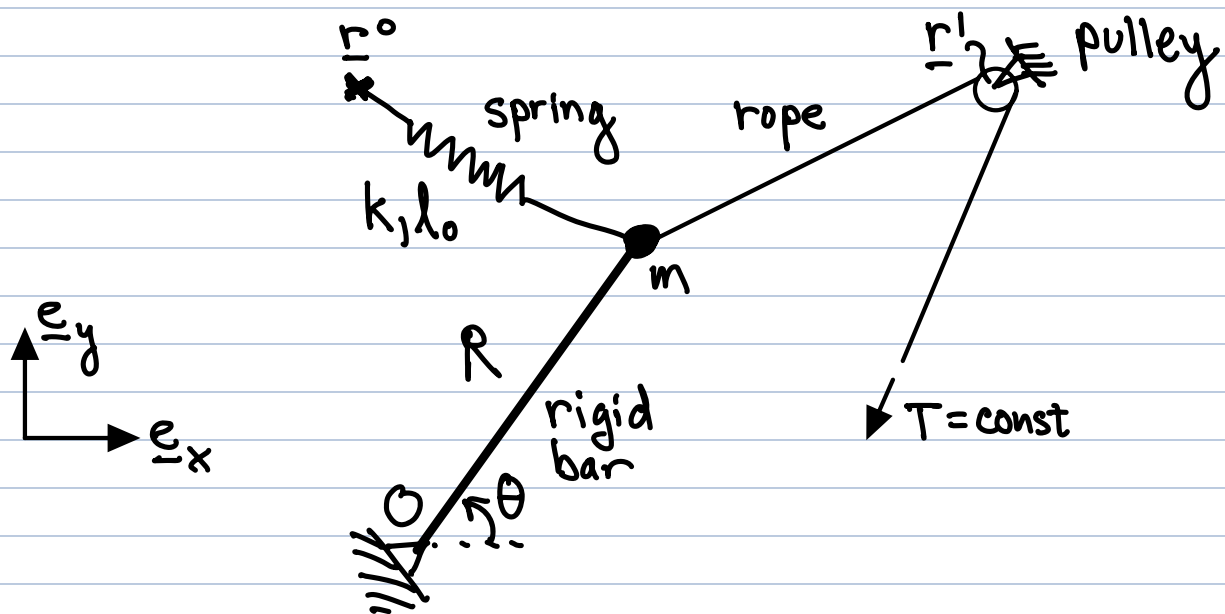
- solve first for (powerless) constraint forces, e.g. normal wall forces or rigid bar forces,
- solve a system of equations,
- solve differential equations.

Issues using energy balance include:

- hard to use if  $W_{AB}^{nc}$  is difficult to calculate (best if  $W_{AB}^{nc} = 0 = \Delta E$ , i.e. energy conserved)
- since it is just one equation it can only tell you one thing about the motion.

Using energy balance, we can often infer unknown forces after the fact.

Example from last time:



We previously calculated  $|\dot{\theta}_f|$  at some  $\theta_f$  assuming mass started stationary at  $\theta_i$ .  
What is  $\underline{F}_{\text{bar}}$  at  $\theta_f$ ?

Ans:  $\sum \underline{F} = m \underline{a}$  gives

$$\{F_{\text{bar}} \underline{e}_r + \underline{F}_s + \underline{F}_{\text{rope}} = m(-R\dot{\theta}^2 \underline{e}_r + R\ddot{\theta} \underline{e}_\theta)\} \cdot \underline{e}_r$$

$$\Rightarrow F_{\text{bar}} + \underline{F}_s \cdot \underline{e}_r + \underline{F}_{\text{rope}} \cdot \underline{e}_r = -mR\dot{\theta}^2$$

Thus, for  $\theta = \theta_f$  and  $\dot{\theta}^2 = |\dot{\theta}_f|^2$ , we have

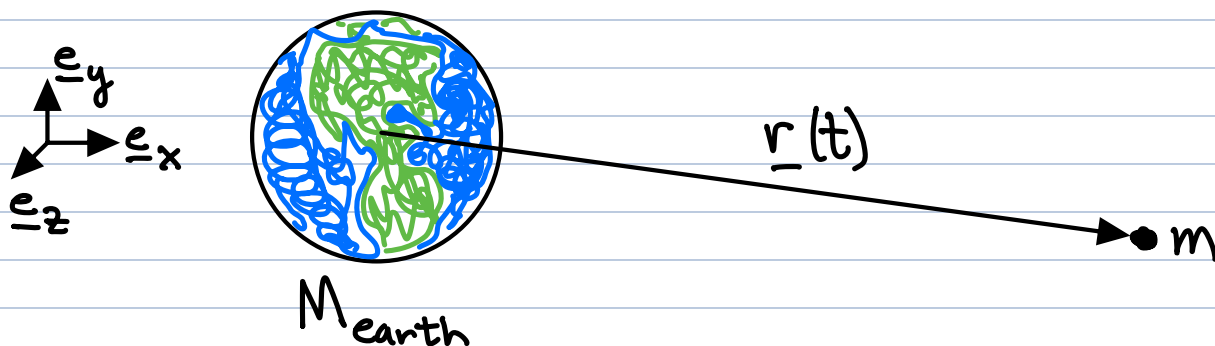
$$F_{\text{bar}} = -mR|\dot{\theta}_f|^2 - [\underline{F}_s(\theta_f, R, \underline{r}^o) + \underline{F}_{\text{rope}}(\theta_f, R, \underline{r}')] \cdot (\cos\theta_f \underline{e}_x + \sin\theta_f \underline{e}_y)$$

Since we know  $R, \theta_f, |\dot{\theta}_f|, \underline{r}^o$ , and  $\underline{r}'$  we can compute boxed formula for  $F_{\text{bar}}$ . ✓

In general, dotting  $\Sigma \underline{F} = m\underline{a}$  into the direction normal to a wall let's you compute the normal force knowing only the particle's speed, position, and path curvature.

## Orbital mechanics

Assume an object interacting with earth's gravity with  $m \ll M_{\text{earth}}$ .  $\Rightarrow$  The center of the earth can be assumed stationary. Let the origin be at earth's center.



$$\textcircled{1} \quad \underline{r} = x \underline{e}_x + y \underline{e}_y + z \underline{e}_z, \quad \underline{v} = \dot{x} \underline{e}_x + \dot{y} \underline{e}_y + \dot{z} \underline{e}_z \\ \underline{a} = \ddot{x} \underline{e}_x + \ddot{y} \underline{e}_y + \ddot{z} \underline{e}_z.$$

$$\textcircled{2} \quad \underline{F}_g = \frac{G M_{\text{earth}} m}{|\underline{r}|^2} \left( \frac{-\underline{r}}{|\underline{r}|} \right)$$

A diagram showing a point labeled  $O$  (the origin) and a mass. A dashed line connects  $O$  to the mass. A vector labeled  $\underline{F}_g$  points from the mass towards  $O$ .

$\textcircled{3}$  Defining  $\beta \equiv G M_{\text{earth}}$ ,  $\Sigma \underline{F} = m \underline{a}$  implies

$$\boxed{\underline{a} = -\beta \underline{r} / |\underline{r}|^3} \quad \text{"Trajectory equation".}$$

In terms of work-energy,  $\underline{F}_g$  is conservative

as a power-law force with  $U_g = -\beta / |\underline{r}|$ .

Since no nc forces, energy is conserved

so :

$$\boxed{\frac{1}{2} m v^2 - \beta m / |\underline{r}| = \text{const}}$$

④ Calculate the "escape velocity"—the speed that the mass must be launched with from earth's surface to escape earth's pull without any extra propulsion.

$$E(\text{right after launch}) = \frac{1}{2}mv_e^2 - \beta m/R_{\text{earth}}$$

$$E(\text{at escape}) = \frac{1}{2}m \cdot 0^2 - \beta m/|r \rightarrow \infty| = 0$$

$$\Rightarrow \frac{1}{2}mv_e^2 - \beta m/R_{\text{earth}} = 0 \Rightarrow v_e = \sqrt{\frac{2\beta}{R_{\text{earth}}}} = \boxed{11,200 \text{ m/s.}}$$

Need more than just energy balance to learn more trajectory behaviors. Must solve trajectory eq directly or use other conservation laws to augment energy balance.

Angular momentum: Let  $\underline{H} = m \underline{r} \times \underline{v}$ .

$$\begin{aligned} \text{Note that } \dot{\underline{H}} &= m \dot{\underline{r}} \times \underline{v} + m \underline{r} \times \dot{\underline{v}} \\ &= m \underbrace{\underline{v} \times \underline{v}}_{=0} + \underline{r} \times (m \underbrace{\underline{a}}_{=\underline{F}_g = -\beta m \underline{r}/|\underline{r}|^3 \propto \underline{r}}) \\ &= \underline{0} \end{aligned}$$

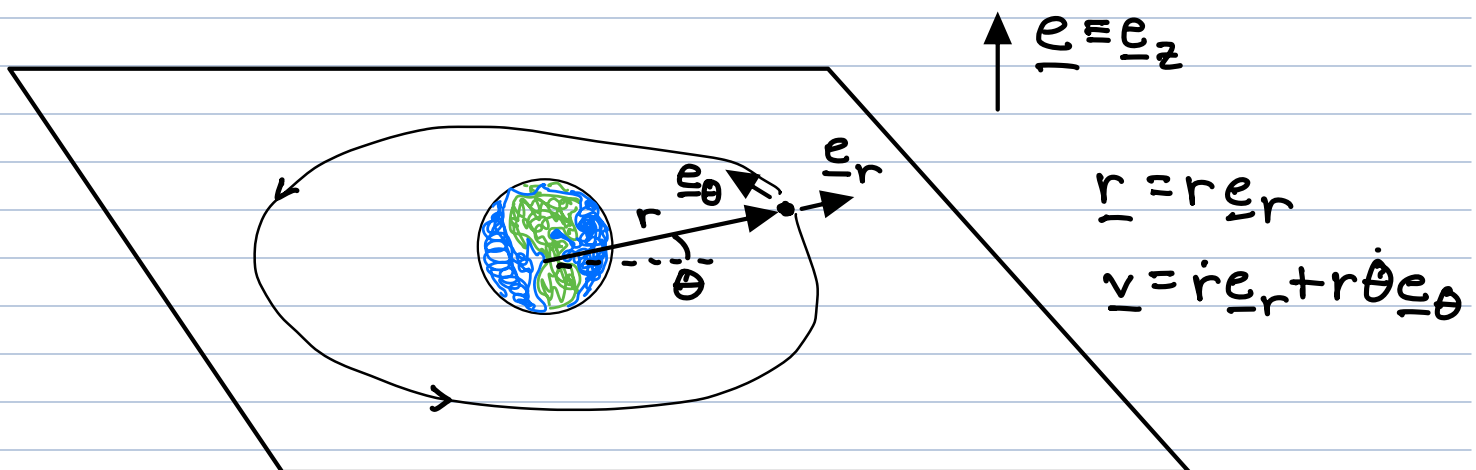
So  $\underline{H}$  = "Angular momentum" =  $\underline{H}_0$  is conserved.

{ Generalizing:  
 $\underline{H}$  is hence conserved whenever a particle is  
subject to a net force  $\parallel \underline{r}$ . 😊 }

$$\underline{H} = m \underline{r} \times \underline{v} = \underline{H}_0 = H_0 \underline{e} \quad \text{for some fixed } H_0 \text{ and } \underline{e}.$$

Fixed  $\underline{e}$  implies  $\underline{r}$  and  $\underline{v}$  must always lie in one fixed plane  $\perp$  to  $\underline{e}$ . But this implies the trajectory of  $\underline{r}(t)$  must lie in a single plane. Orbits always lie in one plane!

Knowing this, the problem is now 2D.



$$\underline{H} = m (r \underline{e}_r) \times (\dot{r} \underline{e}_r + r \dot{\theta} \underline{e}_\theta) = m r^2 \dot{\theta} \underline{e}_z = H_0 \underline{e}_z$$

$$\Rightarrow m r^2 \dot{\theta} = H_0 \quad \text{where } H_0 = m r_i^2 \dot{\theta}_i$$

$$\Rightarrow \dot{\theta} = H_0 / m r^2.$$

Now conserve energy:  $\frac{1}{2}mv^2 - \beta m/r = E_0 = \text{const}$

$$\Rightarrow \frac{1}{2}m(\dot{r}^2 + r^2\dot{\theta}^2) - \beta m/r = E_0$$

$$\Rightarrow \frac{1}{2}m\left(\dot{r}^2 + \frac{H_0^2}{m^2 r^2}\right) - \beta m/r = E_0 \leftarrow \text{ODE for } r(t).$$

Rather than solve for  $r(t)$ , we can use this to learn the trajectory  $r(\theta)$ .

$$\dot{r} = \frac{dr}{d\theta} \dot{\theta} = \frac{dr}{d\theta} \frac{H_0}{mr^2} \quad \text{Plug in above:}$$

$$\Rightarrow \frac{1}{2}m\left(\left(\frac{dr}{d\theta}\right)^2 \left(\frac{H_0}{mr^2}\right)^2 + \frac{H_0^2}{m^2 r^2}\right) - \beta m/r = E_0$$

$$\Rightarrow \frac{dr}{d\theta} = \left( \frac{2}{m} \left( E_0 + \beta m/r \right) - \frac{H_0^2}{m^2 r^2} \right)^{1/2} \frac{m^2 r^4}{H_0^2}$$

Solve by separation of variables. We get:

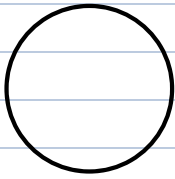
$$r(\theta) = \frac{p}{1 + e \cdot \cos(\theta - \phi)}$$

$$p = \text{"trajectory parameter"} = \frac{H_0^2}{m^2 \beta}$$

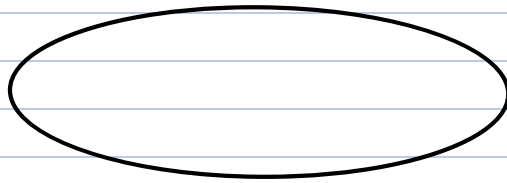
$$e = \text{"eccentricity"} = \left( 1 + \frac{2E_0 H_0^2}{m^3 \beta^2} \right)^{1/2}$$

$$\phi = \text{"offset angle"} = \theta_i - \cos^{-1} \left( \frac{p - r_i}{e r_i} \right)$$

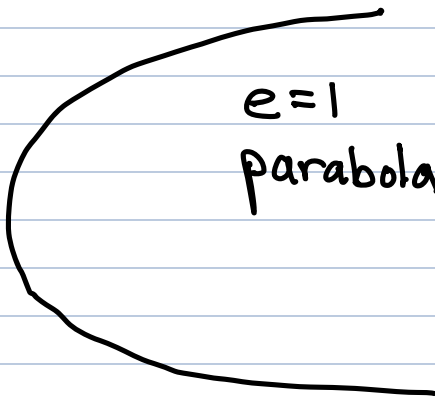
What are these shapes? Conic sections!!



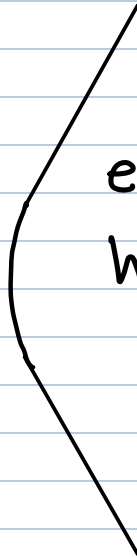
$e=0$   
circle



$0 < e < 1$   
ellipse



$e=1$   
parabola



$e > 1$   
hyperbola