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EXOPLANET DETECTION LIMITS

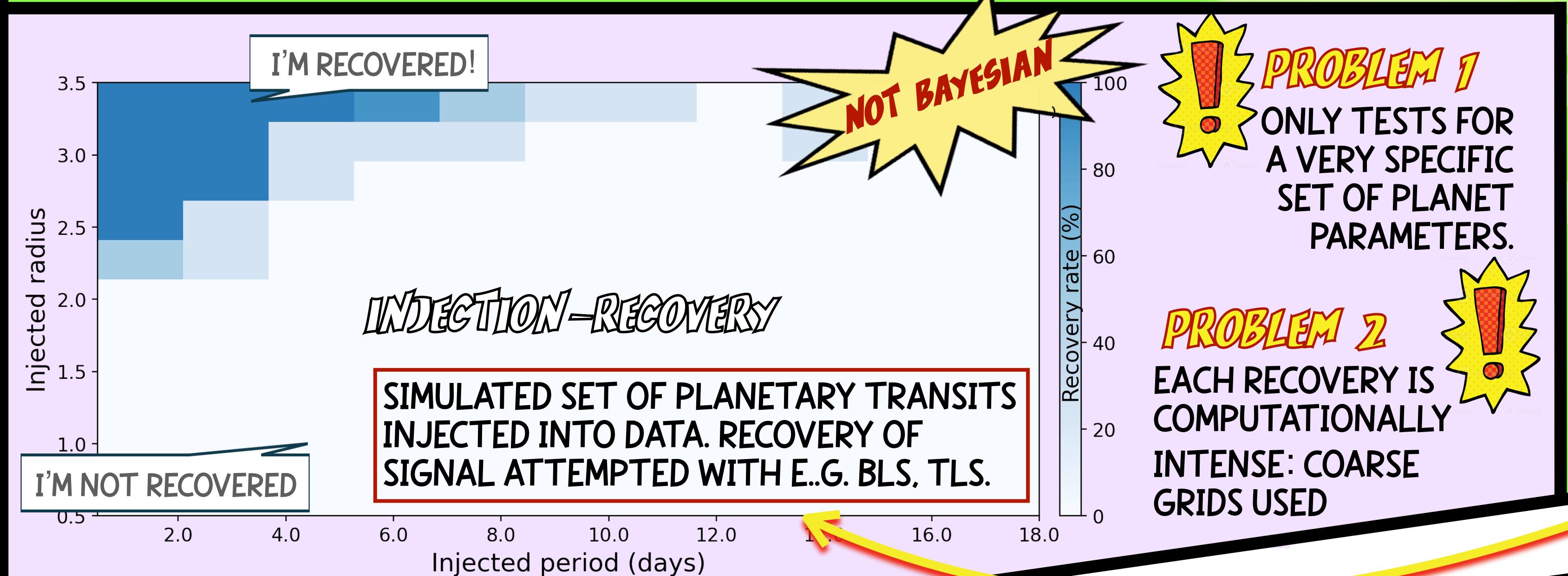
WOW!
OCCURRENCE
RATES!

NEW
METHOD!

WE WANT TO MEASURE
THE OCCURRENCE RATES
OF EXOPLANETS AROUND
LATE-TYPE M DWARFS

WE WANT TO FIND WHERE OUR DETECTION LIMIT LIES FOR A GIVEN PHOTOMETRIC DATASET

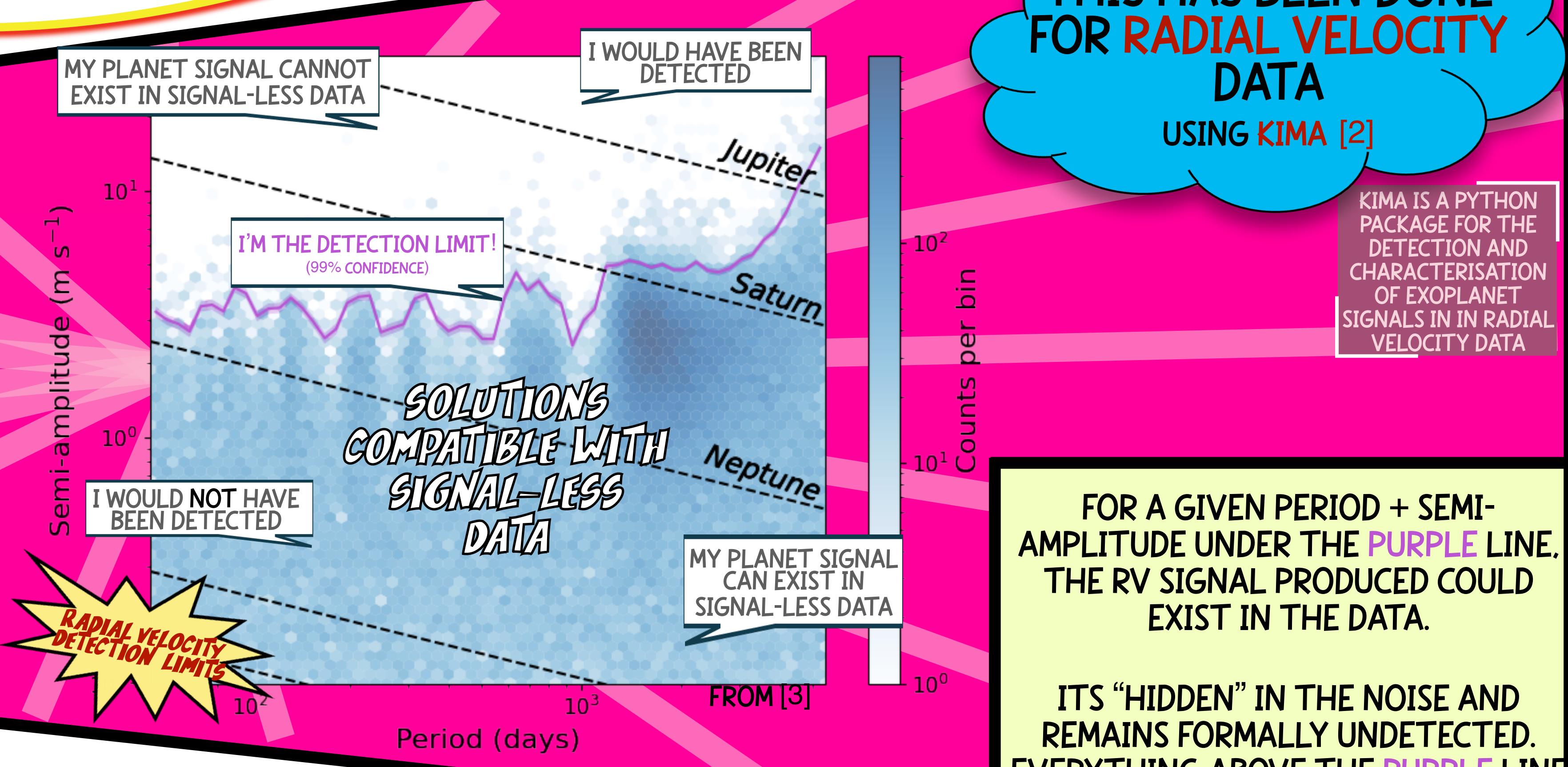
THIS WILL HELP INFORM OCCURRENCE RATES BY ASSESSING OUR COMPLETENESS USING
DETECTION LIMITS



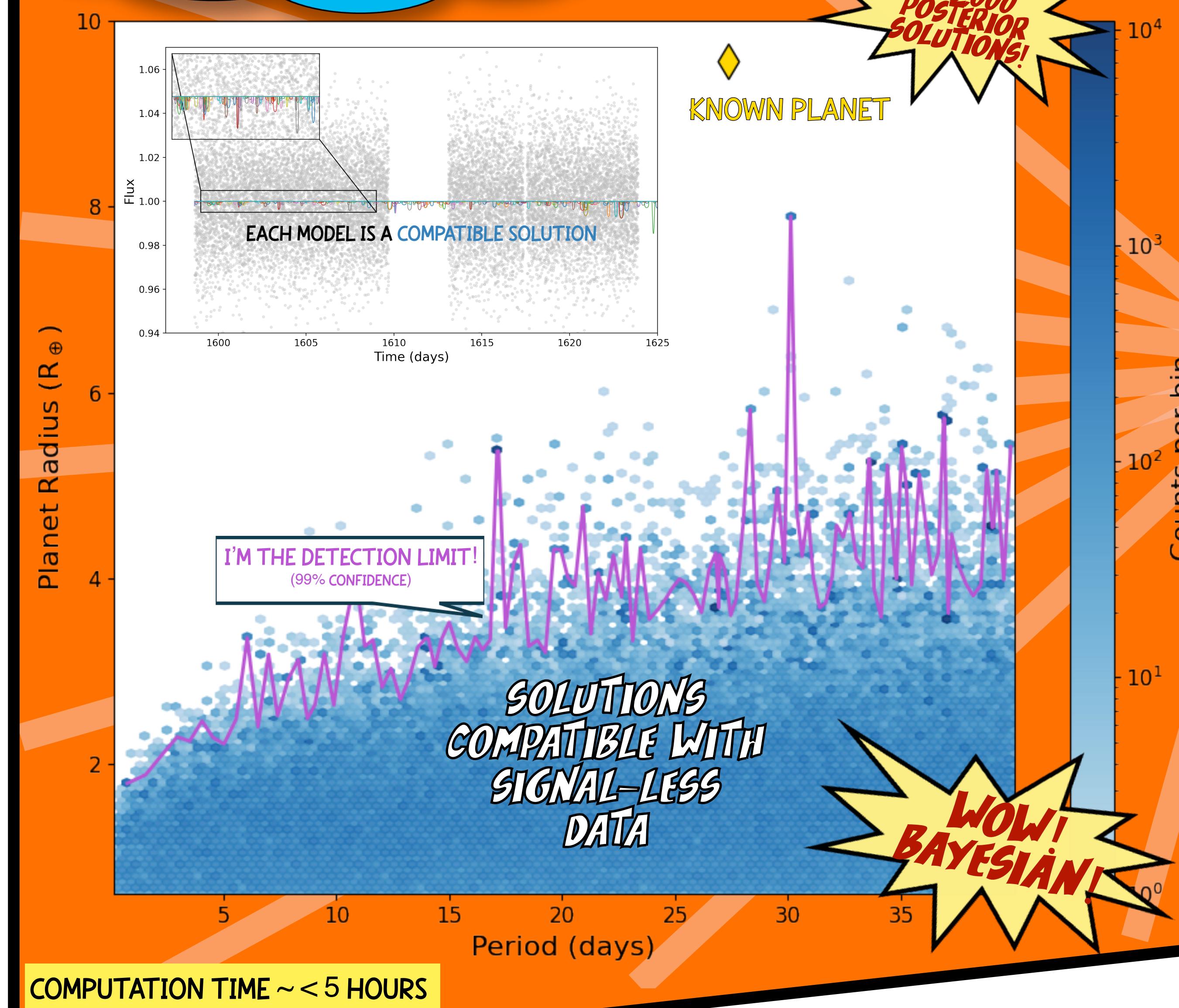
THE APPROACH IS MORE COMPLEX FOR APPLICATIONS TO TRANSITS.

! PROBLEM 3 UNLIKE RV DATA, THERE IS NOT A SIGNAL ACROSS THE ENTIRE ORBIT OF THE PLANET.

! PROBLEM 4 HOLES IN THE DATASET HAVE BIG IMPACTS ON EXOPLANET DETECTIONS.



HOW ARE WE DOING IT FOR PHOTOMETRY?



METHOD

BATMAN [4] + DYNESTY [5] PYTHON PACKAGES.

PRIORS ON MODEL PARAMETERS ARE EITHER **LARGE + UNIFORM** ($t_0, P, R_p/R_{\star}, \omega$) OR **INFORMED** FROM THE CURRENT TRANSITING EXOPLANET POPULATION (e, i)

TENS OF THOUSANDS OF MODELS ARE PROPOSED. MODELS ARE EITHER **REJECTED** IF NOT COMPATIBLE WITH SIGNAL-LESS DATA, OR **ACCEPTED** AS A COMPATIBLE SOLUTION.

ADVANTAGES COMPARED TO INJECTION RECOVERY

- 1 BAYESIAN
- 2 NON-DISCRETE SOLUTION GRID
- 3 4 MORE PARAMETERS DRAWN FROM DISTRIBUTIONS
- 4 100s – 1000s MORE SOLUTIONS
- 5 ~20 × FASTER

SCOTT DAVIES, TRIAUD & IN PREP

COMING 2025/26!

[1] DÉVORA-PAJARES, M. AND POZUELOS, F. J. "MATRIX: MULTI-PHASE TRANSITS RECOVERY FROM INJECTED EXOPLANETS"

[2] FARIA, J. P., SANTOS, N. C., FIGUEIRA, P., AND BREWER, B. J. "KIMA: DETECTION IN RADIAL EXOPLANETS"

[3] BAYCROFT, T. ET AL. "BEBOP VII. SOPHIE DISCOVERY OF BEBOP-3b: A CIRCUMINARY GIANT PLANET ON AN ECCENTRIC ORBIT"

[4] KREIDBERG, L. "BATMAN: BASIC TRANSIT MODEL CALCULATION IN PYTHON"

[5] SPEAGLE, J. S. "DYNESTY: A DYNAMIC SAMPLING PACKAGE FOR ESTIMATING