

Extending the Radial Acceleration Relation using Weak Gravitational Lensing with the Kilo-Degree Survey

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ABSTRACT

TBW

Key words: gravitational lensing; weak – Surveys – methods: statistical – galaxies: haloes – cosmology: dark matter, theory – gravitation.

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1 INTRODUCTION

It has been known for several decades that the outer regions of galaxies rotate faster than would be expected from Newtonian dynamics based on their luminous baryonic mass. This was first discovered by [Rubin \(1983\)](#) through measuring galactic rotation curves of optical disks, and by [Bosma \(1981\)](#) through measuring hydrogen profiles at larger radii. The excess gravity implied by these measurements have been generally attributed to an invisible substance named Dark Matter (DM), a term coined earlier by [Zwicky \(1937\)](#) when he discovered the missing mass problem by measuring the dynamics of galaxies in clusters. Following more recent observations using Weak gravitational Lensing (WL, [Hoekstra et al. 2004](#); [von der Linden et al. 2014](#); [Mandelbaum 2015](#)), Baryon Acoustic Oscillations (BAO's, [Eisenstein et al. 2005](#); [Blake et al. 2011](#)) and the Cosmic Microwave background (CMB, [Spergel et al. 2003](#); [Planck XIII 2016](#)), cold dark matter¹ (CDM) became a key ingredient of the current standard model of cosmology: the Λ CDM model. In this paradigm, CDM accounts for $\Omega_{\text{CDM}} = 0.266$ of the critical density in the Universe, while the cosmological constant Λ used to explain the accelerated expansion of the Universe accounts for $\Omega_{\Lambda} = 0.685$.

Although the Λ CDM model successfully describes the behavior of DM on a wide range of scales, no conclusive evidence for the existence of DM particles has been found so far (despite years of enormous effort; for an overview, see [Bertone et al. 2005](#); [Bertone & Tait 2018](#)). This still leaves some room for alternative theories of gravity, such as Modified Newtonian Dynamics (MOND, [Milgrom 1983](#)) and the more recent theory of Emergent Gravity (EG, [Verlinde 2016](#)). In these theories particle DM does not exist, and all gravity is due to the baryonic matter (or in the case of EG, the interaction between baryonic matter and the entropy associated with dark energy). One of the main properties of these theories, is that the mass discrepancy in galaxies should correlate strongly with their baryonic mass distribution. Such a correlation is indeed observed, first as the baryonic Tully-Fisher relation (BTFR, [Tully & Fisher 1977](#)) between the total luminosity of a spiral galaxy and its asymptotic rotation velocity, then as a strong correlation between the mass dependency as a function of radius and the enclosed baryonic mass ...cite...

2 DATA

2.1 KiDS source galaxies

Write the beginning. Need to know:

- What changes as we go to KiDS-1000 (K1000 paper?).

2.2 GAMA foreground galaxies

Write everything.

¹ DM particles that moved at non-relativistic speeds at the time of recombination, as favoured by measurements of the CMB ([Planck XVI 2014](#)) and the Lyman- α forest ([Viel et al. 2013](#)).

2.3 KiDS foreground selection

Still need to know:

- Maciek's GL-KiDS selection criteria for K1000.
- Angus' stellar mass method for K1000.

2.4 MICE mock galaxies

Write everything.

2.5 Bahamas mock galaxies

Written by Kyle?

3 DATA ANALYSIS

3.1 Isolated galaxy selection

Write the beginning. Still need to know:

- how to test the isolation criterion.

3.2 Lensing measurement

Write the beginning. Still need to know:

- How (if?) the GGL-pipeline changes with K1000.

3.3 Conversion to radial acceleration

Still need to know: whether we will use the SIS assumption or linear interpolation.

- Test both methods using the Bahamas simulation.

4 THEORETICAL PREDICTIONS

4.1 Analytical CDM model

Written by Kyle?

4.2 Modified Newtonian Dynamics

Write everything.

4.3 Emergent Gravity

Write everything.

5 RESULTS

Write when the results are ready. I still need:

- The K1000 lensing catalogues with ANNz redshifts and stellar masses.
- The results from the Bahamas simulation.

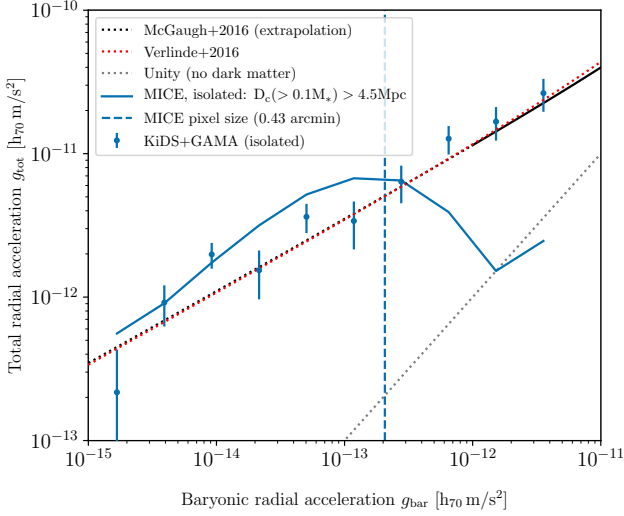


Figure 1. TBW

5.1 Isolated galaxies

5.2 Stellar mass bins

6 DISCUSSION AND CONCLUSION

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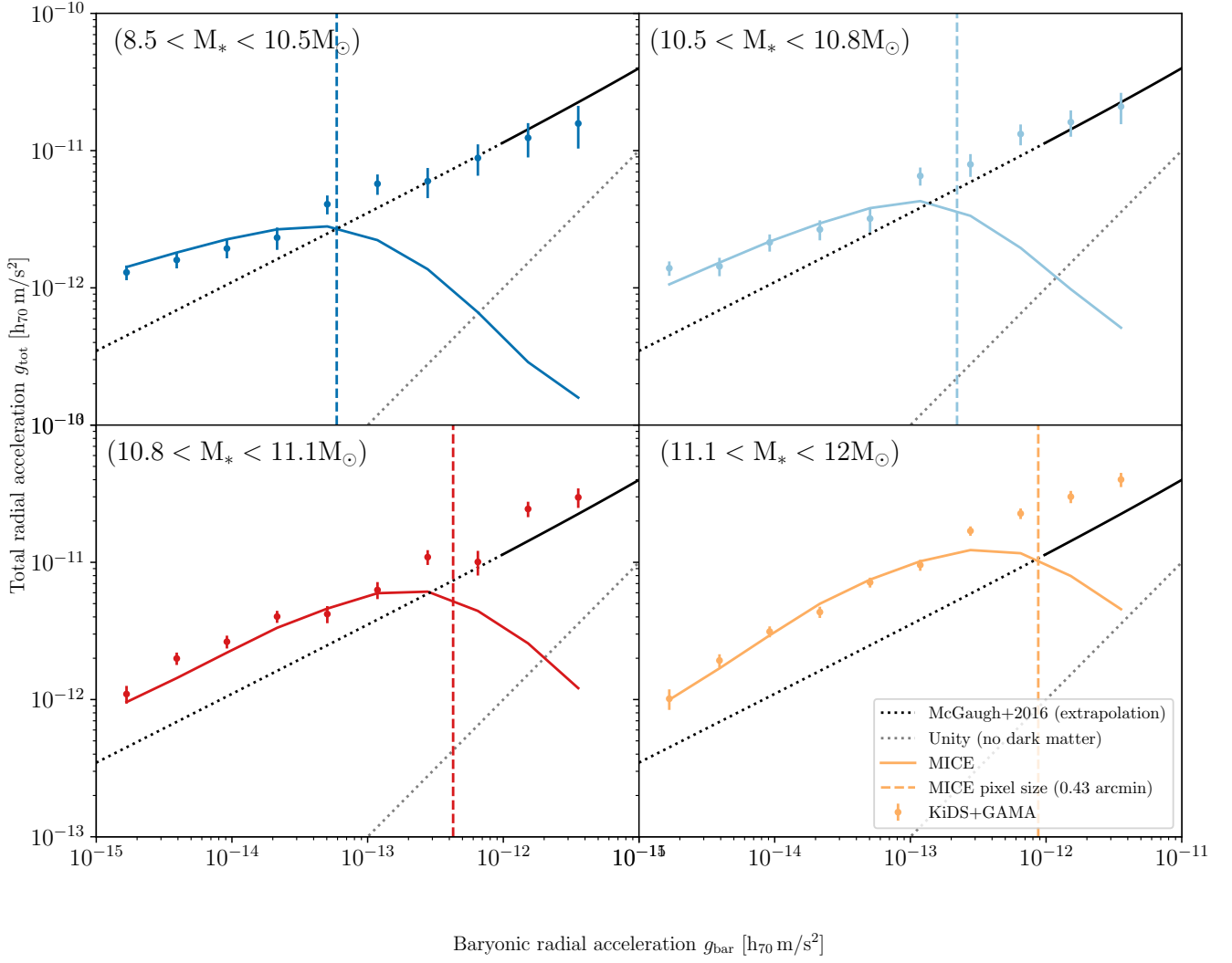
ACKNOWLEDGEMENTS

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**Figure 2.** TBW

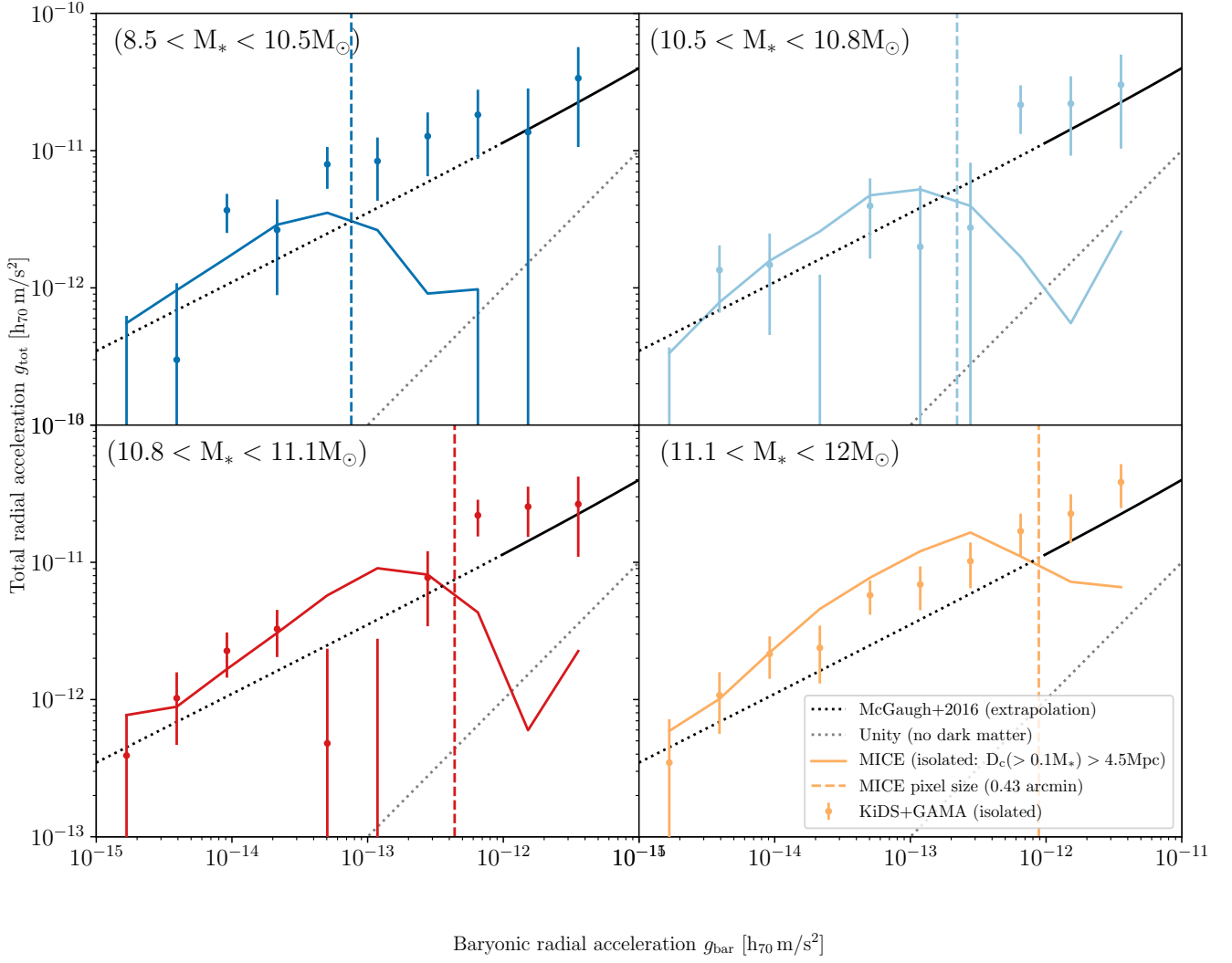


Figure 3. TBW

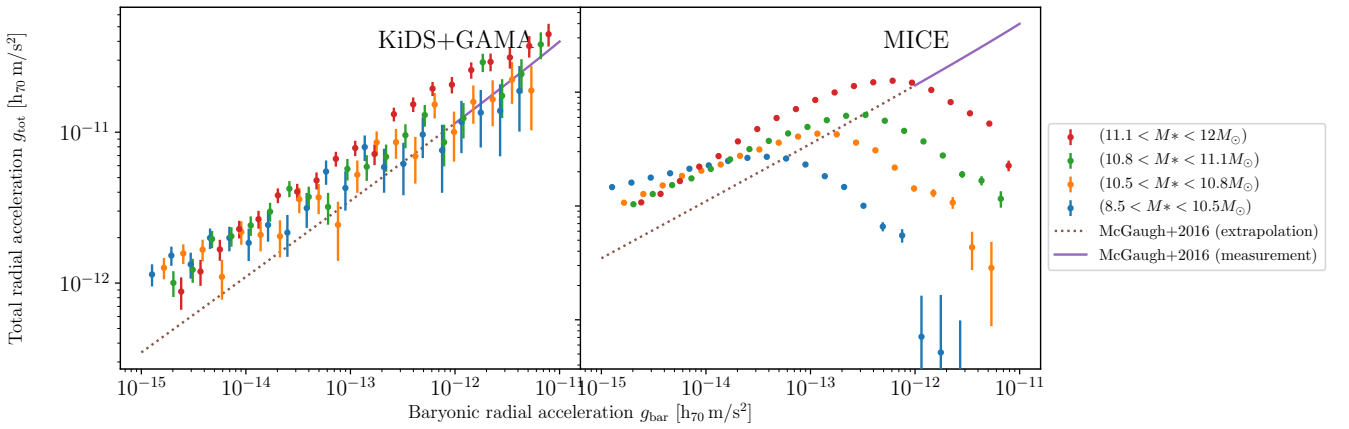


Figure 4. TBW