

Fluid Dynamics: Interesting Problems

Flow down a ramp and Poiseuille flow

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Problem 1

Consider a viscous fluid with kinematic viscosity $\mu = 1$ and density $\rho = 1$ flowing down a ramp with an angle θ of 10 degrees with the horizontal. For simplicity, pick units so that the gravity constant g also equals one.

a) For a uniform flow with height h , derive an expression for the shear $u(y)$ and the corresponding mass flow per unit width

$$Q(h) = \int_0^h \rho u(y) dy.$$

b) Now consider a variable flow, with $h = h(x, t)$ and $u = u(x, y, t)$. Under the assumption that the dependence on x and t is slow, we can assume that the relation you derived before for $Q(h)$ is still valid locally. Then use conservation of mass to write an equation for the evolution of $h(x, t)$.

c) Linearize the equation above by assuming that

$$h = 1 + \eta, \quad \eta \ll 1,$$

and write down the solution $\eta(x, t)$ to the initial value problem

$$\eta(x, 0) = \eta_0(x).$$

d) Returning now to the fully nonlinear equation of part b, find a change of variables $r = r(h)$ so that the equation reduces to the Burgers equation for r

$$r_t + r r_x = 0.$$

– Along which lines in (x, t) space will r (and therefore also h) be constant?

– Consider a situation where the initial data $h_0(x)$ grows from $h = 1$ at $x = -\infty$ to a maximum value $h = 2$ at $x = 0$ and then decays again to $h = 1$ at $x = \infty$. When a shock first forms, will it form on the section where h grows or where it decays?

Solution

Part a

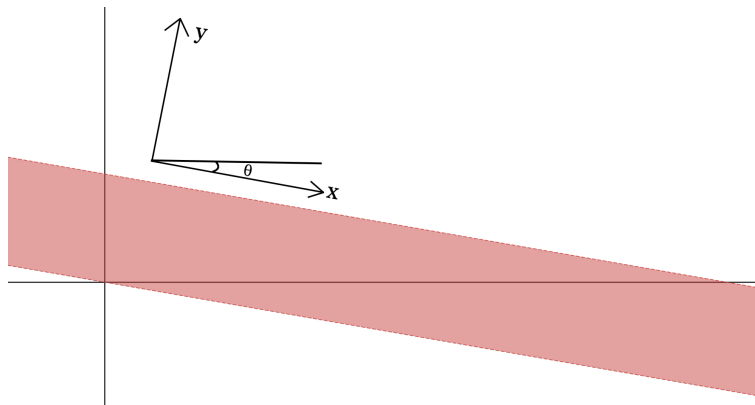


Figure 1: Visualisation of the problem

For this part we need to recall the Navier-Stokes (N-S) equations in both x and y directions, as well as specify what those directions are (see Figure 1).

In the x direction, the N-S equation is written as

$$\frac{\partial u}{\partial t} + (\mathbf{u} \cdot \vec{\nabla}) u + \frac{1}{\rho} \frac{\partial p}{\partial x} = \mu \nabla^2 u + g \sin(\theta). \quad (1)$$

In the y direction, it is

$$\frac{\partial v}{\partial t} + (\mathbf{u} \cdot \vec{\nabla}) v + \frac{1}{\rho} \frac{\partial p}{\partial y} = \mu \nabla^2 v - g \cos(\theta). \quad (2)$$

In order to simplify the equations, we make some assumptions:

1. The flow is time-independent (steady).
2. u does not depend on x , since it extends to $\pm\infty$ in both directions.
3. The total velocity \mathbf{u} is parallel to the x direction. A corollary from this is that $v = 0$.
4. The pressure on the moving boundary ($y = h$) is 0.
5. The flow is incompressible.
6. No-slip: at $y = 0$, $u = 0$.
7. No shear stress due to air: $\left. \frac{\partial u}{\partial y} \right|_{y=h} = 0$.

Applying these assumptions, equations (1) and (2) simplify to:

$$\frac{\partial p}{\partial x} = \eta \frac{\partial^2 u}{\partial y^2} + \rho g \sin(\theta) \quad (3)$$

and

$$\frac{\partial p}{\partial y} = -\rho g \cos(\theta) \quad (4)$$

respectively, where the identity $\eta = \mu\rho$ was used for simplification. From equation (4), we can deduce

$$p = -\rho g y \cos(\theta) + f(x). \quad (5)$$

If we input this expression for p into (3), we can see that on the left hand side, there will be a term with $f'(x)$, while on the right hand side there are no terms which contain an x dependence. We can thus deduce

$$f'(x) = 0 \implies f(x) = Ax + B \quad (6)$$

Furthermore, since we want a reasonable solution according to the assumption 4. (otherwise we would have an instance of infinite pressure), it is evident that $A = 0$, so that $f(x)$ is a constant. Moreover, it holds that

$$p(h) = -\rho g h \cos(\theta) + p_0 = 0 \implies p_0 = \rho g h \cos(\theta). \quad (7)$$

So finally,

$$p(y) = -\rho g \cos(\theta) (h - y). \quad (8)$$

Putting this in equation (3) yields

$$\eta \frac{\partial^2 u}{\partial y^2} = -\rho g \sin(\theta) \implies u = -\frac{\rho g y^2}{2\eta} \sin(\theta) + Ay + B. \quad (9)$$

Implementing assumptions 6. and 7. respectively, we obtain

$$B = 0 \quad \wedge \quad A = \frac{\rho g y}{\eta} \sin(\theta). \quad (10)$$

Putting this in (9) and using the identity $\mu = \frac{\eta}{\rho}$, we obtain

$$u = \frac{g \sin(\theta)}{2\mu} y (2h - y). \quad (11)$$

Substituting given values for μ , ρ and g yields

$$u = \frac{\sin(10^\circ)}{2} y (2h - y). \quad (12)$$

Furthermore, since u is only a function of y we can calculate the mass flow rate per unit width as

$$Q(h) = \int_0^h \rho u(y) dy = \frac{\sin(10^\circ)}{2} \int_0^h (2hy - y^2) dy \quad (13)$$

$$= \frac{h^3 \sin(10^\circ)}{3}. \quad (14)$$

Part b

Conservation of mass tells us

$$\frac{\partial}{\partial t} \iiint_V d\tau + \oint_C \mathbf{u} \cdot d\mathbf{A} = 0. \quad (15)$$

Which, in our 1-dimensional case yields

$$\frac{\partial h}{\partial t} + \frac{\partial Q}{\partial x} = 0 \implies \frac{\partial h}{\partial t} + Q'(h) \frac{\partial h}{\partial x} = 0. \quad (16)$$

Thus, the evolution for $h(x, t)$ can be modelled by the equation

$$\frac{\partial h}{\partial t} + h^2 \sin(10^\circ) \frac{\partial h}{\partial x} = 0. \quad (17)$$

Or equivalently,

$$\frac{dh}{dt} = 0 \quad \text{along the line} \quad \frac{dx}{dt} = h^2 \sin(10^\circ). \quad (18)$$

Part c

If we make the suggested substitution for h , (17) becomes

$$\frac{\partial \eta}{\partial t} + (1 + \eta)^2 \sin(10^\circ) \frac{\partial \eta}{\partial x} = 0. \quad (19)$$

Using a small η approximation yields $(1 + \eta)^2 \approx 1$, so (19) in linearised form becomes

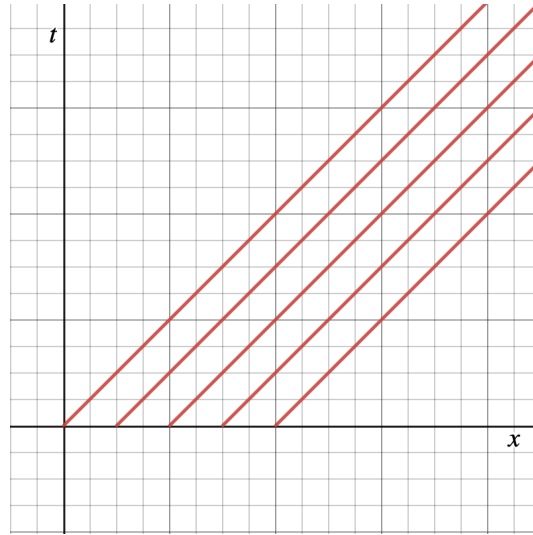
$$\frac{\partial \eta}{\partial t} + \sin(10^\circ) \frac{\partial \eta}{\partial x} = 0. \quad (20)$$

Which in turn implies

$$\frac{d\eta}{dt} = 0 \quad \text{along} \quad \frac{dx}{dt} = \sin(10^\circ). \quad (21)$$

Lines corresponding to the second equation in (21) are linear in nature, this means that along lines depicted by Figure 2, η is constant. According to the graph, it is evident that

$$\eta(x, t) = \eta(x - \sin(10^\circ)t, 0) \implies \boxed{\eta(x, t) = \eta_0(x - \sin(10^\circ)t)} \quad (22)$$

Figure 2: Riemann invariants in the x - t plane (not to scale!)**Part d**

We have to transform (17) into Burgers' equation. To that end, we assume a relationship $h = f(r)$, which transforms (17) into

$$\frac{\partial f}{\partial r} \frac{\partial r}{\partial t} + f^2 \sin(10^\circ) \frac{\partial f}{\partial r} \frac{\partial r}{\partial x} = 0 \implies \frac{\partial r}{\partial t} + f^2 \sin(10^\circ) \frac{\partial r}{\partial x} = 0 \quad (23)$$

which is equivalent to (17). From (23) it is evident that

$$r = h^2 \sin(10^\circ). \quad (24)$$

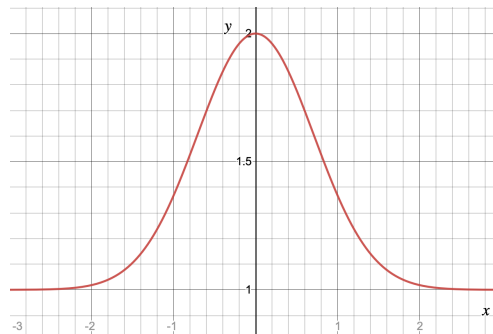
And as a consequence, (17) gets transformed to

$$\frac{\partial r}{\partial t} + r \frac{\partial r}{\partial x} = 0. \quad (25)$$

From (24), it is evident that

$$\frac{dr}{dt} = 0 \quad \text{if} \quad \frac{dx}{dt} = r. \quad (26)$$

Suppose the initial profile from the problem looks like Figure 3.

Figure 3: Example of an initial profile of h

Then, Riemann invariants will not all have a constant slope (like those in part c), but a slope which varies linearly with r (hence quadratically with h !). Hence, to the right of the central initial peak, the height decays and the consequent Riemann invariants have a smaller slope - they do not intersect. From the left of the central peak however, the height increases, so the slopes of the Riemann invariants increase - they intersect. A consequence of Riemann invariants intersecting is a shock creation, so a shock will form on **the section where h grows - to the left of the peak.**

Problem 2

This problem is similar in spirit to the first one, but for Poiseuille's flow down a pipe with slowly varying cross-section. Adopt again, for simplicity, a kinematic viscosity $\mu = 1$ and fluid density $\rho = 1$.

a) For a uniform pipe of radius R and pressure gradient P_x , derive Poiseuille's flow $u(r)$ and the corresponding mass flow rate $Q(R, P_x)$.

b) Consider now a pipe with varying radius $R(x)$, yet varying slowly enough that we can assume that your expression above for Q is still valid. Conservation of mass implies that Q is uniform through the pipe. Assuming that $Q = Q_0$ is also independent of time, derive an ODE for the pressure $P(x)$.

c) Consider a pipe with a contraction near $x = 0$, specifically described by

$$R(x) = 2 - e^{-\frac{x^2}{2}},$$

and $Q_0 = 1$. Solve your ODE for $P(x)$ from part c numerically using the trapezoidal rule:

$$P(x + \Delta x) \approx P(x) + \frac{P_x(x) + P_x(x + \Delta x)}{2} \Delta x$$

for a small enough Δx , and plot the resulting $P(x)$. You may start your integration at an x where $R(x)$ is nearly constant, say $x = -5$, and adopt any value for P there, as it is not the absolute value of the pressure but its gradient that matters for the flow.

Solution

Part a

Given the problem, we must make several assumptions before deriving the velocity of Poiseuille flow down a pipe.

1. The pressure gradient is locally constant - it has no dependence on x in particular.
2. The flow is steady - \mathbf{u} is not a function of time.
3. The flow velocity does not have a component other than that in the x direction - in vectorial form:

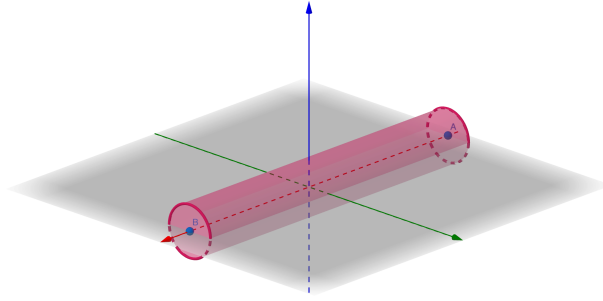
$$\mathbf{u} = \begin{pmatrix} u \\ 0 \\ 0 \end{pmatrix}. \quad (27)$$

4. The flow velocity does not depend on the angle between the central axis of the tube and the line parallel to it, going through a particular point (the flow is cylindrically symmetric - only depends on the distance from the central axis, as shown by Figure 4). This implies $\mathbf{u} = u(r)\hat{x}$.

5. No-slip: $u(R) = 0$.

Then, we recall the Navier-Stokes equation in a single dimensional case:

$$\frac{\partial u}{\partial t} + (\mathbf{u} \cdot \vec{\nabla}) u + \frac{1}{\rho} \frac{\partial p}{\partial x} = \mu \nabla^2 u. \quad (28)$$

Figure 4: A pipe in the x direction facilitating Poiseuille flow

Applying our assumptions reduces (28) to

$$\frac{\partial p}{\partial x} = \eta \nabla^2 u \quad (29)$$

where we used the relationship between the constants $\eta = \mu\rho$. Now, recalling the cylindrical form of the Laplace operator¹ and noting the fact that $\frac{\partial^2 u}{\partial \theta^2} = 0$, we obtain

$$\frac{\partial^2 u}{\partial r^2} + \frac{1}{r} \frac{\partial u}{\partial r} = \frac{P_x}{\eta}. \quad (30)$$

Inputting this PDE into a computing software we obtain a general solution

$$u(r) = \frac{P_x r^2}{4\eta} + a \ln r + b. \quad (31)$$

Applying the no-slip condition at $r = R$, and noting that $\eta = 1$ it must hold that

$$a = 0 \quad \wedge \quad b = -\frac{P_x R^2}{4} \implies \boxed{u(r) = \frac{P_x}{4} (r^2 - R^2)}. \quad (32)$$

We can use (32) to calculate the volume flow rate (hence mass, since $\rho = 1$):

$$Q = \iint_C \mathbf{u} \cdot d\vec{A} = \frac{P_x}{4} \int_0^{2\pi} \int_0^R (r^3 - R^2 r) dr d\theta \quad (33)$$

where we used the fact that the velocity is parallel to the cross-sectional area vector, and that $dA = r dr d\theta$ in polar coordinates. The integral is easily evaluated to be

$$\boxed{Q = -\frac{P_x \pi}{8} R^4} \quad (34)$$

Part b

Conservation of mass tells us that

$$\frac{\partial Q}{\partial x} = 0 \implies Q = Q_0 \quad (35)$$

¹Remark: Here we have a change of coordinates from Cartesian (x, y, z) to cylindrical (x, r, θ) where $y = r \cos \theta$ and $z = r \sin \theta$.

where Q_0 is not a function of time, as given by the problem. Thus, the ODE for P_x is given by

$$\frac{dP}{dx} = -\frac{8Q_0}{(R(x))^4 \pi} \quad (36)$$

Part c

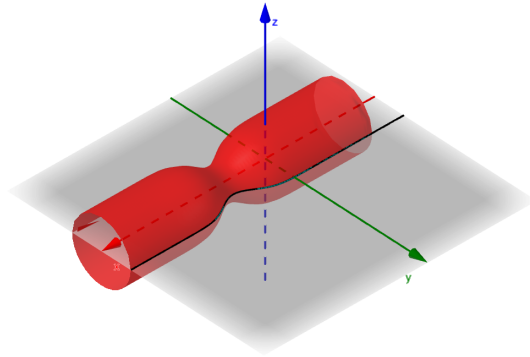


Figure 5: The modified tube around the origin

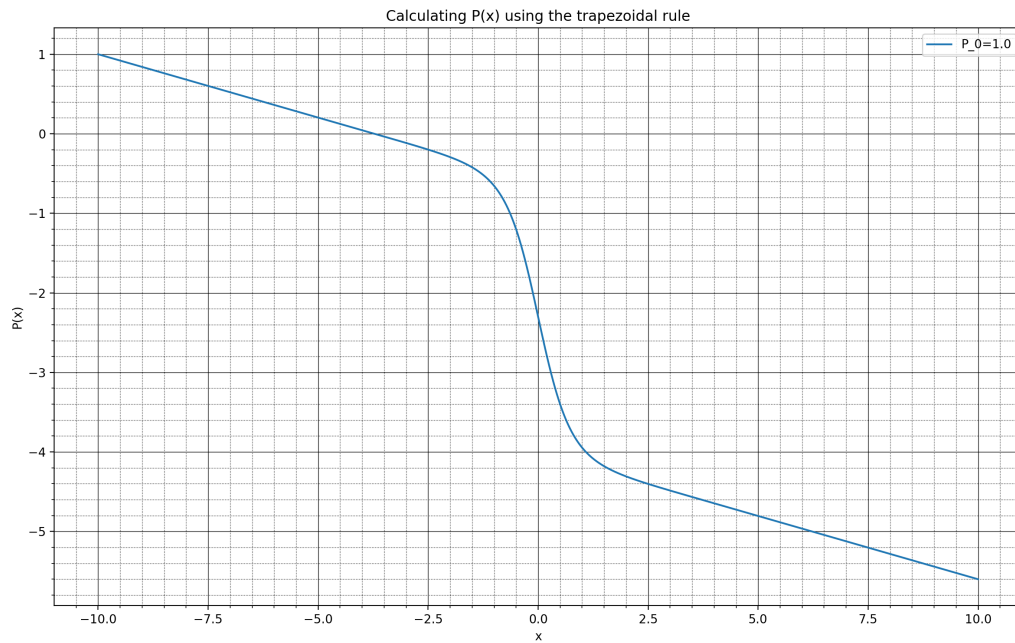


Figure 6: Dependence of P against x

This tube has a slowly varying radius, much like that of a solid of revolution of a Gaussian distribution (see Figure 5). The suggested numerical method was undertaken in a Python script and the results are reported in Figure 6.

In this script, the position interval was taken to be $x \in [-10, 10]$ with a step of $dx = 10^{-4}$, and $P(-10) = P_0 = 1$. Away from the origin, the pressure decays linearly - which is expected for a constant volume flow Q , characteristic for flows influenced by a constant gravitational field².

It is worth noting that the initial condition at $x = -10$ affects $P(x)$ only to the extent of shifting the graph up by P_0 , as shown by Figure 7.

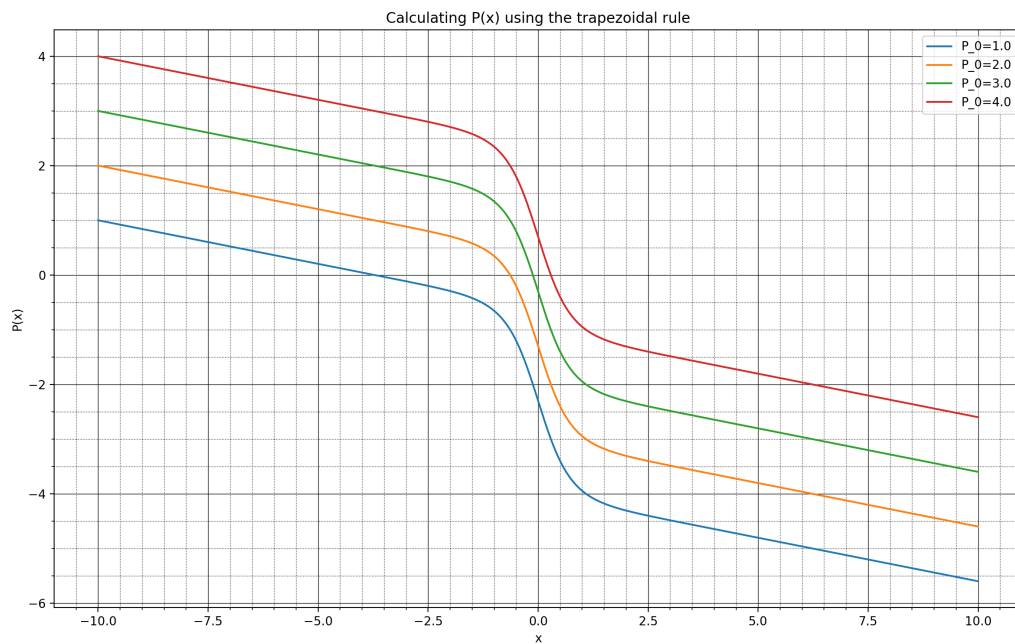


Figure 7: Dependence of P against x for various initial conditions

Furthermore, the code used to obtain Figures 6 and 7 is shown in Figure 8.

²This pressure gradient may not be influenced a constant gravitational field, the comparison is purely a curiosity.

```

trapezium.py
1  import numpy as np
2  import matplotlib.pyplot as plt
3  dx=0.01
4  x=np.arange(-10,10,dx)
5  Q = 1.0
6  P_x = -(8*Q)/((2-np.exp(-(x**2)/2))*4*(np.pi))
7
8  P_0=[1.0,2.0,3.0,4.0]
9
10 P=[]
11 P.append(P_0[0])
12 for i in range (len(P_x)-1):
13     P.append(P[-1]+(P_x[i]+P_x[i+1])*dx/2.0)
14
15 # P_1=[]
16 # P_1.append(P_0[1])
17 # for i in range (len(P_x)-1):
18 #     P_1.append(P_1[-1]+(P_x[i]+P_x[i+1])*dx/2.0)
19
20 # P_2=[]
21 # P_2.append(P_0[2])
22 # for i in range (len(P_x)-1):
23 #     P_2.append(P_2[-1]+(P_x[i]+P_x[i+1])*dx/2.0)
24
25 # P_3=[]
26 # P_3.append(P_0[3])
27 # for i in range (len(P_x)-1):
28 #     P_3.append(P_3[-1]+(P_x[i]+P_x[i+1])*dx/2.0)
29
30
31 plt.title("Calculating P(x) using the trapezoidal rule")
32 plt.minorticks_on()
33 # Customize the major grid
34 plt.grid(which='major', linestyle='-', linewidth='0.5', color='black')
35 # Customize the minor grid
36 plt.grid(which='minor', linestyle=':', linewidth='0.5', color='black')
37 plt.xlabel('x')
38 plt.ylabel('P(x)')
39 plt.plot(x,P)
40 # plt.plot(x,P_1)
41 # plt.plot(x,P_2)
42 # plt.plot(x,P_3)
43 plt.legend(['P_0=1.0'], loc='best')
44 plt.show()

```

Figure 8: Code for trapezoidal rule calculation