

# Mining FIR data with SEDSTACK

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w/

Lorenzo Moncelsi (Caltech), Ryan Quadri (Texas A&M),  
Jason Sun (Caltech), and the HerMES Collaboration

# A workshop?

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- Why are we here? I don't know. So I came up with three parts to this talk, or if you like, three questions:
  1. Moving beyond simple colors to divide galaxy samples, e.g.;
    - ▶ Model SEDs
    - ▶ AGN component, 24um flux, etc.,
  2. Moving beyond single field / single band treatment of stacking algorithms and into more sophisticated treatments, which I unfortunately named;
    - ▶ SEDSTACK
    - ▶ FLUCTFIT
  3. Explaining the weird behaviors that I have found in my stacking work, namely;
    - ▶ The extreme luminosities of high-redshift “quiescent” galaxies

**All codes are open-source, collaboration welcome!**

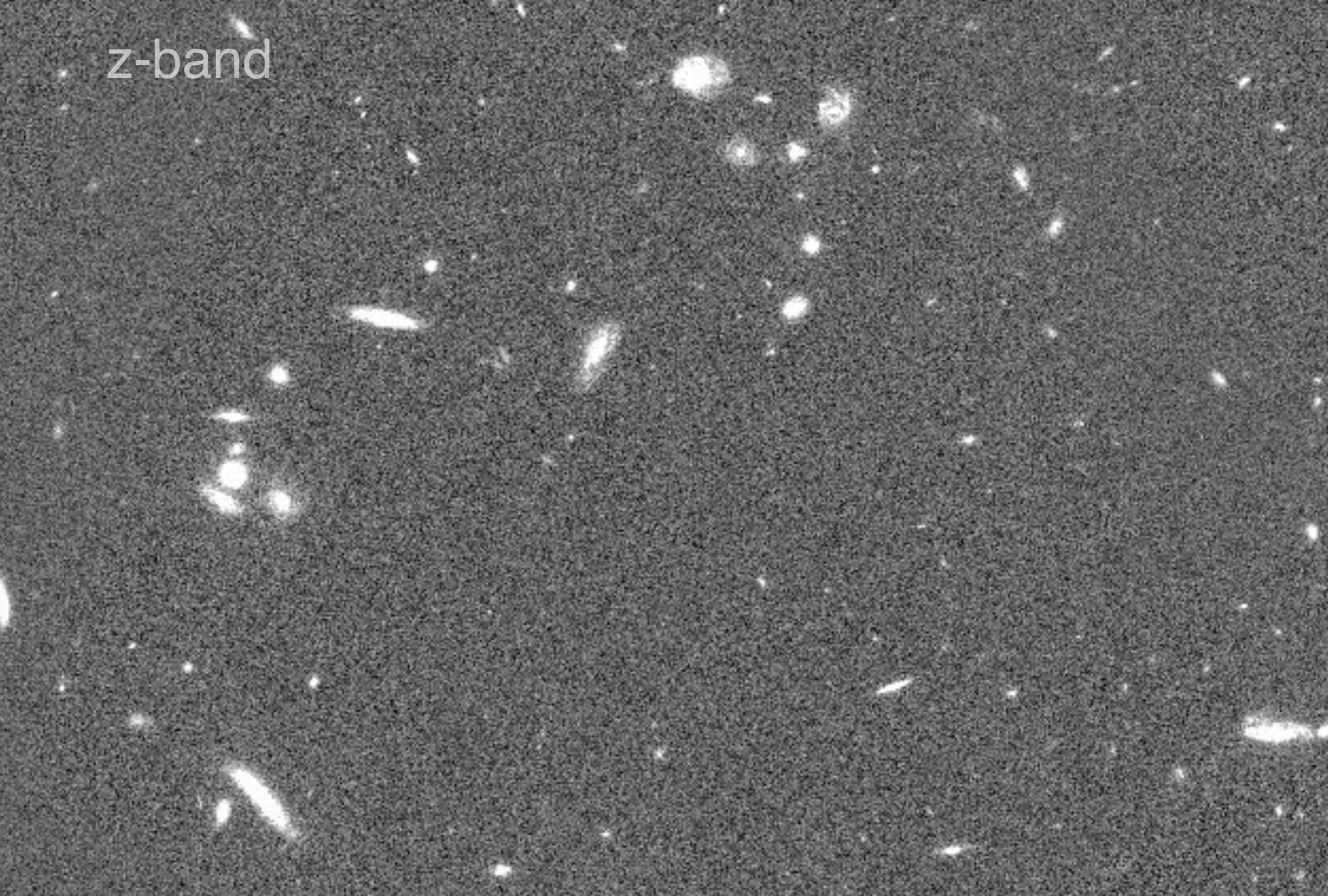
**Python (new) – <https://github.com/marcoviero/simstack>**

**[IDL (old) – <https://web.stanford.edu/~viero/downloads.html>]**

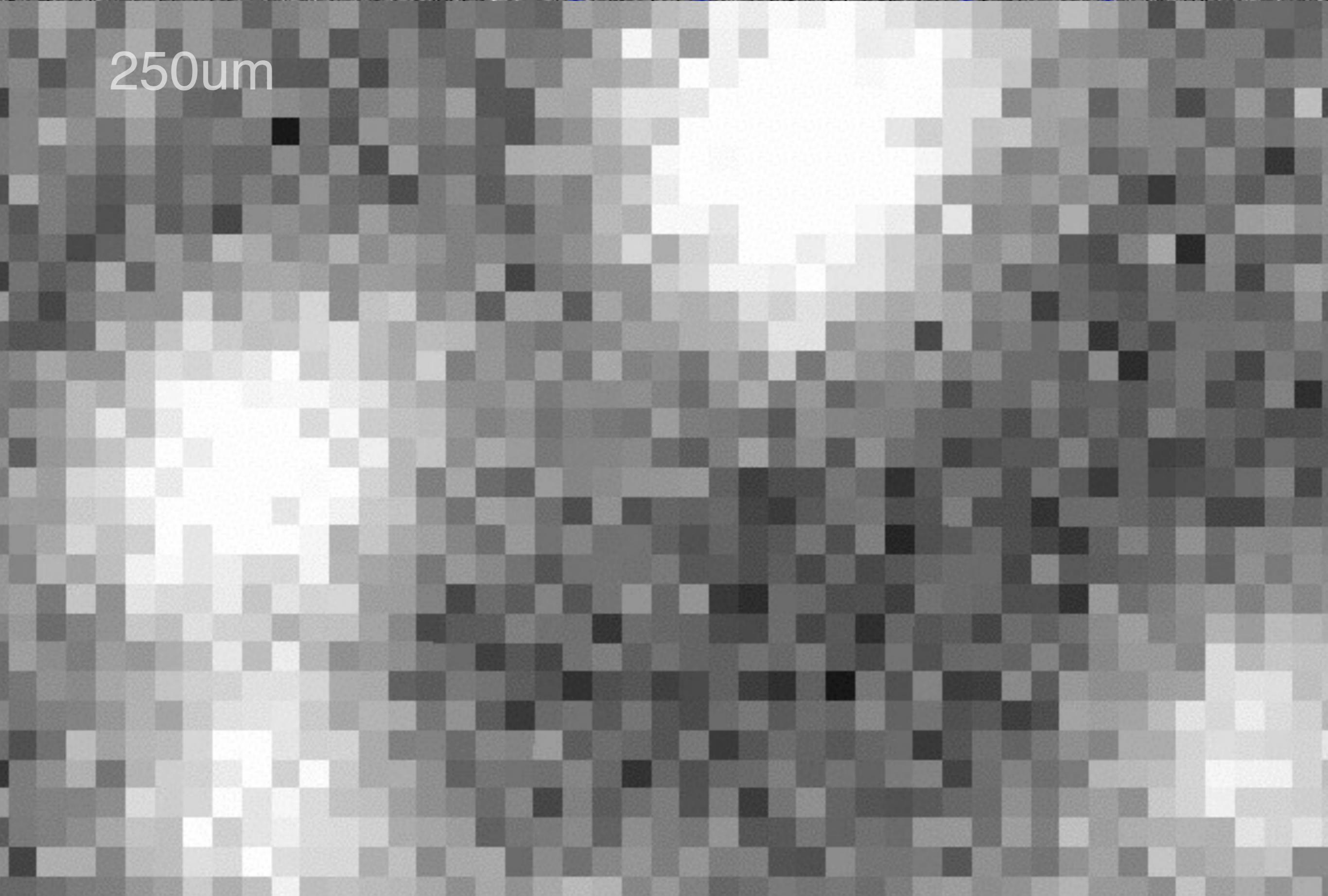
# SIMSTACK

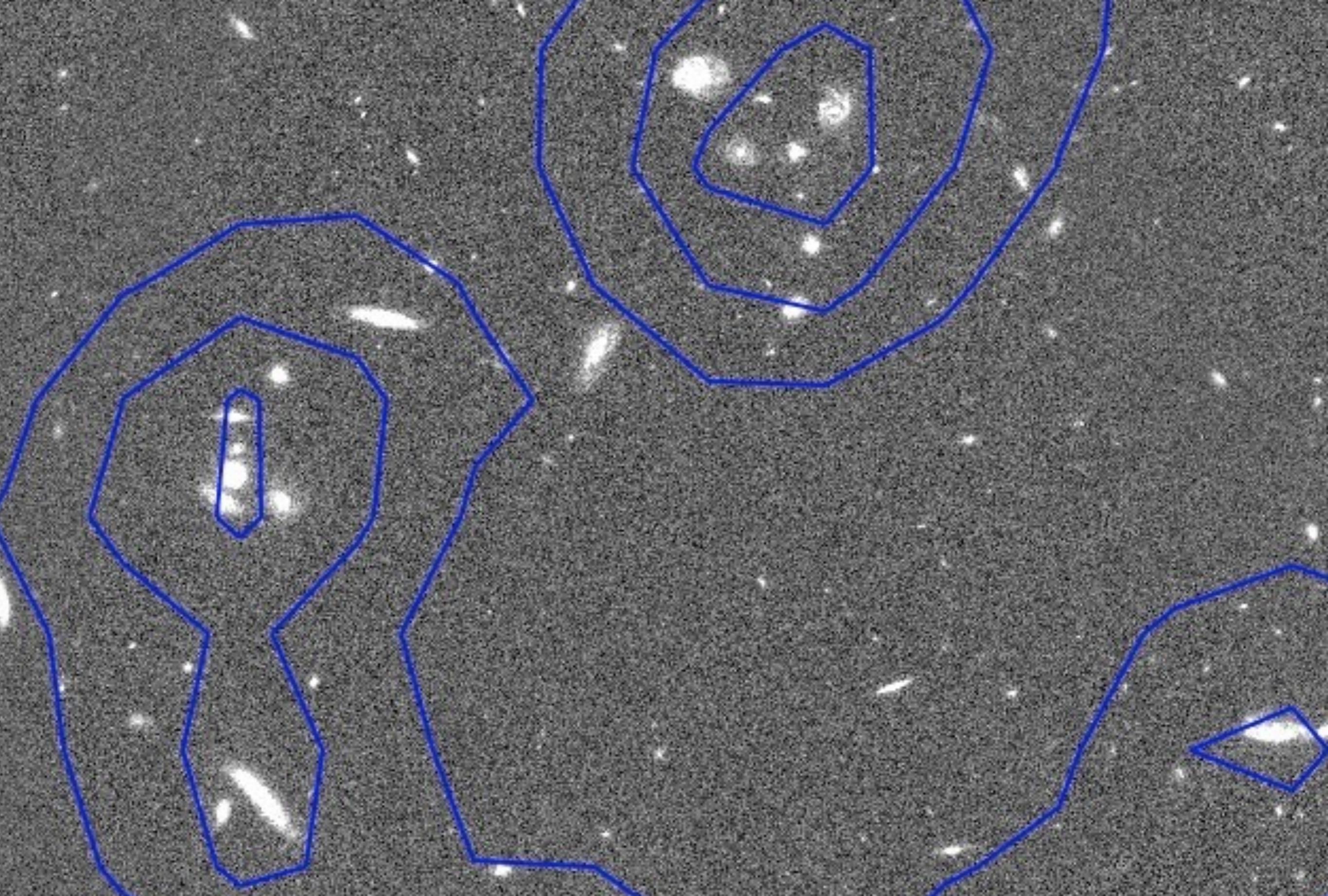
First, an introduction

z-band

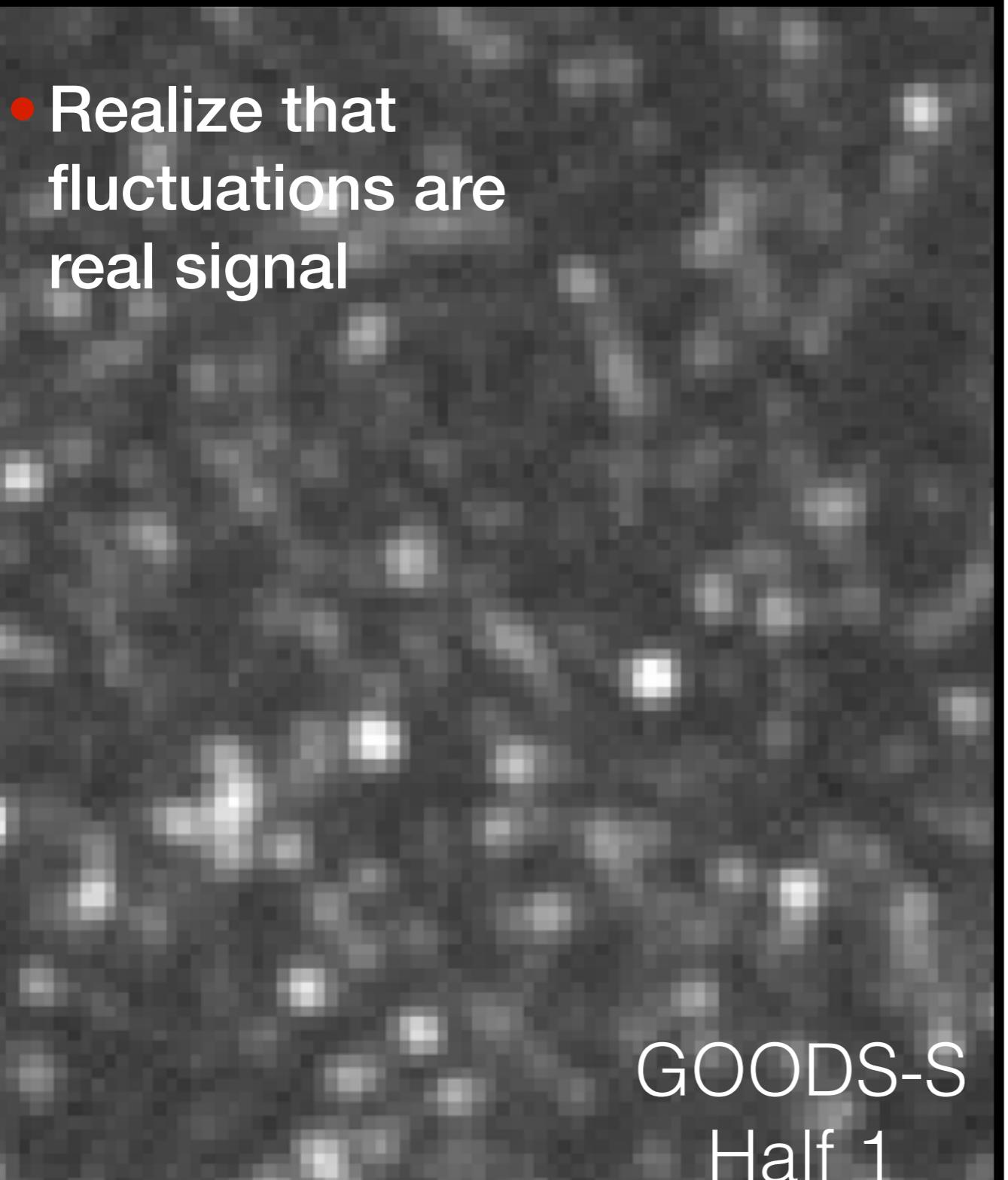


250um

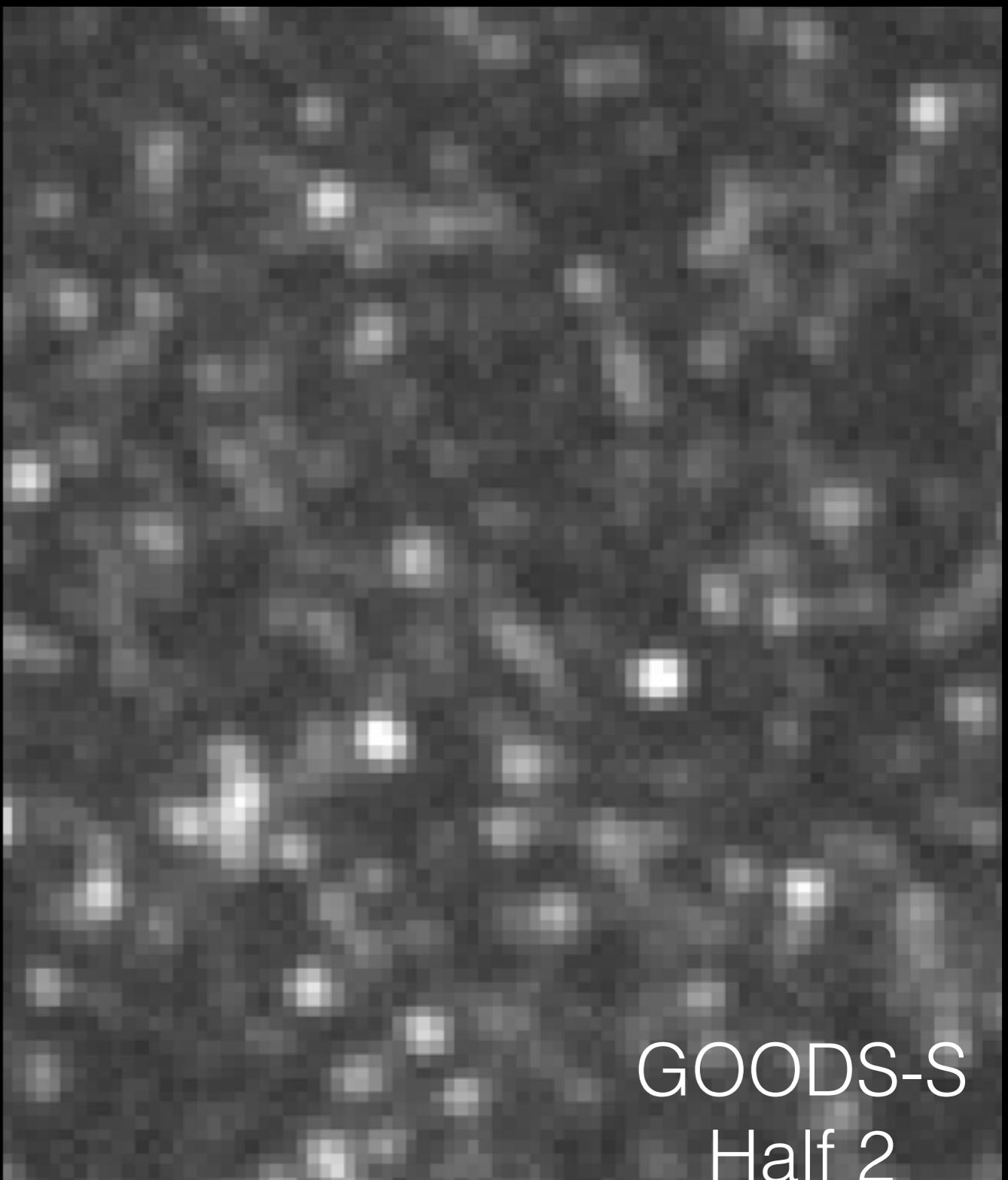




- Realize that fluctuations are real signal



GOODS-S  
Half 1

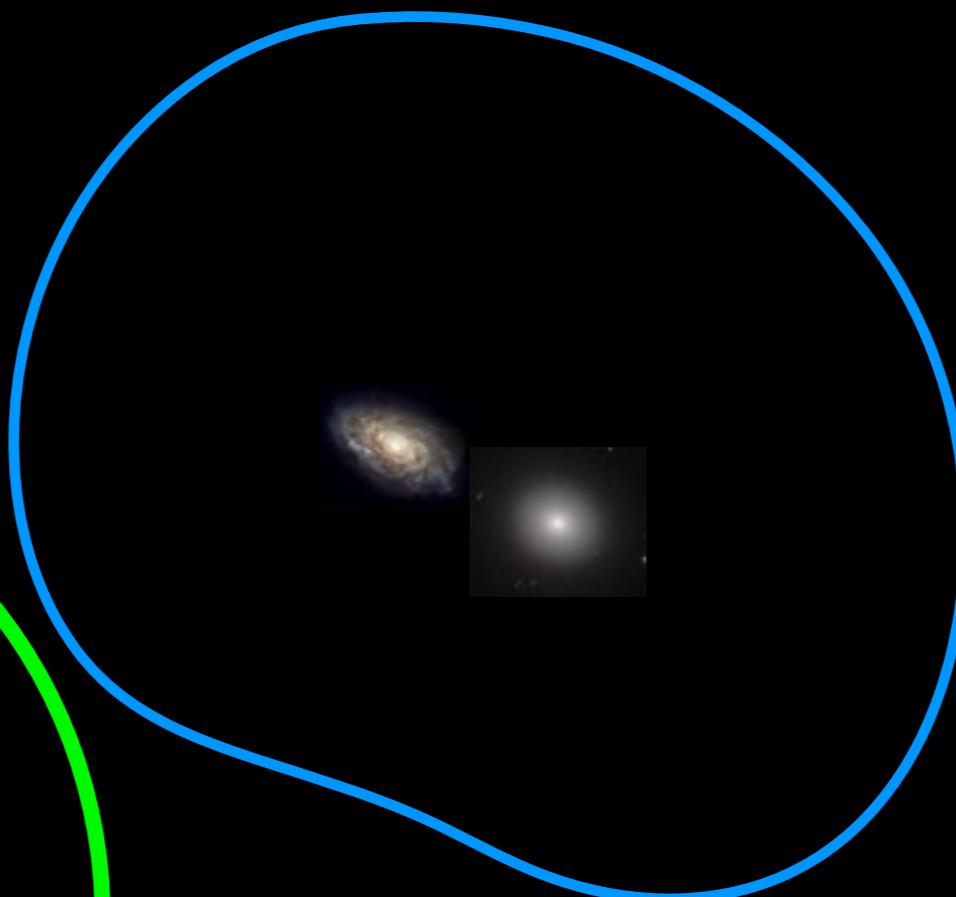
A grayscale astronomical image showing a field of galaxies in the GOODS-S field. The galaxies vary in size and brightness, with some appearing as small points and others as larger, more extended structures. The background is dark, indicating the void between galaxy clusters.

GOODS-S  
Half 2

A grayscale astronomical image showing a field of galaxies in the GOODS-S field. This image appears slightly darker or noisier than the first one, possibly representing a different filter or a different portion of the same field. It shows a similar distribution of galaxies.

SPIRE Contour

SPIRE 250 $\mu$ m  
18" Beam



- Difficult to attribute an individual submillimeter “source” to any single galaxy

SPIRE Contour

SPIRE 250 $\mu$ m  
18" Beam

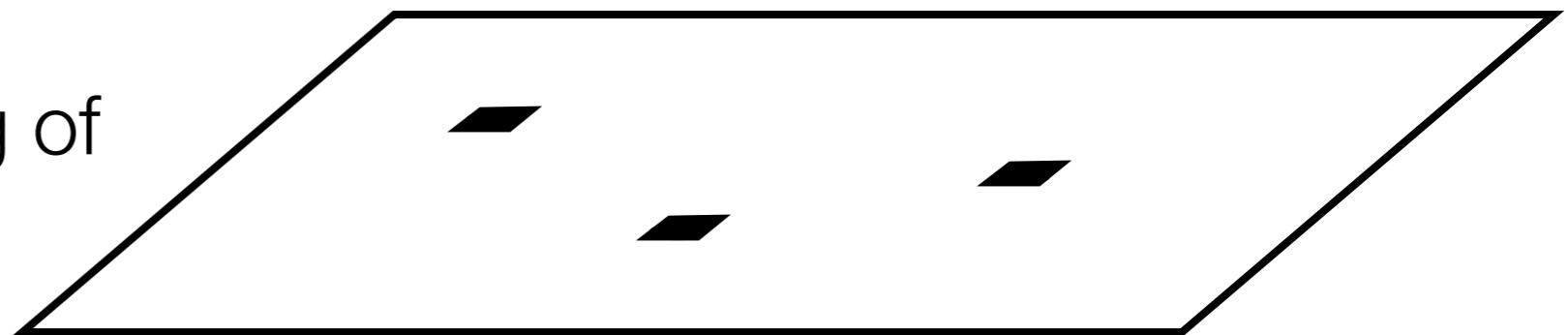


- Key is to identify galaxies with similar *physical* properties, and then rely on **statistics** to fit **fluctuations**

# SIMSTACK: Synthetic Intensity Fitting Algorithm

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make hits map from catalog of similar objects



Formalism developed w/ Lorenzo Moncelsi (Caltech);  
also see Kurczynski & Gawiser (2010), Roseboom et al. (2010)

**SIMSTACK code publicly available (see arXiv:1304.0446):**

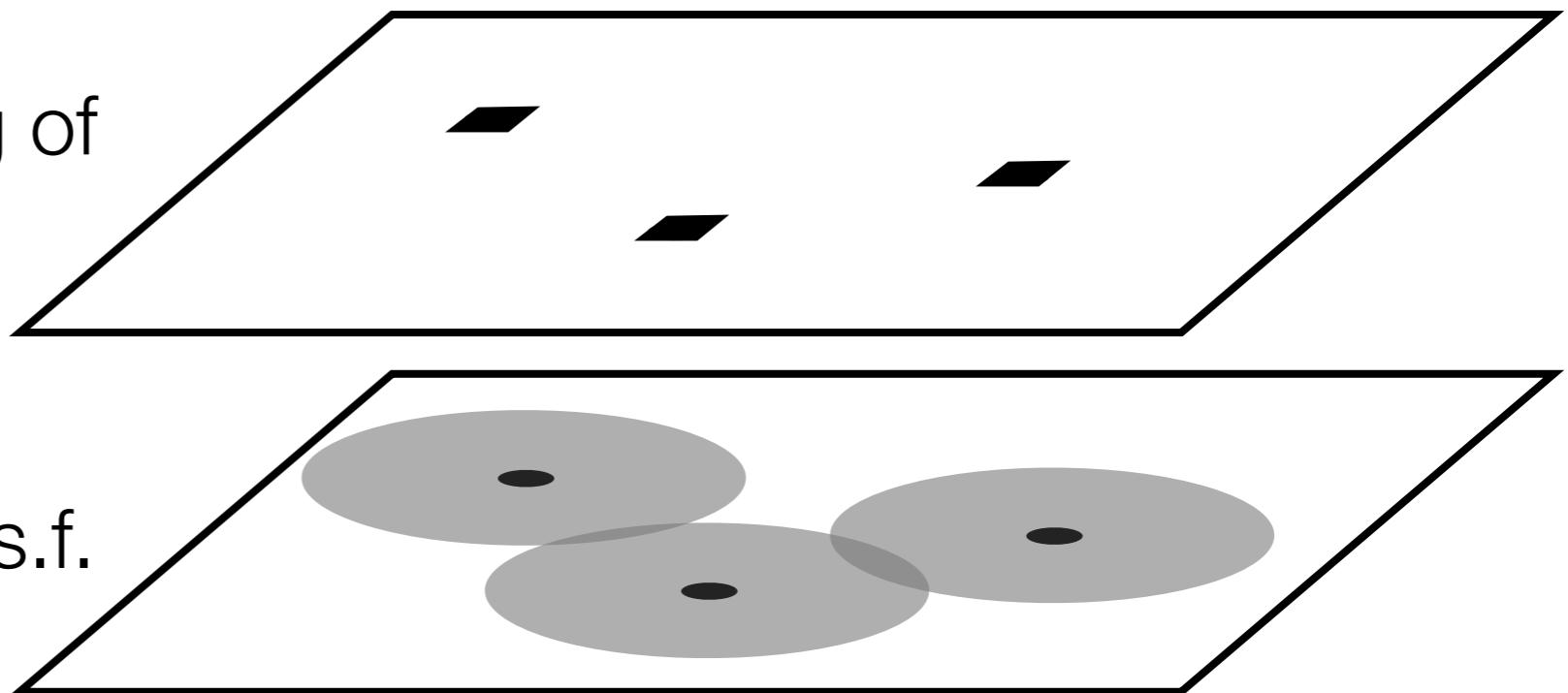
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convolve with instrument p.s.f.

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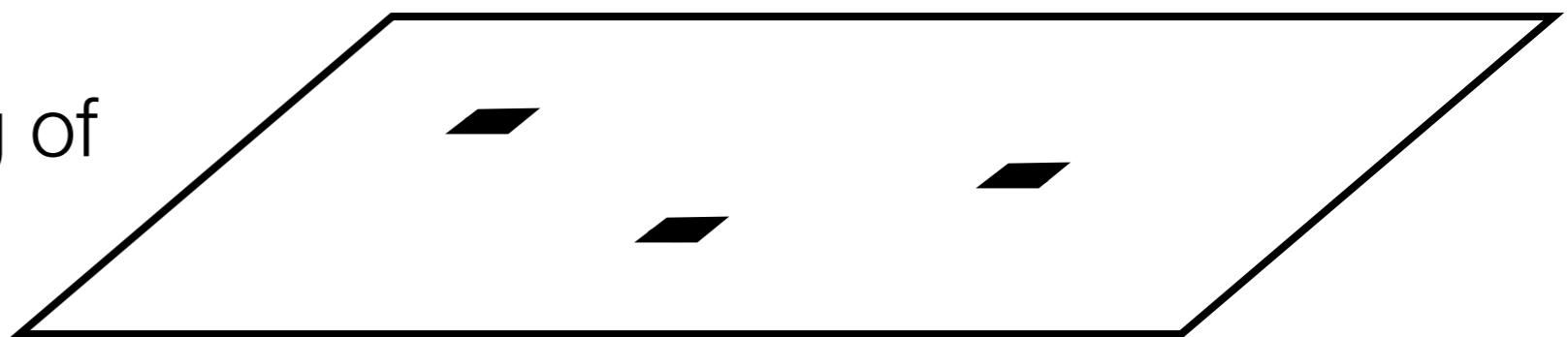
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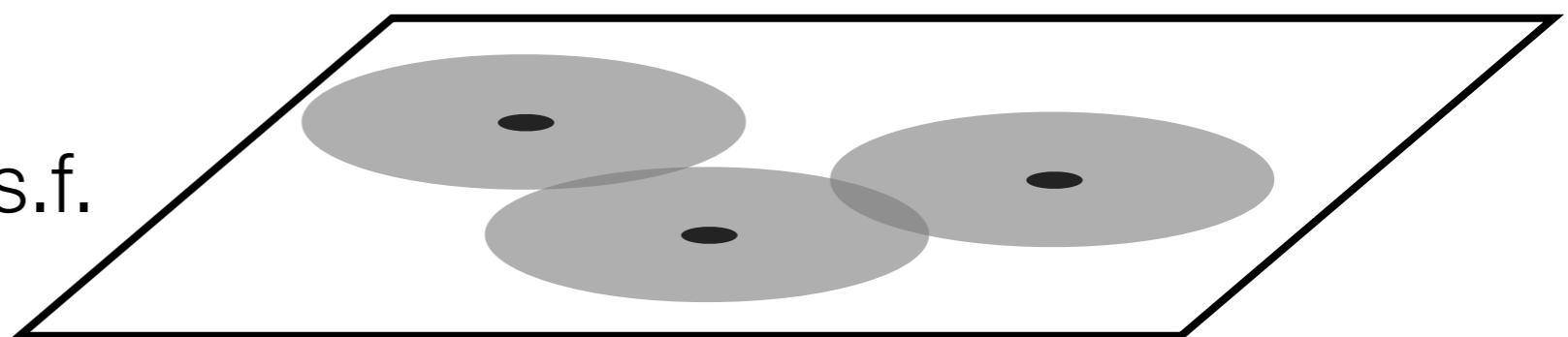
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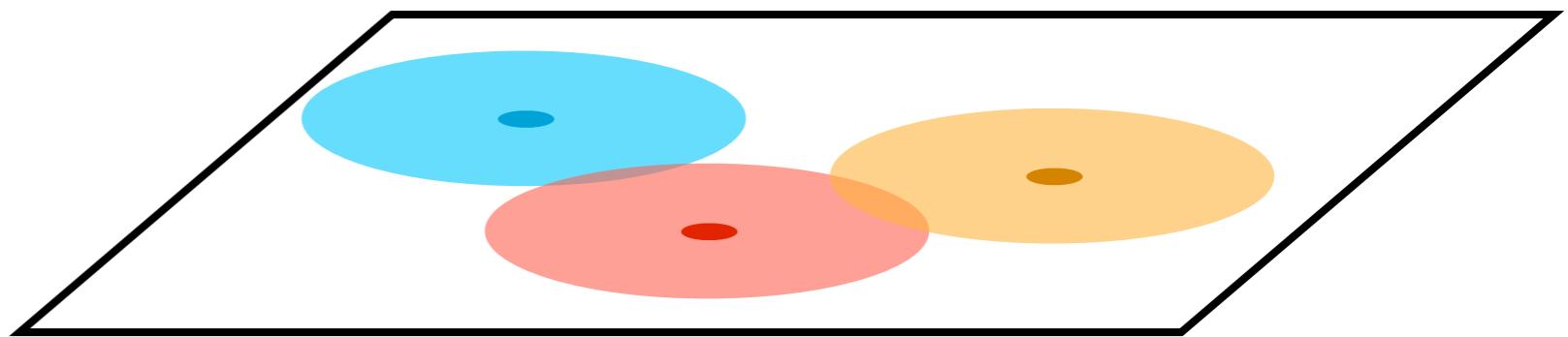
make hits map from catalog of similar objects



convolve with instrument p.s.f.



regress to find *mean* flux density



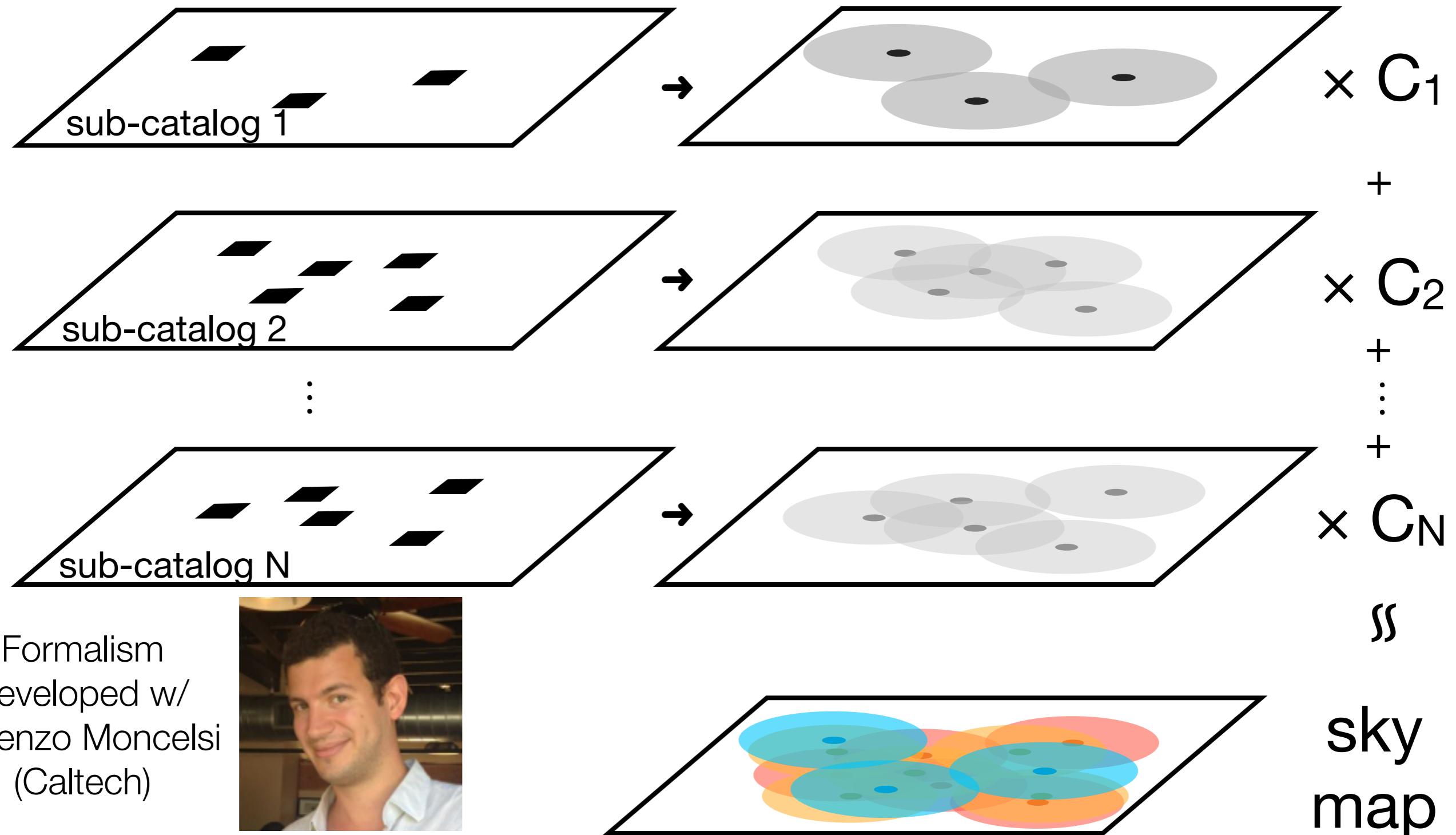
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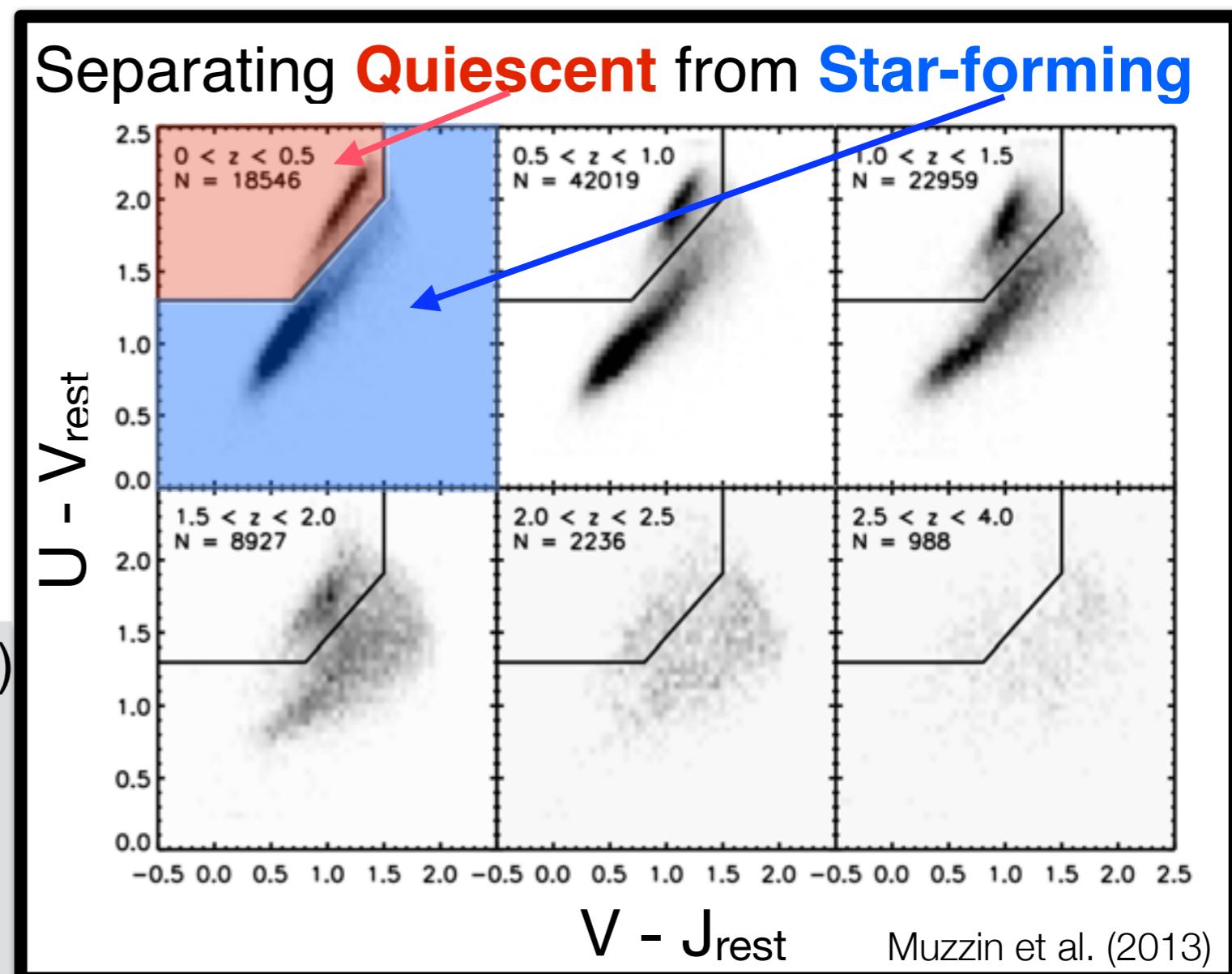
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## Catalogs

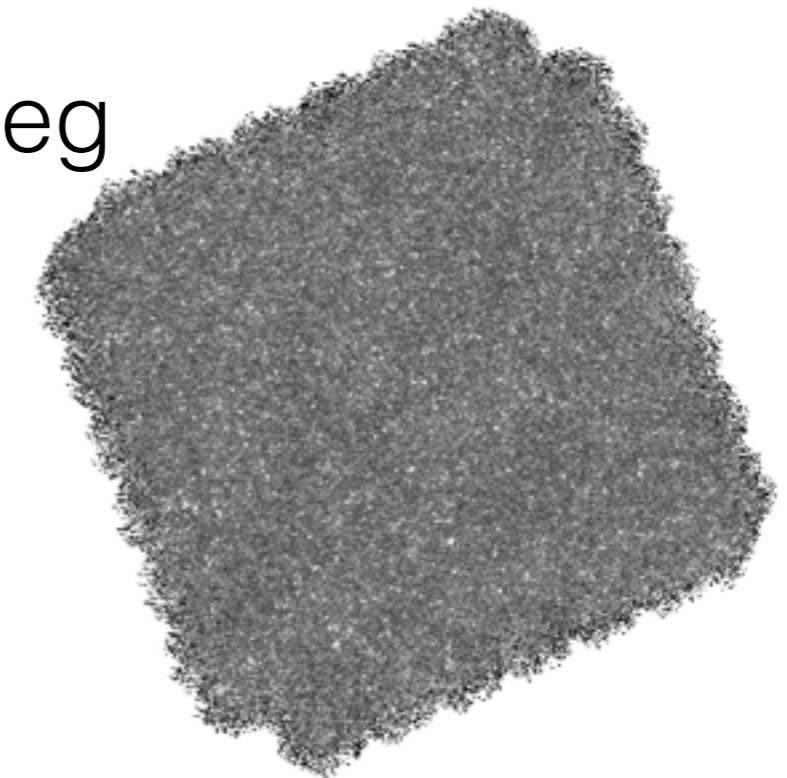
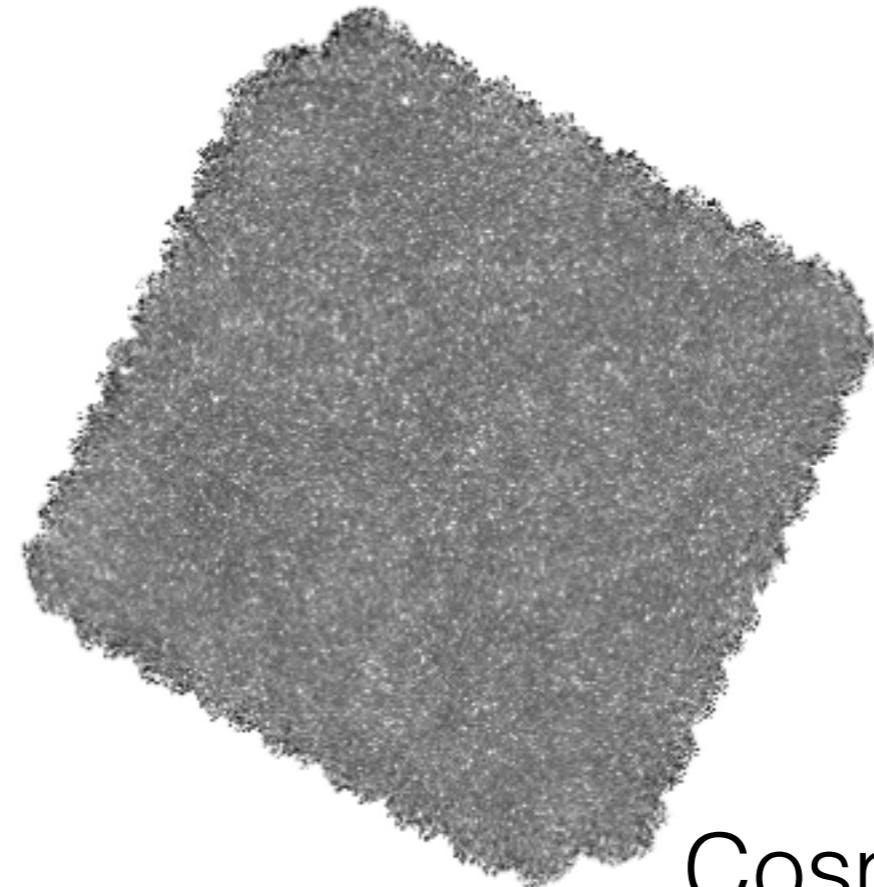
- UKIDSS/UDS [2/3 deg<sup>2</sup>] / COSMOS [1.6 deg<sup>2</sup>]
  - uBVRizJHK + IRAC ch1234
  - K-band cut 23.4 / 24 AB
  - 80,000 / 120,000 sources
- **Redshifts** - EAZY (Brammer 2008)
- **Masses** - FAST (Kriek 2009)
- **Colors** - UVJ (Williams 2009)



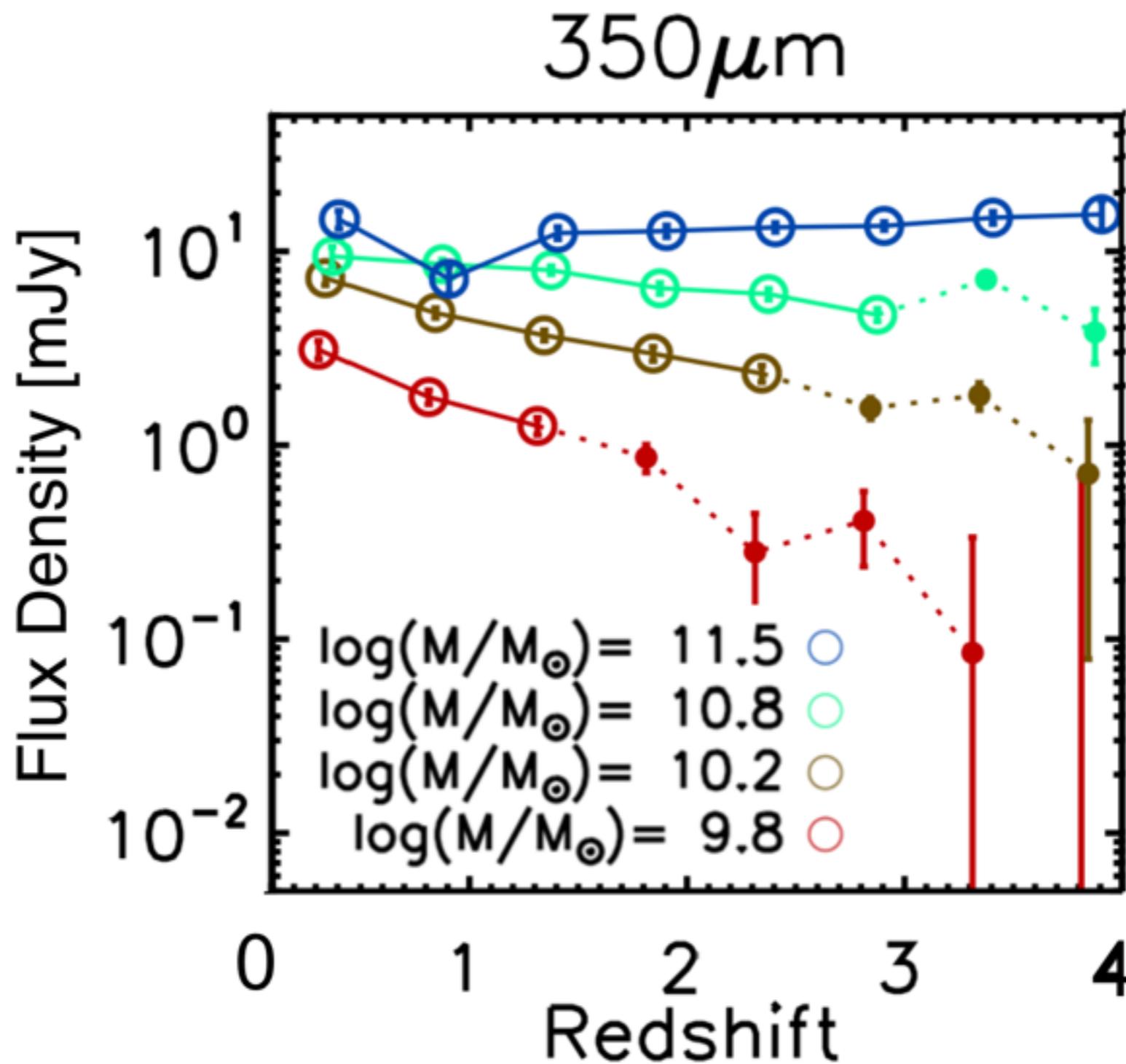
## Maps

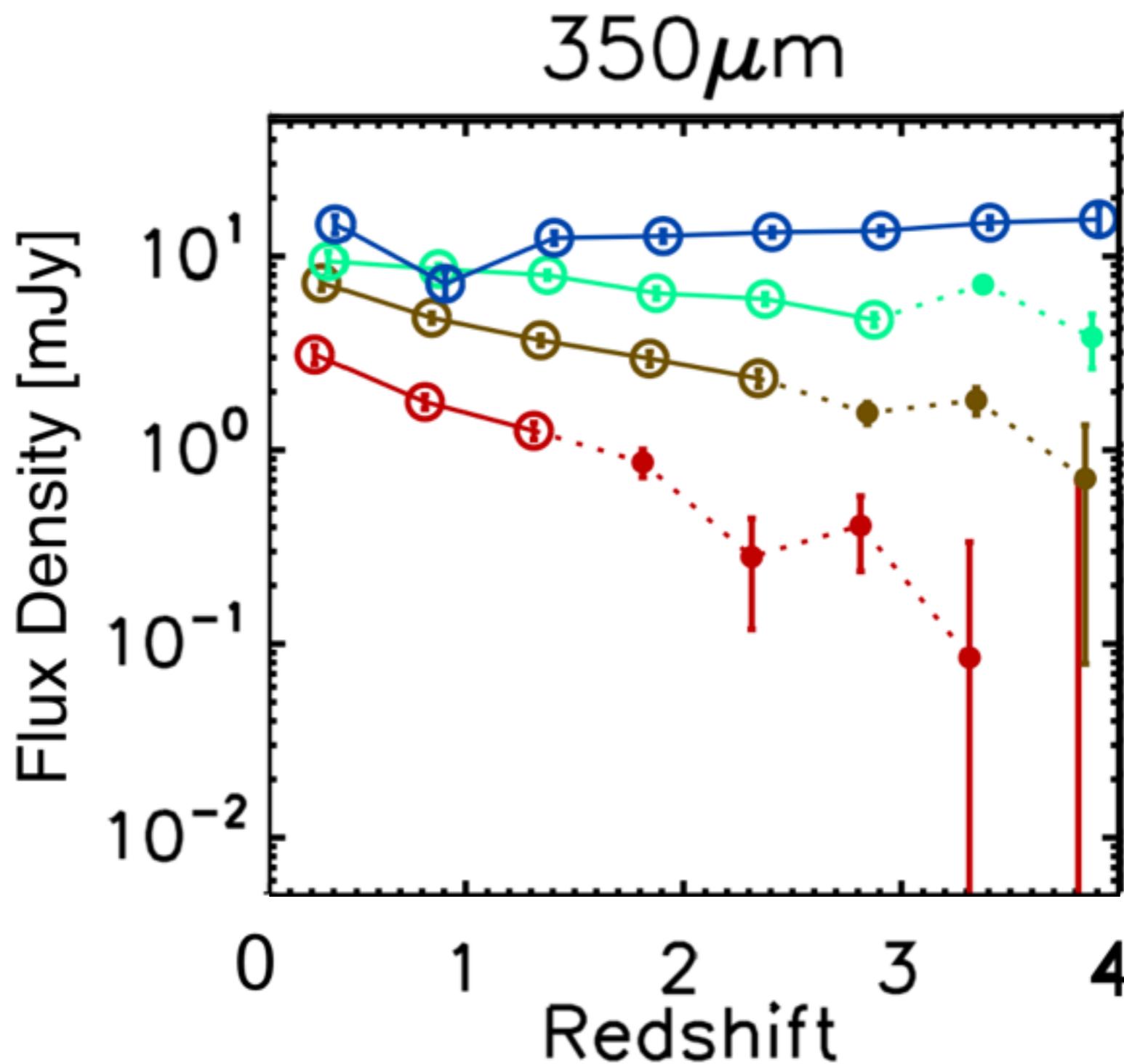
- *Spitzer/MIPS*
  - 24, 70 $\mu$ m
- *Herschel/PACS*
  - 100, 160 $\mu$ m
- *Herschel/SPIRE*
  - 250, 350, 500 $\mu$ m
- ASTE/AzTEC
  - 1100 $\mu$ m

UDS - 1.4 x 1.4 deg



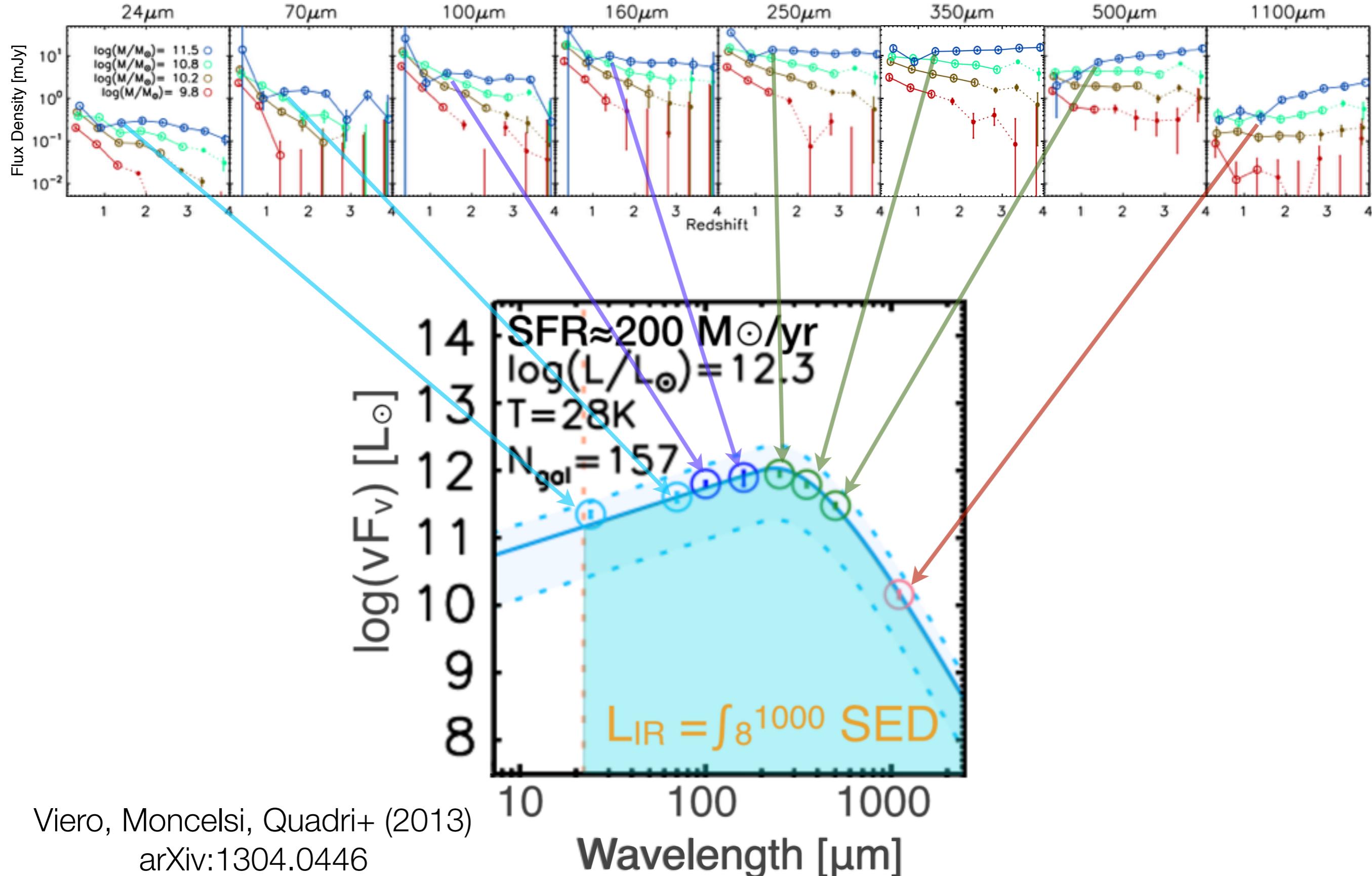
Cosmos - 1.8 x 1.8 deg





# SIMSTACK: Flux Densities (M,z)

 HERMES

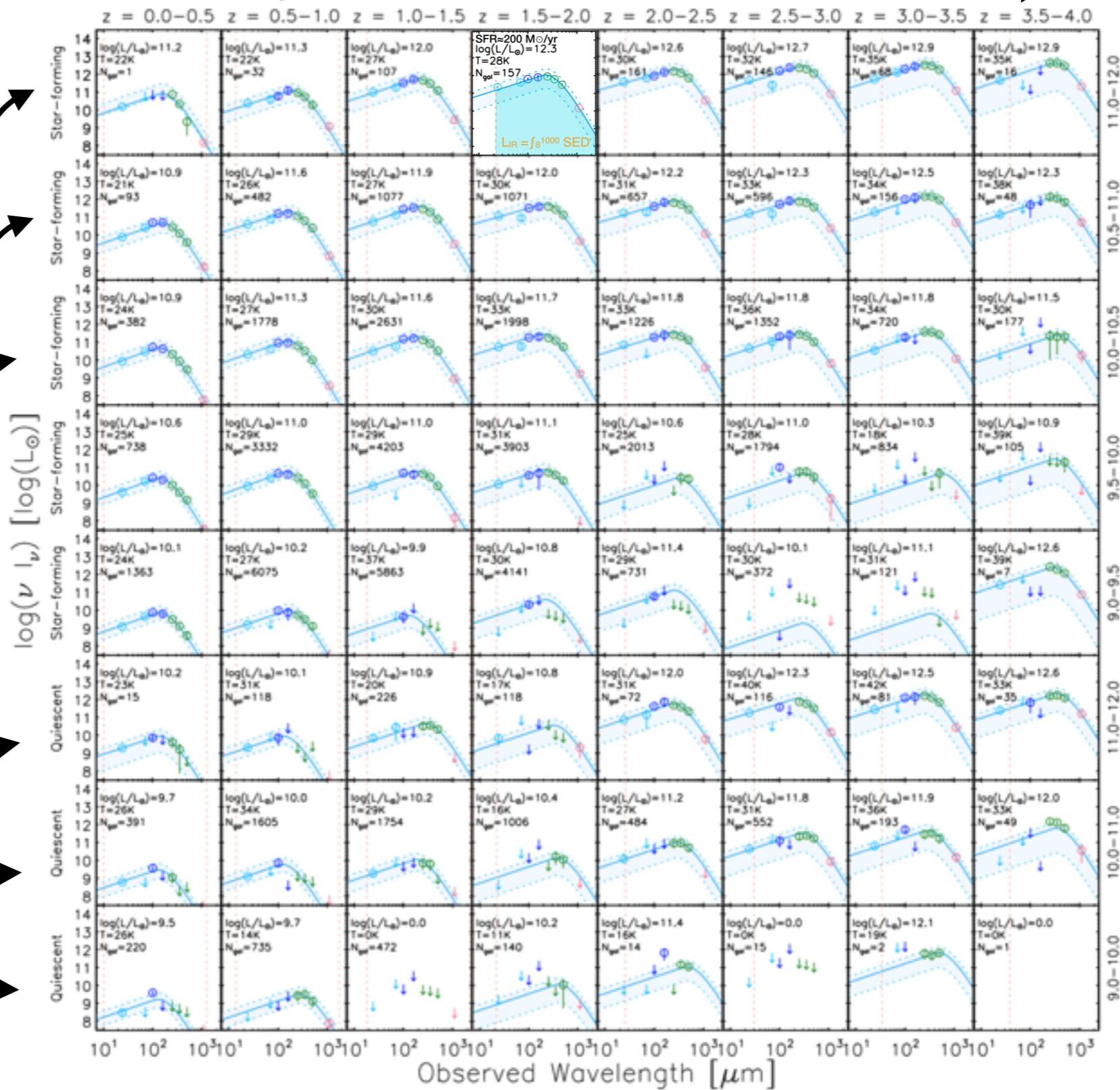


Viero, Moncelsi, Quadri+ (2013)  
arXiv:1304.0446

# SIMSTACK: SEDs

stellar  
mass  
slices

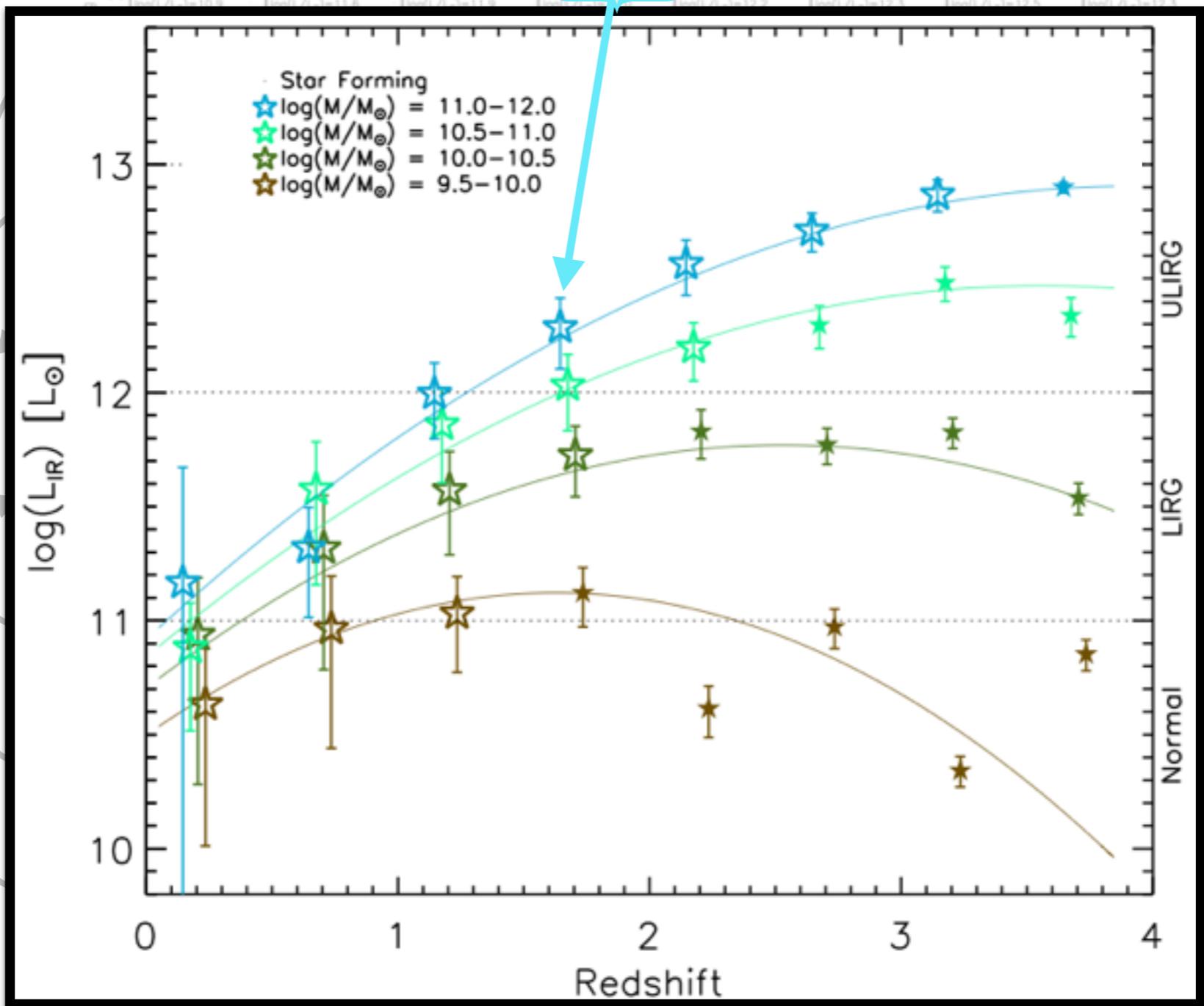
redshift  
slices



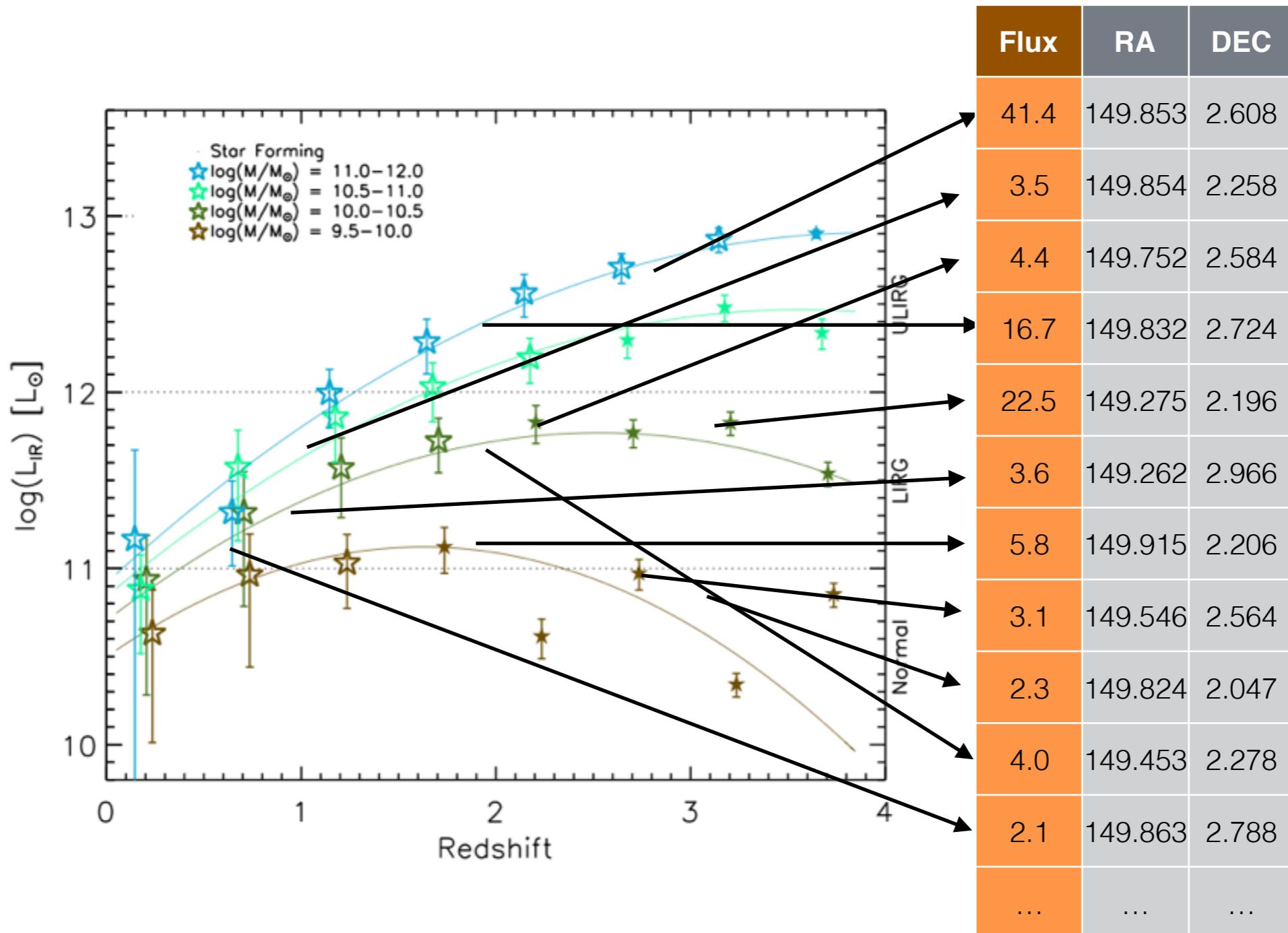
# SIMSTACK: $L_{\text{IR}}(M, z)$

stellar  
mass  
slices

redshift  
slices



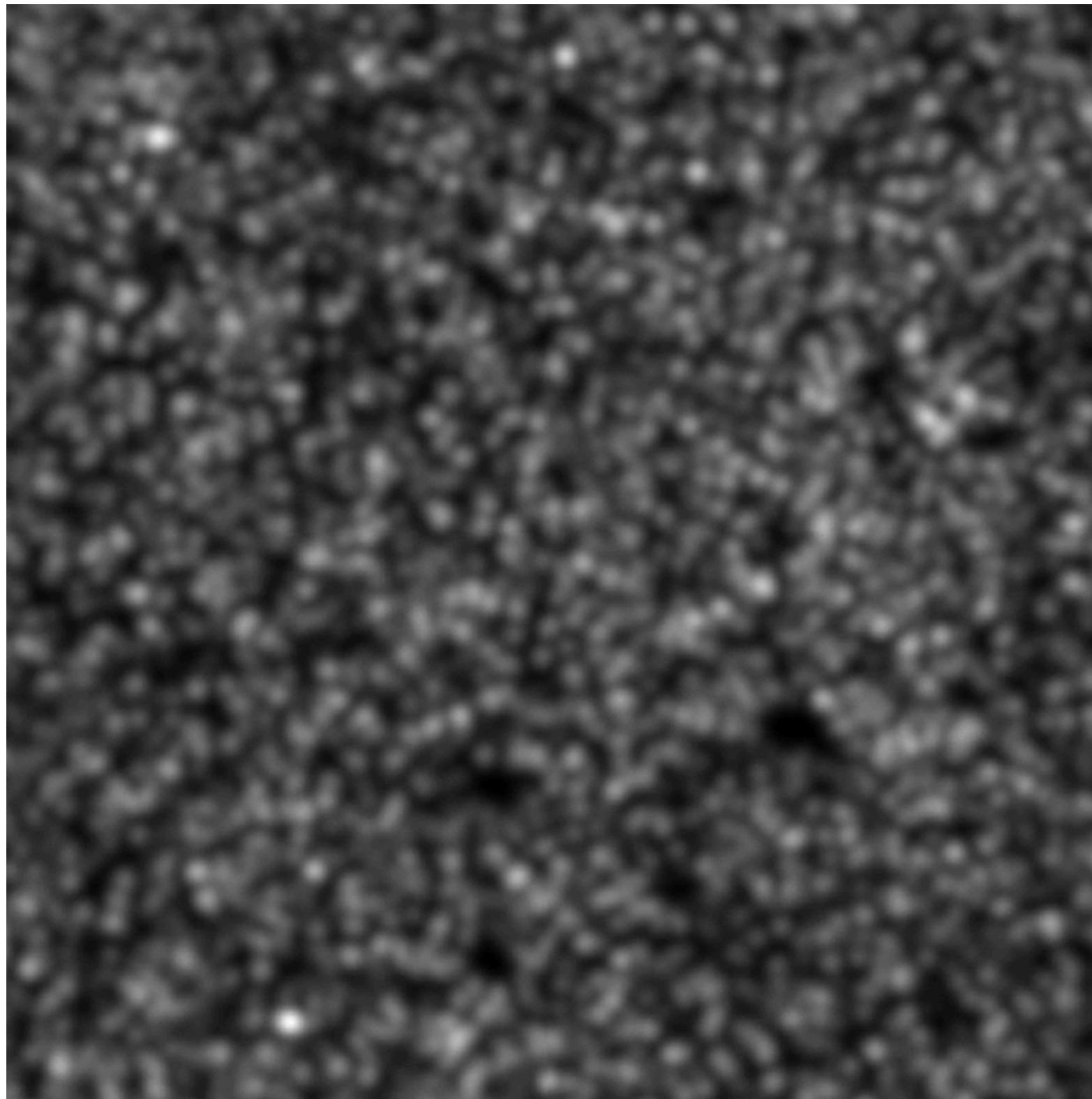
# SIMSTACK: coming full circle



Viero, Moncelsi, Quadri et al. (2013)  
arXiv:1304.0446

# SIMSTACK: simplest results

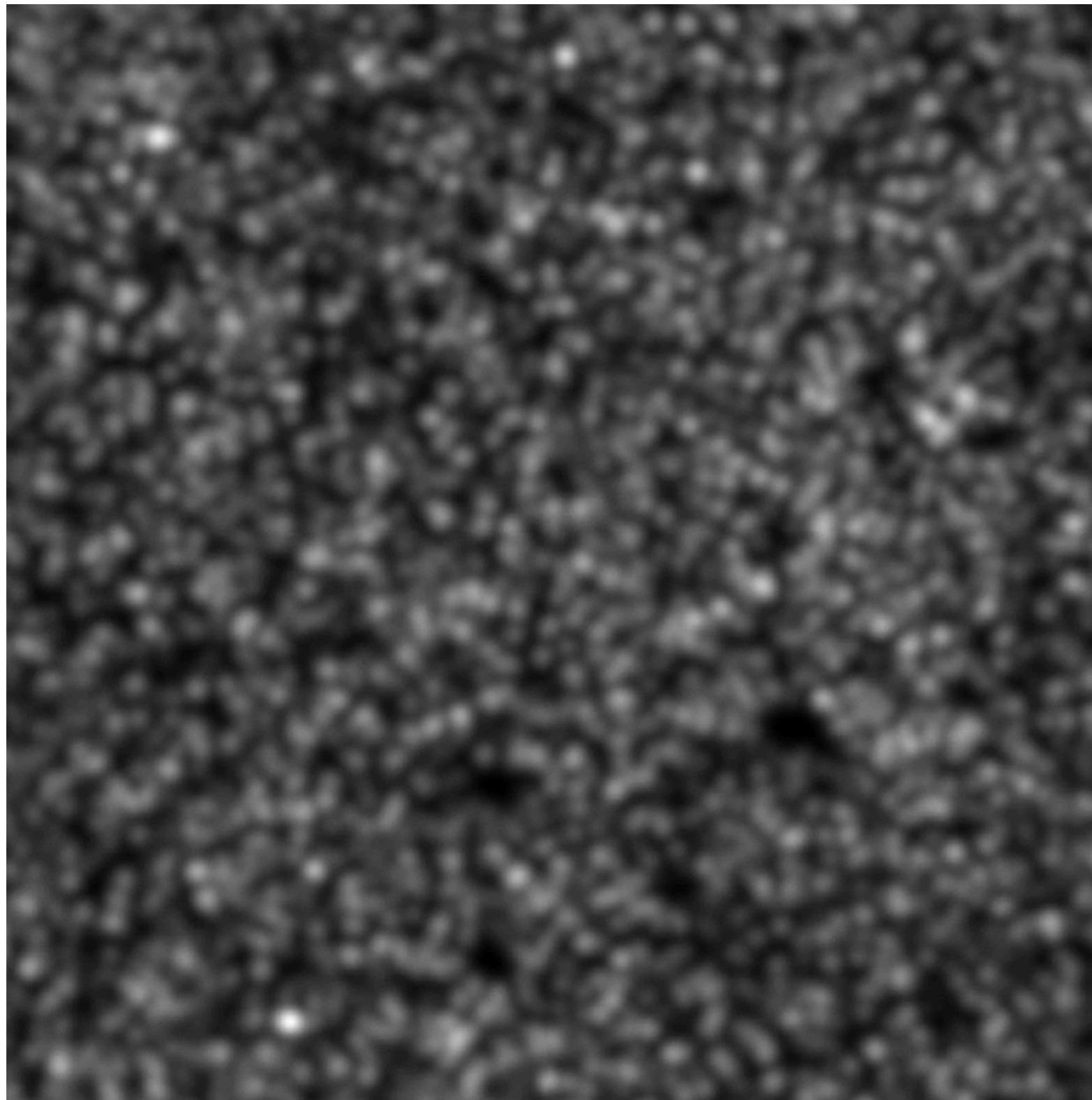
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Flux	RA	DEC
41.4	149.853	2.608
3.5	149.854	2.258
4.4	149.752	2.584
16.7	149.832	2.724
22.5	149.275	2.196
3.6	149.262	2.966
5.8	149.915	2.206
3.1	149.546	2.564
2.3	149.824	2.047
4.0	149.453	2.278
2.1	149.863	2.788
...	...	...

# SIMSTACK: simplest results

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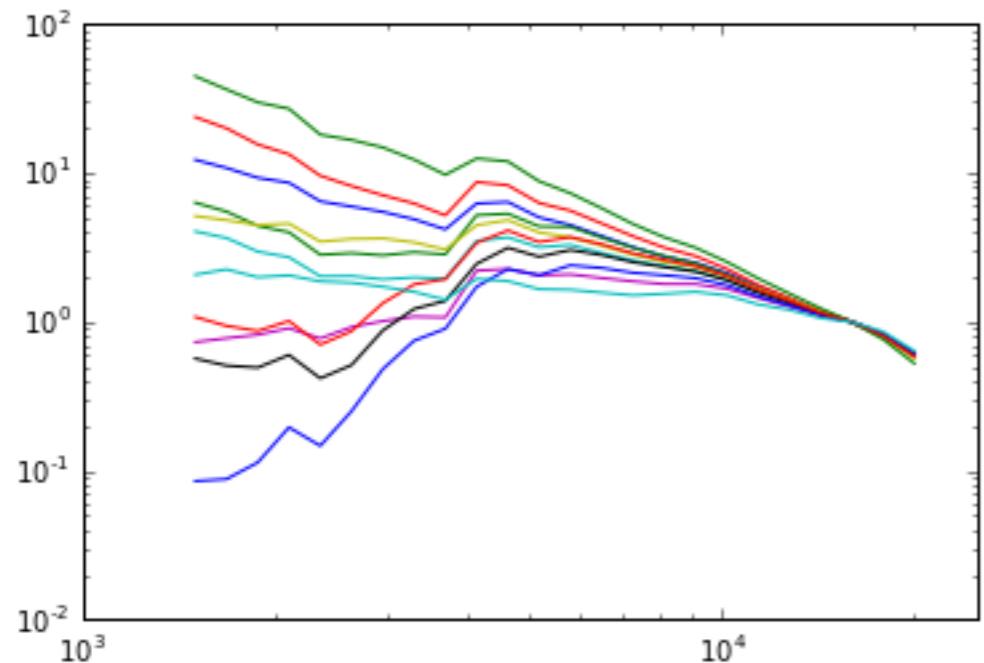
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# Model SEDs

Towards taking full advantage of SIMSTACK

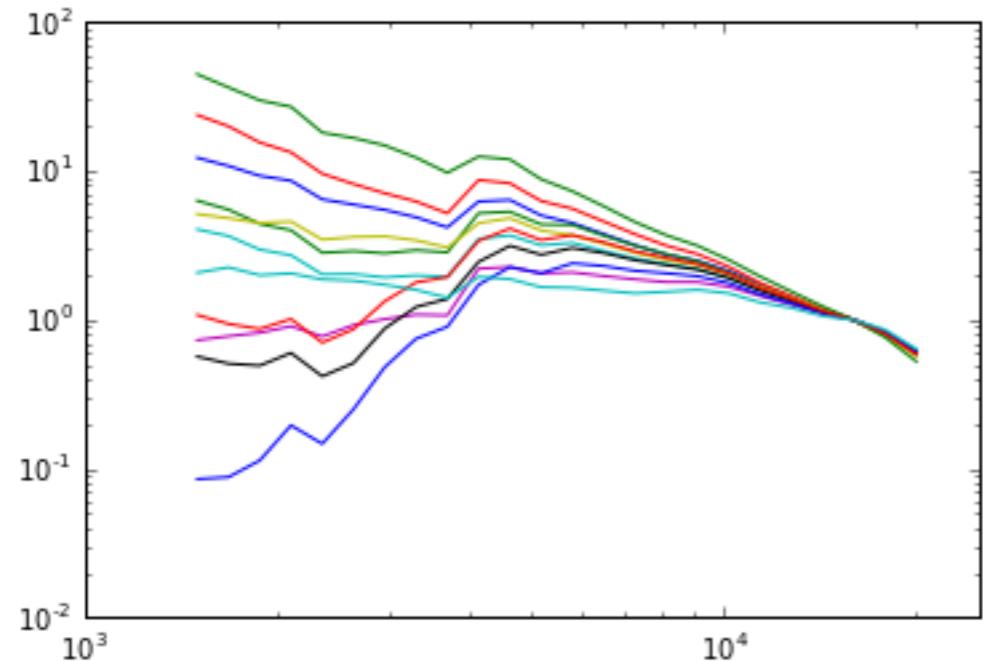
# SIMSTACK: Beyond Colour

- Full SED Categorization  
→ map physical features to FIR flux



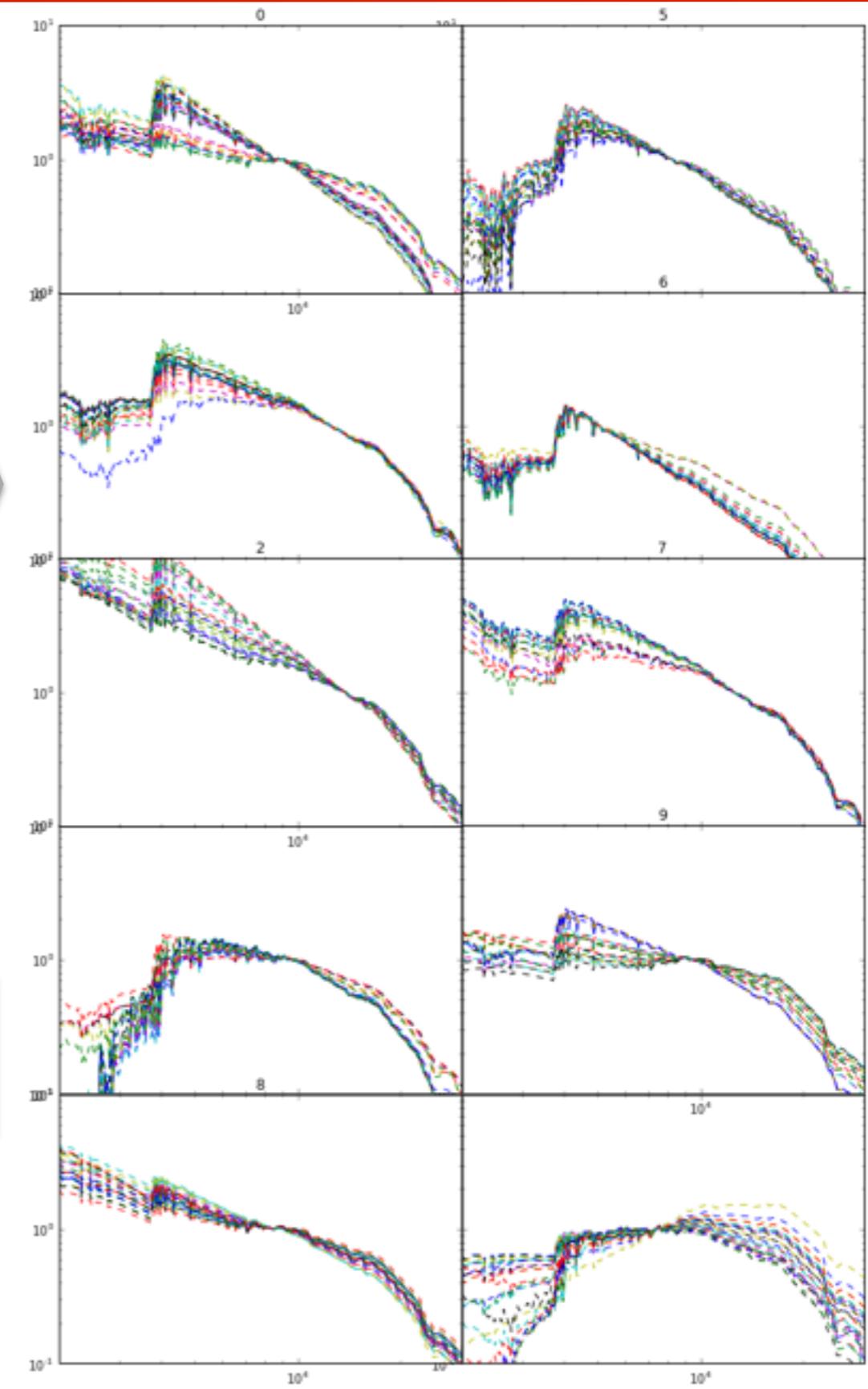
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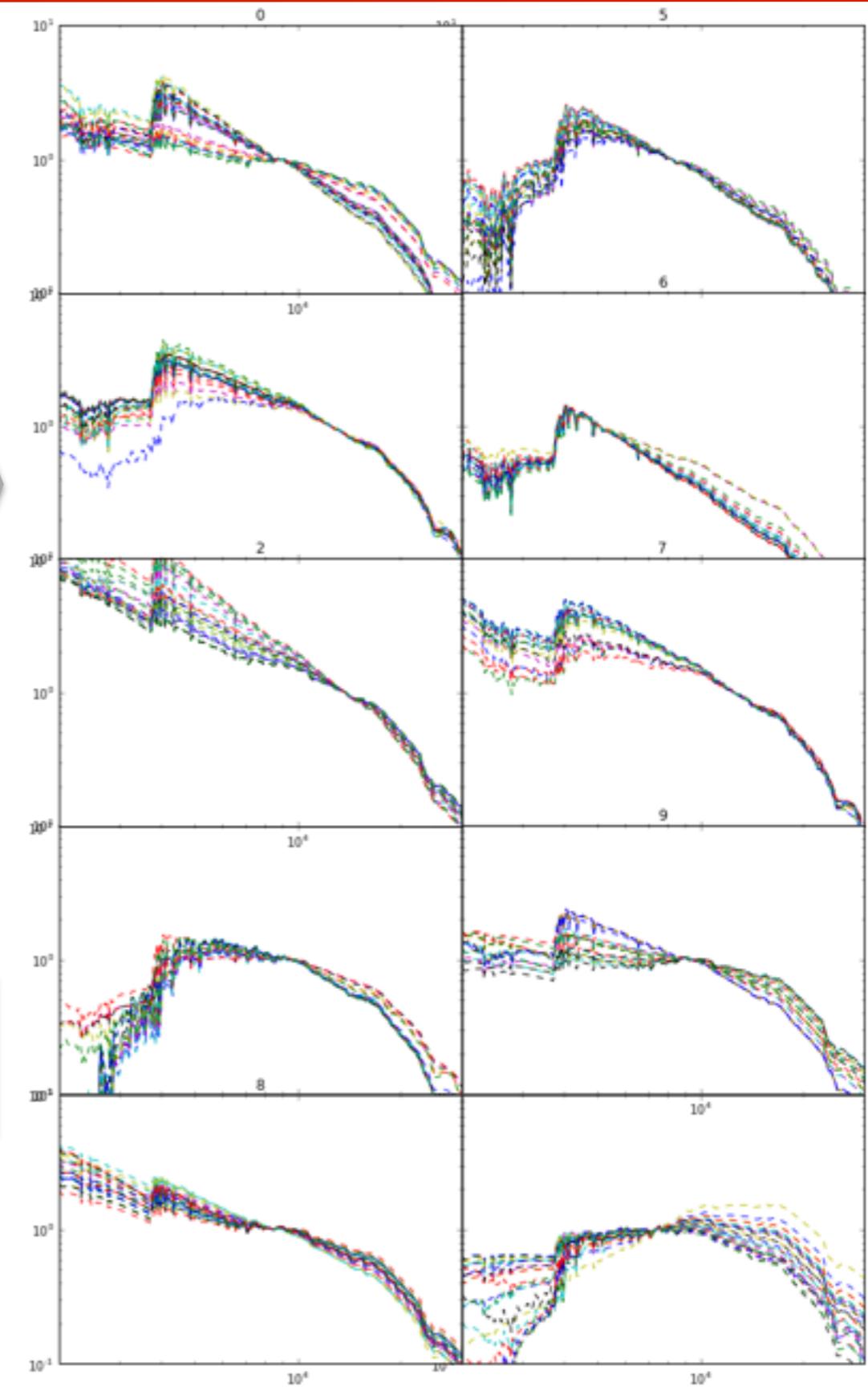
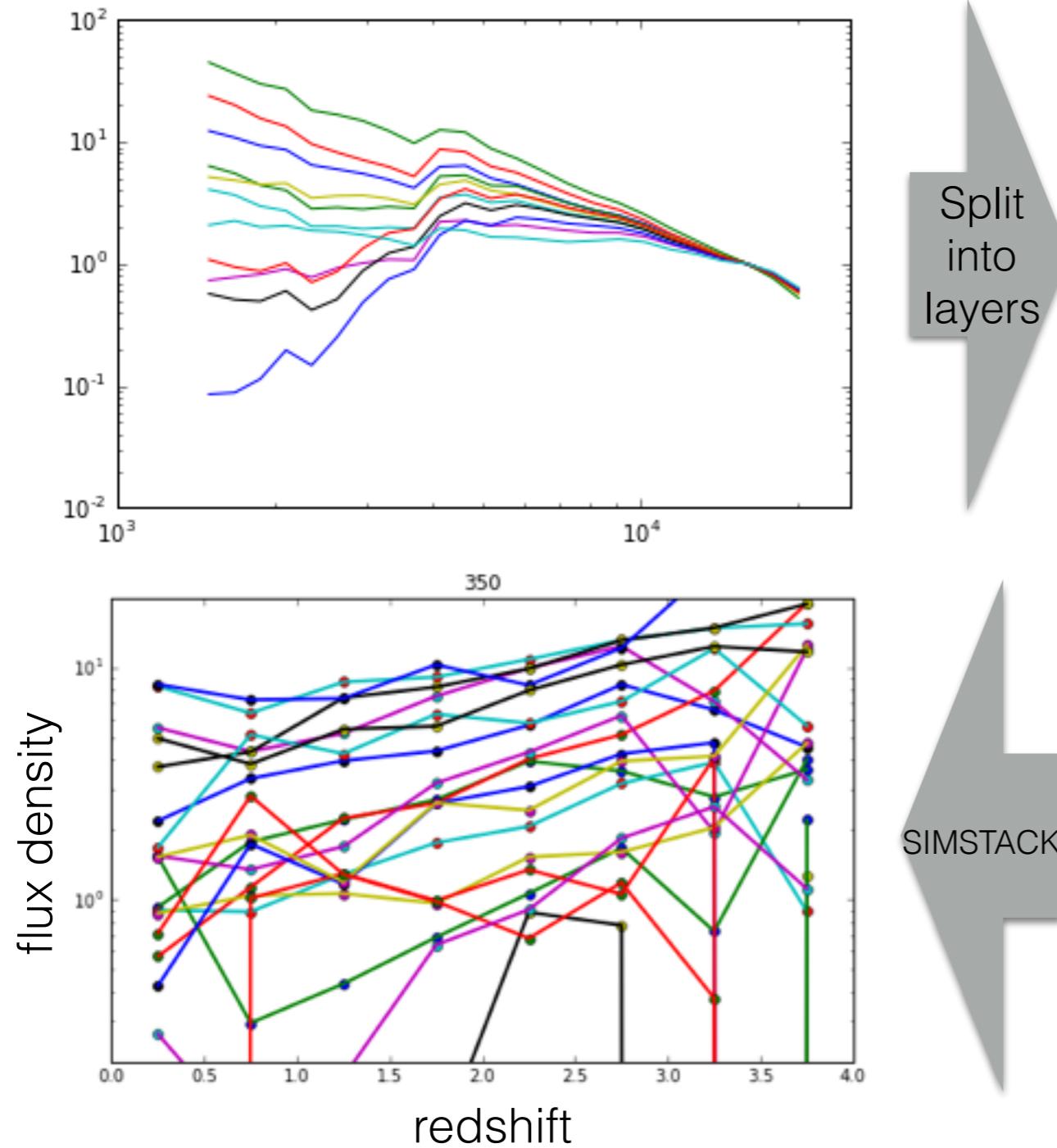
Split  
into  
layers

SIMSTACK



# SIMSTACK: Beyond Colour

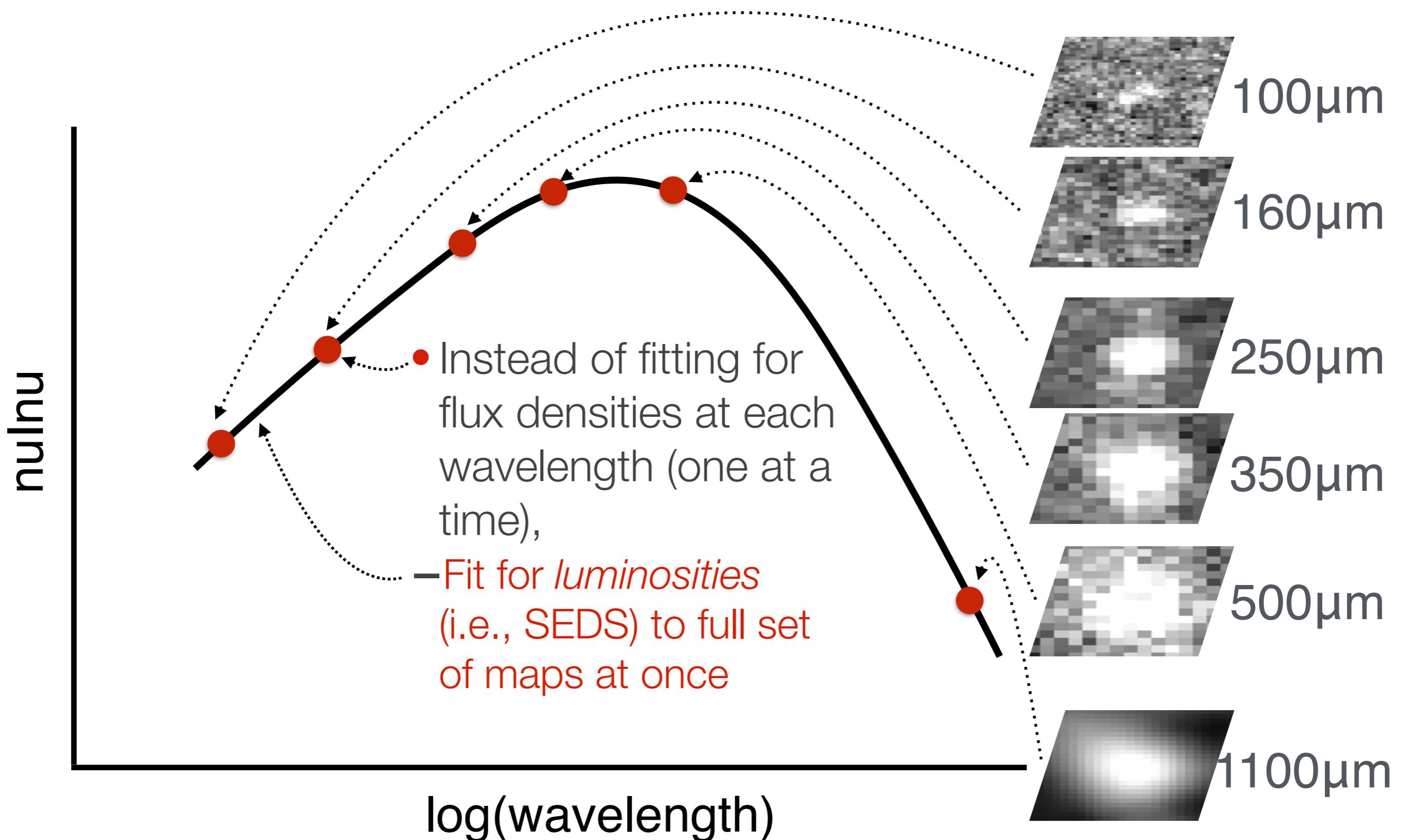
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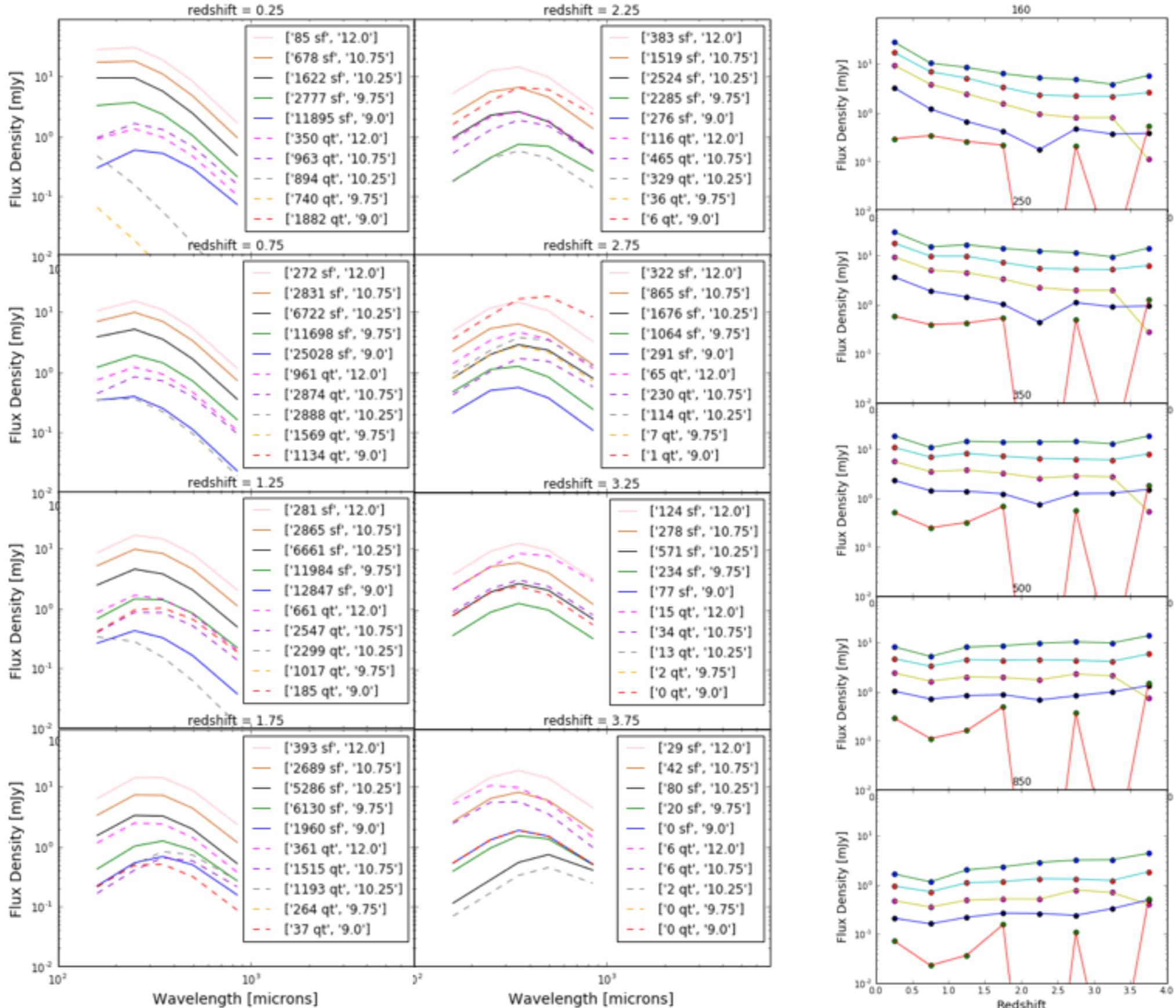
An extension of SIMSTACK

# SEDSTACK: Beyond Flux

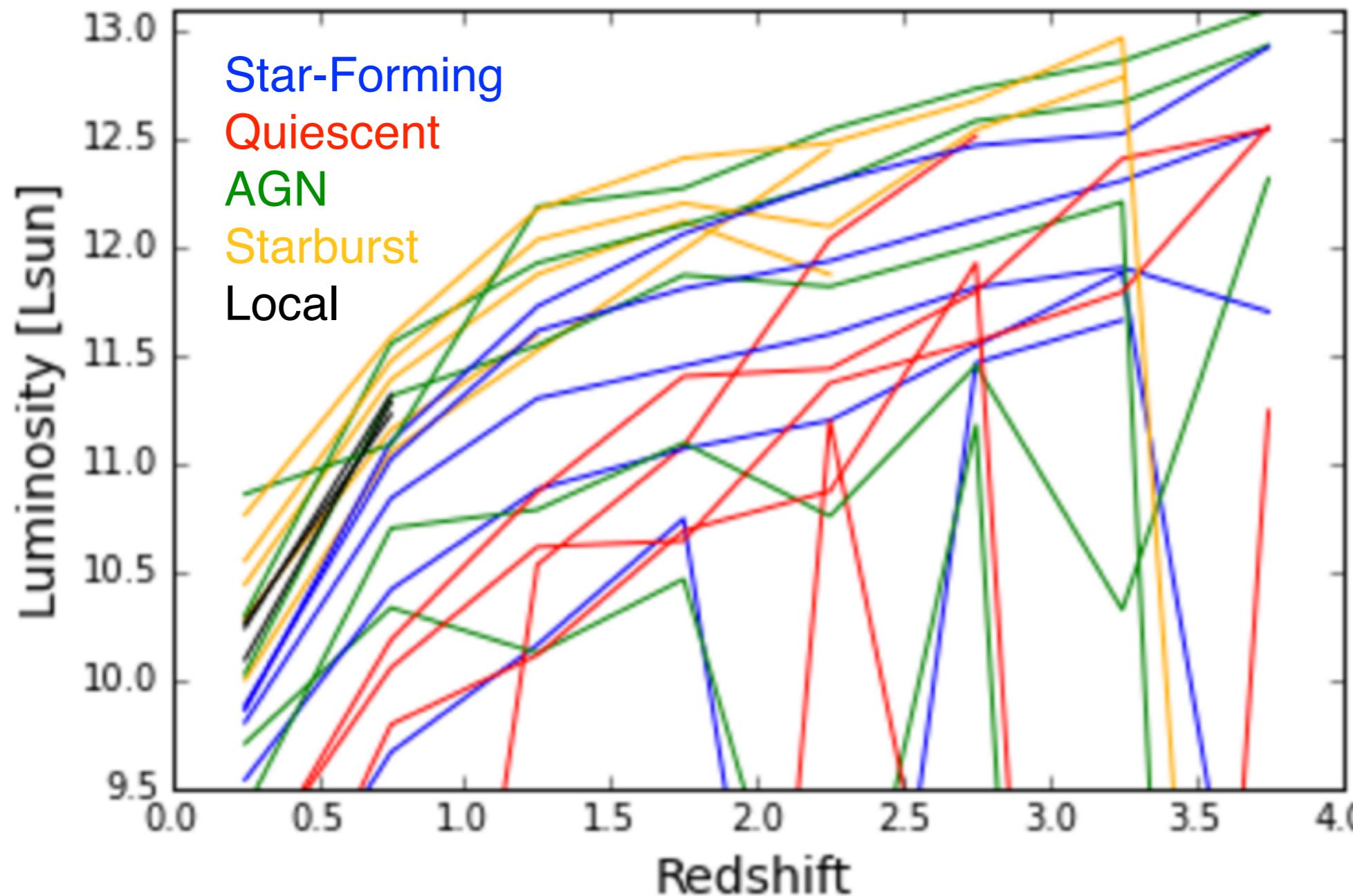


# SEDSTACK in z - M - QT/SF bins

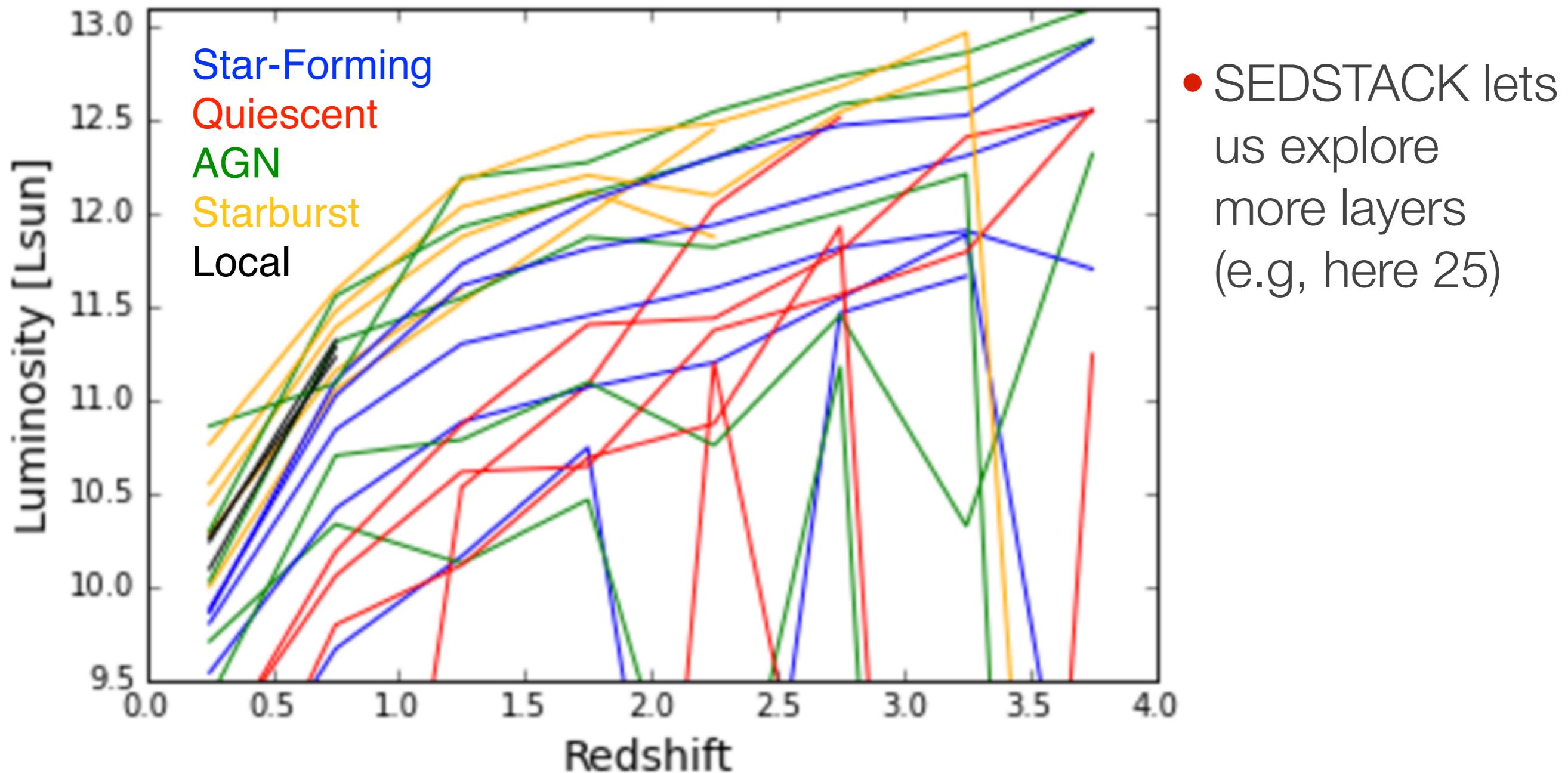
- Advantages:
  - leverage high S/N components to better constrain faint-end
- To-do:
  - quantify improvement
  - speed up
  - Emcee to measure full posterior



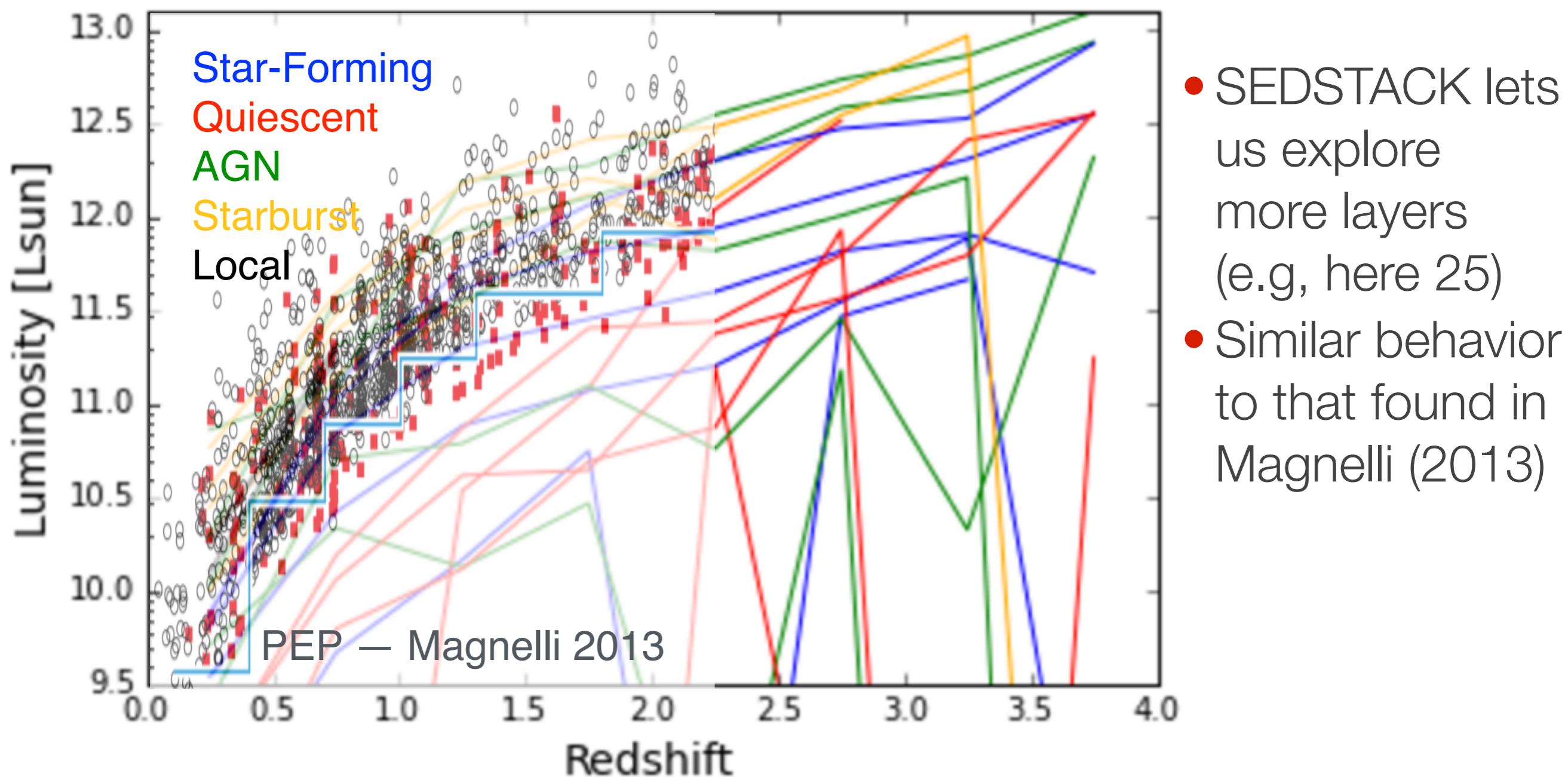
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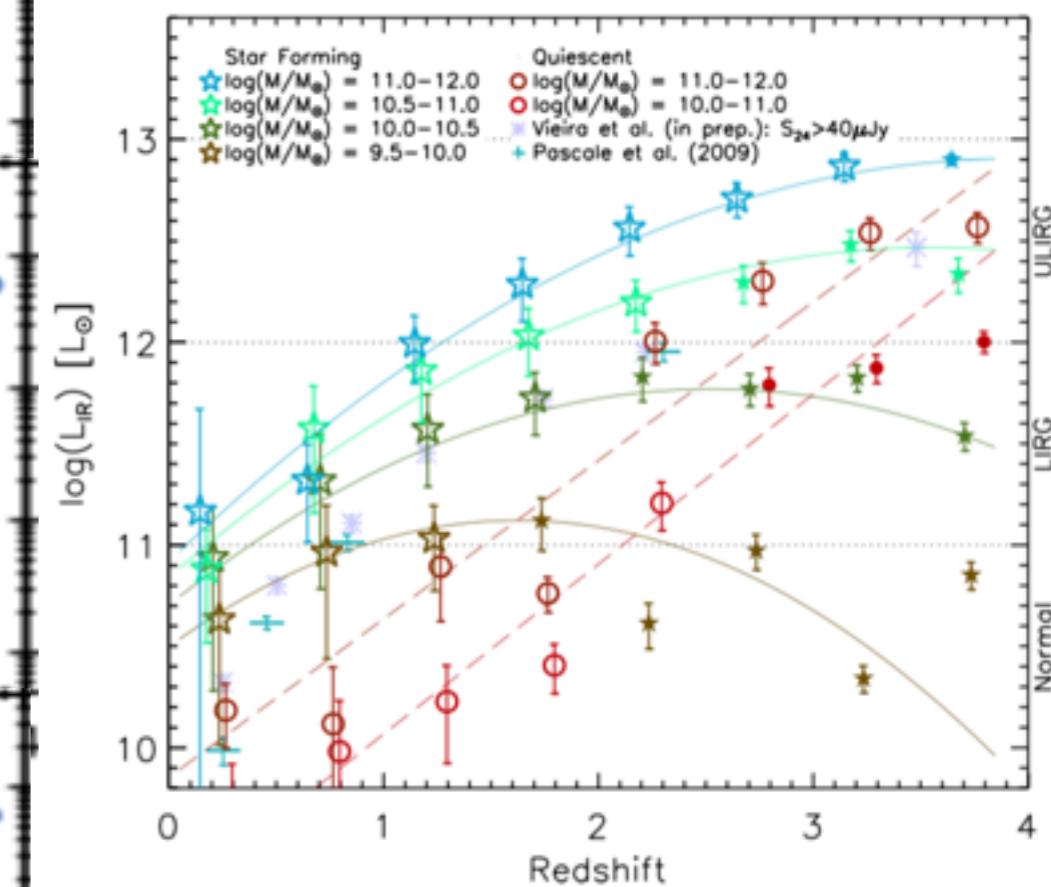
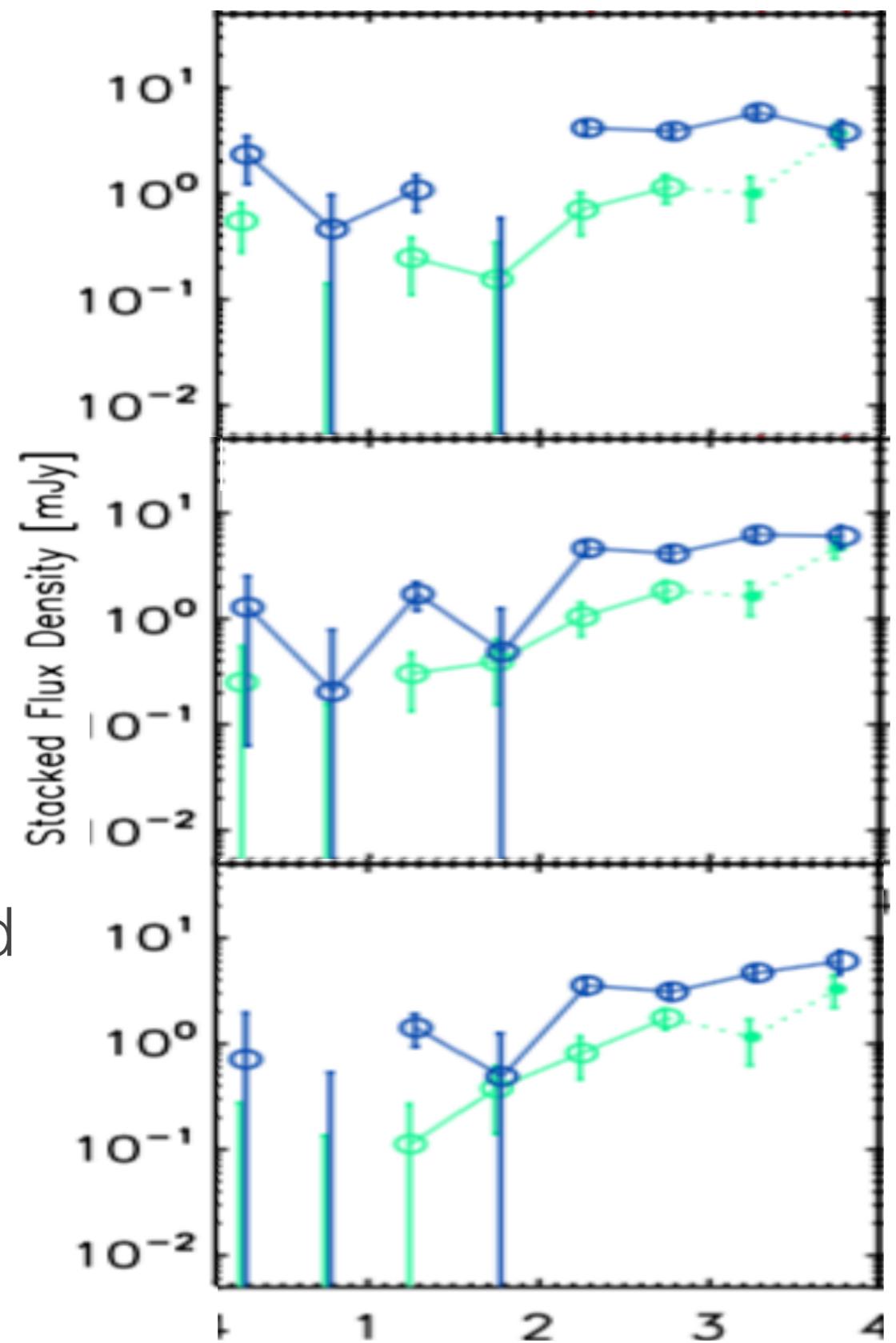


# What is this?

Weird results from SIMSTACK

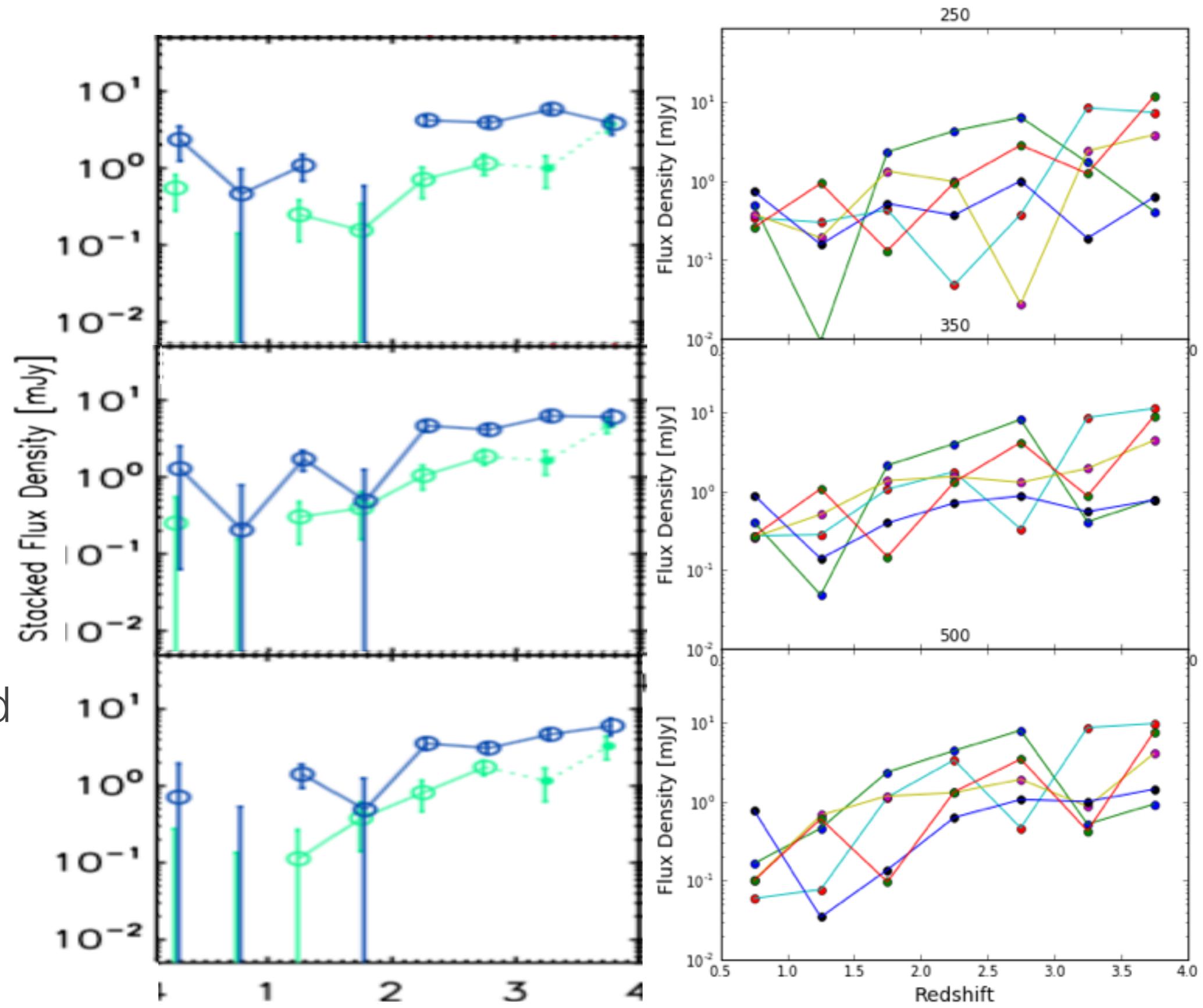
# Luminous High-z Quiescent Galaxies? UVJ failure?

- High-z “quiescent” (UVJ) sources in UDS very luminous!
- Confirm this behavior with alternative catalogs and maps (UVISTA/ COSMOS)



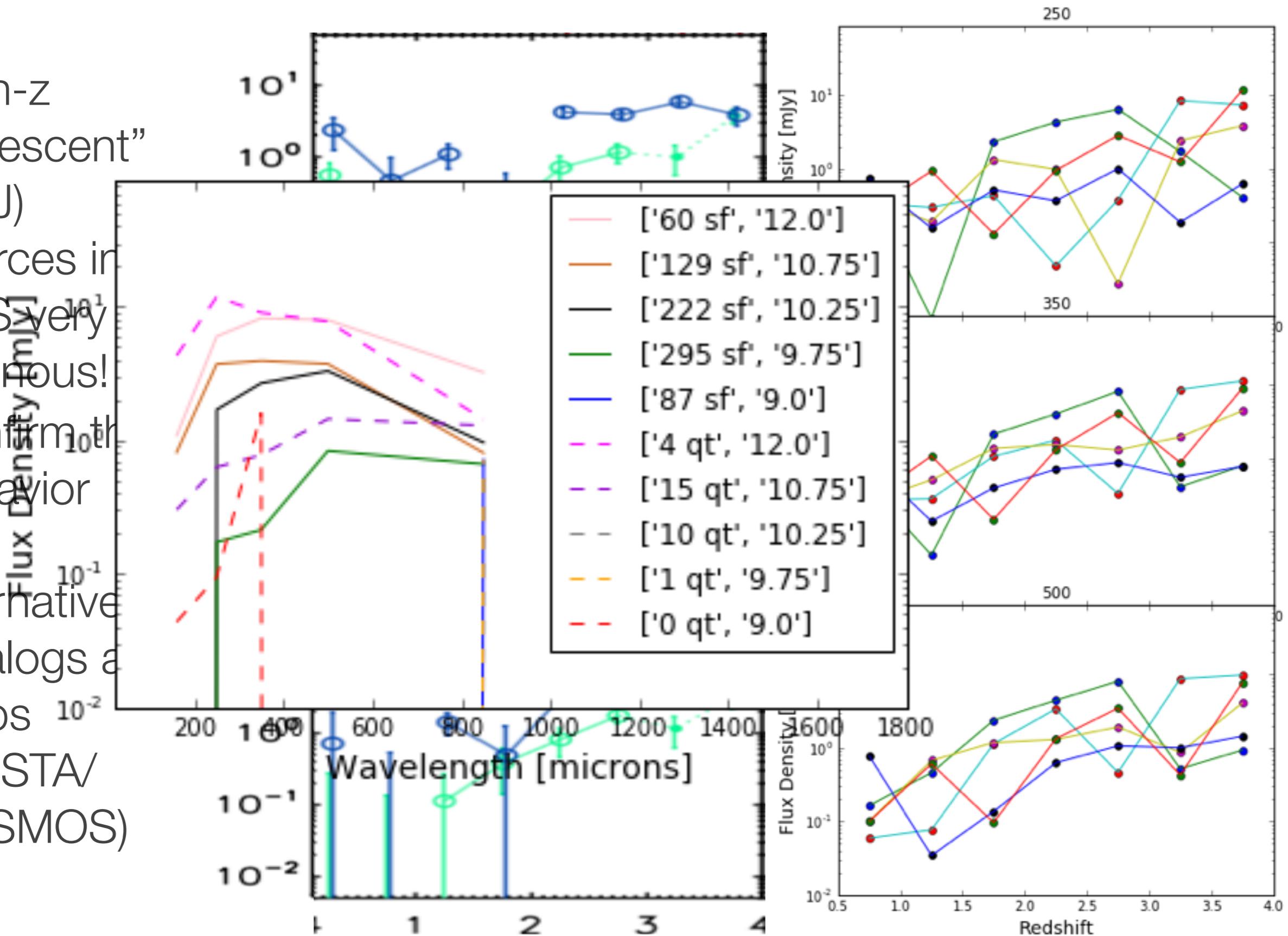
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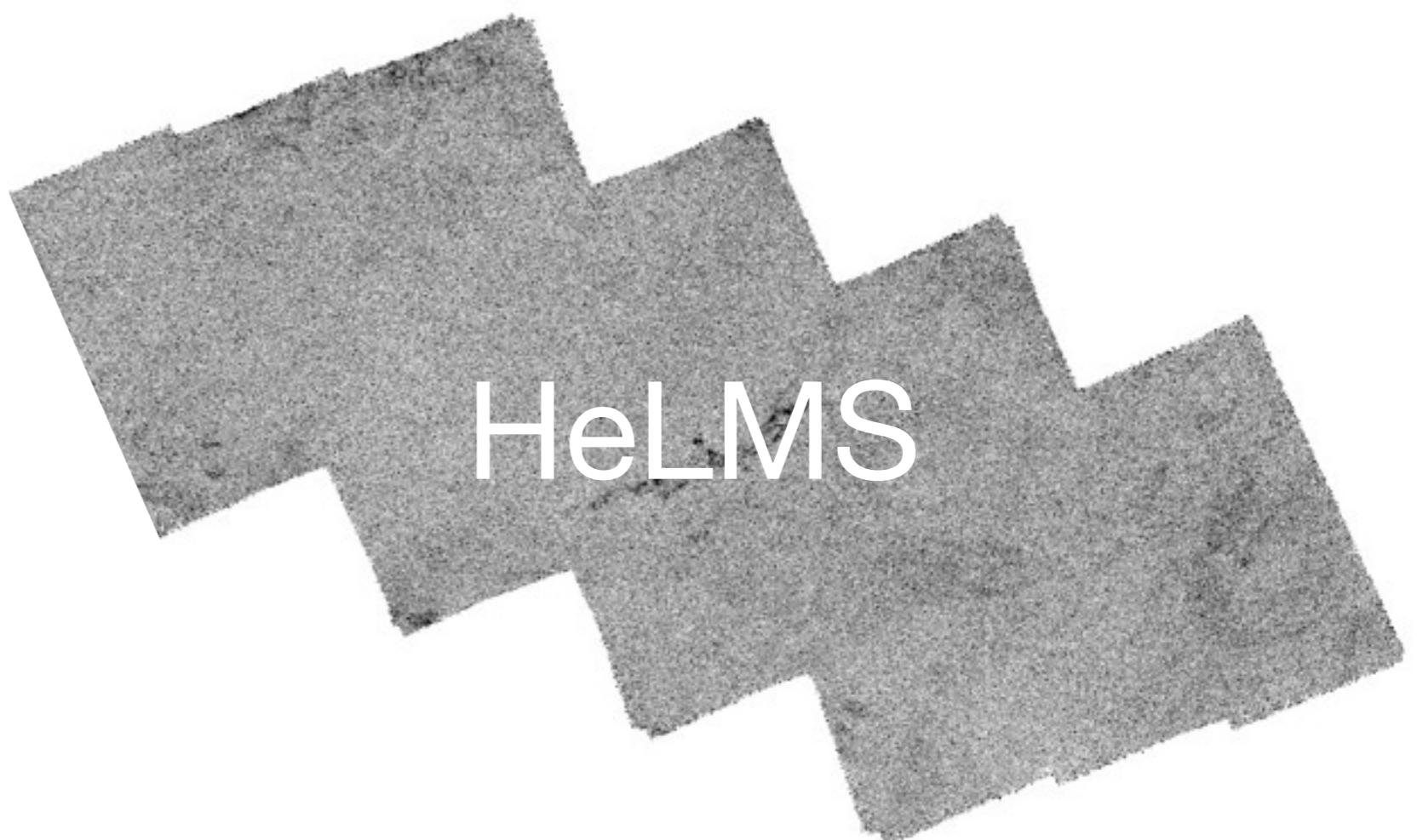
# HeLMS / HerS

The Herschel Surveys in Stripe 82

 HERMES

SDSS Stripe 82

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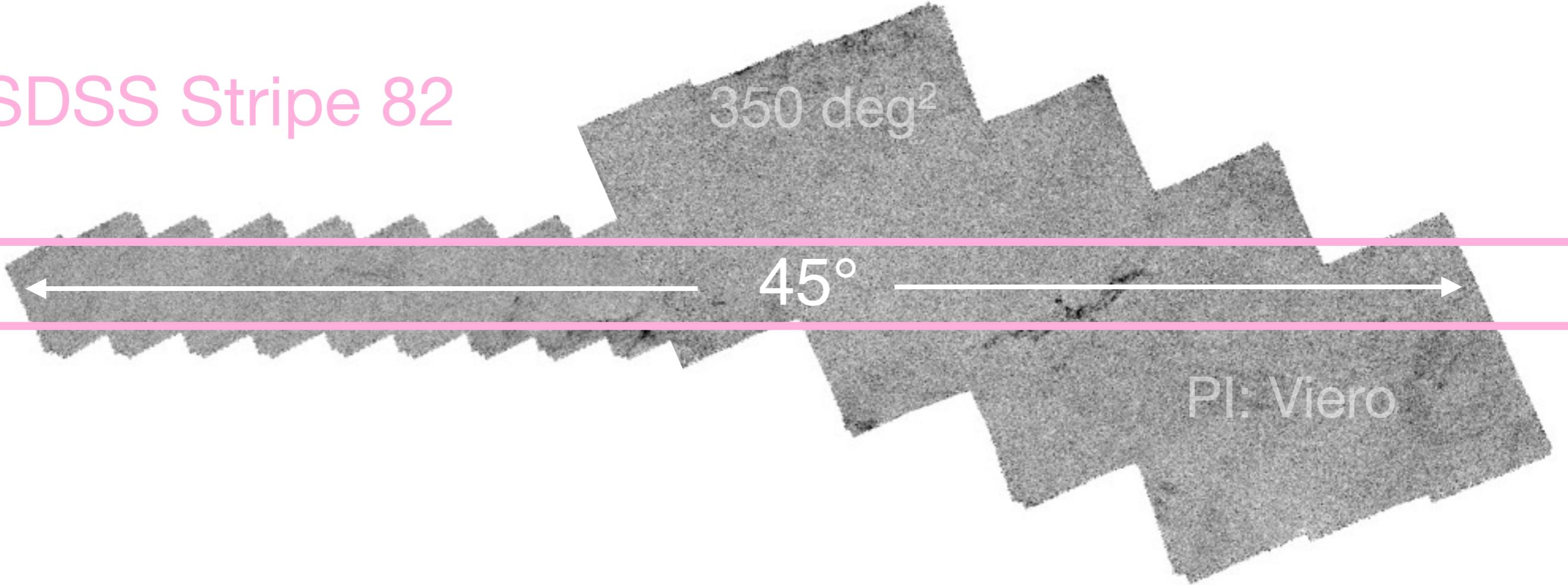


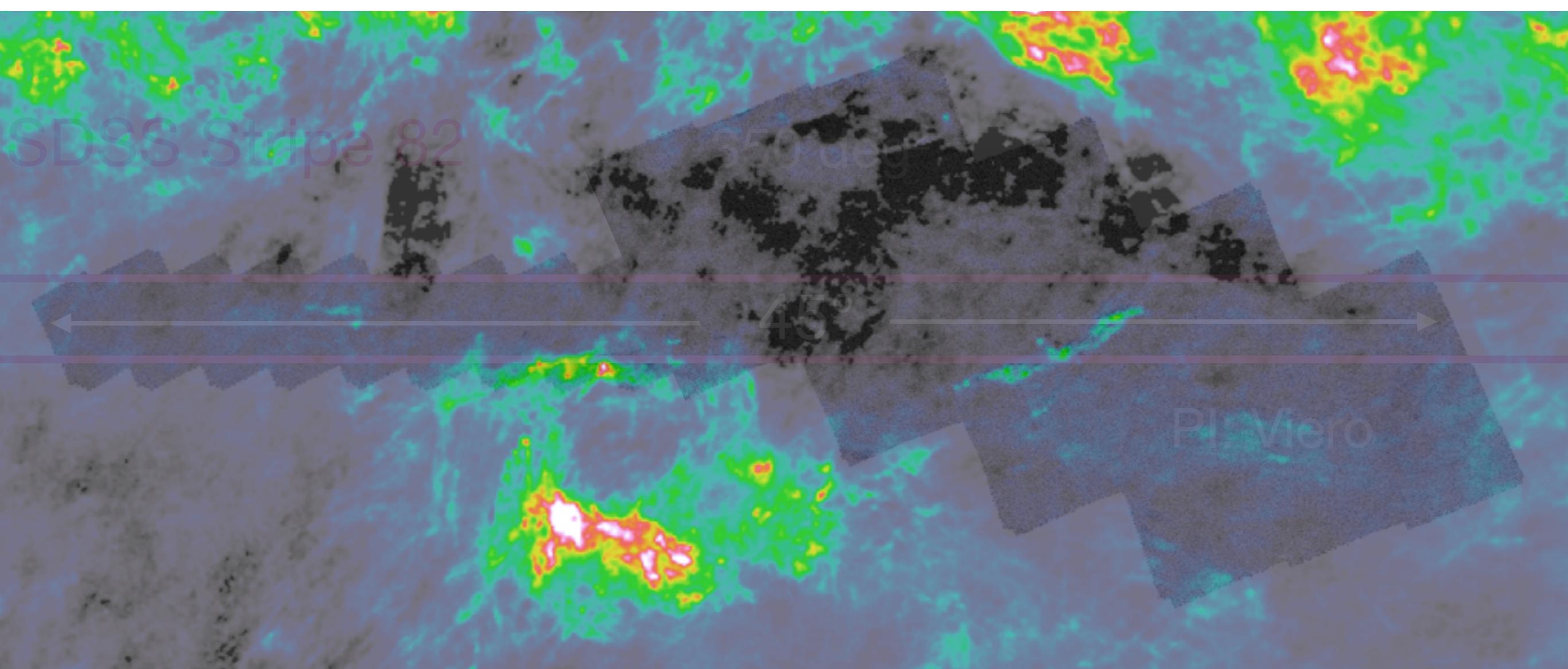
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HerS

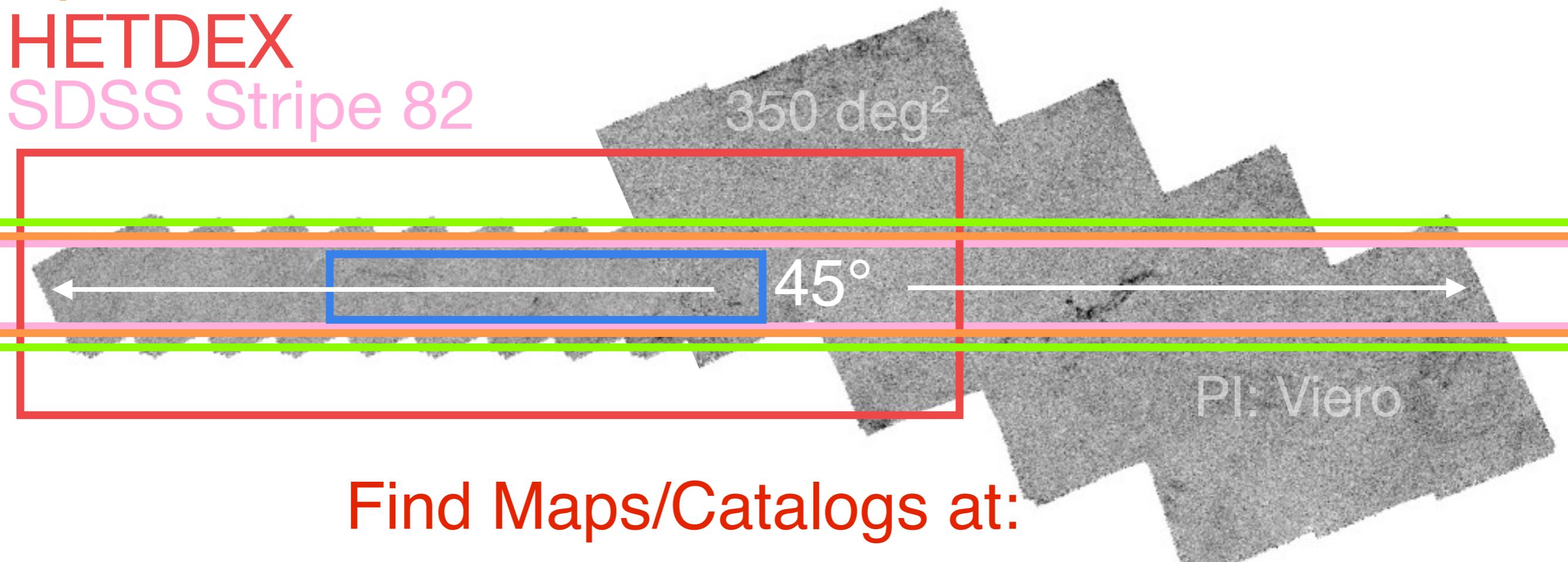


SDSS Stripe 82





ACT  
SHELA  
SpIES  
HETDEX  
SDSS Stripe 82



Find Maps/Catalogs at:

HerS: <http://www.astro.caltech.edu/hers>

HeLMS: <http://hedam.lam.fr/HerMES/>