HALO+EMPIRICAL MODELING

Marco Viero, KIPAC/Stanford COMAP Collaboration Meeting OVRO, January 10-11, 2017

OUTLINE

- ➤ Mixing theory (Tony Li's model) and observation (my SIMSTACK measurements) in new round of sims.
- ➤ Modeling meetings Friday and Tuesday.

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mass function, which Magnelli et al. 2012; hroozi et al. (2013a). a recent comprehenor for high-redshift cope of this paper. In es quoted above, we og $\delta_{\rm MF}=0.0\pm0.3$ prior's $\pm3\sigma$ interval

be written as explicitly accounts as opposed to active simplicity, we have out Equation (1) due be absorbed into σ_{SFR}

Luminosity

minosity, we assume

aforementioned studies, where a scatter has been quoted.

Note that the conversion from L'_{CO} (units of K km s⁻¹ pc²) to funits of L_{\odot}) is

$$L_{\rm CO} = 4.9 \times 10^{-5} L_{\odot} \left(\frac{\nu_{\rm CO, rest}}{115.27 \text{ GHz}} \right)^3 \left(\frac{L'_{\rm CO}}{\text{K km s}^{-1} \text{ pc}^2} \right)$$
 (4)

where $\nu_{\text{CO,rest}} = 115.27 \,\text{GHz}$ is the rest-frame frequency of the CO transition.

To resummarize the model:

- 1. Halos \rightarrow SFR: Get $\overline{SFR}(M, z)$ from the results of Behroozi et al. (2013a)
- 2. Add log-scatter, σ_{SFR}
- 3. SFR $\rightarrow L_{\rm IR}$: Get $L_{\rm IR}$ from SFR $= \delta_{\rm MF} \times 10^{-10} L_{\rm IR}$
- 4. $L_{\rm IR} \to L_{\rm CO}'$: Get $L_{\rm CO}'$ from $\log L_{\rm IR} = \alpha \log L_{\rm CO}' + \beta$
- 5. Add log-scatter, $\sigma_{L_{CO}}$

with fiducial parameter values:

$$\sigma_{\rm SFR} = 0.3, \, \sigma_{L_{\rm CO}} = 0.3, \, \delta_{\rm MF} = 1.0, \, \alpha = 1.37, \, \beta = -1.74.$$

Figure 2 shows the combined result of these steps, plotting the mean $L_{CO}(M_h)$ relation from our fiducial model, as well as the equivalent relation from previous studies. Notably, L_{CO} in this model is not linear in M, a simplifying assumption that has

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To resummarize the model: Stellar Mass - M*

- 1. Halos \rightarrow SFR: Get $\overline{SFR}(M, z)$ from the results of Behroozi et al. (2013a)
- 2. Add log-scatter, USFR Use empirically derived LIR(z,M*)
- 3. SFR $\rightarrow L_{\rm IR}$: Get $L_{\rm IR}$ from SFR $= \delta_{\rm MF} \times 10^{-10} L_{\rm IR}$
- 4. $L_{\rm IR} \rightarrow L_{\rm CO}'$: Get $L_{\rm CO}'$ from $\log L_{\rm IR} = \alpha \log L_{\rm CO}' + \beta$
- 5. Add log-scatter, $\sigma_{L_{CO}}$

with fiducial parameter values:

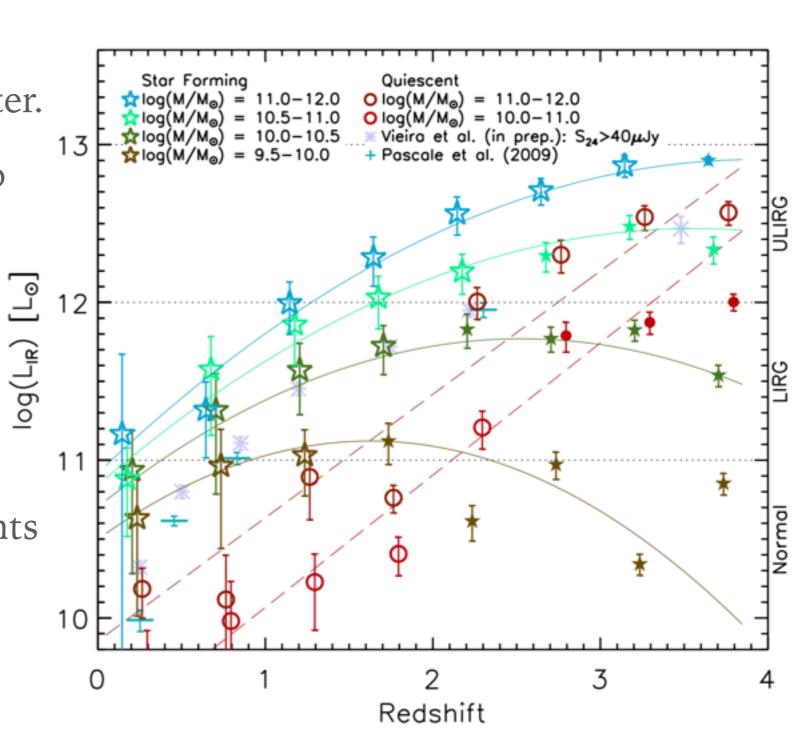
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SIMSTACK — VIERO ET AL. 2013

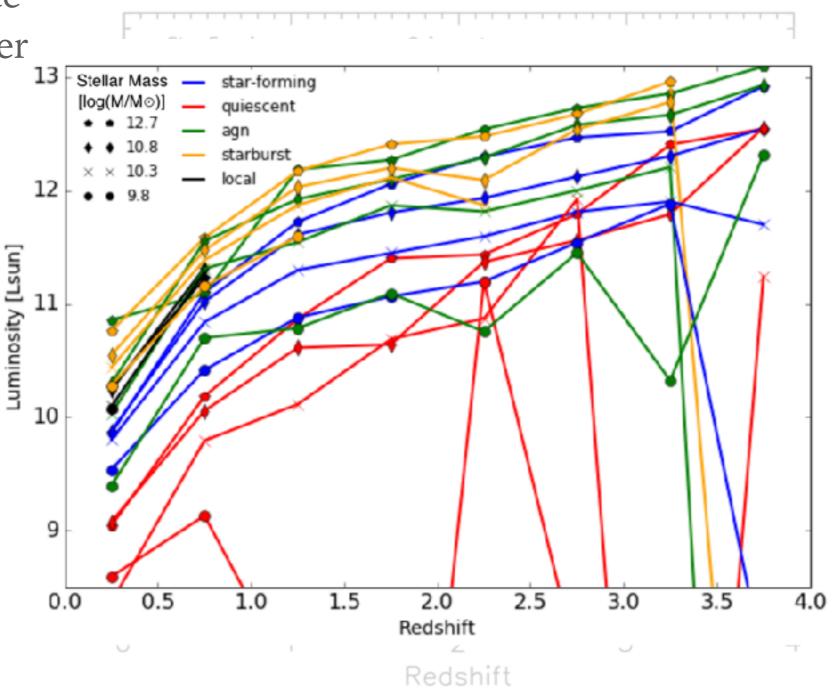
- ➤ SIMSTACK is like familiar thumbnail stacking, but better.
- ➤ Viero 2013 Split sample into stellar-mass, redshift, and star-forming/quiescent.
- Find 70% of the CIB is accounted for by K-selected optical galaxies.
- ➤ Functions fit to measurements (the lines in figure —>) can be used for simulations.



SIMSTACK — VIERO ET AL. 2013

➤ Viero 2017 in prep. update includes larger catalog over more area, and...

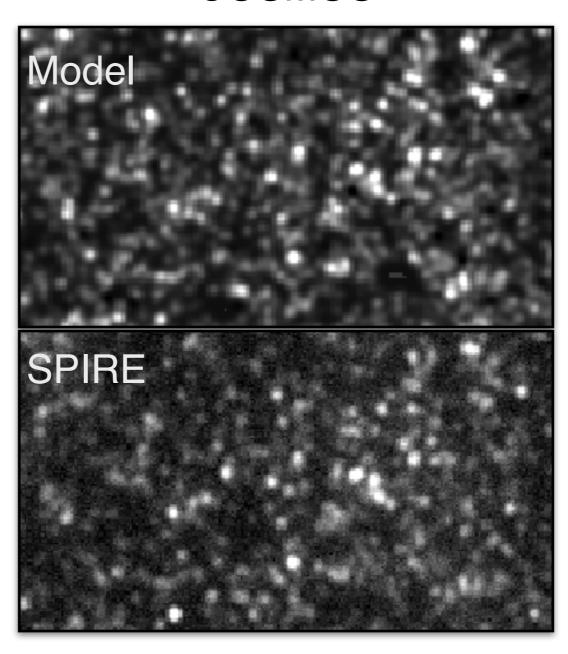
Split further into galaxy types.



SIMSTACK — VIERO ET AL. 2013

- ➤ Works very well for "typical" galaxies that make up background fluctuations, but
- ➤ Fails to find outliers, which may or may not be in the original source catalog.

COSMOS



SCHEDULE OF UPCOMING MODELING MEETINGS (AT CALTECH)

➤ Currently planning to meet Friday 13th and Tuesday 17th. Openended, with *brief* talks to update each other on models, and brainstorming sessions to facilitate collaboration and integration.

➤ Friday

- ➤ Meeting with Phil Hopkins and Norm Murray
- ➤ Working around Phil's teaching, will probably consist of morning and afternoon sessions.
- ➤ Tuesday
 - ➤ Marcelo Alvarez will be here.
- ➤ Hamsa Padmanabhan is giving the Tea Talk on Tuesday at 4pm