

1. The Wave Function

1.1 The Schrödinger Equation

$$i\hbar \frac{\partial \Psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \Psi}{\partial x^2} + V\Psi$$

$$\text{or } i\hbar \frac{\partial \Psi(\vec{r}, t)}{\partial t} = -\frac{\hbar^2}{2m} \vec{\nabla}^2 \Psi(\vec{r}, t) + V(\vec{r}, t) \Psi(\vec{r}, t)$$

$$\text{where } \vec{\nabla}^2 = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$$

Solve for the particle's wave function $\Psi(x, t)$

$$\hbar = \frac{h}{2\pi} = 1.054572 \times 10^{-34} \text{ Js}$$

1.2 The Statistical Interpretation

$$\int_a^b |\Psi(x, t)|^2 dx = \{P \text{ of finding the particle btwn } a \text{ and } b, \text{ at } t\}$$

1.3 Probability

$$\text{Standard deviation: } \sigma = \sqrt{\langle j^2 \rangle - \langle j \rangle^2}$$

$$\text{Expectation value of } x \text{ given } \Psi: \langle x \rangle = \int x |\Psi|^2 dx$$

$$\text{Probability current: } J(x, t) = \frac{i\hbar}{2m} (\Psi \frac{\partial \Psi^*}{\partial x} - \Psi^* \frac{\partial \Psi}{\partial x})$$

1.4 Normalization

$$\int_{-\infty}^{+\infty} |\Psi(x, t)|^2 dx = 1$$

The Schrödinger equation produces unitary time evolution:

$$\vec{J} = -\frac{i\hbar}{2m} (\psi(\vec{r}, t)^* \vec{\nabla} \psi(\vec{r}, t) - \psi(\vec{r}, t) \vec{\nabla} \psi^*(\vec{r}, t)^*)$$

The probability density satisfies the continuity equation,

$$\frac{\partial}{\partial t} \rho + \vec{\nabla} \cdot \vec{J} = 0$$

Because the probability for finding the particle at infinity is 0

(otherwise non-normalizable), $\vec{J} = 0$ at infinity.

Therefore, $\frac{d}{dt} \int_{-\infty}^{\infty} \rho d^3\vec{r} = \frac{d}{dt} P = 0$, where P is the total probability
→ the total probability is constant in time.

1.5 Momentum

For a particle in state Φ , the expectation value of x and p is

$$\langle x \rangle = \int_{-\infty}^{+\infty} x |\Psi(x, t)|^2 dx \quad \langle p \rangle = m \frac{d\langle x \rangle}{dt} = -i\hbar \int (\Psi^* \frac{\partial \Psi}{\partial x}) dx$$

To calculate the expectation value of any quantity, $Q(x, p)$:

$$\langle Q(x, p) \rangle = \int \Psi^* Q(x, \frac{\hbar}{i} \frac{\partial}{\partial x}) \Psi dx$$

Position and momentum operators: $\hat{r} = \vec{r}$, $\hat{p} = -i\hbar \vec{\nabla}$

1.6: The Uncertainty Principle

The wavelength of Ψ is related to the momentum of the particle by the de Broglie formula:

$$p = \frac{h}{\lambda} = \frac{2\pi\hbar}{\lambda}$$

The more precisely determined a particle's position is, the less precisely is its momentum. The Heisenberg's uncertainty principle:

$$\sigma_x \sigma_p \geq \frac{\hbar}{2}$$

Commutation relation btwn position and momentum:

$$\hat{p}_x (\hat{x} \psi(x, t)) = -i\hbar \frac{\partial}{\partial x} [x \psi(x, t)] = -i\hbar \psi(x, t) - i\hbar x \frac{\partial}{\partial x} \psi(x, t)$$

$$\hat{x} (\hat{p}_x \psi(x, t)) = x (-i\hbar \frac{\partial}{\partial x} \psi(x, t))$$

$$\hat{x} \hat{p}_x - \hat{p}_x \hat{x} = [\hat{x}, \hat{p}_x] = i\hbar$$

$$[\hat{x}_i, \hat{p}_j] = i\hbar \delta_{ij}, [\hat{x}_i, \hat{x}_j] = [\hat{p}_i, \hat{p}_j] = 0,$$

where $\delta_{ij} = 1$ for $i = j$ and $\delta_{ij} = 0$ for $i \neq j$

Given three operators $\hat{A}, \hat{B}, \hat{C}$, we have $[\hat{A}, \hat{B}\hat{C}] = [\hat{A}, \hat{B}]\hat{C} + \hat{B}[\hat{A}, \hat{C}]$.

Other: Blackbody Spectrum

$$E = h\nu = \hbar\omega$$

The wave number k is $k = 2\pi/\lambda = \omega/c$

Only two spin states occur (quantum number m is +1 or -1).

$$\rho(\omega) = \frac{\hbar\omega^3}{\pi^2 c^3 (e^{\hbar\omega/k_b T} - 1)}$$

$$\text{Wien displacement law: } \lambda_{\max} = \frac{2.90 \times 10^{-3} \text{ mK}}{T}$$

2. Time-Independent Schrödinger Equation

2.1 Stationary States

Suppose PE is independent of time, $V(\vec{r}, t) = V(\vec{r})$.

Separation of variables: $\Psi(\vec{r}, t) = \psi(\vec{r}) \varphi(t)$

Eq of motion for $\varphi(t)$: $\varphi(t) = e^{-iEt/\hbar}$

Eq of motion for $\psi(\vec{r})$ is the time-independent Schrödinger equation:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(\vec{r})}{dx^2} + V(\vec{r}) \psi(\vec{r}) = E \psi(\vec{r})$$

TD of the wavefunction that corresponds to the constant E is easily written once we solve the TISE: $\Psi_E(\vec{r}, t) = \psi_E(\vec{r}) e^{-iEt/\hbar}$

Properties of solutions for TI potentials:

- **The constant E must be real.**
- **Stationary wavefunction.**
 $\mathcal{P}(\vec{r}, t) = |\psi_E(\vec{r}, t)|^2 = |\psi_E(\vec{r})|^2$ (TD cancels out).
- **Stationary wavefunction is a state of definite energy.**
The total energy (kinetic plus potential) is the Hamiltonian:
 $H(x, p) = \frac{p^2}{2m} + V(x)$.
Hamiltonian operator: $\hat{H} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x)$
Thus the TISE can be written as $\hat{H}\psi = E\psi$
 $\langle \hat{H} \rangle = E$, $\langle \hat{H}^2 \rangle = E^2$, $\Delta E = \sqrt{\langle \hat{H}^2 \rangle - \langle \hat{H} \rangle^2} = 0$
- Spatial part of stationary wavefunction can be chosen to be real.
 $\psi^*(\vec{r})$ is a soln w/ same E
Solns can be chosen to be real: $\psi(\vec{r}) + \psi^*(\vec{r})$ and $\frac{\psi(\vec{r}) - \psi^*(\vec{r})}{i}$.
- **Parity symmetry: even and odd wavefunctions.** Suppose $V(-\vec{r}) = V(\vec{r})$. Then, $\psi_E(-\vec{r})$ is a soln w the same energy.
 $\psi_E(\vec{r}) + \psi_E(-\vec{r})$ is even under reflection, $\psi_E(\vec{r}) - \psi_E(-\vec{r})$ is odd.
When the potential is symmetric under reflection, we can choose the stationary states to be either even or odd under reflection.
- **Orthogonality/orthonormality.**
 $\int \psi_m(\vec{r})^* \psi_n(\vec{r}) d^3\vec{r} = \delta_{mn}$ where δ_{mn} is 0 if $m \neq n$ and 1 if $m = n$.
- **Linearity.**

The SE is linear. Given stationary states, a linear combo of these

$$\psi(\vec{r}, t) = \sum_n c_n \psi_n(\vec{r}, t) = \sum_n c_n \psi_n(\vec{r}) e^{-\frac{i}{\hbar} E_n t}$$

where c_n are complex constants, is a solution to the TDSE

$$i\hbar \frac{\partial \psi(\vec{r}, t)}{\partial t} = \hat{H} \psi(\vec{r}, t)$$

- **Time evolution.** Given
 $\psi(\vec{r}, 0) = \sum_n c_n \psi_n(\vec{r}, 0) = \sum_n c_n \psi_n(\vec{r})$

at time t , the time evolution is

$$\psi(\vec{r}, t) = \sum_n c_n \psi_n(\vec{r}, t) = \sum_n c_n \psi_n(\vec{r}) e^{-\frac{i}{\hbar} E_n t}$$

Once we've expanded a given initial wavefunction in terms of a linear combo of the stationary wavefunctions $\psi_n(\vec{r})$, the time evolution follows simply by putting a factor of $e^{-i/h E_n t}$ to each term containing $\psi_n(\vec{r})$.

- **Normalization.**
The constant coefficients are constrained by $\sum_n |c_n|^2 = 1$
- **Completeness.**

The stationary states form a complete set if

$$\sum_n \psi_n(\vec{r}', t)^* \psi_n(\vec{r}, t) = \delta^3(\vec{r}' - \vec{r})$$

where $\delta^3(\vec{r}' - \vec{r})$ is the Dirac-delta function in 3D defined by

$$\int d^3\vec{r}' \psi(\vec{r}', t)^* \delta^3(\vec{r}' - \vec{r}) = \psi(\vec{r}, t)$$

Euler's formula: $e^{i\theta} = \cos \theta + i \sin \theta$

Delta function: Given $f(x)$, $\delta(x - x')$ is defined as

$$f(x') = \int f(x) \delta(x - x') dx$$

$\int \delta(x - x') dx = 1$, note this is not the area

$$\delta_\alpha(x) = \frac{1}{\alpha\sqrt{\pi}} e^{-\frac{x^2}{\alpha^2}}, \delta_\alpha(x) = \frac{1}{\pi x} \sin\left(\frac{x}{\alpha}\right), \delta_\alpha(x) = \frac{\alpha}{\pi x^2} \sin^2\left(\frac{x}{\alpha}\right)$$

One-dimensional systems

Wavefunction for a system containing a single particle of mass m in 1D with TI potentials.

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(x)}{dx^2} + V(x) \psi(x) = E \psi(x)$$

Once we find the wavefunction $\psi_E(x)$ of energy E , its time dependence follows easily:

$$\psi_E(x, t) = \psi_E(x) e^{-\frac{i}{\hbar} E t}$$

Boundary conditions

1. When the potential $V(x)$ has a finite jump at $x = a$, both $\psi(x)$ and $\psi'(x)$ are continuous across $x = a$.
2. When the potential $V(x)$ has an infinite jump at $x = a$, $\psi(x)$ is continuous but $\psi'(x)$ is discontinuous across $x = a$.

Futhermore, the wavefunction must vanish at $x = \pm\infty$ for a normalizable wavefunction.

2.2 The Infinite Square Well

Suppose

$$V(x) = \begin{cases} 0, & \text{if } 0 \leq x \leq a \\ \infty, & \text{otherwise} \end{cases}$$

$$\psi(x) = 0 \text{ for } x < 0 \text{ and } x > a$$

For $0 \leq x \leq a$, we have $V(x) = 0$ and the Schrödinger equation reduces to

$$\psi''(x) + k^2 \psi(x) = 0, \text{ where } k = \sqrt{\frac{2mE}{\hbar^2}} \text{ and } E > 0$$

Classic simple harmonic oscillator, $\psi(x) = A \sin(kx) + B \cos(kx)$
Boundary conditions

$$E_n = \frac{\hbar^2 k_n^2}{2m} = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

$$\psi_n(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi x}{a}\right)$$

ψ_1 is the ground state, others are excited states.

Properties of $\psi_n(x)$:

1. Alternatively even and odd.
2. As you go up in energy, each successive state has one more node.
3. They are mutually orthogonal, in the sense that $\int \psi_m(x) * \psi_n(x) dx = 0$ whenever $m \neq n$.
 $\int \psi_m(x) * \psi_n(x) dx = \delta_{mn}$ where δ_{mn} (Kronecker delta) is 0 if $m \neq n$ and 1 if $m = n$. We say that the ϕ 's are orthonormal.
4. They are complete, in the sense that any other function, $f(x)$, can be expressed as a linear combination of them (Fourier series),
Dirichlet's theorem:

$$f(x) = \sum_{n=1}^{\infty} c_n \psi_n(x) = \sqrt{\frac{2}{a}} \sum_{n=1}^{\infty} c_n \sin\left(\frac{n\pi x}{a}\right)$$

Fourier's trick: $c_n = \int \psi_n(x)^* f(x) dx$

$$c_m = \frac{2}{a} \int_0^a f(x) \sin\left(\frac{m\pi x}{a}\right) dx$$

$|c_n|^2$ tells you the probability that a measurement of the energy would yield the value E_n .

Sum of these probabilities should be 1:

$$\sum_{n=1}^{\infty} |c_n|^2 = 1$$

The expectation value of the energy is

$$\langle H \rangle = \sum_{n=1}^{\infty} |c_n|^2 E_n$$

Conservation of energy in QM

2.3 The Harmonic Oscillator

Hooke's law: $F = -kx = m \frac{d^2 x}{dt^2}$

Solution is $x(t) = A \sin(\omega t) + B \cos(\omega t)$, where $\omega = \sqrt{\frac{k}{m}}$,

$$V(x) = \frac{1}{2} k x^2.$$

Taylor series:

$$V(x) = V(x_0) + V'(x_0)(x - x_0) + \frac{1}{2} V''(x_0)(x - x_0)^2 + \dots$$

The Schrödinger Equation for the harmonic oscillator:

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} + \frac{1}{2} m \omega^2 x^2 \psi = E \psi$$

Introduce $\xi \equiv \sqrt{\frac{m\omega}{\hbar}} x$, so we have $\frac{d^2 \psi}{d\xi^2} = (\xi^2 - K) \psi$, where

$$K \equiv \frac{2E}{\hbar\omega}.$$

The recursion formula: $a_{j+2} = \frac{(2j+1-K)}{(j+1)(j+2)} a_j$

The complete solution is $h(\xi) = h_{\text{even}}(\xi) + h_{\text{odd}}(\xi)$

$K = 2n + 1$, so $E_n = (n + \frac{1}{2}) \hbar \omega$

Recursion formula for allowed K : $a_{j+2} = \frac{-2(n-j)}{(j+1)(j+2)} a_j$

Hermite polynomials: $H_0 = 1$, $H_1 = 2\xi$, $H_2 = 4\xi^2 - 2$,

$H_3 = 8\xi^3 - 12\xi$, $H_4 = 16\xi^4 - 48\xi^2 + 12$, $H_5 = 32\xi^5 - 160\xi^3 + 120\xi$

The normalized stationary states:

$$\psi_n(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{1}{\sqrt{2^n n!}} H_n(\xi) e^{-\xi^2/2}$$

Rodrigues formula: $H_n(\xi) = (-1)^n e^{\xi^2} \left(\frac{d}{d\xi}\right)^n e^{-\xi^2}$

2.4 The Free Particle

$$\frac{\partial^2 \xi}{\partial x^2} = -k^2 \xi, k = \frac{\sqrt{2mE}}{\hbar}$$

General solution to the TISE: wave packet,

$$\Psi(x, t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \psi(k) e^{i(kx - \frac{\hbar k^2}{2m} t)} dk$$

$$\phi(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} \Psi(x, 0) e^{-ikx} dx$$

Plancherel's theorem:

$$f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} F(k) e^{ikx} dk \leftrightarrow F(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} f(x) e^{-ikx} dx$$

$F(k)$ is the Fourier transform of $f(x)$; $f(x)$ is the inverse Fourier transform of $F(k)$

Phase velocity: speed of individual ripples; group velocity: speed of the envelope

Dispersion relation: the formula for ω as a function of k

2.5 The Delta-Function Potential

2.6 The Finite Square Well