

Università degli Studi di Torino

Corso di Laurea Magistrale in Astrofisica e Fisica Teorica

Quantum Field Theory on a Highly Symmetric Lattice

Tesi di Laurea Magistrale

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Abstract

The regularization on a Euclidean lattice, first proposed by Kenneth G. Wilson in 1974, remains the only approach to study strongly coupled, non-supersymmetric non-Abelian gauge theories (including, in particular, quantum chromodynamics: the fundamental theory of the strong nuclear interaction in the Standard Model of elementary-particle physics) from first principles.

While normally the theory is discretized on a four-dimensional hypercubic grid, this is not the only possible choice, and the fact that the explicit breaking of Lorentz-Poincaré symmetries due to the discretization has an impact on the approach to the continuum limit is a motivation to consider the regularization also on other, more symmetric, lattices. The goal of this thesis project consists in studying Yang-Mills theories based on local SU(N) invariance on the lattice of the roots of the exceptional simple Lie group F_4 , which is a four-dimensional body-centered cubic lattice, and the most symmetric regular lattice that exists in four dimensions.

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Yang-Mills Theories on the Lattice

1.1 The Yang-Mills Continuum Action

The aim of this chapter is to discretize the Yang-Mills action on a hypercubic lattice in 4 dimensions. In order to do so, the action is obtained firstly in the continuum, beginning from the simplest case, Quantum Electrodynamics.

This first section is based on material that can be found in standard textbooks on quantum field theory [1–5], personal notes and computations.

1.1.1 Scalar Fields

In quantum field theory, a real scalar massive field is described, in a 4-dimensional spacetime with metric $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$, by the following Lorentz-covariant action (in natural units, where $c = \hbar = 1$):

$$S[\phi] = \int d^4x \left(-\frac{1}{2} \partial^{\mu} \phi \partial_{\mu} \phi - \frac{1}{2} m^2 \phi^2 + V(\phi) \right)$$
 (1.1.1)

where $V(\phi)$ is any potential, such as $\frac{g}{6}\phi^3$ or $\frac{g}{4!}\phi^4$. Real scalar fields do not describe any real-world elementary particle, though they are useful to learn basic principles of quantum field theory, as they are the simplest fields that can be written.

1.1.2 Dirac Spinor Fields

Let us now take into consideration a (free) quantum field theory describing a fermion, such as a quark or a lepton. Its action can be written as:

$$S_{\psi}[\psi(x), \overline{\psi}(x)] = \int d^4x \left(i\overline{\psi} \partial \psi - m\overline{\psi} \psi \right)$$
 (1.1.2)

from which, upon the application of the variational principle, the Dirac equation follows:

$$(i\partial - m) \psi(x) = 0 \tag{1.1.3}$$

It can now be easily checked by direct computation that this action is invariant under a rigid (global) phase transformation, also called a global U(1) transformation:

$$\psi(x) \to \psi'(x) = e^{-i\alpha}\psi(x)$$

$$\overline{\psi}(x) \to \overline{\psi}'(x) = \overline{\psi}(x)e^{i\alpha}$$
(1.1.4)

where α is a constant that does not depend on the spacetime coordinate x, because if α was a function of x, the kinetic term of the action (1.1.2) would not be invariant under such transformation.

1.1.3 Quantum Electrodynamics

As the free field theory itself is non interacting, it does not provide any real-world prediction, so it is useful to write an interacting action where the spinor field is coupled, for instance, to a vector field A_{μ} , i.e. the photon. One way to implement this interaction is to ask for local, instead of global, invariance of the action (1.1.2) under the phase transformation (1.1.4), where now $\alpha = \alpha(x)$. In order to do so, the covariant derivative has to be defined as follows:

$$D_{\mu} \equiv \partial_{\mu} + igA_{\mu} \tag{1.1.5}$$

where q is the coupling constant.¹

The vector field's kinetic term is written in terms of its field-strength, namely:

$$F_{\mu\nu} \equiv -\frac{\mathrm{i}}{g} \left[D_{\mu}, D_{\nu} \right] =$$

$$= -\frac{\mathrm{i}}{g} \left(D_{\mu} \left(\partial_{\nu} + \mathrm{i} g A_{\nu} \right) - D_{\nu} \left(\partial_{\mu} + \mathrm{i} g A_{\mu} \right) \right) =$$

$$= -\frac{\mathrm{i}}{g} \left(\partial_{\mu} \partial_{\nu} + \mathrm{i} g \partial_{\mu} A_{\nu} - g^{2} A_{\mu} A_{\nu} - \partial_{\nu} \partial_{\mu} - \mathrm{i} g \partial_{\nu} A_{\mu} + g^{2} A_{\nu} A_{\mu} \right) =$$

$$= \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu} + \mathrm{i} g \left[A_{\mu}, A_{\nu} \right] =$$

$$= \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu}$$

$$(1.1.6)$$

where $[A_{\mu}, A_{\nu}] = A_{\mu}A_{\nu} - A_{\nu}A_{\mu} = 0$ in the abelian theory.

Two different fields A_{μ} and A'_{μ} describe the same physics if one can be obtained from

¹Usually, in QED, g is called e, the electron charge, though g will be used in analogy to nonabelian gauge theories.

another through a gauge transformation:

$$A'_{\mu}(x) = A_{\mu}(x) + \frac{1}{g}\partial_{\mu}\alpha(x)$$

$$F'_{\mu\nu} = F_{\mu\nu} + \frac{1}{g}(\partial_{\mu}\partial_{\nu} - \partial_{\nu}\partial_{\mu})\alpha(x) = F_{\mu\nu}$$

$$(1.1.7)$$

Thus, the free action for the vector field is:

$$S_{EM} = -\frac{1}{4} \int d^4x F_{\mu\nu} F^{\mu\nu}$$
 (1.1.8)

That is also gauge invariant, i.e. invariant under (1.1.7), as $F_{\mu\nu}$ is gauge invariant. The term that broke the local phase invariance of the action (1.1.2) can now be "absorbed" by A_{μ} through a gauge transformation (1.1.7), thus making the full action gauge invariant:

$$S_{QED} = \int d^4x \left(i\overline{\psi} \not \!\!\!/ \psi - m\overline{\psi}\psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \right) =$$

$$= \int d^4x \left(i\overline{\psi} \not \!\!\!/ \psi - m\overline{\psi}\psi - g\overline{\psi} \not \!\!\!/ \psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \right)$$

$$S_{QED} \to S'_{QED} = \int d^4x \left(i\overline{\psi} \not \!\!/ \psi + \overline{\psi} \not \!\!\!/ \omega \psi - m\overline{\psi}\psi - g\overline{\psi} \not \!\!\!/ \psi - \overline{\psi} \not \!\!\!/ \omega \psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \right) =$$

$$= \int d^4x \left(i\overline{\psi} \not \!\!/ \psi - m\overline{\psi}\psi - g\overline{\psi} \not \!\!\!/ \psi - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} \right) = S_{QED}$$

$$(1.1.9)$$

1.1.4 Non-Abelian Gauge Theories

Let us now consider a theory with N fermions, all with the same mass m, described by the spinorial fields $\psi_i(x)$ with $i=1,\ldots,N$. These N fermions represent the N possible charges of the same particle² and are not to be confused with, for instance, the different possible flavors of the quarks, that describe different particles with different masses. Its free action is:

$$S_{\psi}[\psi_{i}(x), \overline{\psi}_{i}(x)] = \sum_{i=1}^{N} \int d^{4}x \left(i\overline{\psi}_{i} \partial \psi_{i} - m\overline{\psi}_{i} \psi_{i} \right)$$

$$(1.1.10)$$

From now on, the sum over i (and all other repeated latin indexes) will be omitted, unless differently specified. This action is invariant under the global transformation:

$$\psi_i(x) \to \psi_i'(x) = U_{ij}\psi_j(x)$$

$$\overline{\psi}_i(x) \to \overline{\psi}_i'(x) = \overline{\psi}_i(x)U_{ii}^{\dagger}$$
(1.1.11)

if U is any (constant) $N \times N$ matrix such that $UU^{\dagger} = U^{\dagger}U = \mathbb{1} \Leftrightarrow U^{\dagger} = U^{-1}$, or in other words, if $U \in U(N)$. For this reason, this transformation is also called a global U(N) transformation. The phase transformation (1.1.4) is the particular case where $U = e^{-\mathrm{i}\alpha} \in U(1)$, that is the only Abelian (commutative) unitary group.

As $U(N) = SU(N) \otimes U(1) \ \forall N > 1, \ U \in SU(N)$ instead of $U \in U(N)$ can be imposed,

²For example, if N=3 the 3 possible charges are the color charges of QCD, as will be shown later.

and will be from now on, without loss of generality.

In an analogous way to what has been done in Section 1.1.3, this invariance can be made local by implementing a proper covariant derivative, similar to (1.1.5). In order to do so, the infinitesimal SU(N) transformation has to be considered:

$$U_{ij}(x) = \delta_{ij} + i\theta^{a}(x) (T^{a})_{ij} + O(\theta^{2})$$
(1.1.12)

where the indices i and j run from 1 to N (as before) and the index a runs from 1 to N^2-1 (the dimension of the group SU(N)). The matrixes T^a are the N^2-1 generators of $\mathfrak{su}(N)$ (the Lie algebra of SU(N)), thus they are $N\times N$ hermitean and traceless, which obey the commutation relations:

$$\left[T^a, T^b\right] = if^{abc}T^c \tag{1.1.13}$$

where f^{abc} are called *structure constants* of $\mathfrak{su}(N)$. The normalization of these matrices can be chosen such that they obey the condition:

$$Tr(T^a T^b) = \frac{1}{2} \delta^{ab} \tag{1.1.14}$$

some examples are:

- $N=2,\,T^a=rac{\sigma^a}{2},\,{
 m with}\,\,\sigma^a$ the Pauli matrices and $f^{abc}=\varepsilon^{abc};$
- $N=3,\,T^a=\frac{\lambda^a}{2},\,$ with λ^a the Gell-Mann matrices.

where ε^{abc} is the completely antisymmetric Levi-Civita symbol.

The covariant derivative, therefore, is written as:

$$D_{\mu} \equiv \partial_{\mu} + ig\mathbf{A}_{\mu}(x) \tag{1.1.15}$$

where an $N \times N$ identity matrix 1 multiplying ∂_{μ} has to be understood, and $\mathbf{A}_{\mu}(x)$ is a gauge field of SU(N), i.e. a traceless, hermitean $N \times N$ matrix, or, in other words, $\mathbf{A}_{\mu}(x) \in \mathfrak{su}(N)$.

The covariant derivative can be written more explicitly acting on the set of spinors ψ_i :

$$(D_{\mu})_{ij} \psi_j = \partial_{\mu} \mathbb{1}_{ij} \psi_j + ig \left(\mathbf{A}_{\mu}(x) \right)_{ij} \psi_j$$

In order for the action to be gauge invariant, the field \mathbf{A}_{μ} must satisfy the gauge transformation property

$$\mathbf{A}_{\mu}(x) \to \mathbf{A}'_{\mu}(x) = U(x)\mathbf{A}_{\mu}(x)U^{\dagger}(x) - \frac{\mathrm{i}}{q}U(x)\partial_{\mu}U^{\dagger}(x)$$
(1.1.16)

This expression is a little more complicated than (1.1.7), due to the fact that \mathbf{A}_{μ} is now a non-commuting matrix. However if the Abelian case U(1) is taken into consideration, where $U(x) = e^{-i\alpha(x)}$, (1.1.7) follows directly from (1.1.16).

Now, it can be easily checked that the kinetic term of the Lagrangian

$$\mathcal{L}_K = i\overline{\psi}_i D\!\!\!/ \psi_i = i\overline{\psi}_i \partial\!\!\!/ \psi_i - g\overline{\psi}_i A\!\!\!/ \psi_i$$

is gauge invariant (i.e. invariant under (1.1.11) and (1.1.16)) through direct computation:

$$\mathcal{L}_{K} \to \mathcal{L}_{K}' = i\overline{\psi}_{i}U^{\dagger} \partial (U\psi_{i}) - g\overline{\psi}_{i} \underbrace{U^{\dagger}U}_{1} A \underbrace{U^{\dagger}U}_{1} \psi_{i} + i\overline{\psi}_{i} \underbrace{U^{\dagger}U}_{1} (\partial U^{\dagger}) U\psi_{i} =$$

$$= i\overline{\psi}_{i}U^{\dagger} (\partial U) \psi_{i} + i\underline{\psi}_{i} \partial \psi_{i} - g\overline{\psi}_{i} A \psi_{i} + i\overline{\psi}_{i} (\partial U^{\dagger}) U\psi_{i} =$$

$$= \mathcal{L}_{K} + i\overline{\psi}_{i}\gamma^{\mu} (U^{\dagger}\partial_{\mu}U + \partial_{\mu}U^{\dagger}U) \psi_{i} =$$

$$= \mathcal{L}_{K} + i\overline{\psi}_{i}\gamma^{\mu}\partial_{\mu} (U^{\dagger}U) \psi_{i} =$$

$$= \mathcal{L}_{K} + i\overline{\psi}_{i}\gamma^{\mu} \partial_{\mu} (1) \psi_{i} = \mathcal{L}_{K}$$

Because of this fact, it is directly implied that the covariant derivative (1.1.15) must transform, under a gauge transformation, in the adjoint representation:

$$D_{\mu} \to D_{\mu}' = U D_{\mu} U^{\dagger} \tag{1.1.17}$$

The field-strength for the field \mathbf{A}_{μ} is obtained, as for the Abelian case, through the commutator of two covariant derivatives. The computation is the same as (1.1.6), but this time the commutator term is non-vanishing:

$$F_{\mu\nu} \equiv -\frac{\mathrm{i}}{g} [D_{\mu}, D_{\nu}] = \partial_{\mu} \mathbf{A}_{\nu} - \partial_{\nu} \mathbf{A}_{\mu} + \mathrm{i}g[\mathbf{A}_{\mu}, \mathbf{A}_{\nu}]$$
 (1.1.18)

This expression can be simplified a little by considering that \mathbf{A}_{μ} and $F_{\mu\nu}$ are elements of $\mathfrak{su}(N)$, thus writing them in terms of their components w. r. t. the basis T^a :

$$\mathbf{A}_{\mu}(x) = A_{\mu}^{a}(x)T^{a} \tag{1.1.19}$$

$$F_{\mu\nu}(x) = F^a_{\mu\nu}(x)T^a \tag{1.1.20}$$

and by considering the relation (1.1.13):

$$\begin{split} F_{\mu\nu}^{a}T^{a} &= \left(\partial_{\mu}A_{\nu}^{a} - \partial_{\nu}A_{\mu}^{a}\right)T^{a} + \mathrm{i}g\left[A_{\mu}^{b}T^{b}, A_{\nu}^{c}T^{c}\right] = \\ &= \left(\partial_{\mu}A_{\nu}^{a} - \partial_{\nu}A_{\mu}^{a}\right)T^{a} + \mathrm{i}gA_{\mu}^{b}A_{\nu}^{c}\underbrace{\left[T^{b}, T^{c}\right]}_{if^{bca}T^{a}} = \\ &= \left(\partial_{\mu}A_{\nu}^{a} - \partial_{\nu}A_{\mu}^{a} - gf^{abc}A_{\mu}^{b}A_{\nu}^{c}\right)T^{a} \\ F_{\mu\nu}^{a} &= \partial_{\mu}A_{\nu}^{a} - \partial_{\nu}A_{\mu}^{a} - gf^{abc}A_{\mu}^{b}A_{\nu}^{c} \end{split} \tag{1.1.21}$$

In order to write a kinetic action for the field \mathbf{A}_{μ} , a term proportional to $F_{\mu\nu}F^{\mu\nu}$, like in (1.1.8), is not enough: because of (1.1.17) and the definition (1.1.18), it must transform as $F_{\mu\nu}F^{\mu\nu} \to UF_{\mu\nu}F^{\mu\nu}U^{\dagger}$, therefore it would not be gauge invariant. In fact a gauge invariant action, called Yang-Mills action, is:

$$S_{YM} = -\frac{1}{2} \int d^4x \, \text{Tr} \left(F_{\mu\nu} F^{\mu\nu} \right) \tag{1.1.22}$$

because of the ciclic property of the trace.³ This action can be written in components, using (1.1.20) and the trace property (1.1.14):

$$S_{YM} = -\frac{1}{2} \int d^4x \, \text{Tr} \left(F^a_{\mu\nu} F^{b\mu\nu} T^a T^b \right) = -\frac{1}{4} \int d^4x F^a_{\mu\nu} F^{a\mu\nu}$$
 (1.1.23)

³Actually, a term proportional to det $(F_{\mu\nu}^a F^{a\mu\nu})$ would be gauge invariant as well, but it would not be a suitable kinetic term as it would involve terms of higher order than 2 in the components $F_{\mu\nu}^a$

Here, there are two remarks that need to be done. The first one is that, if the gauge group is taken to be U(1), the action (1.1.23) reduces to (1.1.8), as a=1 because the group U(1) has only 1 generator. The second one is that, if non-Abelian gauge groups are taken into consideration, this action naturally introduces self-interacting cubic and quartic terms, because the structure constants f^{abc} are non-vanishing. This is, for example, the case for the group SU(3), that is used to describe gluon interaction, i.e. Quantum Chromodynamics (QCD). These self-interactions make the Yang-Mills action interesting to be studied even alone, without any other fermionic or bosonic interacting field, as it will be shown later.

1.1.5 Wick Rotation

Up to now, actions were written in Minkowskian spacetime, where $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$. In order to have a positive-defined metric $\eta_{\mu\nu} = \delta_{\mu\nu}$ a Wick rotation can be performed by re-defining the time coordinate to be $\tau = \mathrm{i}t = \mathrm{i}x^0$, as is rigorously proven in [6]. The principle behind Wick rotation is that the time coordinate is promoted to be a complex number and the integration (in the action) from $-\infty$ to ∞ is made to be from $-\mathrm{i}\infty$ to $+\mathrm{i}\infty$ through a rotation of $\frac{\pi}{2}$ in the complex plane, assuming that correlators and other physical quantities do not present singularities in the first and third quadrant of the complex plane of t. This rotation is not only a matter of convenience, in fact it is needed in perturbative QFT for the computation of (otherwise oscillating) functional integrals and in lattice field theory because a Minkowskian lattice cannot be rigorously defined.

The relation between the Minkowski actions written before and the Euclidean actions is $S^M = iS^E$, because $d^4x^E = id^4x^M$.

Thus, the Euclidean actions are⁴:

$$S^{E}[\phi] = \int d^{4}x \left(\frac{1}{2} \partial^{\mu}\phi \partial_{\mu}\phi + \frac{1}{2} m^{2}\phi^{2} + V(\phi) \right)$$
 (1.1.24)

$$S_F^E = \int d^4x \overline{\psi} \left(\gamma^{\mu} \left(\partial_{\mu} + ig \mathbf{A}_{\mu} \right) + m \right) \psi \tag{1.1.25}$$

$$S_{YM}^{E} = \frac{1}{4} \int d^{4}x F_{\mu\nu}^{a} F^{a\mu\nu}$$
 (1.1.26)

where the superscript E, meaning that the action is written in the Euclidean spacetime, will be omitted from now on.

1.2 Lattice Field Theory

1.2.1 Why Lattice Field Theory

After the Wick rotation, quantum field theory can be studied perturbatively through the saddle-point evaluation of the path integral, where the partition function

$$Z = \int \mathcal{D}\varphi e^{\mathrm{i}S^M[\varphi]} = \int \mathcal{D}\varphi e^{-S^E[\varphi]}$$

⁴Note that they are not the same as their Minkowskian counterpart, as some redefinitions of the fields and the γ matrices have been implicitly made, in order to have the respective theories non ill-defined.

has become non-oscillating. This approach, however, works well only if the coupling constants are *small enough* and does not allow to obtain non-perturbative results, i.e. physical quantities that depend on essential singularities in the coupling constant, like glueballs for Yang-Mills theories.

For this reason, a non-perturbative approach has to be taken into consideration and one possibility is Lattice Field Theory: the spacetime is discretized to a lattice, with lattice spacing a, and the fields can assume different values only on particular parts of the lattice. More in detail, a scalar field lives on lattice sites, a vector field lives on links between sites and objects with k indices live on k-simplexes.

If the limit $a \to 0$ is taken, the theory must reproduce results obtained in the continuum, like the ones obtained in perturbation theory in the regime where it can be applied, however the presence of a discrete spacetime provides a natural cutoff for the momenta, allowing ultraviolet divergencies to be kept under control, although they need to be taken into consideration when approaching the continuum limit.

In this section, based on standard quantum field theory textbooks cited before and lattice field theory textbooks [7–9], the regularization of the previous quantum field theories on a Simple Hypercubic (SH) lattice is presented. A more detailed description of the geometric properties of the figures mentioned below can be found in [10].

1.2.2 Lattice

A lattice Λ in \mathbb{R}^D is defined as the set of all possible linear combinations with integer coefficients of the vectors $\{v_i\}$, which form a basis of \mathbb{R}^D . Formally:

$$\Lambda = \left\{ \sum_{i=1}^{D} c_i v_i \mid c_i \in \mathbb{Z} \right\}$$
 (1.2.1)

1.2.2.1 Simple Hypercubic Lattice

If the basis $\{v_i\}$ is taken to be orthonormal, the lattice is called $Simple\ Hypercubic^5$. In Figure (1.2.1) is represented a portion of a Simple Hypercubic lattice in D=3.

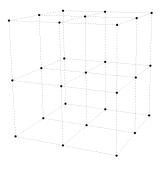


Figure 1.2.1: A cubic lattice.

For the SH lattice in D=4, the fundamental region (the smallest D-dimensional polyhedron) is a 4-dimensional hypercube, also called a tesseract. This means that a SH lattice

⁵The word *simple* is used to make a distinction with the Body-Centered Hypercubic lattice.

can be seen as a tassellation of the spacetime with tesseracts as elementary cells. Each point of the lattice has 8 nearest neighbours that are identified by the vectors obtained through all possible permutations of position and sign of $(\pm 1, 0, 0, 0)$.

The plaquette (the simplest bidimensional figure) is a square. This will be important when implementing gauge theories on the lattice.

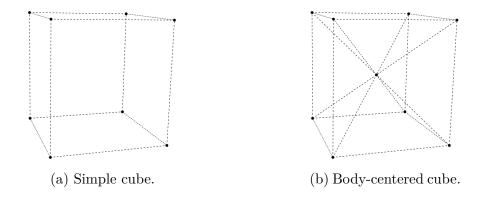


Figure 1.2.2: Tridimensional representation of a simple cubic cell (1.2.2a) and a body-centered one (1.2.2b).

1.2.2.2 Body-Centered Tesseract

The SH lattice is not, of course, the only possible choice of a lattice in 4 spacetime dimensions. Another common choice, already used to simulate gauge theories [11], is the Body-Centered Tesseract (BCT). It consists of packing the spacetime with tesseracts, as the name suggests, but considering both the corners and the centers of every hypercube as lattice sites. Every site has, therefore, 24 nearest neighbours: 16 are identified by all possible sign permutations of $(\pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2})$, the 8 remaining are the ones of the SH lattice. The cell of this lattice is known as 24-cell, shown in Figure (1.2.3), and the plaquettes are triangular.

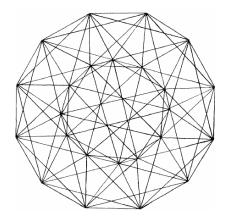


Figure 1.2.3: Bidimensional projection of the 24-cell.

1.2.3 Scalar Fields on the Simple Hypercubic Lattice

Let us consider a simple hypercubic lattice Λ_{SH} extending over $L=L_1=L_2=L_3$ lattice spacings in the spatial directions and over $T=L_0$ lattice spacings in the temporal direction. Let us also consider a scalar field $\phi(x)$, in order to introduce some useful tools that will be needed later. As anticipated in the beginning of this section, a scalar field can only assume values on the sites of the lattice, therefore $x \in \Lambda_{SH}$, and let us assume, for simplicity, periodic boundary conditions on the fields, therefore $\phi(x + a\hat{\mu}L_{\mu}) = \phi(x)$. The Fourier transform of the field becomes

$$\tilde{\phi}(p) = \sum_{x} a^4 e^{-ip \cdot x} \phi(x) \tag{1.2.2}$$

and the allowed momenta are given by

$$p_{\mu} = \frac{2\pi}{aL_{\mu}}n_{\mu}, \qquad n_{\mu} = 0, \dots, L_{\mu}$$
 (1.2.3)

This ensures that the momenta can take only the assigned values of the Brillouin zone (1.2.3) and therefore cannot go to infinity, providing a natural cutoff given by the lattice spacing a.

The discretization also implies that derivatives of the fields cannot be computed, therefore the lattice forward derivative is used:⁶

$$\hat{c}_{\mu}\phi(x) \to \nabla_{\mu}\phi(x) = \frac{\phi(x+a\hat{\mu}) - \phi(x)}{a}$$
(1.2.4)

Assuming a self-interaction potential $V(\phi) = \frac{\lambda}{4!}\phi^4$, the action (1.1.1) can be written in terms of the lattice in the following way:

$$S = a^4 \sum_{x} \left(\frac{1}{2a^2} \left[\phi(x + a\hat{\mu}) - \phi(x) \right]^2 + \frac{1}{2} m^2 \phi^2(x) + \frac{\lambda}{4!} \phi^4(x) \right)$$
(1.2.5)

A similar approach will be used in the following section to obtain the action for Yang-Mills theories.

1.2.4 Gauge Fields on the Simple Hypercubic Lattice

Let us now take into consideration a Yang-Mills theory, as in Section 1.1.4, with gauge group SU(N). As anticipated before, gauge fields, being vector fields, can assume value only on links between sites of the lattice.

One could naively think that putting gauge vectors $A_{\mu}(x)$ on the links is enough, but this would explicitly break gauge invariance. For this reason, Wilson's idea was to put the gauge group (and not algebra) elements, namely $U_{\mu}(x) = e^{iagA_{\mu}(x)}$, on the links.

The field $U_{\mu}(x)$ lives on the link connecting the site x with the site $x + a\hat{\mu}$, therefore the link is oriented and link variables in negative directions can also be defined: $U_{\mu}(x) = U^{\dagger}(x - a\hat{\mu})$

$$U_{-\mu}(x) \equiv U_{\mu}^{\dagger}(x - a\hat{\mu}).$$

⁶Note that, in the continuum limit, the two derivatives coincide.

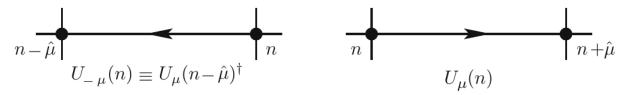


Figure 1.2.4: Schematic visualization of link variables.

Under a gauge transformation, the fields U_{μ} transform according to the following relation:

$$U_{\mu}(x) \to \Omega(x)U_{\mu}(x)\Omega^{\dagger}(x+a\hat{\mu})$$
 (1.2.6)

where $\Omega(x)$ is any SU(N) matrix at the point x on the lattice. As a consequence of this relation, the trace of any product of links forming a closed path is a gauge-invariant quantity, thanks to the cyclic property of the trace.

With this in mind, Wilson's idea was to choose the simplest closed path possible: the plaquette, that in the simple hypercubic lattice is a square. The product of link variables along a square plaquette is defined in the following way (the lattice spacing a is set = 1 for brevity of notation):

$$U_{\mu\nu}(x) \equiv U_{\mu}(x)U_{\nu}(x+\hat{\mu})U_{-\mu}(x+\hat{\mu}+\hat{\nu})U_{-\nu}(x+\hat{\nu}) =$$

$$= U_{\mu}(x)U_{\nu}(x+\hat{\mu})U_{\mu}^{\dagger}(x+\hat{\nu})U_{\nu}^{\dagger}(x)$$
(1.2.7)

Thus a gauge-invariant action, called Wilson action, can be written as follows:

$$S_W[U] = \frac{\beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \operatorname{Re} \operatorname{Tr} [\mathbb{1} - U_{\mu\nu}(x)]$$
 (1.2.8)

where β is a parameter that is going to be set in the following passages, N the number of colors (the same N in SU(N)), the real part is needed to preserve unitarity and $U_{\mu\nu}$ is the plaquette of (1.2.7).

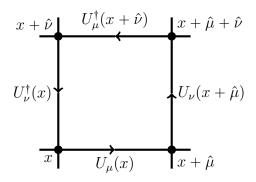


Figure 1.2.5: Link variables building a plaquette $U_{\mu\nu}(x)$.

In the following part, the naive continuum limit is taken, showing that Wilson action reproduces Yang-Mills action of (1.1.26) when $a \to 0$. Let us start by defining an auxiliary field $B_{\mu}(x)$:

$$B_{\mu}(x) = agA^a_{\mu}(x)T^a \tag{1.2.9}$$

thus the gauge group element becomes $U_{\mu}(x) = e^{iB_{\mu}(x)}$. The expansion for small a of the plaquette $U_{\mu\nu}$ is:

$$U_{\mu\nu} = U_{\mu}(x)U_{\nu}(x+\hat{\mu})U_{\mu}^{\dagger}(x+\hat{\nu})U_{\nu}^{\dagger}(x) =$$

$$= e^{iB_{\mu}(x)}e^{iB_{\nu}(x+\hat{\mu})}e^{-iB_{\mu}(x+\hat{\nu})}e^{-iB_{\nu}(x)} =$$

$$= applying (1.2.4), B_{\mu}(x+\hat{\nu}) = a\nabla_{\nu}B_{\mu}(x)$$

$$= e^{iB_{\mu}(x)}e^{iB_{\nu}(x)+a\nabla_{\mu}B_{\nu}(x)}e^{-iB_{\mu}(x)-a\nabla_{\nu}B_{\mu}(x)}e^{-iB_{\nu}(x)} \simeq$$
From now on, the dependance from x will be omitted.

Using BCH formula, $e^{X}e^{Y} = e^{X+Y+\frac{1}{2}[X,Y]}$, up to $O(a^{2})$:
$$\simeq \exp\left\{iB_{\mu}+iB_{\nu}+ia\nabla_{\mu}B_{\nu}-iB_{\mu}-ia\nabla_{\nu}B_{\mu}-iB_{\nu}+\frac{1}{2}(-[B_{\mu},B_{\nu}]+[B_{\mu},B_{\nu}]+[B_{\nu},B_{\mu}]-[B_{\mu},B_{\nu}])\right\} =$$

$$= \exp\{ia(\nabla_{\mu}B_{\nu}-\nabla_{\nu}B_{\mu})-[B_{\mu},B_{\nu}]\}$$

Now, substituting back (1.2.9), the discretized electromagnetic tensor can be found:

$$U_{\mu\nu} \simeq \exp\{ia^{2}g \left(\nabla_{\mu}\mathbf{A}_{\nu} - \nabla_{\nu}\mathbf{A}_{\mu} + ig[\mathbf{A}_{\mu}, \mathbf{A}_{\nu}]\right)\} = \exp\{ia^{2}gF_{\mu\nu}\} \simeq$$

$$\simeq \mathbb{1} + ia^{2}gF_{\mu\nu} - \frac{1}{2}a^{4}g^{2}F_{\mu\nu}F^{\mu\nu} + O(a^{6})$$
(1.2.10)

Finally, the small-a approximation for the Wilson action (1.2.8) can be obtained:

$$S_{W} \simeq \frac{\beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \operatorname{Re} \operatorname{Tr} \left(\mathcal{X} - \mathcal{X} - ia^{2}gF_{\mu\nu}^{i}T^{i} + \frac{1}{2}a^{4}g^{2}F_{\mu\nu}^{i}F^{j\mu\nu}T^{i}T^{j} \right) =$$

$$= -i\frac{a^{2}\beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \operatorname{Re} \left(F_{\mu\nu}^{i} \operatorname{Tr} \left(T^{i} \right) \right) + \frac{g^{2}a^{4}\beta}{4N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \operatorname{Re} \left(\underbrace{F_{\mu\nu}^{i}F^{j\mu\nu}\operatorname{Tr} \left(T^{i}T^{j} \right)}_{= \frac{1}{2}F_{\mu\nu}^{i}F^{i\mu\nu}\operatorname{for} (1.1.14)}^{\text{for} (1.1.12)} \right) =$$

$$= \frac{g^{2}\beta}{8N} a^{4} \sum_{x,\mu,\nu} F_{\mu\nu}^{i}F^{i\mu\nu} \xrightarrow[a \to 0]{} \frac{g^{2}\beta}{8N} \int d^{4}x F_{\mu\nu}^{i}F^{i\mu\nu}$$

$$(1.2.11)$$

where the Yang-Mills action (1.1.26) can be recognized if $\beta = \frac{2N}{g^2}$.

1.2.4.1 Wilson Loops

Now that the lattice action has been defined, observables need to be taken into account. As mentioned below equation (1.2.6), the trace of the product of link variables along any closed path is a gauge-invariant quantity. This means that any quantity of the sort is a good candidate for an observable and, in fact, given any closed path γ on the lattice, the Wilson loop is defined as follows:

$$W[\gamma] \equiv \text{Tr} \left[\prod_{(x,\mu) \in \gamma} U_{\mu}(x) \right]$$
 (1.2.12)

It is worthwile to mention that this definition is does not depend on the type of lattice and of boundary conditions chosen for such lattice. If the path γ does not include any

temporal link, the loop is called spacelike Wilson loop, otherwise it is called timelike Wilson loop. If all the links in γ lie on the same plane, the Wilson loop is planar, otherwise it is nonplanar.

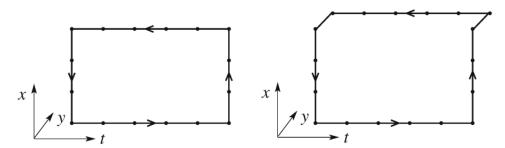


Figure 1.2.6: Example of a planar (left) and nonplanar (right) Wilson loop.

Wilson loops are important for a variety of reasons, that will not be elaborated further as this is not the focus of this project, however it is worth knowing that they can be used as operators for purely gluonic bound states, the so-called *glueballs*: the mass spectrum of these states can be obtained through the exponential decay of Wilson loops' correlation functions.

1.2.4.2 Polyakov Loops

Another important observable can be defined in lattices where periodic boundary conditions in the time direction are assumed. Thanks to periodicity, a whole new class of closed paths arises: all paths that wind around the time direction. A Polyakov loop is thus defined as a line at a fixed spatial position $\vec{x} = (x^1, x^2, x^3)$ that runs along the whole time direction:

$$P(\vec{x}) \equiv \text{Tr} \left[\prod_{t=0}^{T-1} U_0(t, \vec{x}) \right]$$
 (1.2.13)

The Polyakov loop has a lot of interesting properties: its physical interpretation is the world-line of a single static quark, therefore its expectation value is an order parameter of the deconfinement transition, as well as of a new global symmetry, called *center symmetry*, arising from the periodicity of the time direction.

This means that, if the system is confined (quarks and gluons cannot move around freely), its expectation value is 0, while if the system is deconfined, $\langle P(\vec{x}) \rangle \neq 0$.

Another property, that will be used in the simulations, is that the expectation value of the product of two Polyakov loops at distance $r = a|\vec{x} - \vec{y}|$ is a function of the static quark-antiquark potential V(r), as indicated below:

$$\langle P(\vec{x})P^{\dagger}(\vec{y})\rangle \propto e^{-TaV(r)} \left(1 + O(e^{-Ta\Delta E})\right)$$
 (1.2.14)

where ΔE is the difference between V(r) and the first excited energy level of the quark-antiquark pair.

The value of the potential V(r) will be used, in this project, to study the recovery of rotational invariance when approaching the continuum limit.

Various methods of performing computer simulations of lattice field theory and how to obtain these quantities, such as plauette values, Wilson and Polyakov loops, is shown in the following chapter.

Computer Simulation of Pure Gauge Theories

2.1 Introduction

Monte Carlo simulations are a powerful tool that can be used to evaluate observables in lattice field theory. Given an observable O (such as a plaquette, a Wilson or Polyakov loop, etc.), its vacuum expectation value is formally given by the functional integral

$$\langle O \rangle = \frac{1}{Z} \int \mathcal{D}[U] e^{-S[U]} O[U] \quad \text{with} \quad Z = \int \mathcal{D}[U] e^{-S[U]} O[U]$$
 (2.1.1)

where $\mathcal{D}[U]$ is to be intended as a Haar measure.

This expression, however, cannot be evalued analytically except for very small lattices, therefore (2.1.1) is approximated by an average of the observable evaluated on N sample gauge field configurations $\{U_n\}^1$, distributed according a probability density $\propto e^{-S[U_n]}$.

¹Here the subscript n distinguishes the different configurations.

The expectation value is then obtained by computing the following sum for a sufficient number of configurations generated by the proper Monte Carlo algorithm(s):

$$\langle O \rangle \simeq \frac{1}{N} \sum_{\{U_n\}} O[U_n]$$
 (2.1.2)

Because of the probability density $\propto e^{-S[U_n]}$, only configurations $\{U_n\}$ that minimize the action are good configurations, as all other configurations are exponentially suppressed. For this reason, generating totally random gauge fields on the lattice links is not an efficient way to evaluate (2.1.2), as most of the $\{U_n\}$ will have a very small Boltzmann factor (the $e^{-S[U_n]}$) and expression (2.1.2) will give incorrect results unless a huge number (orders of magnitude higher than what is reasonably possible) of different configurations is tried.

In order to avoid the evaluation of a great number of configurations that would contribute little-to-nothing to the observables' values, a sequence of configurations $\{U_n\}$ is generated through a Markov chain process, built such that its stationary distribution minimizes the action S[U]. This process is usually called *thermalization*.

Every process of this type must begin from a starting configuration, usually chosen by the user. The main starting configurations are usually two: if the simulation begins with the fields in an ordered way (for example, all the gauge links set to the identity), the initial configuration is called *cold start*, if the simulation begins with random gauge fields in every link, the initial configuration is called *hot start*. Of course, the Markov chain must have the same stationary distribution if the hot or cold start is chosen.

In this chapter the main Monte Carlo algorithms used to thermalize the starting configuration are presented, then the method for evaluating some observables of interest, such as plaquettes and Polyakov loops is explained.

2.2 Markov Processes

As anticipated before, a Markov chain is needed to evolve an arbitrary starting configuration U_0 up to a region of the configuration space with a relatively large Boltzmann factor e^{-S} , therefore with high probability. A Markov process is charaterized by a probability of transitioning from a configuration α to another configuration β depending only on α and β and not on the history of the process, namely the previous transitions that already occourred. In formulae, this is expressed as

$$P(U_n = \beta | U_{n-1} = \alpha) = T(\beta | \alpha)$$
(2.2.1)

That is to say that the transition matrix T does not depend on the index n, representing the computer time.

Being a transition probability, the matrix T must obey the following equations

$$0 \leqslant T(\beta|\alpha) \leqslant 1 \qquad \forall \alpha, \beta \tag{2.2.2}$$

$$\sum_{\beta} T(\beta | \alpha) = 1 \qquad \forall \alpha$$
 (2.2.2)

where (2.2.2) is a consequence of $T(\beta|\alpha)$ representing a probability, and (2.2.3) means that the probability of transitioning to any configuration must be 1 (of course, the case when $\alpha = \beta$ is included as well).

In order for the stochastic process to not have any sink or source of probability, the following balance equation must be satisfied:

$$\sum_{\alpha} T(\beta|\alpha)P(\alpha) = \sum_{\alpha} T(\alpha|\beta)P(\beta)$$
 (2.2.4)

This equation states that the probability of transitioning to the configuration β , written in the left-hand side as the sum of the transition probability from the configuration α weighted by the probability $P(\alpha)$ that the system is actually in that configuration, must be equal to the probability of transitioning out of the configuration β , given by the probability of finding the system in the configuration β times the transition probability $T(\alpha|\beta)$ over all the final configurations, in the right-hand side. Thanks to (2.2.3), the right-hand side is easily proven to be equal to $P(\beta)$.

A sufficient (but not necessary) condition to obey the balance equation (2.2.4) is obtained by requiring that it holds true term-by-term, thus obtaining the detailed balance condition:

$$T(\beta|\alpha)P(\alpha) = T(\alpha|\beta)P(\beta) \tag{2.2.5}$$

Although it is not a necessary condition, most algorithms, including the ones discussed in the following section, satisfy it.

2.3 Monte Carlo Algorithms

- 2.3.1 Metropolis Algorithm
- 2.3.2 Heat Bath Algorithm
- 2.3.3 Overrelaxation Algorithm

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Simulation Results

Conclusions

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