

Homework 2 Answer Key

Problem 1

Given:

Component	Value (meters)
Position X	326151.080726
Position Y	6077471.251787
Position Z	2944583.918767
Velocity X	-7455.178720
Velocity Y	-482.482572
Velocity Z	1910.883434

MATLAB solution:

Week 2 Homework Solutions

VECTOR 1

```
-----
r:      326151.080726   6077471.251787   2944583.918767  meters
rd:     -7455.178720    -482.482572     1910.883434 m/sec

h:      1.3034049558e+10 -2.2575636068e+10 4.5151272135e+10 m^2/sec

|h|: 5.2136198243e+10 m^2/sec

p:      6819317.99903984 m

E:      -2.9222906294e+07 m^2/s^2

a:      6819999.99902560 m

TP:      5605.15391191 sec

B:      2.1265080535e+12 3.2207421496e+12 9.9650107635e+11

|B|: 3.9860043767e+12

e:      0.0099999999

nu:      0.5337080028 rad
nu:      30.5792160524 deg
Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2
E0:      0.5286424011 rad
E0:      30.2889784571 deg
M:      0.5235987859 rad
n:      0.0011209657 rad/sec
```

```

dt:      467.0961685125 sec
r65: 6790619.6008190252 m
E65:      1.1254198828 rad
E65:      64.4818094624 deg
dt:      528.8267149214 sec

```

Kepler Equation Solutions

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Echeck:      1.1254198828 rad
Echeck:      64.4818094624 deg
nuCheck:     1.1344640138 rad
nuCheck:     65.0000000000 deg

```

```

E2700:      3.5462689773 rad
E2700:      203.1862454165 deg
Perigee crossings after 2700 seconds: 0

```

```

nu2700:     3.5423496855 rad
nu2700:     202.9616865415 deg

```

```

E2TP:      13.0950130155 rad
E2TP:      750.2889784571 deg
Perigee crossings after 1.121031e+04 seconds: 2

```

```

In range [0, 2*pi]
E2TP:      0.5286424011 rad
E2TP:      30.2889784571 deg
nu2TP:     0.5337080028 rad
nu2TP:     30.5792160524 deg

```

```

E15000:     17.3280965310 rad
E15000:     992.8267982226 deg
Perigee crossings after 15000 seconds: 2

```

```

In range [0, 2*pi]
E15000:     4.7617259167 rad
E15000:     272.8267982226 deg
nu15000:    4.7517354548 rad
nu15000:    272.2543869211 deg

```

1. Find v_0

From the Trajectory equation (Vallado Eq. 1-23), the direction of perigee is given by \vec{B} , and the true anomaly is computed from

$$\vec{r} \cdot \vec{B} = r B \cos v_0$$

Solving for v_0 :

$$\begin{aligned}\vec{B} &= [2126508042101.570312, 3220742199864.750000, 996501104661.750000] \\ \vec{r} &= [326151.080726 \quad 6077471.251787 \quad 2944583.918767]\end{aligned}$$

$$v_0 = 0.533708 \text{ radians}$$

30.579215 degrees

e = 0.01

2. Find E₀

$$\sin E_0 = \frac{\sin v_0 \sqrt{1 - e^2}}{(1 + e \cos v_0)}$$

$$\cos E_0 = \frac{e + \cos v_0}{1 + e \cos v_0}$$

E₀ = atan2(sin E₀, cos E₀) → Note use of ATAN2 to return proper quadrant

E₀ = 0.528642 radians
30.288978 degrees

3. M₀ = E₀ - e sin E₀

M₀ = 0.528642 - 0.01 sin(0.528642)

M₀ = 0.523599 radians
30.000000 degrees

Find Time of Flight from Perigee to v₀

$$n = \sqrt{\frac{\mu}{a^3}} = 0.001121 \text{ radians/sec}$$

$$n(t - T_0) = n \Delta t = M_0$$

So

$$\Delta t = \frac{M_0}{n} = 467.096 \text{ sec}$$

Find Time of Flight from v₀ to v = 65°

From Vallado (eq 2-7)

$$t - t_0 = \sqrt{\frac{a^3}{\mu}} \{2\pi k + (E - e \sin E) - (E_0 - e \sin E_0)\}$$

Where k is the number of perigee passages between t₀ and t. In this case, k=0.

As above,

$$M_0 = E_0 - e \sin E_0 = 0.523599 \text{ radians}$$

Find E in the same manner as above,

cos E	0.430798
sin E	0.902449
E	1.125420 radians 64.481809 degrees

M = E - e sinE	1.116395 radians 63.964745 degrees
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$$t - t_0 = \frac{1}{0.001121} \{0 + 1.116395 - 0.523599\} = 528.826724 \text{ sec}$$

Kepler Equation Solutions

The Kepler equation is an iterative solution to find the eccentric anomaly after a specified time of flight, starting at an initial eccentric anomaly, E_o . The iteration algorithm is as follows.

Given an initial eccentric anomaly E_o and time of flight Δt , define

$$M_o = E_o - e \sin E_o$$

The general iteration equation is then

$$E_{k+1} = E_k + \frac{n\Delta t + M_0 - (E_k - e \sin E_k)}{1 - e \cos E_k}$$

As written, the iteration will return an eccentric anomaly that is not bounded within the range $[0, 2\pi]$. In Matlab, to express the solved-for value within that range, use the modulo function:

$$E_{\text{bounded}} = \text{mod}(E, 2\pi)$$

The number of perigee crossings between E_o and the final time of flight point can be computed in MATLAB as

$$k = \text{floor}((E - E_o) / 2\pi)$$

1. Verify: Starting v_0 , compute the true anomaly after the time of flight computed above.

$$\begin{aligned}\Delta t &= 528.8267149214 \text{ sec} \\ E_o &= 0.5286424011 \text{ rad}\end{aligned}$$

After iteration,

E_{check}: 1.1254198828 rad
E_{check}: 64.4818094624 deg
v_{check}: 65.0 deg

This is the expected result.

2. Starting at the same E_o, Find the true anomaly after a time of flight of 2700 seconds. After iteration:

E₂₇₀₀: 3.5462689773 rad
E₂₇₀₀: 203.1862454165 deg
nu₂₇₀₀: 3.5423496855 rad
nu₂₇₀₀: 202.9616865415 deg

3. Starting at the same E_o, Find the true anomaly after exactly two orbit periods?

E_{2TP}: 13.0950130155 rad
E_{2TP}: 750.2889784571 deg
Perigee crossings after 1.121031e+04 seconds: 2 (as expected)

In range [0, 2*pi]
E_{2TP}: 0.5286424011 rad
E_{2TP}: 30.2889784571 deg
v_{2TP}: 0.5337080028 rad
v_{2TP}: 30.5792160524 deg

As expected, the results are exactly the same as the initial conditions.

4. Starting at the same E_o, Find the true anomaly after a time of flight of 15000 seconds. Because the Δt for this case exceeds the orbit period, this iteration solution should have multiple perigee crossings.

E₁₅₀₀₀: 17.3280965310 rad
E₁₅₀₀₀: 992.8267982226 deg
Perigee crossings after 15000 seconds: 2

In range [0, 2*pi]
E₁₅₀₀₀: 4.7617259167 rad
E₁₅₀₀₀: 272.8267982226 deg
nu₁₅₀₀₀: 4.7517354548 rad
nu₁₅₀₀₀: 272.2543869211 deg

Problem 2

Given:

Component	Value (meters)
Position X	572461.711228
Position Y	-1015437.194396
Position Z	7707337.871302
Velocity X	-6195.262945
Velocity Y	-3575.889650
Velocity Z	-5.423283

1. Find perifocal position and velocity, \vec{r}_0 and \vec{v}_0

The full set of Keplerian elements for this position and velocity is:

```
a:      7800000.00120126 m
e:      0.00100000
inc:    98.60000000 deg
raan:   30.00000000 deg
wp:     40.00000696 deg
nu:     50.08784582 deg
```

Also, the perifocal position and velocity are computed from the above using,

$$\begin{aligned}x &= r \cos v \\y &= r \sin v \\\dot{x} &= -\sqrt{\frac{\mu}{p}} \sin v \\\dot{y} &= \sqrt{\frac{\mu}{p}} (e + \cos v) \\p &= a(1 - e^2) = 7799992.2\end{aligned}$$

Plugging in the known values, the perifocal position and velocity are given by:

$$\begin{aligned}\vec{r}_0 &= [5001362.438709 \quad 5978984.522949 \quad 0.000000] \\ \vec{v}_0 &= [-5483.194150 \quad 4593.787225 \quad 0.000000]\end{aligned}$$

Note that the Z-components of position and velocity are zero, as the position and velocity vectors in the perifocal frame are by definition in the orbit plane.

Find f, g, \dot{f} , \dot{g} for $\Delta v = 33^\circ$.

Using the equations from Vallado (2-63)

f:	0.838690
g:	593.813828
fdot:	-0.000499
gdot:	0.838774

2. Find \vec{r} and \vec{v} corresponding to Δv

Using:

$$\vec{r} = f \vec{r}_0 + g \vec{v}_0$$

$$\vec{v} = \dot{f} \vec{r}_0 + \dot{g} \vec{v}_0$$

$$\begin{aligned}\vec{r} &= [938596.059563 \quad 7742368.795881 \quad 0.000000] \\ \vec{v} &= [-7096.655979 \quad 867.465852 \quad 0.000000]\end{aligned}$$

Note: you can check your work for this by computing the vectors directly:

Compute $v=v_0 + \Delta v$ 83.087853 degrees

Compute r using the ellipse equation, then apply the position and velocity equations above.

3. Using energy equation and ellipse equation, find an expression for the orbit speed as a function of true anomaly

Energy equation:

$$\xi = -\frac{\mu}{2a} = \frac{v^2}{2} - \frac{\mu}{r}$$

Rearrange:

$$\frac{\mu}{r} - \frac{\mu}{2a} = \frac{v^2}{2}$$

Multiply through by 2

$$v^2 = \frac{2\mu}{r} - \frac{\mu}{a}$$

Recall from the ellipse equation that

$$r = \frac{a(1-e^2)}{1+e\cos v}$$

Substitute into the energy equation for r

$$v^2 = \frac{2\mu(1+e\cos v)}{a(1-e^2)} - \frac{\mu}{a}$$

Collect terms

$$v^2 = \frac{\mu}{a} \left(\frac{2 + 2e\cos v}{(1-e^2)} - 1 \right)$$

Find a common denominator

$$v^2 = \frac{\mu}{a} \left(\frac{2 + 2e\cos v - (1-e^2)}{(1-e^2)} \right)$$

Simplify

$$v^2 = \frac{\mu}{a} \left(\frac{1 + 2e\cos v + e^2}{(1-e^2)} \right)$$

Factor the denominator and take square root

$$v = \sqrt{\frac{\mu}{a} \left(\frac{1 + 2e\cos v + e^2}{(1-e)(1+e)} \right)}$$

4. Show that

$$v_{\text{perigee}} = \sqrt{\frac{\mu}{a} \left(\frac{1+e}{1-e} \right)}$$

$$v_{\text{apogee}} = \sqrt{\frac{\mu}{a} \left(\frac{1-e}{1+e} \right)}$$

At perigee, $\nu=0^\circ$, so $\cos\nu=1$, and

$$v_{\text{perigee}} = \sqrt{\frac{\mu}{a} \left(\frac{1+2e+e^2}{(1-e)(1+e)} \right)} = \sqrt{\frac{\mu}{a} \left(\frac{(1+e)^2}{(1-e)(1+e)} \right)} = \sqrt{\frac{\mu}{a} \left(\frac{(1+e)}{(1-e)} \right)}$$

At apogee, $\nu=180^\circ$, so $\cos\nu=-1$, and

$$v_{\text{apogee}} = \sqrt{\frac{\mu}{a} \left(\frac{1-2e+e^2}{(1-e)(1+e)} \right)} = \sqrt{\frac{\mu}{a} \left(\frac{(1-e)^2}{(1-e)(1+e)} \right)} = \sqrt{\frac{\mu}{a} \left(\frac{(1-e)}{(1+e)} \right)}$$

QED

MATLAB Solutions for Homework 2

Problem 1:

MATLAB code for this problem was adapted and extended from the code used for Problem 1 of the Week 1 homework. Additional code was added to achieve the HW2 results.

Results

Week 2 Homework Solutions

```
VECTOR 1
-----
r:      326151.080726   6077471.251787   2944583.918767   meters
rd:     -7455.178720      -482.482572      1910.883434 m/sec

h:      1.3034049558e+10  -2.2575636068e+10  4.5151272135e+10 m^2/sec

|h|: 5.2136198243e+10 m^2/sec

p:      6819317.99903984 m

E:      -2.9222906294e+07 m^2/s^2

a:      6819999.99902560 m

TP:      5605.15391191 sec

B:      2.1265080535e+12  3.2207421496e+12  9.9650107635e+11

|B|: 3.9860043767e+12

e:      0.0099999999

nu:      0.5337080028 rad
nu:      30.5792160524 deg
Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2
E0:      0.5286424011 rad
E0:      30.2889784571 deg
M:      0.5235987859 rad
n:      0.0011209657 rad/sec
dt:      467.0961685125 sec
r65: 6790619.6008190252 m
sin(E65): 0.9024485581 rad
cos(E65): 0.4307976322 rad
E65: 1.1254198828 rad
E65: 64.4818094624 deg
dt: 528.8267149214 sec
```

Kepler Equation Solutions

```
Echeck: 1.1254198828 rad
Echeck: 64.4818094624 deg
```

```

nuCheck:      1.1344640138 rad
nuCheck:      65.0000000000 deg

E2700:       3.5462689773 rad
E2700:       203.1862454165 deg
Perigee crossings after 2700 seconds: 0
nu2700:      3.5423496855 rad
nu2700:      202.9616865415 deg

E2TP:        13.0950130155 rad
E2TP:        750.2889784571 deg
Perigee crossings after 1.121031e+04 seconds: 2

In range [0, 2*pi]
E2TP:        0.5286424011 rad
E2TP:        30.2889784571 deg
nu2TP:       0.5337080028 rad
nu2TP:       30.5792160524 deg

E15000:      17.3280965310 rad
E15000:      992.8267982226 deg
Perigee crossings after 15000 seconds: 2

In range [0, 2*pi]
E15000:      4.7617259167 rad
E15000:      272.8267982226 deg
nu15000:     4.7517354548 rad
nu15000:     272.2543869211 deg

```

Week 2 Problem 2 Homework Solutions

VECTOR 2

```

-----
R0:      572461.711228 -1015437.194396    7707337.871302 meters
R0d:     -6195.262945      -3575.889650      -5.423283 m/sec

h:      2.7566096726e+10 -4.7745880097e+10 -8.3379603316e+09 m^2/sec

|h|: 5.5759127840e+10 m^2/sec

p:      7799992.201199780 m

E:      -2.5551310368e+07 m^2/s^2

a:      7800000.00120126 m

TP:      6855.71704370 sec

B:      2.8359372629e+11 1.1949255982e+11 2.5333469724e+11

|B|: 3.9860047952e+11

e:      0.0010000001

```

nu: 0.8741978248 rad
nu: 50.0878458185 deg

Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2
R0p: 5001362.438709 5978984.522949 0.000000 meters
V0p: -5483.194150 4593.787225 0.000000 m/sec

f: 0.8386899812 deg
g: 593.8138283683 deg
gdot: 0.8387740125 deg
fdot1: -0.0004993630 deg
fdot2: -0.0004993630 deg

Rp: 938596.059564 7742368.795881 0.000000 meters
Vp: -7096.655979 867.465852 0.000000 m/sec

Test
Rp: 938596.059564 7742368.795881 0.000000 meters
Vp: -7096.655979 867.465852 0.000000 m/sec

MATLAB Code for Problem 1

```
fid = fopen('c:\temp\HW2results_2024.txt','wt');

fprintf(fid,'Week 2 Homework Solutions\n\n');

mu = 3.986004418e14;

fprintf(fid, 'VECTOR 1-----\n');
r = [326151.080726; 6077471.251787; 2944583.918767];
rd = [-7455.178720; -482.482572; 1910.883434];

fprintf(fid,'r: %16.6f %16.6f %16.6f meters\n', r(1), r(2), r(3));
fprintf(fid,'rd: %16.6f %16.6f %16.6f m/sec\n', rd(1), rd(2), rd(3));
fprintf(fid,'\n');

rnorm = norm(r);

h=cross(r,rd);
fprintf(fid,'h: %16.10e %16.10e %16.10e m^2/sec\n', h(1), h(2), h(3));
fprintf(fid,'\n');

hnorm = norm(h);
fprintf(fid,'|h|: %16.10e m^2/sec\n', hnorm);
fprintf(fid,'\n');

p = hnorm^2/mu;
fprintf(fid,'p: %16.8f m\n', p);
fprintf(fid,'\n');

energy = dot(rd,rd)/2 - mu/rnorm;
fprintf(fid,'E: %16.10e m^2/s^2\n', energy);
fprintf(fid,'\n');

a = -mu/(2*energy);
fprintf(fid,'a: %16.8f m\n', a);
fprintf(fid,'\n');

TP=2*pi*sqrt(a^3/mu);
fprintf(fid,'TP: %16.8f sec\n', TP);
fprintf(fid,'\n');

B = cross(rd,h)-(mu/rnorm)*r;
fprintf(fid,'B: %16.10e %16.10e %16.10e \n', B(1), B(2), B(3));
fprintf(fid,'\n');

Bnorm = norm(B);

fprintf(fid,'|B|: %16.10e \n', Bnorm);
fprintf(fid,'\n');

e = Bnorm/mu;
fprintf(fid,'e: %16.10f \n', e);
fprintf(fid,'\n');

cosnu = dot(r,B)/(rnorm*Bnorm);
nu = acos(cosnu);
fprintf(fid,'nu: %16.10f rad\n', nu);
nud = acosd(cosnu);
fprintf(fid,'nu: %16.10f deg\n', nud);

%Check sign
sign = dot(r, rd);
if(sign<0)
    fprintf(fid,'Dot product dot(r, rd) < 0. Sign is negative. Quadrant 3 or 4\n');
else
    fprintf(fid,'Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2\n');
end

%Adjust for quadrant if sign is negative
if sign < 0.
```

```

nu = 2*pi - nu;
fprintf(fid,'nu:      %16.8f (adjusted) radians\n', nu);
nud = 360. - nud;
fprintf(fid,'nud:      %16.8f (adjusted) degrees\n', nud);
end

% Begin computing initial eccentric anomaly
sinE = sin(nu)*sqrt(1-e^2)/(1 + e*cos(nu));
cosE = (e+cos(nu))/(1+e*cos(nu));

% Use ATAN2 to get E0 in range of [-pi,pi]
E0 = atan2(sinE, cosE);

% Logic to put it in range [0,2pi]
if(E0<0)
    E0 = E0 + 2*pi;
end

fprintf(fid,'E0:  %16.10f rad\n', E0);
fprintf(fid,'E0:  %16.10f deg\n', E0*180/pi);

% Initial mean anomaly
M0 = E0 - e*sin(E0);
fprintf(fid,'M:  %16.10f rad\n', M0);

% Mean motion
n = sqrt(mu/a^3);
fprintf(fid,'n:  %16.10f rad/sec\n', n);

% Time of flight from perigee to initial NU
dt_per = M0/n;
fprintf(fid,'dt:  %16.10f sec\n', dt_per);

% Now compute radius, eccentric anomaly, and other
% parameters for nu = 65 degrees
nu65 = 65*pi/180.;

% Need to compute radius of orbit at nu=65 deg
r65 = a*(1-e^2)/(1+e*cos(nu65));
fprintf(fid,'r65: %16.10f m\n', r65);

% Compute eccentric anomaly for nu=65
% sinE65 = r65*sin(nu65)/(a*sqrt(1-e^2));
sinE65 = sin(nu65)*sqrt(1-e^2)/(1 + e*cos(nu65));
cosE65 = (e+cos(nu65))/(1+e*cos(nu65));

% Use ATAN2 to get E65 in range of [-pi,pi]
E65 = atan2(sinE65, cosE65);

% Logic to put it in range [0,2pi]
if(E65<0)
    E65 = E65 + 2*pi;
end
fprintf(fid,'sin(E65): %16.10f rad\n', sinE65);
fprintf(fid,'cos(E65): %16.10f rad\n', cosE65);

fprintf(fid,'E65: %16.10f rad\n', E65);
fprintf(fid,'E65: %16.10f deg\n', E65*180/pi);

% compute time of flight from initial nu, to nu=65 deg
% Here we assume no perigee crossing between the two points
dt65 = 1/n*(E65-e*sin(E65))-1/n*(E0-e*sin(E0));
fprintf(fid,'dt:  %16.10f sec\n', dt65);

fprintf(fid,'\n\nKepler Equation Solutions\n\n');

% Solution using time of flight computed above
dt = dt65;
Echeck = Kepler_Equation(mu, a, e, M0, dt);
fprintf(fid,'Echeck:  %16.10f rad\n', Echeck);
fprintf(fid,'Echeck:  %16.10f deg\n', Echeck*180/pi);

```

```

nuCheck = nu_from_E(Echeck, e);
fprintf(fid,'nuCheck: %16.10f rad\n', nuCheck);
fprintf(fid,'nuCheck: %16.10f deg\n', nuCheck*180/pi);

% Solution using time of flight = 2700 seconds
dt = 2700;
E2700 = Kepler_Equation( mu, a, e, M0, dt);
fprintf(fid,'\nE2700: %16.10f rad\n', E2700);
fprintf(fid,'E2700: %16.10f deg\n', E2700*180/pi);

numPerigeeCrossings = floor((E2700 - E0) / (2*pi));
fprintf(fid,'Perigee crossings after %d seconds: %d\n', dt, numPerigeeCrossings);

nu2700 = nu_from_E(E2700, e);
fprintf(fid,'nu2700: %16.10f rad\n', nu2700);
fprintf(fid,'nu2700: %16.10f deg\n', nu2700*180/pi);

% Solution using time of flight = exactly two orbit periods
dt = 2 * TP;
E2TP = Kepler_Equation( mu, a, e, M0, dt);
fprintf(fid,'\nE2TP: %16.10f rad\n', E2TP);
fprintf(fid,'E2TP %16.10f deg\n', E2TP*180/pi);

numPerigeeCrossings = floor((E2TP - E0) / (2*pi));
fprintf(fid,'Perigee crossings after %d seconds: %d\n', dt, numPerigeeCrossings);

fprintf(fid,'\nIn range [0, 2*pi]\n');
E2TP = mod(E2TP, 2*pi);
fprintf(fid,'E2TP: %16.10f rad\n', E2TP);
fprintf(fid,'E2TP: %16.10f deg\n', E2TP*180/pi);

nu2TP = nu_from_E(E2TP, e);
fprintf(fid,'nu2TP: %16.10f rad\n', nu2TP);
fprintf(fid,'nu2TP: %16.10f deg\n', nu2TP*180/pi);

% Solution using time of flight = 1500 seconds
dt = 15000;
E15000 = Kepler_Equation( mu, a, e, M0, dt);
fprintf(fid,'\nE15000: %16.10f rad\n', E15000);
fprintf(fid,'E15000: %16.10f deg\n', E15000*180/pi);

numPerigeeCrossings = floor((E15000 - E0) / (2*pi));
fprintf(fid,'Perigee crossings after %d seconds: %d\n', dt, numPerigeeCrossings);

fprintf(fid,'\nIn range [0, 2*pi]\n');
E15000 = mod(E15000, 2*pi);
fprintf(fid,'E15000: %16.10f rad\n', E15000);
fprintf(fid,'E15000: %16.10f deg\n', E15000*180/pi);

nu15000 = nu_from_E(E15000, e);
fprintf(fid,'nu15000: %16.10f rad\n', nu15000);
fprintf(fid,'nu15000: %16.10f deg\n', nu15000*180/pi);

```

Solution of Kepler Equation

```
function [ E1 ] = Kepler_Equation( mu, a, e, M0, dt)
% Solves Kepler's equation to find E1 given an initial MA and dt

n = sqrt(mu/a^3); % Mean motion
tolerance = 1e-10;
update = 100000; % Initialize to a big number

Ek = M0; % As a first estimate, set Ek to input mean anomaly

iteration = 0;

% The logic uses the absolute value of the update, which could be positive
% or negative.
while (abs(update)>tolerance && (iteration<=10));

% Newtonian iteration has the following form:
%
%     Ek+1 = Ek - f(E)/f' (E)
%
% Function and derivative are as shown (see, e.g., Vallado sect 2.2.5)
f=n*dt+M0-(Ek-e*sin(Ek));
fp= -(1-e*cos(Ek));

% We compute update as a separate variable for use in the "While" logic
update = f/fp;

% Compute updated estimate
Ekp1=Ek-update;
iteration = iteration+1;

% Set up Ek for the next iteration
Ek=Ekp1;
end

% Set the return value to the converged
E1=Ekp1;
end
```

Compute True Anomaly from Eccentric Anomaly

```
function [ nu ] = nu_from_E( E, e)
% Compute nu given E and e
% E - eccentric anomaly in RADIANS
cosnu = (cos(E)-e)/(1-e*cos(E));
sinu = sin(E)*sqrt(1-e^2)/(1-e*cos(E));

nu = atan2(sinu, cosnu);
if(nu<0)
    nu = nu + 2*pi;
end

end
```

Problem 2:

MATLAB code for this problem was adapted and extended from the code used for Problem 2 of the Week 1 homework. Additional code was added to achieve the HW2 results.

Results

Week 2 Homework Solutions

VECTOR 2

```
-----
R0:      572461.711228   -1015437.194396    7707337.871302  meters
R0d:     -6195.262945      -3575.889650      -5.423283 m/sec

Keplerian Elements for Vector 2
a:      7800000.00120126 m
e:          0.00100000
inc:      98.60000000 deg
raan:      30.00000000 deg
wp:       40.00000696 deg
nu:       50.08784582 deg

h:      2.7566096726e+10  -4.7745880097e+10  -8.3379603316e+09 m^2/sec

|h|: 5.5759127840e+10 m^2/sec

p:      7799992.201199780 m

E:      -2.5551310368e+07 m^2/s^2

a:      7800000.00120126 m

TP:      6855.71704370 sec

B:      2.8359372629e+11  1.1949255982e+11  2.5333469724e+11

|B|: 3.9860047952e+11

e:          0.0010000001

nu:       0.8741978248 rad
nu:       50.0878458185 deg

Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2
R0p:      5001362.438709   5978984.522949      0.000000  meters
V0p:     -5483.194150      4593.787225      0.000000 m/sec

f:          0.8386899812 deg
g:          593.8138283683 deg
gdot:        0.8387740125 deg
fdot1:      -0.0004993630 deg
fdot2:      -0.0004993630 deg

Rp:      938596.059564    7742368.795881      0.000000  meters
Vp:     -7096.655979      867.465852      0.000000 m/sec

Test
Rp:      938596.059564    7742368.795881      0.000000  meters
Vp:     -7096.655979      867.465852      0.000000 m/sec
```

MATLAB Code for Problem 2

```
fprintf(fid, '\n\nWeek 2 Homework Solutions\n\n');

mu = 3.986004418e14;
fprintf(fid, '\n\nVECTOR 2\n-----\n');
R0=[572461.711228; -1015437.194396; 7707337.871302];
R0d=[-6195.262945; -3575.889650; -5.423283];

fprintf(fid,'R0: %16.6f %16.6f %16.6f meters\n', R0(1), R0(2), R0(3));
fprintf(fid,'R0d: %16.6f %16.6f %16.6f m/sec\n', R0d(1), R0d(2), R0d(3));
fprintf(fid, '\n');

% Compute the Keplerian elements
[a, e, inc, raan, wp, nu] = KeplerianElements(mu, R0, R0d);

WriteKeplerianElements(fid, 2, a, e, inc, raan, wp, nu);
r0 = norm(R0);

h=cross(R0,R0d);
fprintf(fid, '\n');
fprintf(fid,'h: %16.10e %16.10e %16.10e m^2/sec\n', h(1), h(2), h(3));
fprintf(fid, '\n');

hnorm = norm(h);
fprintf(fid,'|h|: %16.10e m^2/sec\n', hnorm);
fprintf(fid, '\n');

p = hnorm^2/mu;
fprintf(fid,'p: %16.9f m\n', p);
fprintf(fid, '\n');

energy = dot(R0d,R0d)/2 - mu/r0;
fprintf(fid,'E: %16.10e m^2/s^2\n', energy);
fprintf(fid, '\n');

a = -mu/(2*energy);
fprintf(fid,'a: %16.8f m\n', a);
fprintf(fid, '\n');

TP=2*pi*sqrt(a^3/mu);
fprintf(fid,'TP: %16.8f sec\n', TP);
fprintf(fid, '\n');

B = cross(R0d,h)-(mu/r0)*R0;
fprintf(fid,'B: %16.10e %16.10e %16.10e \n', B(1), B(2), B(3));
fprintf(fid, '\n');
Bnorm = norm(B);

fprintf(fid,'|B|: %16.10e \n', Bnorm);
fprintf(fid, '\n');

e = Bnorm/mu;
fprintf(fid,'e: %16.10f \n', e);
fprintf(fid, '\n');

cosnu = dot(R0,B)/(r0*Bnorm);
nu = acos(cosnu);
fprintf(fid,'nu: %16.10f rad\n', nu);
nud = acosd(cosnu);
fprintf(fid,'nu: %16.10f deg\n', nud);
fprintf(fid, '\n');

%Check sign
sign = dot(R0, R0d);
if(sign<0)
    fprintf(fid,'Dot product dot(r, rd) < 0. Sign is negative. Quadrant 3 or 4\n');
else
```

```

        fprintf(fid,'Dot product dot(r, rd) > 0. Sign is positive. Quadrant 1 or 2\n');
end

%Adjust for quadrant if sign is negative
if sign < 0.
    nu = 2*pi - nu;
    fprintf(fid,'nu:      %16.8f (adjusted) radians\n', nu);
    nud = 360. - nud;
    fprintf(fid,'nud:      %16.8f (adjusted) degrees\n', nud);
end

% Compute perifocal vector corresponding to the input vector
x = r0*cos(nu);
y = r0*sin(nu);
xdot = -sqrt(mu/p)*sin(nu);
ydot = sqrt(mu/p)*(e+cos(nu));

% form into vectors r0 and v0
R0p=[x; y; 0.];
V0p=[xdot; ydot; 0.];
fprintf(fid,'R0p: %16.6f %16.6f %16.6f meters\n', R0p(1), R0p(2), R0p(3));
fprintf(fid,'V0p: %16.6f %16.6f %16.6f m/sec\n', V0p(1), V0p(2), V0p(3));
fprintf(fid,'\n');

% Find f, g, fdot, gdot for dnu=33 deg
dnu = 33*pi/180.;

% To find f, we need the radius at nu+33 deg
nuplus33=nu+dnu;
r = p/(1+e*cos(nuplus33));
f = 1-(r/p)*(1-cos(dnu));
g = (r*r0)/sqrt(mu*p)*sin(dnu);
gdot = 1-(r0/p)*(1-cos(dnu));

% compute fdot using both available methods
fdot1 = (f*gdot-1)/g;
fdot2 = sqrt(mu/p)*tan(dnu/2)*((1-cos(dnu))/p - 1/r - 1/r0);
fprintf(fid,'f:      %16.10f deg\n', f);
fprintf(fid,'g:      %16.10f deg\n', g);
fprintf(fid,'gdot:   %16.10f deg\n', gdot);
fprintf(fid,'fdot1:  %16.10f deg\n', fdot1);
fprintf(fid,'fdot2:  %16.10f deg\n', fdot2);
fprintf(fid,'\n');

% Now use the f and g functions to compute R and V at dnu=33 deg
% Remember to use the perifocal vectors R0p and V0p
Rp = f*R0p + g*V0p;
Vp = fdot1*R0p + gdot*V0p;
fprintf(fid,'Rp:  %16.6f %16.6f %16.6f meters\n', Rp(1), Rp(2), Rp(3));
fprintf(fid,'Vp:  %16.6f %16.6f %16.6f m/sec\n', Vp(1), Vp(2), Vp(3));
fprintf(fid,'\n');

% Check answers using the x, y, xdot, ydot equations
x = r*cos(nuplus33);
y = r*sin(nuplus33);
xdot = -sqrt(mu/p)*sin(nuplus33);
ydot = sqrt(mu/p)*(e+cos(nuplus33));

% form into vectors r0 and v0
Rp=[x; y; 0.];
Vp=[xdot; ydot; 0.];
fprintf(fid,'Test\n');
fprintf(fid,'Rp:  %16.6f %16.6f %16.6f meters\n', Rp(1), Rp(2), Rp(3));
fprintf(fid,'Vp:  %16.6f %16.6f %16.6f m/sec\n', Vp(1), Vp(2), Vp(3));

fclose(fid);

```