1 a a k	Sund	2/-	
Lecture 2	3441.	<i>&amp;</i> _	

1.	Summany	a.P	Lectme	1

Planck's Law Wien's displacement land Stefan-Boltemann law

emissive power internity Sold angle.

2. Correction to intensity



$$I_{\lambda} = \frac{P_{\lambda}'}{dA_{\perp}d\lambda} I_{\perp} = \frac{P_{\lambda}'}{dA_{\alpha} \circ oDd}$$

EGIN THOU

Pr = In asodAdada

PA = Sa In wodAdada

$$E_{\lambda} = \frac{\rho_{\lambda}}{\partial AA} = \int_{\Omega} J_{\lambda} d\sigma s o d\Omega$$
.

3. Comments

I k a scalar

=> distoibuti functi

f(x,y,2, 2x, Uy, Vz, t) phase space

b. How to make a blackbooky

isothermal, diffue

blackbody also best absorber

Hohlraum

T black area

Lecture 2 2/1 dAp = R2 Sino do do ds= sinodody  $\Delta = \int_0^{\frac{\pi}{2}} d\theta \int_0^{2\pi} \sin \theta d\theta = 2\pi$ Stop 41 Blackbody In = Constant (maximum)  $E_{bn} = \frac{\int P_{\lambda}/d\Omega}{dAdn} = \int_{\Omega} I_{bn} d\theta s \theta d\Omega$ = In So disonodo S. B. dg dAdrda Jan 650 Lambert Low (Surface) 4. Radiation fressure

Einstein:  $p = \frac{h}{\lambda}$ mo the & the photon per mit time 10 P. 1/2 1000 DPD  $\frac{I_{\lambda}}{dAd\lambda} = I_{\lambda} \cos \theta d\Omega$ Incosod2

Normal momentum 
$$\int_{\Omega} \frac{h}{\lambda} \cdot \frac{I_{\lambda} \log dx}{h\nu} \cdot \cos \theta =$$

$$= \frac{I_{\lambda}}{c} \int_{0}^{\frac{a}{2}} G_{\lambda}^{2} \operatorname{siddo} \int_{0}^{2\lambda} dy$$

$$= \frac{2\pi I_{\lambda}}{c} \cdot \frac{1}{3} = \frac{2\pi I_{\lambda}}{3\epsilon}$$

$$= \frac{2}{3c} E_{\delta\lambda}.$$

Solar Soul

5. Visible rachation Human.

Certain fraction of intensity is observed as light

Luminous Intensity  $L_{\pi} = k_{\pi} I_{\pi}$ Turninous efficacy.  $\frac{L_{m}}{W}$ 

AZ DI DAL

[ lun m2 sr um]

L normal to ray

( to 8 addition of

kmax = 683 hor/W

Kmax = 683 hor/W

04 0.565 0.7 µm

Total
$$L = \int_{0}^{\infty} k_{\Lambda} I_{\Lambda} d\Lambda \approx 286 \frac{k_{00}}{W} \int_{0.4}^{0.7} I_{\Lambda} d\Lambda$$

$$60 W \Rightarrow 840 km \quad bight bulb \Rightarrow \int_{0.4}^{0.7} I_{\Lambda} d\Lambda = 193W \Rightarrow 5\%$$

## 6. Properties of Surfaces (ch. 1 & 3) Emissivity energy emitted from a surface energy emitted by a black hot surface at some temp Ex time directional spectral directional En spectral hemisphenial total, directional total, hemisphenical $e_{\lambda}' = I_{\lambda}(0.9,T) dA GOSO dA d\lambda d\lambda$ Ibr(0.9.T) asodAd>d2 $\epsilon_n = \int I_{\lambda} \cos d\Omega = \int_{\Delta} \epsilon_{\lambda}' I_{b\lambda} \cos d\Omega$ SIDA WOOD dr TIBA 走 Sen' asod 12 = Similarly E'= 15T4 Jo E' EDD (T) d) $E(T) = \frac{1}{n^2 6 T^4} \int_0^\infty e_{\lambda}(T, \lambda) E_{b,\lambda}(T, \lambda) d\lambda$ gray emitter: $\epsilon_{\lambda}' = \epsilon'$

diffuse emitter  $\xi \chi = \xi \chi$ diffuse-gray emitter:  $E(T) = E_{\lambda} = E_{\lambda}' = E'$ 

Absorptanty

$$\alpha = \frac{absorbed energy}{incorp energy}$$

Tranomissindy

Define Incoming Ruchation power

$$P_{\lambda} = I_{\lambda} \cos \theta_{i} dA d\lambda d\lambda d\lambda.$$

$$I_{\lambda} (T_{s}, \theta_{i}, \theta_{i}, \lambda)$$

source temp or tower

Reflectively & Travemissivity

bi-directional reflectively

incident reflected

$$S_{n}^{"}(\lambda,\hat{S}_{i},\hat{S}_{r}) = \frac{d \ln \Omega,\hat{S}_{i},\hat{S}_{v}}{\ln \Omega,\hat{S}_{i},\hat{S}_{v}}$$
 $I_{\lambda}(\lambda,\hat{S}_{i},\hat{S}_{r}) = \frac{d \ln \Omega,\hat{S}_{i},\hat{S}_{v}}{\ln \Omega,\hat{S}_{i},\hat{S}_{v}}$ 

directional - hemisphenical reflectively
$$S_{n}^{1D} = \frac{\int_{\Omega} d \ln (\hat{s}_{i}, \hat{s}_{r}) \cos \sigma d \Omega_{r}}{H_{n}^{\prime} d \Omega_{i}}$$

= J.z 5" (7. si. sr) dosord 2r = T. 97"
L'if diffuse reflection other ga' - spectral temispherial directional 8 definitions: 9", 914, 901, 92 8", 910, 901, 9 Tranomiosituly - Similar energy balance absorpti.  $S+\alpha+\tau=1$  $S_{\lambda}^{\prime A} + N_{\lambda}^{\prime} + T_{\lambda}^{\prime A} = 1.$ opaque 7'' = 0S+ × = 1. we will learn how to calculate I &X for ideal surfaces. Prostant Diffuse absorber gray absorber
Diffuse reflector

## 7. Kirchoffls Law

EX(T)= VX(CT)

Sample temperative

Not Rice vigorous prof black area T only man only  $\in_{\lambda'}$   $\times_{\lambda'}$  not zero reflecting

Sample T. all other wave both le guide zero.

The only way to exchange energy is between these de - g . Q . .

Brenzy balance

(2) I by cose ded John cose ded John

 $\forall \lambda' = \in \lambda'$ Principle of detasted bolonce

2nd law of Hornodynamics.

Diffuee-gray Rumface, simple cases

diffuse-gray emitter: \$\$\$\$\$\fixe-gray reflector: \$\$\$\$\$\$\$\$\$\$\$\$\$\$\$\$\$\$

Opaque

$$S + \alpha = 1$$

$$\alpha = F = \epsilon$$

$$E = \Sigma E_{b}$$

Diffuse-gray surfaces, view-factor.

$$I(\vec{r}_i) \text{ isotropic}$$

$$I = \frac{P'}{dA_{i,1} dS_{i,2} dA_{j,1}}$$

P' - all the energy in  $\frac{dA}{R^2}$  the Solid ayle  $P' = I dA_i \cos 0i \frac{dA_7 \cos 0}{R^2}$ 

Total energy leaving dAi

TIDA:

Ratio

Cosoicoso; de dA;

"Intercepted + absorbed"