

MATHEMATICAL PRELIMINARIES, PART 1

SPECIAL AND GENERAL RELATIVITY

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AN IMPORTANT PRELIMINARY requirement for the study of relativity is a good understanding of how co-ordinate systems work. This is best understood in hindsight: general relativity is primarily a geometric description of spacetime and, as a result, relies heavily on co-ordinates.

1 SPHERICAL CO-ORDINATES

Conventionally a spherical polar co-ordinate system in three dimensions consists of a linear distance measure from the origin r , an azimuthal angle swept between the x - and y -axes ϕ , and a polar angle swept up from the xy -plane to the z -axis θ . If such a co-ordinate system were placed with its origin in line with the corner of a room, the simplest way to understand this would be to imagine the x - and y -axes along the floor and the z -axis pointing to the ceiling. The right-hand rule would apply: if the right-hand index finger were extended along the direction of the x -axis, the middle finger would point along the y -axis and the thumb along the z -axis. Every point in space can in this way be represented in terms of (r, θ, ϕ) .

Mathematicians do things differently: their ϕ is the polar angle and is between z and \mathbf{r} while their azimuthal angle θ lies on the xy -plane.

Recall that the relationship between Cartesian (x, y, z) and spherical co-ordinates is as follows:

$$x = r \sin \theta \cos \phi \quad y = r \sin \theta \sin \phi \quad z = r \cos \theta$$

Also, much like how a vector \mathbf{V} may be expressed in terms of its Cartesian components, it may also be expanded using spherical polar components:

$$V_x \hat{x} + V_y \hat{y} + V_z \hat{z} = \mathbf{V} = V_r \hat{r} + V_\theta \hat{\theta} + V_\phi \hat{\phi}$$

It is here that we notice the first change between Cartesian systems and spherical polar systems. For any pair of points in Cartesian space, the co-ordinate origin may be shifted so that \hat{x} , \hat{y} and \hat{z} always point in the same mutually perpendicular set of three directions.

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However, for the same two points their \hat{r} , $\hat{\theta}$ and $\hat{\phi}$ values do *not* point in the same direction. It is easy to see how the direction of \hat{r} changes, so use that to consider that $\hat{\theta}$ and $\hat{\phi}$ are drawn corresponding to \hat{r} in order to keep the three co-ordinates mutually perpendicular. This means the direction of at least two of the three, if not all three, is always different across two points.

1.1 The line element

The first question that arises, having made this observation, has to do with the line element ds —by how much does the line element change if an infinitesimal change dr is made in the orientation of r ?

We may reason as follows:

1. Any change in r is reflected as-is in the r -component of ds i.e. $ds_r = dr$.
2. Any change in the θ -component causes r to sweep an arc over an angle $d\theta$, an arc described by $r d\theta$ since the angle is small i.e. $ds_\theta = r d\theta$.
3. The new r now has a projection $r \sin \theta$ which, in case of any change in the ϕ -component, causes the projection $r \sin \theta$ to sweep an arc over an angle $d\phi$, an arc described by $r \sin \theta d\phi$ since $d\phi$ is a small angle i.e. $ds_\phi = r \sin \theta d\phi$.

Consequently the change ds is given by

$$\begin{aligned} ds &= ds_r \hat{r} + ds_\theta \hat{\theta} + ds_\phi \hat{\phi} \\ &= dr \hat{r} + r d\theta \hat{\theta} + r \sin \theta d\phi \hat{\phi} \end{aligned}$$

1.2 The area and volume elements

The area and volume elements are given by similar reasoning. First, the area element on the surface of a sphere is given by the product of the areas within a pair of curves that sweep the azimuthal angle and a pair of curves that sweep the polar angle. In other words,

$$d\mathbf{A} = ds_\theta ds_\phi = r^2 \sin \theta d\theta d\phi \hat{r}$$

where the \hat{r} indicates that the area is in the direction of \hat{r} , directed away from the centre of the sphere.

The volume element is given by the product of all three possible dimensions (since we are in three-dimensional space) which means

$$d\tau = ds_r ds_\theta ds_\phi = r^2 \sin \theta dr d\theta d\phi$$

1.3 Unit vectors

Next, let us make note of what the typical basis vectors would look in spherical co-ordinate systems. To find these we simply need to consider that $\mathbf{r} =$

$r \sin \theta \cos \phi \hat{x} + r \sin \theta \sin \phi \hat{y} + r \cos \theta \hat{z}$ and $\hat{r} = \sin \theta \cos \phi \hat{x} + \sin \theta \sin \phi \hat{y} + \cos \theta \hat{z}$ as we discussed before and compute each unit vector as follows:

$$\begin{aligned}\hat{i}_r &= \frac{\frac{\partial \mathbf{r}}{\partial r}}{\left| \frac{\partial \mathbf{r}}{\partial r} \right|} = \frac{\sin \theta \cos \phi \hat{x} + \sin \theta \sin \phi \hat{y} + \cos \theta \hat{z}}{\sqrt{\sin^2 \theta \cos^2 \phi + \sin^2 \theta \sin^2 \phi + \cos^2 \theta}} = \sin \theta \cos \phi \hat{x} + \sin \theta \sin \phi \hat{y} + \cos \theta \hat{z} \\ \hat{j}_\theta &= \frac{\frac{\partial \mathbf{r}}{\partial \theta}}{\left| \frac{\partial \mathbf{r}}{\partial \theta} \right|} = \frac{r (\cos \theta \cos \phi \hat{x} + \cos \theta \sin \phi \hat{y} - \sin \theta \hat{z})}{\sqrt{r^2 (\cos^2 \theta \cos^2 \phi + \cos^2 \theta \sin^2 \phi - \sin^2 \theta)}} = \cos \theta \cos \phi \hat{x} + \cos \theta \sin \phi \hat{y} - \sin \theta \hat{z} \\ \hat{k}_\phi &= \frac{\frac{\partial \mathbf{r}}{\partial \phi}}{\left| \frac{\partial \mathbf{r}}{\partial \phi} \right|} = \frac{r (-\sin \theta \sin \phi \hat{x} + \sin \theta \cos \phi \hat{y})}{\sqrt{r^2 (\sin^2 \theta \sin^2 \phi + \sin^2 \theta \cos^2 \phi)}} = -\sin \phi \hat{x} + \cos \phi \hat{y}\end{aligned}$$

These could have been rewritten in terms of \hat{r} , $\hat{\theta}$ and $\hat{\phi}$ as well if we had worked out in the same manner but in reverse:

$$\begin{aligned}\hat{i}_x &= \sin \theta \cos \phi \hat{r} + \cos \theta \cos \phi \hat{\theta} - \sin \phi \hat{\phi} \\ \hat{j}_y &= \sin \theta \sin \phi \hat{r} + \cos \theta \sin \phi \hat{\theta} + \cos \phi \hat{\phi} \\ \hat{k}_z &= \cos \theta \hat{r} - \sin \theta \hat{\theta}\end{aligned}$$

Another way of arriving at this, for example in case of \hat{i}_x , is to consider the dot product of that component with each of the other three i.e. $\hat{x} \cdot \hat{r} = \sin \theta \cos \phi$, $\hat{x} \cdot \hat{\theta} = \cos \theta \cos \phi$ and $\hat{x} \cdot \hat{\phi} = -\sin \phi$ so that, as in the equation for \hat{i}_x above, we get $\hat{i}_x = \sin \theta \cos \phi \hat{r} + \cos \theta \cos \phi \hat{\theta} - \sin \phi \hat{\phi}$. The other two work similarly.

We will use these results to understand some fundamental co-ordinate mathematics. But as one last step before we continue, let us put our results into a beautiful matrix just for fun:

$$\begin{bmatrix} \hat{i}_r \\ \hat{j}_\theta \\ \hat{k}_\phi \end{bmatrix} = \begin{bmatrix} \sin \theta \cos \phi & \sin \theta \sin \phi & \cos \theta \\ \cos \theta \cos \phi & \cos \theta \sin \phi & -\sin \theta \\ -\sin \phi & \cos \phi & 0 \end{bmatrix} \begin{bmatrix} \hat{i}_x \\ \hat{j}_y \\ \hat{k}_z \end{bmatrix}$$

1.4 The gradient

The gradient of a scalar (and divergence of a vector) depend on operators of the form $\partial/\partial x$ in three dimensions. We can work these out as follows:

$$\begin{aligned}\frac{\partial}{\partial x} &= \frac{\partial r}{\partial x} \frac{\partial}{\partial r} + \frac{\partial \theta}{\partial x} \frac{\partial}{\partial \theta} + \frac{\partial \phi}{\partial x} \frac{\partial}{\partial \phi} \\ &= \left(\frac{x}{\sqrt{x^2 + y^2 + z^2}} \right) \frac{\partial}{\partial r} + \left(\frac{xz}{(x^2 + y^2 + z^2) \sqrt{x^2 + y^2}} \right) \frac{\partial}{\partial \theta} + \left(\frac{-y}{x^2 + y^2} \right) \frac{\partial}{\partial \phi} \\ \therefore \frac{\partial}{\partial x} &= (\sin \theta \cos \phi) \frac{\partial}{\partial r} + \left(\frac{\cos \theta \cos \phi}{r} \right) \frac{\partial}{\partial \theta} - \left(\frac{\sin \phi}{r \sin \theta} \right) \frac{\partial}{\partial \phi}\end{aligned}$$

Similarly we can work out

$$\frac{\partial}{\partial y} = (\sin \theta \sin \phi) \frac{\partial}{\partial r} + \left(\frac{\cos \theta \sin \phi}{r} \right) \frac{\partial}{\partial \theta} + \left(\frac{\cos \phi}{r \sin \theta} \right) \frac{\partial}{\partial \phi}$$

and $\frac{\partial}{\partial z} = (\cos \theta) \frac{\partial}{\partial r} - \left(\frac{\sin \theta}{r} \right) \frac{\partial}{\partial \theta}$

We are now in a position to define the gradient operator using $\hat{x} \equiv \hat{i}_x$ etc. from our previous discussion as follows:

$$\nabla = \hat{x} \frac{\partial}{\partial x} + \hat{y} \frac{\partial}{\partial y} + \hat{z} \frac{\partial}{\partial z}$$

where

$$\hat{x} \frac{\partial}{\partial x} = \left[\sin \theta \cos \phi \hat{r} + \cos \theta \cos \phi \hat{\theta} - \sin \phi \hat{\phi} \right] \left[(\sin \theta \cos \phi) \frac{\partial}{\partial r} + \left(\frac{\cos \theta \cos \phi}{r} \right) \frac{\partial}{\partial \theta} - \left(\frac{\sin \phi}{r \sin \theta} \right) \frac{\partial}{\partial \phi} \right]$$

and we similarly have expressions for the y and z components too. Solving this gives us the gradient operator as

$$\nabla = \hat{r} \frac{\partial}{\partial r} + \frac{1}{r} \hat{\theta} \frac{\partial}{\partial \theta} + \frac{1}{r \sin \theta} \hat{\phi} \frac{\partial}{\partial \phi} \quad (1)$$

1.5 The divergence

Out of interest and curiosity, we may employ an alternate method to compute the divergence. In our discussion of the volume element $d\tau$ we combined the three components ds_r , ds_θ and ds_ϕ . Normally these are also labelled h_1 , h_2 and h_3 so that

$$h_1 = 1 \quad h_2 = r \quad h_3 = r \sin \theta$$

Additionally the co-ordinates are also labelled q_1 , q_2 and q_3 so that

$$q_1 = r \quad q_2 = \theta \quad q_3 = \phi$$

This allows for direct substitution into the general divergence formula:

$$\begin{aligned} \nabla \cdot \mathbf{V} &= \frac{1}{h_1 h_2 h_3} \left[\frac{\partial}{\partial q_1} (V_1 h_2 h_3) + \frac{\partial}{\partial q_2} (V_2 h_3 h_1) + \frac{\partial}{\partial q_3} (V_3 h_1 h_2) \right] \\ \therefore \nabla \cdot \mathbf{V} &= \frac{1}{r^2 \sin \theta} \left[\sin \theta \frac{\partial}{\partial r} (r^2 V_r) + r \frac{\partial}{\partial \theta} (\sin \theta V_\theta) + r \frac{\partial V_\phi}{\partial \phi} \right] \end{aligned} \quad (2)$$

The use of h_i and q_i to write general forms of co-ordinate equations is also known as writing the gradient, divergence, laplacian and curl in 'arbitrary co-ordinates' since the cylindrical and Cartesian forms too can be written based on these formulae.

Observe that the general formula

$$\nabla = \sum_i \hat{q}_i \frac{1}{h_i} \frac{\partial}{\partial q_i}$$

could have helped us arrive at eq. (1) too but it was important that we worked at least one of them out from first principles. Also, while we speak of the divergence it is also worth noting the scalar Laplacian:

$$\begin{aligned} \nabla^2 &= \frac{1}{h_1 h_2 h_3} \left[\frac{\partial}{\partial q_1} \left(\frac{h_2 h_3}{h_1} \frac{\partial}{\partial q_1} \right)_{\text{cyc}} \right] \\ \therefore \nabla^2 &= \frac{1}{r^2 \sin \theta} \left[\sin \theta \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) + \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin \theta} \frac{\partial^2}{\partial \phi^2} \right] \end{aligned} \quad (3)$$

1.6 The curl

The curl too can be arrived at using the general formula:

$$\begin{aligned} \nabla \times \mathbf{V} &= \frac{1}{h_1 h_2 h_3} \begin{vmatrix} \hat{q}_1 h_1 & \hat{q}_2 h_2 & \hat{q}_3 h_3 \\ \frac{\partial}{\partial q_1} & \frac{\partial}{\partial q_2} & \frac{\partial}{\partial q_3} \\ h_1 V_1 & h_2 V_2 & h_3 V_3 \end{vmatrix} \\ \therefore \nabla \times \mathbf{V} &= \frac{1}{r^2 \sin \theta} \begin{vmatrix} \hat{r} & r \hat{\theta} & r \sin \theta \hat{\phi} \\ \frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & \frac{\partial}{\partial \phi} \\ V_r & r V_\theta & r \sin \theta V_\phi \end{vmatrix} \end{aligned} \quad (4)$$

These discussions have set us up to properly discuss relativity. A familiarity with them will help us better understand general relativity without letting co-ordinates trip us up.