Introduction to Quantum Mechanics Summary

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Quantum Mechanics Summary created by michavan@student.ethz.ch

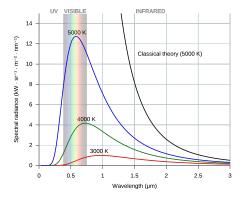
This summary has been written based on the Lecture 151-0966-00 L Introduction to Quantum Mechanics for Engineers by Prof. Dr. D. J. Norris (Spring 21s). There is no guarantee for completeness and/or correctness regarding the content of this summary. This summary is basically a modified and a more completed version of the summary of Lasse Fierz. Use it at your own discretion

Basics

Black Body Radiation and light duality:

A Black body is an ideal object that absorbes all light that hits it. \Rightarrow A Pinhole into a cavity is a good approximation. Classical physics predict: Infinite radiation at high ν at any $T \Rightarrow$ WRONG! Why does classical physics fail?

- 1. Oscillatory electrons in material emit light (electromagnetic waves)
- 2. Only certain standing waves fit in to cavity
- 3. Many nodes fit inside cavity at short wavelength \rightarrow high frequencies.



Note:

- 1. The emission peak shifts to higher frequency with increasing T.
- 2. The number of nodes per unit volume scales with $\nu^2 \Rightarrow$ at high frequencies Infinite nodes Finite energy = Infinite energy
- 3. Low frequency photon carry $k_b \cdot T$ energy
- 4. High frequency photon cannot radiate because $E_{high\nu} = h \cdot \nu >> k_b \cdot T$

Theorem: The Broglie Relation

An electron can act simultaneously as a particle and a wave at small lengths scales: $\lambda = \frac{h}{p} = \frac{h}{mv}$

A general wave in \mathbb{C} notation is written as: Ae^{ikx} , where $k = \frac{2\pi}{\lambda} = \frac{p}{\hbar}$.

Definitions

Avogadro:
$$N_A = 6.022 \cdot 10^{23} \ mol^{-1}$$

Planck: $h = 6.626 * 10^{-34} Js$, $\hbar = \frac{h}{2\pi}$
Boltzmann: $k_b = 1.38 * 10^{-23} \frac{J}{K}$

Wavefunction Ψ represents a state (of a particle) and provides statistical information:

 $|\Psi|^2$ Is the prob. of finding the particle at place x and time t. Prob. density, $\int_a^b |\Psi|^2 dx$ yields probability of finding the particle between a and b. If it's 0 or ∞ it is invalid. (Unphysical)

Normalize Ψ :

Find $A \in \mathbb{C}$ so that $\int_a^b |A\Psi|^2 dx = 1$ Note: The wavefunction, once normalized stays that way for all t.

Expectation value of Q: $\langle \hat{Q} \rangle = \int_{-\infty}^{\infty} \Psi^* \hat{Q} \Psi dx = \int_{-\infty}^{\infty} \hat{Q} |\Psi|^2 dx$

Theorem: Ehrenfest's theorem

Expectation values follow classical laws:

$$m\frac{d}{dt}\langle x\rangle = \langle p\rangle, \quad \frac{d}{dt}\langle p\rangle = -\langle V'(x)\rangle$$

$\text{Variance/Standard Deviation:} \ \frac{\sigma_{\hat{Q}}^2 = \langle \hat{Q}^2 \rangle - \langle \hat{Q} \rangle^2, \quad \sigma_{\hat{Q}} = \sqrt{\langle \hat{Q}^2 \rangle - \langle \hat{Q} \rangle^2}$

The variance of a discrete variable j is given by $\sum_{j=0}^{\infty} (j - \langle j \rangle)^2 P(j)$, with $P(j) \to P(x) dx \Rightarrow$ probability of finding particle in [x, x + dx].

Theorem: Heisenberg's uncertainty principle

Incompatible observables such as \hat{x} and \hat{p} or **energy and time** can not be measured simultaneously without uncertainty $(\Delta E = \sigma_H, \Delta t = \frac{T}{2\pi})$

$$\begin{cases} \sigma_x \sigma_p \geq \frac{\hbar}{2}, \Delta E \Delta t \leq \frac{\hbar}{2} \text{ or in general: } \sigma_A^2 \sigma_B^2 \geq (\frac{1}{2i} \langle [\hat{A}, \hat{B}] \rangle)^2 \\ \sigma_{\hat{H}}^2 \sigma_{\hat{Q}}^2 \geq (\frac{\hbar}{2})^2 (\frac{d}{dt} \langle \hat{Q} \rangle)^2 \quad \text{ and } \quad \frac{d \langle Q \rangle}{dt} = \frac{i}{\hbar} \langle \left[\hat{H}, \hat{Q} \right] \rangle + \langle \frac{\partial \hat{Q}}{\partial t} \rangle \end{cases}$$

Example: Rivisited Ideal Gas Law

 $pV = nRT \Rightarrow pV = Nk_bT$, N = number of molecules. $\frac{V}{N} = \frac{k_bT}{p} \Rightarrow d = (\frac{V}{N})^{1/3}$, d = diameter of molecules.

Operators (1D):

$$\hat{x} = x \quad | \quad \hat{p} = -i\hbar \frac{\partial}{\partial x} = \frac{\hbar}{i} \frac{\partial}{\partial x} \quad | \quad \hat{H} = -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + V(x) = \frac{\hat{p}^2}{2m} + V(x)$$

 \hat{H} being the operator of the <u>Hamiltonian</u> = sum of potential and kinetic energy (classical mechanics). V(x) is the potential energy. $F = -\nabla V(x)$

Good to know: We can allways approximate v(x) as a polynomial of x. (Locally it might look like a parabola even if the potential is arbitrairy)

Theory: O.E

If V(x)=0 \Rightarrow the energy consists only of kinetic energy $\to T=\frac{1}{2m}\langle p^2\rangle$. The kinetic energy of a wavefunction is related to its curvature.

Schrödinger Equation:

Time dependent Schrödinger Equation(TDSE):

$$i\hbar\frac{\partial\Psi}{\partial t} = -\frac{\hbar^2}{2m}\frac{\partial^2\Psi}{\partial x^2} + V(x,t)\Psi$$

For non-time-dependent potentials $V \neq V(t)$ we can use separation of variables $(\Psi(x,t)=\psi(x)\phi(t))$ and arrive at two parts (Attention: $\psi(x)$ must be continuous and so its first derivative as well as square integrable):

Time dependent part: $\varphi(t) = \exp(-iEt/\hbar)$ (most often the same) Second part is the solution of the Time Independent Schrödinger Equation (TISE) with the coefficient c_n :

$$-\frac{\hbar^2}{2m}\frac{d^2\psi}{dx^2} + V(x)\psi = E\psi \quad or \quad \hat{H}\psi = E\psi, \quad \hat{H}^2\psi = E^2\psi$$

So that in the end: $\Psi(x,t) = \psi(x)\varphi(t) = \psi(x)\exp\left(-\frac{iEt}{\hbar}\right)$

Stationary states:

Most solution to the TDSE **can't** be seperated as described above. But the general solution can **always** be written as a linear combination of our stationary solutions.:

$$\Psi_{gen}(x,t) = \sum_{n=1}^{\infty} c_n \psi_n \exp\left(-\frac{iE_n t}{\hbar}\right)$$

Procedure:

Given V(x) and $\Psi(x,0)$: \to TISE $\to \Psi(x,t) = \sum_{n=0}^{\infty} c_n \psi_n \varphi_n$

Before measuring the system any E_n of the ψ_n is a possible result with a probability of $|c_n|^2$. $\sum_{n=1}^{\infty} |c_n|^2 = 1 \Rightarrow \text{Probabilistic measurement!}$

But: when measuring the same prepared system repeatedly, the following measurements will always yield the same result.

$$c_n = \int_{-\infty}^{\infty} \psi_n^*(x) \Psi_{gen}(x, 0) dx$$

Seperable solutions describe stationary states $\Psi(\Psi_n)$ meaning that although Ψ depends on x and t, the prob. density $|\Psi|^2$ does not $\Rightarrow |\psi(x,t)|^2 = |\Psi(x)|^2$, Expectation values of stationary states are constant in time: $\langle Q \rangle = \langle Q(x,p) \rangle$

Note:

- If $\langle x \rangle$ is time-independent $\Rightarrow \langle p \rangle = m \frac{d\langle x \rangle}{dt} = 0$
- The expectation value of \hat{H} : $\langle H \rangle$ is E, $\langle H^2 \rangle = E^2 \Rightarrow \sigma_H = 0$. **Note:** if in linear combination $(\Psi = c_1 \psi_1 \varphi_1(t) + c_2 \psi_2 \varphi_2(t) : \langle H \rangle = |c_1|^2 E_1 + |c_2|^2 E_2), \langle H^2 \rangle = |c_1|^2 E_1^2 + |c_2|^2 E_2^2$

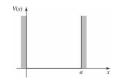
Every measurement of the total energy yields E for stationairy states.

In detail:
$$\langle H \rangle = \int_{n=1}^{\infty} |c_n|^2 E_n = \int \Psi^* \hat{H} \Psi dx$$

Quantum mechanical systems

Infinite square well (ISW):

Infinite potential well:



$$V(x) = \begin{cases} 0, & 0 \le x \le a. \\ \infty, & otherwise \end{cases}$$

SE:
$$-\frac{\hbar^2}{2m}\frac{d^2\psi}{dx^2} = E\psi \Rightarrow \psi_n(x) = \sqrt{\frac{2}{a}}\sin(\frac{n\pi}{a}x)$$

TISE

Characteristics of the infinite square well:

- The solution alternates between even and odd.
- $n\uparrow \Rightarrow$ more nodes, n = 0,1,2,3...
- ψ_n are mutually orthogonal $\Rightarrow \int \psi_m^* \psi_n dx = \delta_{mn}$ (not $\forall \Psi_{qen} =$
- Eigenstates are spaced as 2n+1 in energy
- The solutions are complete.

Completeness (for any f(x) between a and b):

$$f(x) = \sum_{n=1}^{\infty} c_n \psi_n(x) = \sqrt{\frac{2}{a}} \sum_{n=1}^{\infty} c_n \sin\left(\frac{n\pi}{a}x\right)$$

The quantized E level: $E_n = \frac{n^2 \pi^2 \hbar^2}{2ma^2} = \frac{\hbar^2 k_n^2}{2m}, \quad k_n = \frac{n\pi}{a}$

Note: For N non interacting bosons/fermions: $\frac{N \cdot (1)^2 \pi^2 h^2}{2ma^2}, E_{0,Fermions} = \frac{\pi^2 h^2}{2ma^2} \cdot 2 \sum_{n=1}^{N/2} n^2$ Yielding for the general solution:

$$\Psi_{gen}(x,t) = \sum_{n=1}^{\infty} c_n \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi}{a}x\right) \exp\left(-i\left(\frac{n^2\pi^2\hbar}{2ma^2}\right)t\right)$$

$$c_n = \sqrt{\frac{2}{a}} \int_0^a \sin\left(\frac{n\pi}{a}x\right) \Psi_{gen}(x,0) dx$$

Example: Sin- Integral

- $\langle \psi_n | x^2 \psi_m \rangle = \frac{1}{2} \int_0^a x^2 (\sin(\frac{n\pi x}{a}) \sin(\frac{m\pi x}{a})) dx = (n \neq m)$ $\frac{2a^2}{\pi^2} \left[\frac{(-1)^{n-m}}{(n-m)^2} \frac{(-1)^{n+m}}{(n+m)^2} \right]$ $\int_0^a x \sin(\frac{n\pi x}{a}) = \frac{a^2}{4}$

HO's apply to important particles (photons, phonons, molecular vibrations..)

$$V(x) = \frac{1}{2}m\omega^2 x^2 \Rightarrow \hat{H} = -\frac{\hbar^2}{2m}\frac{d^2}{dx^2} + \frac{1}{2}m\omega^2 x^2, \omega = \sqrt{\frac{k_s}{m}}$$

Note: In this case V(x) is simple. An arbitrairy potential could allways locally be approximated an HO.

Total number of Quanta: total quanta $=\frac{E_{tot}}{E_{spacing}} = \frac{E_{kin} + E_{pot}}{\hbar \cdot \omega}$

After sep. of var. we arrive at various states ψ_n of which the ground state is:

$$\psi_0 = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}} \exp\left(\frac{-m\omega}{2\hbar}x^2\right) (normalized) \Rightarrow \hat{a}_- \cdot \psi_0 = 0$$

We can ascend and descend the energy level by applying the raising operator \hat{a}_{+} and the lowering operator \hat{a}_{-} :

$$\hat{a}_{\pm} = \frac{1}{\sqrt{2\hbar m\omega}} (\mp i\hat{p} + m\omega\hat{x})$$
 , $[\hat{a}_{+}, \hat{a}_{-}] = 1$

 \hat{a}_{\pm} is not a hermitian operator but $\hat{a}_{+}\hat{a}_{-}$ is.

$$\hat{H} = \hbar\omega(\hat{a}_{+}\hat{a}_{-} + \frac{1}{2}) = \hbar\omega(\hat{a}_{-}\hat{a}_{+} - \frac{1}{2})$$

 $\hat{H}\psi_n = E_n\psi_n \Rightarrow \psi_n$ is a solution of the TISE.

But $\hat{a}_{\pm}\psi_n$ is also a solution: $\hat{H}\hat{a}_{\pm}\psi_n = (E_n \pm \hbar\omega)\hat{a}_{\pm}\psi_n$

Per application of \hat{a}_{\pm} we rise/fall one energylevel. $(\hbar\omega)$ Therefore concluding:

$$\psi_n = \frac{1}{\sqrt{n!}} \hat{a}_+^n \psi_0, \quad E_n = \langle \psi_n | \hat{H} | \psi_n \rangle = (n + \frac{1}{2}) \hbar \omega, n = 0, 1, \dots$$

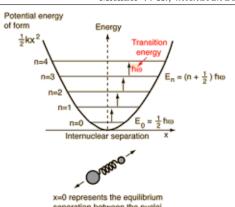
$$\psi_n(x) = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} \frac{1}{\sqrt{2^n n!}} H_n\left(\sqrt{\frac{m\omega}{\hbar}}\right) \cdot \exp\left\{\left(\frac{-m\omega}{2\hbar}x^2\right)\right\}$$

with $H_n(x) = (-1)^n \exp\{(x^2)\} \frac{d^n}{dx^n} [\exp(-x^2)]$

Note: We can't go lower in energy than the ground level $E_0 = \frac{1}{2}\hbar\omega \Rightarrow$ therefore the oscillator is allways vibrating!

Some definitions:

- n = number of energy quanta, n = 0,1,2,3... (starts at 0) $\hat{N} = \hat{a}_{+}\hat{a}_{-}$ 'number operator', is hermitian $\langle N \rangle = n$
- $\hat{a}_-\psi_n = \sqrt{n}\psi_{n-1}$, $\hat{a}_+\hat{a}_-\psi_n = n\psi_n$, $\hat{a}_+\psi_n = \sqrt{n+1}\psi_{n+1}$
- ψ_n is even (odd) for n even (odd)
- Same characteristics as the ISW.



To be classical: $\Delta E_{n,n+1} \to 0$ and the region outside parabola are vorbidden!

Note: V(x) = 0 everywhere, $[\hat{H}, \hat{p}] = 0$ and $\langle H \rangle = \langle K \rangle$, with $\hat{K} = \frac{\hat{p}}{2m}$ kinetic energy. Since $\hat{p} = \hbar \hat{k} \rightarrow \hat{K} = \frac{\hbar^2 k^2}{2m}$:

$$\langle H \rangle = \int_{-\infty}^{\infty} g^*(k,0) \cdot \hat{K} \cdot g(k,0) dk$$

Wavefunction:

$$\psi(x) = A \exp(ikx) + B \exp(-ikx)$$

Therefore for the wavefunction Ψ_k :

$$\Psi_k(x,t) = A \exp i(kx - \frac{\hbar k^2}{2m}t)$$

$$k = \pm \frac{\sqrt{2mE}}{\hbar}$$
 with $\begin{cases} k > 0, & \Rightarrow \text{ wave travelling to the right.} \\ k < 0, & \Rightarrow \text{ wave travelling to the left.} \end{cases}$

The stationairy states of the free particle are propagating waves with wavelength $\lambda = \frac{2\pi}{|k|}$ and momentum $p = \hbar k$.

Problem: Ψ_k is not normalizable: $|A|^2 \int_{-\infty}^{\infty} \Psi_k^* \Psi_k dx = |A|^2 \infty$

For the free particle the separable solutions don't represent physically realizable states.

There is no free particle with a definite energy. [continuous]

The general solution is still a linear combination (Just here an integral over the continuous k instead of a sum):

$$\Psi_{gen}(x,t) = \int_{-\infty}^{\infty} g(k) \exp i(kx - \frac{\hbar k^2}{2m}t) dk$$

Shape function:

$$g(k,t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \Psi(x,t) \exp(-ikx) dx$$

Note: g(k) := g(k, 0).

Wave packages are normalizable but energy not always defined.

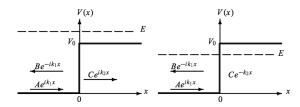
Bound and scattering states:

Bound particles: V(x) > E The particle is trapped in the potential well.(Unless additional energy is provided)

Scattering particles V(x < E) The particle can't be trapped by the potential.

$$\begin{cases} E < V(-\infty) \text{ and } V(\infty) & \Rightarrow bound \quad state \\ E > V(-\infty) \text{ or } V(\infty) & \Rightarrow scattering \quad state \end{cases}$$

Tunneling is the phenomenon of a wavefunction leaking through a finite barrier. It can be quantified by the transmission coefficient T.(Prob. of particle being transmitted)

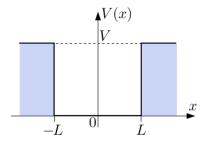


If E > V(x = 0): The wave will be transmitted and reflected.

If 0 < E < V(x = 0) The wave will be reflected but not really transmitted (exponential decay).

For E < 0 there are no physical solutions.

Well with finite Potential barrier

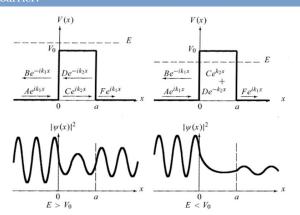


- Refl. and Transm. (E > 0)
- Bounded particle with penetr. into wall $(0 \le E \le V)$

Theory: O.E

- If a single particle approaches a potential barrier, its wavefunction is NOT always completely transmitted if it has kinetic energy above the height of the barrier.
- For any finite square well, there is always at least one bound state.
- The finite potential barrier allows only scattering states.
- If an electron has an energy E that is higher than the minimum potential energy, V_{min} , of a finite well, then the particle could be in bound state or scattering state

Finite barrier:



If E > V(x = 0): Transm. and refl. as for the finite step.

If E < V(x = 0): Both transm. and refl. but with an exp. decay through the barrier. Lighter particles penetrate deeper. Energy is conserved.

Tunneling: $T \approx \frac{16E(V_0 - E)}{V_0^2} \exp\left\{-\frac{4a}{\hbar}\sqrt{2m(V_0 - E)}\right\}$

Small **spike** barrier in the potential, not in Hilbert space. $\Rightarrow V_{barrier} = a\delta$ Dirac delta function: $\delta(x-a)$ defines a spike at position x=a.

$$\int_{-\infty}^{\infty} f(x)\delta(x-a)dx = f(a)$$

Formalism

We imagine our wavefunctions to be n-dimensional vectors

represented by n tuples in reference to an orthonormal basis except that they are actually infinite. We talk linear algebra.

The Hilbert space is the 'vector' space of all square-integrable (normalizable) functions. Therefore all wavefunctions of physically relizable states are in the Hilbert space.

Theorem: Dirac Notation

We define the *inner product* of two functions f(x) and g(x):

$$\langle f|g\rangle = \int_a^b f(x)^*g(x)dx, \quad \langle f|g\rangle = \langle g|f\rangle^*$$

- 1. Wavefunctions are represented by the 'Ket' notation: $|\Psi\rangle$
- 2. 'Bra' = $\langle B| = (b_1^*, b_2^*, \dots, b_n^*)$, it is always complex conj.'
 3. 'Ket' = $|A\rangle = (a_1, a_2, \dots, a_n)^T$ (Note: $\langle A| = (a_1^*, a_2^*, \dots, a_n^*)$)
- 4. $\langle B|A\rangle = b_1^* a_1 + \dots + b_n^* a_n$
- 5. Normalisation: $\langle \chi | \chi \rangle = 1$, with i.e $| \chi \rangle = A \begin{bmatrix} a_1 & a_2 & a_3 \dots \end{bmatrix}^T$.

For example \hat{p} is usually defined as $-\hbar \frac{\partial}{\partial x}$ but this is with respect to the **basis** \hat{x} , in the momentum space \hat{p} would just be p.

Theory: O.E

For a general one-dimensional wavefunction $\Psi(x)$, the wavevector $|\Psi\rangle$ is

$$\hat{Q}$$
 is Hermitian if: $\langle \psi | \hat{Q} \psi \rangle = \langle \hat{Q}^* \psi | \psi \rangle$

For Hermitian operators: $\langle \psi | \hat{Q} \psi \rangle = \langle \psi | \hat{Q}^* \psi \rangle = \langle \hat{Q} \psi | \psi \rangle$

Observables are represented by Hermitian operators.

Hermitian conjugate: $\langle f|\hat{Q}g\rangle = \langle \hat{Q}^{\dagger}f|g\rangle$

For Hermitian operators holds $\hat{Q} = \hat{Q}^{\dagger} \hat{Q}^{\dagger} = (\hat{Q}^T)^*$

For **unitairy** operators: $\hat{Q}^{-1} = \hat{Q}^{\dagger} \equiv (\hat{Q}^T)^*$

Note: EF of a hermitian operator are dirac-orthogonal. In non relativistic QM, the observable time is an independent variable.

$$\bullet \begin{cases}
\left[\hat{A}, \hat{B}\right] = \hat{A}\hat{B} - \hat{B}\hat{A}, \quad \left[\hat{A}, \hat{B}\right] = -\left[\hat{B}, \hat{A}\right] \\
\left[\hat{A}, \hat{B}^{2}\right] = \left[\hat{A}, \hat{B}\right]\hat{B} + \hat{B}\left[\hat{A}\hat{B}\right] \\
\left[\hat{A}\hat{B}, \hat{C}\hat{D}\right] = \hat{A}\hat{D}\left[\hat{B}, \hat{C}\right] + \hat{D}\hat{A}\left[\hat{B}, \hat{C}\right]
\end{cases}$$

- \hat{A}, \hat{B} commute. $\Leftrightarrow [\hat{A}, \hat{B}] = 0 \Leftrightarrow \hat{A}, \hat{B}$ are compatible.
- \hat{A}, \hat{B} do not commute. $\Leftrightarrow \left[\hat{A}, \hat{B}\right] \neq 0 \Leftrightarrow \hat{A}, \hat{B}$ are incompatible osserva-
- If \hat{A} and \hat{B} are compatible (osservable), they share Eigenfunctions.

Example: Some compatibility operators

$$\begin{split} &[\hat{x},\hat{p}]=i\hbar, \quad [\hat{x},V(x)]=0, \ (V(x)=\sum_n a_n \hat{x}^n) \\ &[\hat{Q},\hat{Q}]=0, \quad \left[\hat{Q},\hat{Q}^n\right]=0, \quad [\hat{a}_-,\hat{a}_+]=1 \\ &[\hat{x},\hat{p}^2]=2i\hbar\hat{p}, \quad \left[\hat{x},\hat{H}\right]=\frac{1\hbar\hat{p}}{m}, \quad [\hat{a}_\pm,c(const)]=0 \end{split}$$

Example: Compatibilty calculous

$$\begin{array}{l} \left[\frac{d^{2}}{dx^{2}},x\right] \, \Rightarrow \, \frac{d^{2}}{dx^{2}}[xf(x)] \, - \, x \frac{d^{2}f(x)}{dx^{2}} \, = \, \frac{d}{dx}[f(x) \, + \, x \frac{df(x)}{dx}] \, - \, x \frac{d^{2}f(x)}{dx^{2}} \, = \\ \frac{df(x)}{dx} \, + \, \frac{df(x)}{dx} \, - \, x \frac{d^{2}f(x)}{dx^{2}} \, - \, x \frac{d^{2}f(x)}{dx^{2}} \, = 2 \frac{d}{dx}f(x) \, \to \left[\frac{d^{2}}{dx^{2}},x\right] \, = 2 \frac{d}{dx} \end{array}$$

Theory: Compatibility experiment

Given are 3 Observables $\hat{A}, \hat{B}, \hat{C}$ with $[\hat{A}, \hat{B}] = [\hat{B}, \hat{C}] = 0$ and

Any inital state Ψ_{init} can be written as $|\Psi_{init}\rangle = \sum_n \tilde{c}_n |\Psi_{A,n}\rangle$ (Due to completeness), $|\Psi_{A,n}\rangle$ is A's EF

We have an initial state $|\Psi_{init}\rangle$ and measure three observables \hat{A}, \hat{B} and \hat{C} in a row.

 $\hat{A}|\Psi_{A,n}\rangle=a_n|\Psi_{A,n}\rangle,\quad \hat{B}|\Psi_{B,n}\rangle=b_n|\Psi_{B,n}\rangle,\quad \hat{C}|\Psi_{C,n}\rangle=c_n|\Psi_{C,n}\rangle$ Then during the first measurement the initial state collapses:

$$\begin{split} \Psi_{init} \rightarrow \begin{cases} \Psi_{A,1}(measure & a_1), & w/ & probability & |\tilde{c}_1|^2 \\ \Psi_{A,2}(measure & a_2), & w/ & probability & |\tilde{c}_2|^2 \end{cases} \\ \text{Second measurement is when compatibility comes into play:} \end{split}$$

Since
$$[\hat{A}, \hat{B}] = 0 \Rightarrow |\Psi_{A,n}\rangle = |\Psi_{B,n}\rangle$$
 and thus

$$\begin{array}{l} |\langle \Psi_{B,1}| \Psi_{A,1}\rangle = 1 \quad and \quad |\langle \Psi_{B,2}| \Psi_{A,1}\rangle = 0 \\ \rightarrow \text{ if } \Psi_{init} \rightarrow \Psi_{A,1} \end{array}$$

We know in \hat{B} we get $\begin{cases} b1, & 1(100\%) \\ b2, & 0(0\%) \end{cases}$ vice versa for $\Psi_{A,2}$

 \Rightarrow The second measurement is deterministic.

Given that \hat{C} and \hat{A} are incompatible, they do not share the same eigenfunctions and so the measurement of \hat{C} is probabilistic: we write EF of \hat{A} as a linear combination of EF of \hat{C} to determine probabilities. (Es: given $\to \Psi_{A,1} = k_{1a}\Psi_{C,1} + k_{1b}\Psi_{C,2}$ and $\Psi_{A,2} = k_{2a}\Psi_{C,1} - k_{2b}\Psi_{C,2}$ if Ψ_{gen} collapsed to $\Psi_{A,1}$ after \hat{A} we measure $(c_1;c_2)$ with probabilities (k_{1a}^2, k_{1b}^2) . Analog hold for the case when we collapsed into $\Psi_{A,2} \to \mathcal{P}(c_1; c_2) = (k_{2a}^2; k_{2b}^2))$

Determinate states are states that yield the same result for measurements on equally prepared systems in the same state.

One example are stationairy states for which we will in a certain state allways measure E_n .

Measurement always yields same result q, $\Rightarrow \sigma_Q = 0$.

$$\hat{Q}\Psi = q\Psi$$

 Ψ is the **Eigenfunction** of \hat{Q} and q its **Eigenvalue**.

Operators essentially perform linear transformations on a vector. Operators always expressed w.r.t. a basis.(usually \hat{x})

- $1. \ \,$ The TISE is an example of such an Eigenvalue equation.
- 2. The **EF** of the momentum operator $\hat{p} \to f_p(x) = \frac{1}{\sqrt{2\pi\hbar}} \exp\{(ipx/\hbar)\}$
- 3. Multiplying an Eigenfunction by a constant yields still an Eigenfunction.
- 4. The **Spectrum** of an operator are all its Eigenvalues.
- 5. Eigenvalues can be shared by different Eigenfunctions \Rightarrow **Degenerate** spectrum. Any linear combination of these Eigenfunctions is an Eigenfunction to the same Eigenvalue itself.
- 6. **Discrete spectrums** have separated Eigenvalues (\hat{H} in HO),
- 7. Continuous spectrums have continuous Eigenvalues (\hat{H} of FP) and are not normalizable.(ISW has both discrete and continuous.)

Discrete spectra:

- Real Eigenvalues
- Eigenfunctions of distinct Eigenvalues are orthogonal. (HO, ISW)

Continuous spectra: (e.g. free particle)

- Continuous Eigenvalues
- \bullet Eigenfunctions not normalizable \Rightarrow not in Hilbert space
- Linear combination must be used

The Eigenfunctions of an observable operator are complete.

Note: Measuring an observable \hat{Q} : State $\Psi(x,t) = \sum_{n} c_n(t) f_n(x)$, with c_n describing how much of f_n there is in Ψ . Of course $\sum_n |c_n|^2 = 1 \Rightarrow \hat{Q}f_n = q_n f_n$, $\langle \hat{Q} \rangle = \sum_{n} q_n |c_n|^2$

Theory: Degeneracy

Symmetry is the source of most degeneracy in Q.M. Consider \hat{H} invariant to symmetry operation \hat{Q} where $|\psi'_n\rangle = \hat{Q}|\psi_n\rangle$. Invariance says

$$[\hat{H}, \hat{Q}] = 0 \Rightarrow \begin{cases} \hat{H}|\psi_n\rangle = E_n|\psi_n\rangle \\ \hat{H}|\psi_n'\rangle = E_n|\psi_n'\rangle \end{cases}$$

Compatible observables share same eigenstates \rightarrow no degeneracy, they are same state.

To have degeneracy we require 2 sets of compatible observables: Let \hat{H} be invariant to at least 2 symmetry operators $\hat{Q}_1, \hat{Q}_2 \Rightarrow$

$$[\hat{H},\hat{Q}_1] = 0, [\hat{H},\hat{Q}_2] = 0 \text{ but } [\hat{Q}_1,\hat{Q}_2] \neq 0 \Rightarrow \begin{cases} \hat{H}|\psi\rangle = E_n|\psi\rangle \\ \hat{Q}_1|\psi\rangle = q_n|\psi\rangle \end{cases}$$

But
$$HQ_2|\psi\rangle = Q_2H|\psi\rangle = E_nQ_2|\psi\rangle = E_n|\chi\rangle$$
.

But $\hat{H}\hat{Q}_2|\psi\rangle=\hat{Q}_2\hat{H}|\psi\rangle=E_n\hat{Q}_2|\psi\rangle=E_n|\chi\rangle$. This states $|\psi\rangle,|\chi\rangle$ are distinct but with same $E_n\to$ degeneracy. Multiple non commuting symmetry operators ⇒ energy degeneracy

Theory: O.E.

- If Ψ_{qen} is an eigenstate of an operator \hat{Q} , the measurement is deterministic.
- If Ψ_{gen} is a superposition of eigenstates \Rightarrow probabilistic.

Quantum mechanics in 3 dimensions

3D - Schrödinger equation:

In 3D the hamiltonian is:

$$\hat{H} = \frac{1}{2m}(p_x^2 + p_y^2 + p_z^2), \quad p_x = -i\hbar \frac{\partial}{\partial x}$$

We write $p = -i\hbar\nabla$ and therefore TISE:

$$-\frac{\hbar^2}{2m}\nabla^2\psi+V\psi=E\psi,\quad \nabla^2=\frac{\partial^2}{\partial x^2}+\frac{\partial^2}{\partial y^2}+\frac{\partial^2}{\partial z^2}$$

Spherical coordinates:

In general:

$$\int |\Psi|^2 d^3 r = 1$$

TISE:

$$-\frac{\hbar^2}{2m}\left[\frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial\psi}{\partial r}\right) + \frac{1}{r^2\sin\theta}\frac{\partial}{\partial\theta}\left(\sin(\theta)\frac{\partial\psi}{\partial\theta}\right) + \frac{1}{r^2\sin^2\theta}\left(\frac{\partial^2\psi}{\partial r^2}\right)\right] + V\psi = E\psi$$

Sep. of var.(again since we already split $\psi(r)\epsilon(t)$): $\psi(r,\theta,\varphi)=R(r)V(\theta,\varphi)$ The normalized spherical part is called Spherical Harmonics:

$$Y_l^{m_l}(\theta,\varphi) = \epsilon \sqrt{\frac{(2l+1)}{4\pi} \frac{(l-m_l)!}{(l+m_l)!}} e^{im_l \varphi} P_l^{m_l}(\cos \theta)$$

 $\epsilon = (-1)^{m_l}$ for $m_l > 0$, $\epsilon = 1$ for $m_l \leq 0$

- For a spherically symmetric potential V=V(r) the spherical part is the Spherical harmonics $Y_l^{m_l}$, which is orthogonal
- $n = \text{principle QN}, l = \text{azimutal quantum number}, m_l = \text{magnetic quan-}$
- The degeneracy of n energy level will be 2n + 1 due to m_l .

- ullet We could approximate the H atom as a single particle because m_{n^+} +
- The ground state has no angular dependence
- Electron Electron interaction: $V(\vec{r}) = \frac{1}{4\pi\epsilon_0} \frac{e^2}{|\vec{r_1} \vec{r_2}|}$
- Total orbital states up to n: $\# = 2 \cdot \sum_{\tilde{n}=1}^{n} \tilde{n}^2 = 2 \cdot \sum_{l=0}^{l=n-1} 2l + 1$

	Radiale	Drehimpuls-	Magnetische
	Quantenzahl	quantenzahl	Quantenzahl
Symbol	n	l	m_l
Mögliche Werte	$n = 1, 2, 3, \dots$	$l=0,1,\ldots,n-1$	$m_l=-l,-l+1,\ldots$
			$\ldots, l-1, l$
Assoziierte	r	ϑ	φ
Polarkoordinate	100		
Assoziierte	Energie:	Drehimpuls (Betrag):	Drehimpuls
physikalische	$E_n = -E_R Z^2 / n^2$	$L = \sqrt{l(l+1)}\hbar$	(z-Komponente):
Grössen		W 100 B	$L_z = m_l \hbar$
			Magnetischer Dipol:
		ter muneaveer	$m_z = -\mu_B m_l$
Merkmal in	Ist gleich der	Ist gleich der	Bestimmt die Lage
Wellenfunktion	totalen Anzahl	Anzahl	der Knotenebene(n)
	Knoten plus eins	Knotenebenen	` '

Theory: Normalisation/othogonality and expectation values

- Normalisation of angular part: $\int_0^{2\pi} \int_0^\pi |Y_l^{m_l}|^2 \sin\theta \ d\theta \ d\varphi = 1$ Orthogonality: $\int_0^{2\pi} \int_0^\pi (Y_{l_1}^{m_{l_1}})^* Y_{l_2}^{m_{l_2}} \sin\theta \ d\theta \ d\varphi = 0$ Expectation value of $r: \langle r \rangle = \int_0^\infty \int_0^{2\pi} \int_0^\pi \Psi^* r \Psi \sin\theta \ d\theta \ d\varphi \ dr$ Expectation value of $x: \langle x \rangle = 0$ for symmetry

- Expectation value of x^2 :
 - 1. Ground state: $\langle x^2 \rangle \to \text{use that } \langle r^2 \rangle = \langle x^2 \rangle + \langle y^2 \rangle + \langle z^2 \rangle =$ $3\langle (x|y|z)^2\rangle$
 - 2. Else: use that $x = r \cos \varphi \sin \theta$

Radial part:

$$R_{nl}(r) = \sqrt{\left(\frac{2}{na}\right)^3 \frac{(n-l-1)!}{2n[(n+l)!]^3}} \exp\left(-\frac{r}{na}\right) \left(\frac{2r}{na}\right)^l \cdot \left(L_{(n+l)-(2l+1)}^{2l+1}\left(\frac{2r}{na}\right)\right)$$

With a being the **Bohr radius**: $a = \frac{4\pi\epsilon_0 \hbar^2}{me^2} = 0.529 * 10^{-10} m$ Final solution:

$$\Psi_{n,l,m_l}(r,\theta,\varphi) = R_{n,l}(r)Y_l^{m_l}(\theta,\varphi)$$

$$E_n = -\left[\frac{m}{2\hbar^2} \left(\frac{e^2 Z}{4\pi\epsilon_0}\right)^2 \frac{1}{n^2}\right] = E_1 \frac{Z^2}{n^2}, E_1 = -13.6eV$$

$$a_z = \frac{a}{Z} \Rightarrow \text{ The general radius: } r_n = a_z \cdot n^2$$

 E_1 is the energy of the ground state (Rhydberg energy). **Note:** For hydrogen atoms the energys depend only on $n \Rightarrow$ degeneracys, for multielectron atoms this is NOT true.

Theory: Hydrogenic Atom and Screening

- 1. A hydrogen atom has 1 proton and 1 electron. A hydrogenic Atom has Z protons and 1 electron \rightarrow the binding energy of a hydrogenic Atom $(E_1(z))$ will be much higher, as $Z \uparrow$. This is due to the fact that the single electron will be bound more tightly to multiple protons in nucleus of hydrogenic atom.
- 2. What happen when you try to solve the TISE for non Hydrogenic Atom with (A) and without (B) electron-eletron forces? Using (A) the electrons in higher orbitals now feel the repelling force eletrons sitting in inner orbitals. (es: electron in 1s and 2s repell electron in 2p orbital). The "pull" from the protons in the nulcleons gets therefore reduced which leads to a higher total energy respect to (B). This effect is called screening.

$$\hat{L}_x = y\hat{p}_z - z\hat{p}_y, \quad \hat{L}_y = z\hat{p}_x - x\hat{p}_z, \quad \hat{L}_z = x\hat{p}_y - y\hat{p}_x$$

Example: Some compatibility operators and calcolous

$$\begin{bmatrix} \hat{L}_x, \hat{L}_y \end{bmatrix} = i\hbar \hat{L}_z, \quad \begin{bmatrix} \hat{L}_y, \hat{L}_z \end{bmatrix} = i\hbar \hat{L}_x, \quad \begin{bmatrix} \hat{L}_z, \hat{L}_x \end{bmatrix} = i\hbar \hat{L}_y$$

$$\begin{bmatrix} \hat{L}_z, x \end{bmatrix} = i\hbar y, \quad \begin{bmatrix} \hat{L}_z, y \end{bmatrix} = -i\hbar x, \quad \begin{bmatrix} \hat{L}_z, z \end{bmatrix} = 0$$

$$\begin{bmatrix} \hat{L}_z, \hat{p}_x \end{bmatrix} = i\hbar \hat{p}_y, \quad \begin{bmatrix} \hat{L}_z, \hat{p}_y \end{bmatrix} = -i\hbar \hat{p}_x, \quad \begin{bmatrix} \hat{L}_z, \hat{p}_z \end{bmatrix} = 0$$

$$\begin{bmatrix} \hat{L}_z, \hat{p}_x \end{bmatrix} = [x\hat{p}_y - y\hat{p}_x, \hat{p}_x] = [x\hat{p}_y, \hat{p}_x] = i\hbar \hat{p}_y$$

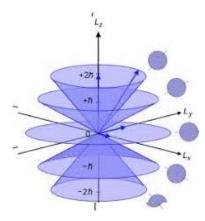
- 1. \hat{L}_x , \hat{L}_y and \hat{L}_x are **incompatible** \Rightarrow The can't be measured simultaneously. Their uncertainty being: $\sigma_x \sigma_y \geq \frac{\hbar}{2} |\langle L_z \rangle|$, analogous for $\sigma_x \sigma_z$
- 2. Consider $\hat{L}^2 = L_x^2 + L_y^2 + L_z^2 \Rightarrow [L^2, L_i] = 0 \Rightarrow$ they share EF (to get EV see below).
- 3. Since $\left[\hat{L}_z,\hat{r}^2\right]=\left[\hat{L}_z,\hat{p}^2\right]=0$ for spheric symmetrical potential we get $\Rightarrow \left[\hat{L}_{x/y}, \hat{r}^2\right] = \left[\hat{L}_{x/y}, \hat{p}^2\right] = 0$. So, writing $r = \sqrt{r^2}$ we get (recalling the spherically symmetric Hamiltonian previously defined): $\left[\hat{L}_{x/y/z}, \sqrt{r^2}\right] = 0$. So all three components of \hat{L} commutes with

Note: Since $\hat{H}, \hat{L}^2, \hat{L}_z$ all commute, they are compatible observables. This implies that the energy, angular momentum squared, and projection of the angular momentum along the z-axis can be determined simultaneously.

To get the **EV of the momentum operators** we use L_{\pm} :

$$\hat{L}^2 f_l^{m_l} = \hbar^2 l(l+1) f_l^{m_l}, \quad \hat{L}_z f_l^{m_l} = \hbar m_l f_l^{m_l}$$

Diagonal and z-projection are known resulting in this appearance. (L_z defined



$$\hat{L}_{z,spherical} = -i\hbar \frac{\partial}{\partial \varphi}$$

$$\hat{L}_{x,spherical} = -i\hbar(-\sin\varphi\frac{\partial}{\partial\theta} - \cos\varphi\cot\theta\frac{\partial}{\partial\varphi}),$$

$$\hat{L}_{y,spherical} = -i\hbar(+\cos\varphi\frac{\partial}{\partial\theta} - \sin\varphi\cot\theta\frac{\partial}{\partial\varphi})$$

Note: Z was chosen by convention.

$$\hat{L}^2 = -\hbar^2 \left[\frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial}{\partial \theta} \right) + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \varphi^2} \right]$$

The EF of \hat{L}_z and \hat{L}^2 are the Spherical Harmonics $Y_l^{m_l}$ from the

$$\hat{L}_{\pm} = \hat{L}_x \pm i\hat{L}_y \pm \hbar e^{\pm i\varphi} \left(\frac{\partial}{\partial \theta} \pm i \cot \theta \frac{\partial}{\partial \varphi} \right)$$

$$\hat{L}_{+}\hat{L}_{-} = -\hbar^{2}\left(\frac{\partial^{2}}{\partial\theta^{2}} + \cot\theta\frac{\partial}{\partial\theta} + \cot^{2}\theta\frac{\partial^{2}}{\partial\varphi^{2}} + i\frac{\partial}{\partial\varphi}\right)$$

Note: $\hat{L}_{\pm}|lm_{l1}\rangle = \sqrt{2}\hbar|lm_{l1\pm1}\rangle$, if $m_l = m_{l,max} \Rightarrow \hat{L}_+|lm_{l1}\rangle = 0$ Similarly for \hat{L}_{-} and $m_{l,min}$

Spin of a particle is not really a spin. Characteristic of the part. The total angular momentum comes from the orbiting of the electron plus the AM of its

The spin has the same interactions across multiple dimensions:

$$\[\hat{S}_x, \hat{S}_y\] = i\hbar \hat{S}_z, \quad \left[\hat{S}_y, \hat{S}_z\right] = i\hbar \hat{S}_x, \quad \left[\hat{S}_z, \hat{S}_x\right] = i\hbar \hat{S}_y$$

And for the EV's as before: (here $m = m_s$)

$$\hat{S}^2|s,m\rangle = \hbar^2 s(s+1)|s,m\rangle, \hat{S}_z|s,m\rangle = \hbar m|s,m\rangle \text{ with } |s,m\rangle = EF$$

$$\hat{S}_{\pm} = \hat{S}_x \pm i\hat{S}_y = \hbar\sqrt{s(s+1) - m(m\pm 1)}|s, (m\pm 1)\rangle$$

This time the EF are **NOT** $Y_l^{m_l}$. They are new EF's depending on s and m_s . We have no mathematical function for $|s, m_s\rangle$

 $s = 0, \frac{1}{2}, 1, \frac{3}{2}, m_s = -s, -(s-1), ..., \neq 0, ..., (s-1), s$

Theorem: Immutability of the spin

The value of s is **immutable** for any elementar particle. For electrons $s = \frac{1}{2}$, for photons s = 1.

 $(\pi$ - mesons: s = 0, baryons: $s = \frac{3}{2}$, gravitons: s = 2...)

One fundamental difference between s and l is, that l can change for a particle but s can't.

For **electrons** there are two states: $|s, m_s\rangle = |\frac{1}{2}; \frac{1}{2}\rangle$ and $|\frac{1}{2}; (-\frac{1}{2})\rangle$ Representing spin up $(\frac{1}{2})$ and spin down $(-\frac{1}{2})$ for the m_s

Theory: Spin and 3/2 basis

- For given $s \to m_s = [-s, -s + 1, \dots, 0, \dots, s]$
- s is always $\frac{1}{2}$ for electrons! It is never negative!
- If we apply \hat{S}_z to a free electron, and then immediately measure \hat{S}_x on the same electron, we DO NOT always obtain the same
- For the basis 3/2 we have: $|s, m_s\rangle = |\frac{3}{2}; \frac{1}{2}\rangle, |\frac{3}{2}; -\frac{1}{2}\rangle, |-\frac{3}{2}; \frac{1}{2}\rangle$ and $|-\frac{3}{2};-\frac{1}{2}\rangle$ [4 x 4 Matrix]

General state of spin: we can use a linear combination: $|\chi\rangle = a|\frac{1}{2};\frac{1}{2}\rangle +$ $b|\frac{1}{2};(-\frac{1}{2})\rangle$ which we can write in vector notation: $|\chi\rangle=\begin{vmatrix} a\\b\end{vmatrix}$, therefore e.g. spin up: $\chi_+ = \left| \frac{1}{2}; \frac{1}{2} \right\rangle = \begin{bmatrix} 1\\0 \end{bmatrix}$.

- 1. χ is called 'spinor'. 2. $|a|^2+|b|^2=1\Rightarrow |a|^2=$ probability of measuring \uparrow orientation $(|b|^2$ for

When measuring \hat{S}^2 or $\hat{S}_{x,y,z}$: $\chi \to \chi_{\pm}$ yielding either of the two respective EV's with **probability** $|c_{eig}|^2$

$$c_{eig} = Eigenvector * |\chi\rangle, \quad \sum |c_{eig}|^2 = 1$$

Example: c_{eig} probability

Given $|\chi\rangle=\begin{bmatrix} a\\b \end{bmatrix}$ and an EV of $+\frac{\hbar}{2}$ and an EF of $\binom{1/\sqrt{2}}{i/\sqrt{2}}\Rightarrow c_{+\frac{\hbar}{2}}=\frac{1}{\sqrt{2}}(a-ib).$ So we get $+\frac{\hbar}{2}$ with probability $\frac{1}{2}|a-ib|^2.$ More in general, given $|\chi\rangle \to P(|\uparrow_z\rangle) = |\langle\uparrow,\downarrow\rangle\rangle$

 \hat{S}^2 and \hat{S}_z are 2x2 matrices (they satisfies EV equation in their base with their EV):

$$\hat{S}^2 = \frac{3}{4}\hbar^2 \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \quad \hat{S}_z = \frac{\hbar}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

So are \hat{S}_x and \hat{S}_y :

$$\hat{S}_x = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \hat{S}_y = \frac{\hbar}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

$$\hat{S}^2 = \hat{S}_x^2 + \hat{S}_y^2 + \hat{S}_z^2 \Rightarrow \hat{S}_x^2 = \hat{S}_y^2 = \hat{S}_z^2 = \frac{\hbar^2}{4} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

Eigenfunctions in their respective base:

$$\chi_{\hat{S}^2} : \underbrace{\begin{pmatrix} 1 \\ 0 \end{pmatrix}}_{(1)}, \underbrace{\begin{pmatrix} 0 \\ 1 \end{pmatrix}}_{(1)}$$

$$\chi_{\hat{S}_x} : \underbrace{\begin{pmatrix} 1/\sqrt{2} \\ 1/\sqrt{2} \end{pmatrix}}_{(1)}, \underbrace{\begin{pmatrix} 1/\sqrt{2} \\ -1/\sqrt{2} \end{pmatrix}}_{(1)}$$

$$\chi_{\hat{S}_y} : \underbrace{\begin{pmatrix} 1/\sqrt{2} \\ i/\sqrt{2} \end{pmatrix}}_{(1)}, \underbrace{\begin{pmatrix} i/\sqrt{2} \\ 1/\sqrt{2} \end{pmatrix}}_{(1)}$$

And \hat{S}_{+} and \hat{S}_{-} :

$$\hat{S}_{+}\chi_{-} = \hbar\chi_{+}, \quad \hat{S}_{-}\chi_{+} = \hbar\chi_{-}, \quad \hat{S}_{\pm}\chi_{\pm} = 0$$

$$\hat{S}_{+} = \hbar \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}, \quad \hat{S}_{-} = \hbar \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$$

 \hat{S}_{+} and \hat{S}_{-} are not hermitian and therefore not observable.

Example: Expectation value for spin matrix

Given a spin state of $|\chi\rangle$ (normalized, see subsection Hilbert space and Dirac notation) $\Rightarrow \langle \hat{S}_x \rangle = \langle \chi | \hat{S}_x | \chi \rangle$.

Example: Eigenvalue of spin

Given the spin 1/2 basis i.e. $|\frac{1}{2};\frac{1}{2}\rangle=\begin{bmatrix}1\\0\end{bmatrix}$ and $|\frac{1}{2};-\frac{1}{2}\rangle=\begin{bmatrix}0\\1\end{bmatrix}$, determine EV and EF of \hat{S}_y :

- 1. Solve EV equation i.e $\hat{S}_y|?\rangle = \lambda|?\rangle$ and find λ
- 2. Plug back into \hat{S}_y and find EF \Rightarrow remember that $|a|^2 + |b|^2 = 1$

Multiparticle Systems

Differences and analogy:

The following is the same in multiparticle systems:

- ullet Ψ describes the system
- TDSE and TISE \rightarrow total energy $\Psi = \sum_{n} c_{n} \psi_{n} e^{\frac{-iE_{n}t}{\hbar}}$

The changes are:

- $\bullet \ \Psi = \Psi(\underline{r_1},\underline{r_2},...,\underline{r_n},t)$
- \hat{H} includes particle interaction!!! (through V)

left: kinetic energy, middle: potential energy, right: interaction

Probability that particle 1 is in volume 1 and part. 2 in volume 2:

$$|\Psi(\underline{r_1},\underline{r_2},t)|^2 d^3\underline{r_1}d^3\underline{r_2}$$

Note: Two particles in the same atom WITH electron interactions have never the same energy.

TISE:
$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} \right) + V(\underline{r_1}, \underline{r_2}) \psi = E \psi$$

Not analytically solvable. Unless hydrogenic

- Particle interaction in form of a distance-dependent potenital (Nucleuselectron) ⇒ can be handled like 1-particle problem.
- Two particle problem \rightarrow division into center of mass \underline{R} and relative motion r:

$$\underline{R} = \frac{m_1\underline{r}_1 + m_2\underline{r}_2}{m_1 + m_2}, \quad \underline{r} = \underline{r}_1 - \underline{r}_2, \Rightarrow \psi = \psi_R(\underline{R})\psi_r(\underline{r})$$

- 1. Almost no contribution from $\psi_R \Rightarrow 1$ particle (H-Atom).
- 2. No particle interaction \Rightarrow hydrogenic state:
- 3. Two cases: Distinguishable (e.g. Prot. and electron) or nondistinguishable (e.g. e^- and e^-):

Distinguishable:

$$\psi(\underline{r}_1,\underline{r}_2) = \psi_a(\underline{r}_1)\psi_b(\underline{r}_2) \quad \underline{or} \quad \psi_a(\underline{r}_2)\psi_b(\underline{r}_1)$$

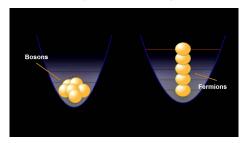
Non-distinguishable:

$$\psi_{\pm}(\underline{r}_1,\underline{r}_2) = C \left[\psi_a(\underline{r}_1)\psi_b(\underline{r}_2) \pm \psi_a(\underline{r}_2)\psi_b(\underline{r}_1) \right]$$

Note: with this superpos. we can build (a)symmetric spacial wavefunctions! For Bosons there is also an other solution for non distinguishable particle: $\psi_{Bos} = \psi_a(r_1)\psi_a(r_2)$ or $\psi_{Bos} = \psi_b(r_1)\psi_b(r_2)$

Exchange operator \hat{P} : $\underline{\hat{P}}f(\underline{r}_1,\underline{r}_2) = f(\underline{r}_2,\underline{r}_1)$

- 1. $[\hat{P}, \hat{H}] = 0$ they share EF's $(\psi_{\pm}) \Rightarrow \hat{P}\psi = \pm \psi$
- 2. For 2 identical particle \hat{P} commute and \hat{H} too
- 3. ψ_{\pm} is $\begin{cases} symmetric, & +1 \\ antisymmetric, & -1 \end{cases}$ with respect to exchange.
- 4. (Anti-)symmetry **doesn't** change over time.
- 5. ψ_{\pm} are EF's of \hat{H} and represent stationary states.



Bosons:

- Integer spin
- Can inhabit same energy levels ⇒ collaps into ground-state
- Unlimited bosons can be in the ground state simultaneously.
- Symmetric $\rightarrow \psi_+$
- E.g photons, gravitons
- Non degenerate state \rightarrow if degeneracy \rightarrow linear combination

Fermions:

- Half integer spin
- Only two fermions can inhabit the same state(Pauli principle)
- Antisymmetric $\rightarrow \psi_{-}$
- E.g electrons, protons, neutrons

$$\text{Now spin: } \chi(s) = \begin{cases} \uparrow \uparrow, & symmetric \Rightarrow \text{ triplet } (+) \\ \downarrow \downarrow, & \text{triplet} \\ \frac{1}{\sqrt{2}} (\uparrow \downarrow + \downarrow \uparrow), & \text{triplet} \\ \frac{1}{\sqrt{2}} (\downarrow \uparrow - \uparrow \downarrow), & antisymmetric \Rightarrow \text{ singlet } (-) \end{cases}$$

$$\psi_{final} = \psi_{\pm}(\underline{r}_1, \underline{r}_2)\chi(s)$$

Theorem: Symmetry requirements

The symmetry of the overall wavefunction Ψ for identical (1. bosons, 2. fermions) is (1. symmetric, 2. antiymmetric) with respect to exchange of any two particle

Boson and Fermion: **combined state** of spacial ψ_{\pm} and spin.

An electron and a photon are always distinguishable. Aall electrons in the universe are indistinguishable.

Note: generally ψ_+ has less energy than ψ_- so two electrons in ground state of a H_2 molecule would rather be in the ψ_+ state.

3D infinite square well with multiple particles:

3D infinite square well neglecting electron interactions inbetween each other:

$$E_{n_x,n_y,n_z} = \frac{\hbar^2 \pi^2}{2m} \left(\frac{n_x^2}{l_x^2} + \frac{n_y^2}{l_y^2} + \frac{n_z^2}{l_z^2} \right) = \frac{\hbar^2 k^2}{2m}$$

Where k is the absolute value of the vector $\vec{k} = (k_x, k_y, k_z)$.

$$k_x = \frac{n_x \pi}{l_x}, \quad k_y = \frac{n_y \pi}{l_y}, \quad k_y = \frac{n_z \pi}{l_z}$$

$$\psi_{n_x,y,z} = \sqrt{\frac{8}{l_x l_y l_z}} \sin\left(\frac{n_x \pi x}{l_x}\right) \sin\left(\frac{n_y \pi y}{l_y}\right) \sin\left(\frac{n_z \pi z}{l_z}\right)$$

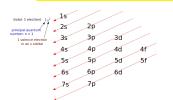
Accidental degeneracy: For the case of a cubic box in the next table are reported the first 14 "n-Energy parts". E_{14} is the first energy level where we get two different combinations of n_x, n_y, n_z that are degenerate. This is known as an äccidental degeneracy". The degeneracy in this case is 1[333] + 3[115] = 4.

·		E_1	E_2	E_3	E_4	E_5	E_6	E_7	E_8	E_9	E_{10}	E_{11}	E_{12}	E_{13}	E_{\cdot}	14
	n_x n_y n_z $\sum_i n_i^2$	1	1	1	1	2	1	2	1	1	1	2	2	1	3	1
	n_y	1	1	2	1	2	2	2	1	3	2	3	2	3	3	1
	n_z	1	2	2	3	2	3	3	4	3	4	3	4	4	3	5
	$\sum_i n_i^2$	3	6	9	11	12	14	17	18	19	21	22	24	26	2	7

Note: For every column the energy level are permutation of the quantum number reported (these degeneracies come about by simply changing the quantum numbers among n_x , n_y , $n_z \to [322], [232], [223]).$

Multi-particle atoms:

$$\hat{H} = \sum_{j=1}^{Z} \left[-\frac{\hbar^2}{2m_j} \nabla_j^2 - \frac{1}{4\pi\epsilon_0} \frac{Ze^2}{r_j} + \frac{1}{2} \frac{1}{4\pi\epsilon_0} \sum_{k \neq j}^{Z} \frac{e^2}{|r_j - r_k|} \right]$$



Note: For hydrogenic atoms each subshell is on the same energy level.

- Z: number of electrons
- n: number of shells (1,2,3,4)
- l: number of subshells (0=s,1=p,2=d,3=f...)
- m_l : which instance of a subshell group (e.g. which dumbbell for p)(-1,...,0,..,l

Exceptions due to electron interaction:

- Chromium($[Ar]4s^13d^5$)
- Copper($[Ar]4s^13d^{10}$)
- Niobium($[Kr]5s^14d^4$)
- Molybdenum($[Kr]5s^14d^5$)

Whe still have too little information for determining the order of "filling" the atom. \Rightarrow angular momentum:

- 1. $\mathbb{L} = \text{total orbital angular momentum}$,
- 2. S = total spin angular momentum,
- 3. $\mathbb{J} = \text{total combined angular momentum}$,
- 4. $M_{\mathbb{L}} = \sum m_l$, $M_{\mathbb{S}} = \sum m_s$

Filled subshells don't contribute to the Term symbol!

Term symbol: ${}^{2\mathbb{S}+1}\mathbb{L}_{\mathbb{T}}$

For two particles the way of adding angular momenta is:

- 1. $\mathbb{L} = (l_1 + l_2, l_1 + l_2 1, ..., |l_1 l_2|) \Rightarrow \text{(with more } l, \text{ see example)},$ L is represented by S,P,D,F.. (for 0,1,2,3...) 2. $\mathbb{S} = (s_1 + s_2, s_1 + s_2 - 1, ..., |s_1 - s_2|)$ 3. $\mathbb{J} = (\mathbb{S} + \mathbb{L}, \mathbb{S} + \mathbb{L} - 1, ..., |\mathbb{L} - \mathbb{S}|)$

Note: for more particle (es: 4 particle) $\mathbb{L} = \mathbb{L}_{12} + \mathbb{L}_{23}$ and we do all combinations possible.

Example: 4 distinguishable particle

Given are two particle with spin 1/2 and two particle with spin 1. Find

- 1. $\mathbb{S}_{1/2;1/2} = 0, 1$
- 2. $S_{1,1} = 0, 1, 2$
 - For $\mathbb{S}_{1/2;1/2} = 1$ and $\mathbb{S}_{1;1} = 0$ we have $\mathbb{S} = 1$
 - For $\mathbb{S}_{1/2;1/2}=1$ and $\mathbb{S}_{1;1}=0$ we have $\mathbb{S}=1$ For $\mathbb{S}_{1/2;1/2}=1$ and $\mathbb{S}_{1;1}=1$ we have $\mathbb{S}=0,1,2$ For $\mathbb{S}_{1/2;1/2}=1$ and $\mathbb{S}_{1;1}=2$ we have $\mathbb{S}=1,2,3$

 - $\bullet \ \, \text{For} \, \, \mathbb{S}_{1/2;1/2} = 0$ and $\mathbb{S}_{1;1} = 0,1,2$ we have $\mathbb{S} = 0,1,2$ respectively. tively

So we have the possible values of $\mathbb{S} = 0, 1, 2, 3$ with multiplicity $\mathbb{S} = 0$ (x 2), S = 1 (x 4), S = 2 (x 3), S = 3 (x 1)

Example: Nitrogen Term symbols

We have $2p^3$ as valence band $\rightarrow 3p$ electrons.

- 1. $l_1 = 1, l_2 = 1, l_3 = 1$ can combine to give $\mathbb{L} = 3, 2, 1, 0$
- 2. $s_1 = \frac{1}{2}, s_2 = \frac{1}{2}, s_3 = \frac{1}{2}$ can combine to give $\mathbb{S} = \frac{3}{2}, \frac{1}{2}$ 3. For $\mathbb{J} = \mathbb{L} + \mathbb{S}$ we have 19 combinations:
- For $\mathbb{J} = \mathbb{L} + \mathbb{S}$ we have 19 combinat (a) $\mathbb{L} = 3, \mathbb{S} = \frac{3}{2} \Rightarrow \mathbb{J} = \frac{9}{2}, \frac{7}{2}, \frac{5}{2}, \frac{3}{2}$ (b) $\mathbb{L} = 3, \mathbb{S} = \frac{1}{2} \Rightarrow \mathbb{J} = \frac{7}{2}, \frac{5}{2}$ (c) $\mathbb{L} = 2, \mathbb{S} = \frac{3}{2} \Rightarrow \mathbb{J} = \frac{7}{2}, \frac{5}{2}, \frac{3}{2}, \frac{1}{2}$ (d) $\mathbb{L} = 2, \mathbb{S} = \frac{1}{2} \Rightarrow \mathbb{J} = \frac{5}{2}, \frac{3}{2}$ (e) $\mathbb{L} = 1, \mathbb{S} = \frac{3}{2} \Rightarrow \mathbb{J} = \frac{5}{2}, \frac{3}{2}, \frac{1}{2}$ (f) $\mathbb{L} = 1, \mathbb{S} = \frac{1}{2} \Rightarrow \mathbb{J} = \frac{3}{2}, \frac{1}{2}$ (g) $\mathbb{L} = 0, \mathbb{S} = \frac{3}{2} \Rightarrow \mathbb{J} = \frac{3}{2}$ (h) $\mathbb{L} = 0, \mathbb{S} = \frac{1}{2} \Rightarrow \mathbb{J} = \frac{1}{2}$

Note: The term-symbol also includes symmetric states which is impossible for electrons. They defy the Pauli exclusion principle.

Theorem: Hund's rule

For determining the lowest energy state (neglecting exchange symmetry):

- 1. State with largest S is most stable.
- 2. States with the same $\mathbb{S} :\Rightarrow \text{largest } \mathbb{L} \text{ is most stable.}$
- 3. States with the same $\mathbb S$ and $\mathbb L$ are most stable for:

smallest \mathbb{J} , subshells not more than half full (\leq)

subshells more than half full (>) largest \mathbb{J} ,

Note: Singlet $\mathbb{S} = 0$, Triplet $\mathbb{S} = 1$

Solids

Fermi energy and solid model:

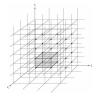
Fermi energy: The energy level of the energeticly highest electron **Fermi level:** The energy at which f(E)(T > 0) = 0.5

$$f(E) = \begin{cases} T = 0 \Rightarrow \begin{cases} 1, & \text{falls } E \leq E_F \\ 0 & \text{falls } E > E_F \end{cases} \\ T > 0 \Rightarrow \frac{1}{1 + e^{\frac{E - E_F}{k_B T}}} \end{cases}$$

Two ways to describe solids:

1. Free electron model: ignore electron-electron interactions, material is metal, V(r) = 0 inside box and so outside, material = box with electrons in

k- space:





Electronic states are filled up to k_f (right picture), $k_{x,y,z} > 0$

Every solution in the k-space occupies the same space: $\frac{\pi^3}{|x|^2 |x|^2} = \frac{\pi^3}{V}$

Occupied volume in k-space = $\frac{1}{2}$ #of electrons * volume per electronic state

with N being the number of atoms, q their respective number of valence elec-

Fermy energy for 1D, 2D and 3D case: plugging the equation above for V(x) = 0 into $E = \frac{\hbar^2 k^2}{2m}$, $\hat{p} = \hbar \hat{k}$ ($1eV = 1.6 \cdot 10^{-19} J$) we get:

1. **3D**: $\frac{1}{8}V_{sphere} = \frac{1}{2}N * q * \frac{\pi^3}{V} \Rightarrow \text{ where } V_{sphere} = \frac{4}{3}\pi k^3, \ k = \dots$ (a) $E_F = \frac{\hbar^2}{2m} \left(\frac{3\pi^2 Nq}{V} \right)^{2/3}$

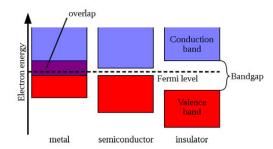
- (b) $Nq = \frac{V}{3\pi^2} \left(\frac{2mE}{\hbar^2}\right)^{3/2}$
- 2. **2D**: $\frac{1}{4}A_{circle} = \frac{1}{2}N*q*\frac{\pi^2}{A} \Rightarrow$ where $A_{circle} = \pi k^2, k = \dots$ (a) $E_F = \frac{\hbar^2 \pi N q}{Am}$ (b) $Nq = \frac{mEA}{\hbar^2 \pi}$

Density of states: $D(E) = \frac{dNq}{dE}$ (represents #e⁻ per unit energy, # of oneelectron-levels)

- 3D: $D(E) = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} E^{1/2}$ 2D: $D(E) = \frac{mA}{\hbar^2\pi}$

Energy spacing: $ES = \frac{1}{D(E)}$

2. Band model: periodic V(x) in atomic lattice, also bosons have Band Gaps, witch only depend on atomic lattice and not on type of particle.



The Fermi energy for metals is inside a band, for semi-conductors and insulators it is in between. Semi conductors can still conduct, the electron jumps. $(E_{gap} \text{ for semi-cond. is smaller than 4eV})$

Semicond. can be made more conductive with thermal and or doping measures:

- 1. Thermal excitation: $n_{e^-} = N_i e^{-\frac{L_g}{k_B T}}$, $N_i =$ undoped carrier concentration, $n_{e^-} = \#$ free electrons per unit Volume. **Note**; # photons absorbed = $n_{e^-} \cdot Volume$, $E_{photon} \geq E_g$
- 2. Doping excitation: $n_{e^-} = N_D e^{-\frac{E_d}{k_B T}}$, $N_d = \text{Concentration of Donor}$

Note: Dopants have low ionization energy so can release easily its extra electron or hole.

Theorem: Bloch's

For periodic potentials:

$$V(x) = V(x+a) \Rightarrow \psi(x+a) = e^{iKa}\psi(x)$$

Note: $|\psi(x)|^2 = |\psi(x+a)|^2 \Rightarrow |\psi|^2$ is periodic but the wavefunction is not the same for every atom in the lattice!.

Theory: O.E.

- If we add more electrons to a crystalline metallic solid, such that additional electronic states are occupied, the Fermi energy rises.
- The electronic density of states at the Fermi energy affects the ability of a solid to conduct electricity.

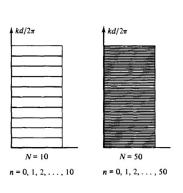


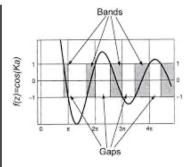
The potential is of the form: $V(x) = \alpha \sum_{j=0}^{N-1} \delta(x-ja)$.

Consider a one dimensionale solid of \overline{N} atoms. The allowed energies for the electron are:

$$cos(Ka) = cos(ka) + \beta \frac{sin(ka)}{ka}$$

where K is a real constant, a is the periodicity of the potential, $\beta = \frac{m\alpha a}{k^2}$, m is the particle mass and α is the strength of the Dirac comb.



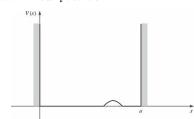


Since the LHS of the above equation is a cosine, it has to be between ± 1 . Also, the energy is proportional to k. Thus, the energy at the bottom of the lowest band will be at the point where the curve first met f(z) = 1. Note:

- For higher β , the bands exists even for larger ka (i.e) the strength of the delta function makes the band gaps arise.
- For larger ka, bands start vanishing.
- Though the bands are discrete ('n' is an integer) it has N number of solutions that makes the band forming a continuoum (For very large

Approximation methods

Problem: perturbation in our potential:



- 1. $\hat{H} = \hat{H}^0 + \lambda \hat{H}'$, \hat{H}^0 being the usual potential,
- 2. \hat{H}' the perturbation and λ an aiding parameter.
- 3. $\psi_n = \psi_n^0 + \lambda \psi_n^1 + \lambda^2 \psi_n^2 \dots$, $E_n = E_n^0 + \lambda E_n^1 \dots$ (when I calculate the error correction then the total Energy is $E = E^{(0)} + E^{(1)} + \dots$ with $E^{(0)}$ from the unperturbed Hamiltonian diagonal for example).
- 4. ψ^p_p, E^p_n being the corrections of order p. 5. ψ^0_n from ISW or from EV of unperturbed Hamiltonian \hat{H} (as example).

First and second order correction (can yield both to exact answer!):

•
$$E_n^1 = \langle \psi_n^0 | \hat{H}' \psi_n^0 \rangle$$
 \Rightarrow $\psi_n^1 = \sum_{m \neq n} \frac{\langle \psi_n^0 | \hat{H}' \psi_n^0 \rangle}{E_n^0 - E_m^0} \psi_m^0$

•
$$E_n^2 = \sum_{m \neq n} \frac{|\langle \psi_m^0 | \hat{H}' \psi_n^0 \rangle|^2}{E_n^0 - E_m^0} = \sum_{m \neq n} \frac{\langle \psi_m^0 | \hat{H}' \psi_n^0 \rangle \langle \psi_n^0 | \hat{H}' \psi_m^0 \rangle}{E_n^0 - E_m^0}$$

If $E_m = E_n \Rightarrow$ degeneration: the two identical E^0 get split by perturbation.

$$\begin{cases} E_{\pm}^{1} = \frac{1}{2} \left[W_{aa} + W_{bb} \pm \sqrt{(W_{aa} - W_{bb})^{2} + 4|W_{ab}|^{2}} \right] \\ W_{ij} = \langle \psi_{i}^{0} | \hat{H}' \psi_{j}^{0} \rangle, \quad \hat{H}' = \frac{\hat{H}(\lambda_{0} + \partial \lambda) - \hat{H}(\lambda_{0})}{\partial \lambda} = \frac{\partial \hat{H}}{\partial \lambda}, \lambda_{0} = \text{unperturbed state} \end{cases}$$

Theorem: Feynman-Hellman

Suppose the Hamiltonian H is a function of λ ; let $E_n(\lambda)$ and $\psi_n(\lambda)$ be the EV and EF of $H(\lambda)$ so:

$$\frac{\partial E_n}{\partial \lambda} = \langle \psi_n | \frac{\partial H}{\partial \lambda} | \psi_n \rangle$$

Boundary for E_{ground} . Can be a boundary for first excitation too. We choose any trial function and see (can yield to exact answer):

$$E_{ground} \le \langle \psi_{trial} | \hat{H} \psi_{trial} \rangle = \langle \hat{H} \rangle = \sum_{n=1}^{\infty} E_n |c_n|^2$$

Expectation value of \hat{H} is upper bound for E_{ground} . In practice we try to make $\langle \hat{H} \rangle$ as small as possible. E.g through adjustable parameters in our trial function. \Rightarrow derivation w.r.t that variable. For example for He atom, ψ_{trial} can be analog to 1s orbital $\rightarrow \Psi_{trial} = Ae^{-\alpha \cdot (r_1 + r_2)/a}$

Additions

Wave representation:

States	$ \Psi\rangle = \sum_{n} c_n f_n\rangle$
Basis	f_n
Operations	\hat{Q} differential operators
TISE	$\hat{H}\psi_n = E_n\psi_n$
Process	solve dif. equation
output	$\psi_n(\text{eig.funct.}) \text{ and } E_n \text{ (eig.value)}$

Matrix representation:

States	$ \Psi\rangle = \sum_n c_n f_n\rangle = (c_1, c_2,, c_n)^T$
Basis	f_n
Operations	matrices, elements: $\hat{H}_{ij} = \langle f_i \hat{H} f_j \rangle$
TISE	$\hat{H}\psi_n = E_n\psi_n$
Process	diagonalize \hat{H} (Eigenvalue procedure)
output	$\psi_n(\text{eig.vectors}) \text{ and } E_n \text{ (eig.value)}$

Dirac Notation:

 $\langle a|b\rangle$ is a complex number while $|a\rangle\langle b|$ is a square matrix. $|\psi\rangle = \sum_{i=1}^{n} c_i |f_i\rangle = \sum_{i=1}^{n} |f_i\rangle\langle f_i|\psi\rangle, \sum_{i=1}^{n} |f_i\rangle\langle f_i| = \mathbb{I}$

Symmetry comes for various operations:

- Translation
- Inversion

- Rotation
- reflection...
- 1. Translation:

$$\hat{T}(a)\psi(x) = \psi'(x) = \psi(x - a)$$

 $\hat{T}(a) = exp\left(-\frac{ia}{\hbar}\hat{p}\right) \Rightarrow \text{generated by momentum.}$

 \hat{T} is unitary (recall operators) $\hat{Q}'(\hat{x}, \hat{p}) = \hat{Q}(\hat{x} + a, \hat{p})$

Translation symmetry: $V(x) = V(x+a) \leftrightarrow \left[\hat{H}, \hat{T}\right] = 0$ (Linear momentum conserved!)

2. Parity:

$$\hat{\Pi}\psi(r) = \psi(-r)$$

Inversion of coordinates

 $\hat{\Pi} = \hat{\Pi}' \dagger$ (hermitian) $\hat{\Pi}^{-1} = \hat{\Pi}$ (unitairy)

 $\underline{\hat{Q}}'(\underline{r},\underline{\hat{p}}) = \underline{\hat{Q}}'(-\underline{r},-\underline{\hat{p}})$

Inversion symmetry: $V(\underline{r}) = V(-\underline{r}) \leftrightarrow |\hat{H}, \hat{\Pi}| = 0$

Implications: \hat{H} and $\hat{\Pi}$ share eigenstates (even states remain even and odd stav odd).

Selection rules whether photon can excite electron from ψ_{nlm} to $\psi_{n'l'm'}$

3. Rotations: $\hat{R}_z(\varphi)\psi(r,\theta,\phi) = \psi'(r,\theta,\phi) = \psi(r,\theta,\phi-\varphi)$

Here rotation around p-axis

$$\hat{R}_n(\varphi) = exp\left(-\frac{i\varphi}{\hbar}\hat{L}_n\right) \Rightarrow \text{generated by } \hat{L}_n$$

For small rotations: $\hat{R}_n(\delta) = 1 - \frac{i\delta}{\hbar}\hat{L}_n$

Continuous rotation symmetry: \leftrightarrow If \hat{H} is rotationally invariant \Rightarrow $\hat{V}(\vec{r}) = \hat{V}(r)$ (spherically symmetric). This implies $\Rightarrow [\hat{H}, \hat{R}_n(\varphi)] = 0$.

Implications: Since $\hat{R}_n(\delta) \approx 1 - \frac{\delta}{\hbar} \hat{L}_n \Rightarrow [\hat{H}, \hat{L}_n] = 0 \Rightarrow \frac{d}{dt} \langle \hat{L}_n \rangle = 0$. This mean that Angular momentum conservation is a consequence of rotational invariance.

For central potential eigenstates Ψ_{nlm_l} uniquely specified by $\hat{H}, \hat{L}_n, \hat{L}^2$. This is known as complete set of compatible observables. If only two operators, eigenstates are not uniquely specified.