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**Quantum Field Theory on a Highly  
Symmetric Lattice**

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# Abstract

The regularization on a Euclidean lattice, first proposed by Kenneth G. Wilson in 1974, remains the only approach to study strongly coupled, non-supersymmetric non-Abelian gauge theories (including, in particular, quantum chromodynamics: the fundamental theory of the strong nuclear interaction in the Standard Model of elementary-particle physics) from first principles.

While normally the theory is discretized on a four-dimensional hypercubic grid, this is not the only possible choice, and the fact that the explicit breaking of Lorentz-Poincaré symmetries due to the discretization has an impact on the approach to the continuum limit is a motivation to consider the regularization also on other, more symmetric, lattices.

The goal of this thesis project consists in studying Yang-Mills theories based on local  $SU(N)$  invariance on the lattice of the roots of the exceptional simple Lie group  $F_4$ , which is a four-dimensional body-centered cubic lattice, and the most symmetric regular lattice that exists in four dimensions.



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# Yang-Mills Theories on the Lattice

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## 1.1 The Yang-Mills Continuum Action

The aim of this chapter is to discretize the Yang-Mills action on a hypercubic lattice in 4 dimensions. In order to do so, the action is obtained firstly in the continuum, beginning from the simplest case, Quantum Electrodynamics.

This first section is based on material that can be found in standard textbooks on quantum field theory [1–5], personal notes and computations.

### 1.1.1 Scalar Fields

In quantum field theory, a real scalar massive field is described, in a 4-dimensional space-time with metric  $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$ , by the following Lorentz-covariant action (in natural units, where  $c = \hbar = 1$ ):

$$S[\phi] = \int d^4x \left( -\frac{1}{2} \partial^\mu \phi \partial_\mu \phi - \frac{1}{2} m^2 \phi^2 + V(\phi) \right) \quad (1.1.1)$$

where  $V(\phi)$  is any potential, such as  $\frac{g}{6} \phi^3$  or  $\frac{g}{4!} \phi^4$ . Real scalar fields do not describe any real-world elementary particle, though they are useful to learn basic principles of quantum field theory, as they are the simplest fields that can be written.

### 1.1.2 Dirac Spinor Fields

Let us now take into consideration a (free) quantum field theory describing a fermion, such as a quark or a lepton. Its action can be written as:

$$S_\psi[\psi(x), \bar{\psi}(x)] = \int d^4x (\bar{\psi} \not{\partial} \psi - m \bar{\psi} \psi) \quad (1.1.2)$$

from which, upon the application of the variational principle, the Dirac equation follows:

$$(i \not{\partial} - m) \psi(x) = 0 \quad (1.1.3)$$

It can now be easily checked by direct computation that this action is invariant under a rigid (global) phase transformation, also called a global  $U(1)$  transformation:

$$\begin{aligned} \psi(x) &\rightarrow \psi'(x) = e^{-i\alpha} \psi(x) \\ \bar{\psi}(x) &\rightarrow \bar{\psi}'(x) = \bar{\psi}(x) e^{i\alpha} \end{aligned} \quad (1.1.4)$$

where  $\alpha$  is a constant that does not depend on the spacetime coordinate  $x$ , because if  $\alpha$  was a function of  $x$ , the kinetic term of the action (1.1.2) would not be invariant under such transformation.

### 1.1.3 Quantum Electrodynamics

As the free field theory itself is non interacting, it does not provide any real-world prediction, so it is useful to write an interacting action where the spinor field is coupled, for instance, to a vector field  $A_\mu$ , i.e., the photon. One way to implement this interaction is to ask for local, instead of global, invariance of the action (1.1.2) under the phase transformation (1.1.4), where now  $\alpha = \alpha(x)$ . In order to do so, the covariant derivative has to be defined as follows:

$$D_\mu \equiv \partial_\mu + ig A_\mu \quad (1.1.5)$$

where  $g$  is the coupling constant.<sup>1</sup>

The vector field's kinetic term is written in terms of its field-strength, namely:

$$\begin{aligned} F_{\mu\nu} &\equiv -\frac{i}{g} [D_\mu, D_\nu] = \\ &= -\frac{i}{g} (D_\mu (\partial_\nu + ig A_\nu) - D_\nu (\partial_\mu + ig A_\mu)) = \\ &= -\frac{i}{g} (\cancel{\partial_\mu \partial_\nu} + ig \partial_\mu A_\nu - g^2 A_\mu A_\nu - \cancel{\partial_\nu \partial_\mu} - ig \partial_\nu A_\mu + g^2 A_\nu A_\mu) = \\ &= \partial_\mu A_\nu - \partial_\nu A_\mu + ig [A_\mu, A_\nu] = \\ &= \partial_\mu A_\nu - \partial_\nu A_\mu \end{aligned} \quad (1.1.6)$$

where  $[A_\mu, A_\nu] = A_\mu A_\nu - A_\nu A_\mu = 0$  in the abelian theory.

Two different fields  $A_\mu$  and  $A'_\mu$  describe the same physics if one can be obtained from

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<sup>1</sup>Usually, in QED,  $g$  is called  $e$ , the electron charge, though  $g$  will be used in analogy to nonabelian gauge theories.



another through a gauge transformation:

$$\begin{aligned} A'_\mu(x) &= A_\mu(x) + \frac{1}{g} \partial_\mu \alpha(x) \\ F'_{\mu\nu} &= F_{\mu\nu} + \frac{1}{g} (\partial_\mu \partial_\nu - \partial_\nu \partial_\mu) \alpha(x) = F_{\mu\nu} \end{aligned} \quad (1.1.7)$$

Thus, the free action for the vector field is:

$$S_{EM} = -\frac{1}{4} \int d^4x F_{\mu\nu} F^{\mu\nu} \quad (1.1.8)$$

That is also gauge invariant, i.e. invariant under (1.1.7), as  $F_{\mu\nu}$  is gauge invariant. The term that broke the local phase invariance of the action (1.1.2) can now be “absorbed” by  $A_\mu$  through a gauge transformation (1.1.7), thus making the full action gauge invariant:

$$\begin{aligned} S_{QED} &= \int d^4x \left( i\bar{\psi} \not{D} \psi - m\bar{\psi} \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) = \\ &= \int d^4x \left( i\bar{\psi} \not{\partial} \psi - m\bar{\psi} \psi - g\bar{\psi} \not{A} \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) \\ S_{QED} \rightarrow S'_{QED} &= \int d^4x \left( i\bar{\psi} \not{\partial} \psi + \cancel{\bar{\psi} \not{\partial} \alpha \psi} - m\bar{\psi} \psi - g\bar{\psi} \not{A} \psi - \cancel{\bar{\psi} \not{\partial} \alpha \psi} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) = \\ &= \int d^4x \left( i\bar{\psi} \not{\partial} \psi - m\bar{\psi} \psi - g\bar{\psi} \not{A} \psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) = S_{QED} \end{aligned} \quad (1.1.9)$$

### 1.1.4 Non-Abelian Gauge Theories

Let us now consider a theory with  $N$  fermions, all with the same mass  $m$ , described by the spinorial fields  $\psi_i(x)$  with  $i = 1, \dots, N$ . These  $N$  fermions represent the  $N$  possible charges of the same particle<sup>2</sup> and are not to be confused with, for instance, the different possible flavors of the quarks, that describe different particles with different masses.

Its free action is:

$$S_\psi[\psi_i(x), \bar{\psi}_i(x)] = \sum_{i=1}^N \int d^4x (i\bar{\psi}_i \not{\partial} \psi_i - m\bar{\psi}_i \psi_i) \quad (1.1.10)$$

From now on, the sum over  $i$  (and all other repeated latin indexes) will be omitted, unless differently specified. This action is invariant under the global transformation:

$$\begin{aligned} \psi_i(x) &\rightarrow \psi'_i(x) = U_{ij} \psi_j(x) \\ \bar{\psi}_i(x) &\rightarrow \bar{\psi}'_i(x) = \bar{\psi}_j(x) U_{ji}^\dagger \end{aligned} \quad (1.1.11)$$

if  $U$  is any (constant)  $N \times N$  matrix such that  $UU^\dagger = U^\dagger U = \mathbf{1} \Leftrightarrow U^\dagger = U^{-1}$ , or in other words, if  $U \in U(N)$ . For this reason, this transformation is also called a global  $U(N)$  transformation. The phase transformation (1.1.4) is the particular case where  $U = e^{-i\alpha} \in U(1)$ , that is the only Abelian (commutative) unitary group.

As  $U(N) = SU(N) \otimes U(1) \forall N > 1$ ,  $U \in SU(N)$  instead of  $U \in U(N)$  can be imposed,

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<sup>2</sup>For example, if  $N = 3$  the 3 possible charges are the color charges of QCD, as will be shown later.

and will be from now on, without loss of generality.

In an analogous way to what has been done in Section 1.1.3, this invariance can be made local by implementing a proper covariant derivative, similar to (1.1.5). In order to do so, the infinitesimal  $SU(N)$  transformation has to be considered:

$$U_{ij}(x) = \delta_{ij} + i\theta^a(x) (T^a)_{ij} + O(\theta^2) \quad (1.1.12)$$

where the indices  $i$  and  $j$  run from 1 to  $N$  (as before) and the index  $a$  runs from 1 to  $N^2 - 1$  (the dimension of the group  $SU(N)$ ). The matrixes  $T^a$  are the  $N^2 - 1$  generators of  $\mathfrak{su}(N)$  (the Lie algebra of  $SU(N)$ ), thus they are  $N \times N$  hermitean and traceless, which obey the commutation relations:

$$[T^a, T^b] = if^{abc}T^c \quad (1.1.13)$$

where  $f^{abc}$  are called *structure constants* of  $\mathfrak{su}(N)$ . The normalization of these matrixes can be chosen such that they obey the condition:

$$\text{Tr}(T^a T^b) = \frac{1}{2} \delta^{ab} \quad (1.1.14)$$

some examples are:

- $N = 2$ ,  $T^a = \frac{\sigma^a}{2}$ , with  $\sigma^a$  the Pauli matrixes and  $f^{abc} = \varepsilon^{abc}$ ;
- $N = 3$ ,  $T^a = \frac{\lambda^a}{2}$ , with  $\lambda^a$  the Gell-Mann matrixes.

where  $\varepsilon^{abc}$  is the completely antisymmetric Levi-Civita symbol.

The covariant derivative, therefore, is written as:

$$D_\mu \equiv \partial_\mu + ig\mathbf{A}_\mu(x) \quad (1.1.15)$$

where an  $N \times N$  identity matrix  $\mathbf{1}$  multiplying  $\partial_\mu$  has to be understood, and  $\mathbf{A}_\mu(x)$  is a gauge field of  $SU(N)$ , i.e., a traceless, hermitean  $N \times N$  matrix, or, in other words,  $\mathbf{A}_\mu(x) \in \mathfrak{su}(N)$ .

The covariant derivative can be written more explicitly acting on the set of spinors  $\psi_i$ :

$$(D_\mu)_{ij} \psi_j = \partial_\mu \mathbf{1}_{ij} \psi_j + ig (\mathbf{A}_\mu(x))_{ij} \psi_j$$

In order for the action to be gauge invariant, the field  $\mathbf{A}_\mu$  must satisfy the gauge transformation property

$$\mathbf{A}_\mu(x) \rightarrow \mathbf{A}'_\mu(x) = U(x) \mathbf{A}_\mu(x) U^\dagger(x) - \frac{i}{g} U(x) \partial_\mu U^\dagger(x) \quad (1.1.16)$$

This expression is a little more complicated than (1.1.7), due to the fact that  $\mathbf{A}_\mu$  is now a non-commuting matrix. However if the Abelian case  $U(1)$  is taken into consideration, where  $U(x) = e^{-i\alpha(x)}$ , (1.1.7) follows directly from (1.1.16).

Now, it can be easily checked that the kinetic term of the Lagrangian

$$\mathcal{L}_K = i\bar{\psi}_i \not{D} \psi_i = i\bar{\psi}_i \not{\partial} \psi_i - g\bar{\psi}_i \not{A} \psi_i$$

is gauge invariant (i.e., invariant under (1.1.11) and (1.1.16)) through direct computation:

$$\begin{aligned}
 \mathcal{L}_K &\rightarrow \mathcal{L}'_K = i\bar{\psi}_i U^\dagger \not{\partial} (U\psi_i) - g\bar{\psi}_i \underbrace{U^\dagger U}_1 \underbrace{\mathbf{A}}_1 \underbrace{U^\dagger U}_1 \psi_i + i\bar{\psi}_i \underbrace{U^\dagger U}_1 (\not{\partial} U^\dagger) U\psi_i = \\
 &= i\bar{\psi}_i U^\dagger (\not{\partial} U) \psi_i + \underbrace{i\bar{\psi}_i \not{\partial} \psi_i - g\bar{\psi}_i \mathbf{A} \psi_i + i\bar{\psi}_i (\not{\partial} U^\dagger) U\psi_i}_{\mathcal{L}_K} = \\
 &= \mathcal{L}_K + i\bar{\psi}_i \gamma^\mu (U^\dagger \partial_\mu U + \partial_\mu U^\dagger U) \psi_i = \\
 &= \mathcal{L}_K + i\bar{\psi}_i \gamma^\mu \partial_\mu (U^\dagger U) \psi_i = \\
 &= \mathcal{L}_K + i\bar{\psi}_i \gamma^\mu \underbrace{\partial_\mu (1)}_{=0} \psi_i = \mathcal{L}_K
 \end{aligned}$$

Because of this fact, it is directly implied that the covariant derivative (1.1.15) must transform, under a gauge transformation, in the adjoint representation:

$$D_\mu \rightarrow D'_\mu = U D_\mu U^\dagger \quad (1.1.17)$$

The field-strength for the field  $\mathbf{A}_\mu$  is obtained, as for the Abelian case, through the commutator of two covariant derivatives. The computation is the same as (1.1.6), but this time the commutator term is non-vanishing:

$$F_{\mu\nu} \equiv -\frac{i}{g}[D_\mu, D_\nu] = \partial_\mu \mathbf{A}_\nu - \partial_\nu \mathbf{A}_\mu + ig[\mathbf{A}_\mu, \mathbf{A}_\nu] \quad (1.1.18)$$

This expression can be simplified a little by considering that  $\mathbf{A}_\mu$  and  $F_{\mu\nu}$  are elements of  $\mathfrak{su}(N)$ , thus writing them in terms of their components w.r.t. the basis  $T^a$ :

$$\mathbf{A}_\mu(x) = A_\mu^a(x) T^a \quad (1.1.19)$$

$$F_{\mu\nu}(x) = F_{\mu\nu}^a(x) T^a \quad (1.1.20)$$

and by considering the relation (1.1.13):

$$\begin{aligned}
 F_{\mu\nu}^a T^a &= (\partial_\mu A_\nu^a - \partial_\nu A_\mu^a) T^a + ig[A_\mu^b T^b, A_\nu^c T^c] = \\
 &= (\partial_\mu A_\nu^a - \partial_\nu A_\mu^a) T^a + ig A_\mu^b A_\nu^c \underbrace{[T^b, T^c]}_{if^{bca} T^a} = \\
 &= (\partial_\mu A_\nu^a - \partial_\nu A_\mu^a - gf^{abc} A_\mu^b A_\nu^c) T^a \\
 F_{\mu\nu}^a &= \partial_\mu A_\nu^a - \partial_\nu A_\mu^a - gf^{abc} A_\mu^b A_\nu^c \quad (1.1.21)
 \end{aligned}$$

In order to write a kinetic action for the field  $\mathbf{A}_\mu$ , a term proportional to  $F_{\mu\nu} F^{\mu\nu}$ , like in (1.1.8), is not enough: because of (1.1.17) and the definition (1.1.18), it must transform as  $F_{\mu\nu} F^{\mu\nu} \rightarrow U F_{\mu\nu} F^{\mu\nu} U^\dagger$ , therefore it would not be gauge invariant. In fact a gauge invariant action, called Yang-Mills action, is:

$$S_{YM} = -\frac{1}{2} \int d^4x \operatorname{Tr} (F_{\mu\nu} F^{\mu\nu}) \quad (1.1.22)$$

because of the cyclic property of the trace.<sup>3</sup> This action can be written in components, using (1.1.20) and the trace property (1.1.14):

$$S_{YM} = -\frac{1}{2} \int d^4x \operatorname{Tr} (F_{\mu\nu}^a F^{\mu\nu a}) = -\frac{1}{4} \int d^4x F_{\mu\nu}^a F^{\mu\nu a} \quad (1.1.23)$$

---

<sup>3</sup>Actually, a term proportional to  $\det(F_{\mu\nu}^a F^{\mu\nu a})$  would be gauge invariant as well, but it would not be a suitable kinetic term as it would involve terms of higher order than 2 in the components  $F_{\mu\nu}^a$

Here, there are two remarks that need to be done. The first one is that, if the gauge group is taken to be  $U(1)$ , the action (1.1.23) reduces to (1.1.8), as  $a = 1$  because the group  $U(1)$  has only 1 generator. The second one is that, if non-Abelian gauge groups are taken into consideration, this action naturally introduces self-interacting cubic and quartic terms, because the structure constants  $f^{abc}$  are non-vanishing. This is, for example, the case for the group  $SU(3)$ , that is used to describe gluon interaction, i.e., Quantum Chromodynamics (QCD). These self-interactions make the the Yang-Mills action interesting to be studied even alone, without any other fermionic or bosonic interacting field, as it will be shown later.

### 1.1.5 Wick Rotation

Up to now, actions were written in Minkowskian spacetime, where  $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$ . In order to have a positive-defined metric  $\eta_{\mu\nu} = \delta_{\mu\nu}$  a Wick rotation can be performed by re-defining the time coordinate to be  $\tau = it = ix^0$ , as is rigorously proven in [6]. The principle behind Wick rotation is that the time coordinate is promoted to be a complex number and the integration (in the action) from  $-\infty$  to  $\infty$  is made to be from  $-i\infty$  to  $+i\infty$  through a rotation of  $\frac{\pi}{2}$  in the complex plane, assuming that correlators and other physical quantities do not present singularities in the first and third quadrant of the complex plane of  $t$ . This rotation is not only a matter of convenience, in fact it is needed in perturbative QFT for the computation of (otherwise oscillating) functional integrals and in lattice field theory because a Minkowskian lattice cannot be rigorously defined.

The relation between the Minkowski actions written before and the Euclidean actions is  $S^M = iS^E$ , because  $d^4x^E = id^4x^M$ .

Thus, the Euclidean actions are<sup>4</sup>:

$$S^E[\phi] = \int d^4x \left( \frac{1}{2} \partial^\mu \phi \partial_\mu \phi - \frac{1}{2} m^2 \phi^2 + V(\phi) \right) \quad (1.1.24)$$

$$S_F^E = \int d^4x \bar{\psi} (\gamma^\mu (\partial_\mu + ig\mathbf{A}_\mu) + m) \psi \quad (1.1.25)$$

$$S_{YM}^E = \frac{1}{4} \int d^4x F_{\mu\nu}^a F^{a\mu\nu} \quad (1.1.26)$$

where the superscript  $E$ , meaning that the action is written in the Euclidean spacetime, will be omitted from now on.

## 1.2 Lattice Field Theory

### 1.2.1 Why Lattice Field Theory

After the Wick rotation, quantum field theory can be studied perturbatively through the saddle-point evaluation of the path integral, where the partition function

$$Z = \int \mathcal{D}\varphi e^{iS^M[\varphi]} = \int \mathcal{D}\varphi e^{-S^E[\varphi]}$$

---

<sup>4</sup>Note that they are not the same as their Minkowskian counterpart, as some redefinitions of the fields and the  $\gamma$  matrixes have been implicitly made, in order to have the respective theories non ill-defined.

has become non-oscillating. This approach, however, works well only if the coupling constants are *small enough* and does not allow to obtain non-perturbative results, i.e., physical quantities that depend on essential singularities in the coupling constant, like glueballs for Yang-Mills theories.

For this reason, a non-perturbative approach has to be taken into consideration and one possibility is Lattice Field Theory: the spacetime is discretized to a lattice, with lattice spacing  $a$ , and the fields can assume different values only on particular parts of the lattice. More in detail, a scalar field lives on lattice sites, a vector field lives on links between sites and objects with  $k$  indices live on  $k$ -simplexes.

If the limit  $a \rightarrow 0$  is taken, the theory must reproduce results obtained in the continuum, like the ones obtained in perturbation theory in the regime where it can be applied, however the presence of a discrete spacetime provides a natural cutoff for the momenta, allowing ultraviolet divergencies to be kept under control, although they need to be taken into consideration when approaching the continuum limit.

In this section, based on standard quantum field theory textbooks cited before and lattice field theory textbooks [7–9], the regularization of the previous quantum field theories on a Simple Hypercubic (SH) lattice is presented. A more detailed description of the geometric properties of the figures mentioned below can be found in [10].

## 1.2.2 Lattice

A lattice  $\Lambda$  in  $\mathbb{R}^D$  is defined as the set of all possible linear combinations with integer coefficients of the vectors  $\{v_i\}$ , which form a basis of  $\mathbb{R}^D$ . Formally:

$$\Lambda = \left\{ \sum_{i=1}^D c_i v_i \mid c_i \in \mathbb{Z} \right\} \quad (1.2.1)$$

### 1.2.2.1 Simple Hypercubic Lattice

If the basis  $\{v_i\}$  is taken to be orthonormal, the lattice is called *Simple Hypercubic*<sup>5</sup>.

In Figure (1.2.1) is represented a portion of a Simple Hypercubic lattice in  $D = 3$ .

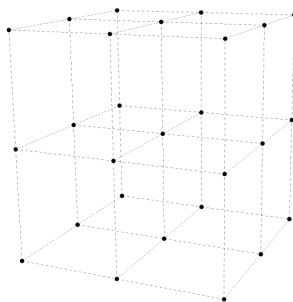


Figure 1.2.1: A cubic lattice.

For the SH lattice in  $D = 4$ , the fundamental region (the smallest  $D$ -dimensional polyhedron) is a 4-dimensional hypercube, also called a tesseract. This means that a SH lattice

<sup>5</sup>The word *simple* is used to make a distinction with the Body-Centered Hypercubic lattice presented in Section 3.2 .

can be seen as a tassellation of the spacetime with tesseracts as elementary cells. Each point of the lattice has 8 nearest neighbours that are identified by the vectors obtained through all possible permutations of position and sign of  $(\pm 1, 0, 0, 0)$ .

The plaquette (the simplest bidimensional figure) is a square. This will be important when implementing gauge theories on the lattice.

### 1.2.3 Scalar Fields on the Simple Hypercubic Lattice

Let us consider a simple hypercubic lattice  $\Lambda_{SH}$  extending over  $L = L_1 = L_2 = L_3$  lattice spacings in the spatial directions and over  $T = L_0$  lattice spacings in the temporal direction. Let us also consider a scalar field  $\phi(x)$ , in order to introduce some useful tools that will be needed later. As anticipated in the beginning of this section, a scalar field can only assume values on the sites of the lattice, therefore  $x \in \Lambda_{SH}$ , and let us assume, for simplicity, periodic boundary conditions on the fields, therefore  $\phi(x + a\hat{\mu}L_\mu) = \phi(x)$ . The Fourier transform of the field becomes

$$\tilde{\phi}(p) = \sum_x a^4 e^{-ip \cdot x} \phi(x) \quad (1.2.2)$$

and the allowed momenta are given by

$$p_\mu = \frac{2\pi}{aL_\mu} n_\mu, \quad n_\mu = 0, \dots, L_\mu \quad (1.2.3)$$

This ensures that the momenta can take only the assigned values of the Brillouin zone (1.2.3) and therefore cannot go to infinity, providing a natural cutoff given by the lattice spacing  $a$ .

The discretization also implies that derivatives of the fields cannot be computed, therefore the lattice forward derivative is used:<sup>6</sup>

$$\partial_\mu \phi(x) \rightarrow \nabla_\mu \phi(x) = \frac{\phi(x + a\hat{\mu}) - \phi(x)}{a} \quad (1.2.4)$$

Assuming a self-interaction potential  $V(\phi) = \frac{\lambda}{4!} \phi^4$ , the action (1.1.1) can be written in terms of the lattice in the following way:

$$S = a^4 \sum_x \left( \frac{1}{2a^2} [\phi(x + a\hat{\mu}) - \phi(x)]^2 + \frac{1}{2} m^2 \phi^2(x) + \frac{\lambda}{4!} \phi^4(x) \right) \quad (1.2.5)$$

A similar approach will be used in the following section to obtain the action for Yang-Mills theories.

### 1.2.4 Gauge Fields on the Simple Hypercubic Lattice

Let us now take into consideration a Yang-Mills theory, as in Section 1.1.4, with gauge group  $SU(N)$ . As anticipated before, gauge fields, being vector fields, can assume value only on links between sites of the lattice.

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<sup>6</sup>Note that, in the continuum limit, the two derivatives coincide.

One could naively think that putting gauge vectors  $A_\mu(x)$  on the links is enough, but this would explicitly break gauge invariance. For this reason, Wilson's idea was to put the gauge group (and not algebra) elements, namely  $U_\mu(x) = e^{iagA_\mu(x)}$ , on the links.

The field  $U_\mu(x)$  lives on the link connecting the site  $x$  with the site  $x + a\hat{\mu}$ , therefore the link is oriented and link variables in negative directions can also be defined:

$$U_{-\mu}(x) \equiv U_\mu^\dagger(x - a\hat{\mu}).$$

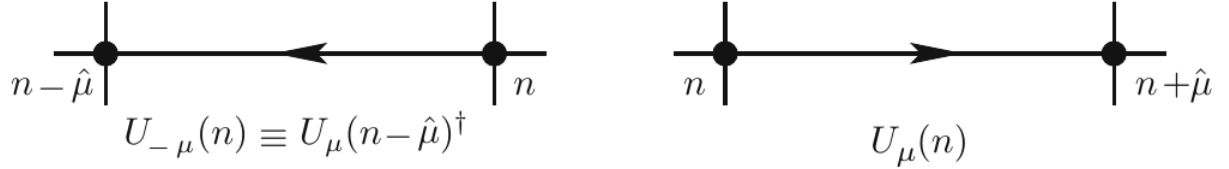


Figure 1.2.2: Schematic visualization of link variables.

Under a gauge transformation, the fields  $U_\mu$  transform according to the following relation:

$$U_\mu(x) \rightarrow \Omega(x)U_\mu(x)\Omega^\dagger(x + a\hat{\mu}) \quad (1.2.6)$$

where  $\Omega(x)$  is any  $SU(N)$  matrix at the point  $x$  on the lattice. As a consequence of this relation, the trace of any product of links forming a closed path is a gauge-invariant quantity, thanks to the cyclic property of the trace.

With this in mind, Wilson's idea was to choose the simplest closed path possible: the plaquette, that in the simple hypercubic lattice is a square. The product of link variables along a square plaquette is defined in the following way (the lattice spacing  $a$  is set = 1 for brevity of notation):

$$\begin{aligned} U_{\mu\nu}(x) &\equiv U_\mu(x)U_\nu(x + \hat{\mu})U_{-\mu}(x + \hat{\mu} + \hat{\nu})U_{-\nu}(x + \hat{\nu}) = \\ &= U_\mu(x)U_\nu(x + \hat{\mu})U_\mu^\dagger(x + \hat{\nu})U_\nu^\dagger(x) \end{aligned} \quad (1.2.7)$$

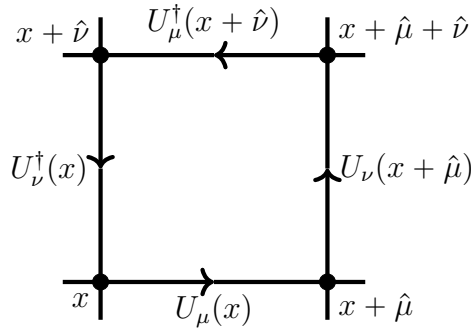
Thus a gauge-invariant action, called Wilson action, can be written as follows:

$$S_W[U] = \frac{\beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \text{Re Tr}[\mathbb{1} - U_{\mu\nu}(x)] \quad (1.2.8)$$

where  $\beta$  is a parameter that is going to be set in the following passages,  $N$  the number of colors (the same  $N$  in  $SU(N)$ ), the real part is needed to preserve unitarity and  $U_{\mu\nu}$  is the plaquette of (1.2.7).

In the following part, the naive continuum limit is taken, showing that Wilson action reproduces Yang-Mills action of (1.1.26) when  $a \rightarrow 0$ . Let us start by defining an auxiliary field  $B_\mu(x)$ :

$$B_\mu(x) = agA_\mu^a(x)T^a \quad (1.2.9)$$

Figure 1.2.3: Link variables building a plaquette  $U_{\mu\nu}(x)$ .

thus the gauge group element becomes  $U_\mu(x) = e^{iB_\mu(x)}$ .

The expansion for small  $a$  of the plaquette  $U_{\mu\nu}$  is:

$$\begin{aligned}
 U_{\mu\nu} &= U_\mu(x) U_\nu(x + \hat{\mu}) U_\mu^\dagger(x + \hat{\nu}) U_\nu^\dagger(x) = \\
 &= e^{iB_\mu(x)} e^{iB_\nu(x + \hat{\mu})} e^{-iB_\mu(x + \hat{\nu})} e^{-iB_\nu(x)} = \\
 &\quad \text{applying (1.2.4), } B_\mu(x + \hat{\nu}) = a \nabla_\nu B_\mu(x) \\
 &= e^{iB_\mu(x)} e^{iB_\nu(x + \hat{\mu})} e^{-iB_\mu(x) - ia \nabla_\nu B_\mu(x)} e^{-iB_\nu(x)} \simeq \\
 &\quad \text{From now on, the dependance from } x \text{ will be omitted.} \\
 &\quad \text{Using BCH formula, } e^X e^Y = e^{X+Y+\frac{1}{2}[X,Y]}, \text{ up to } O(a^2): \\
 &\simeq \exp \left\{ i\cancel{B_\mu} + i\cancel{B_\nu} + ia \nabla_\mu B_\nu - i\cancel{B_\mu} - ia \nabla_\nu B_\mu - i\cancel{B_\nu} + \right. \\
 &\quad \left. + \frac{1}{2} (-[B_\mu, B_\nu] + [\cancel{B_\mu}, \cancel{B_\nu}] + [\cancel{B_\nu}, \cancel{B_\mu}] - [B_\mu, B_\nu]) \right\} = \\
 &= \exp \{ ia (\nabla_\mu B_\nu - \nabla_\nu B_\mu) - [B_\mu, B_\nu] \}
 \end{aligned}$$

Now, substituting back (1.2.9), the discretized electromagnetic tensor can be found:

$$\begin{aligned}
 U_{\mu\nu} &\simeq \exp \{ ia^2 g (\nabla_\mu \mathbf{A}_\nu - \nabla_\nu \mathbf{A}_\mu + ig [\mathbf{A}_\mu, \mathbf{A}_\nu]) \} = \exp \{ ia^2 g F_{\mu\nu} \} \simeq \\
 &\simeq \mathbb{1} + ia^2 g F_{\mu\nu} - \frac{1}{2} a^4 g^2 F_{\mu\nu} F^{\mu\nu} + O(a^6)
 \end{aligned} \tag{1.2.10}$$

Finally, the small- $a$  approximation for the Wilson action (1.2.8) can be obtained:

$$\begin{aligned}
 S_W &\simeq \frac{\beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \text{Re Tr} \left( \cancel{\mathbb{1}} - \cancel{\mathbb{1}} - ia^2 g F_{\mu\nu}^i T^i + \frac{1}{2} a^4 g^2 F_{\mu\nu}^i F^{j\mu\nu} T^i T^j \right) = \\
 &= \underbrace{-i \frac{a^2 \beta}{2N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \text{Re}(F_{\mu\nu}^i \text{Tr}(T^i))}_{=0 \text{ for (1.1.12)}} + \frac{g^2 a^4 \beta}{4N} \sum_{x \in \Lambda} \sum_{\mu < \nu} \underbrace{\text{Re}(F_{\mu\nu}^i F^{j\mu\nu} \text{Tr}(T^i T^j))}_{=\frac{1}{2} F_{\mu\nu}^i F^{i\mu\nu} \text{ for (1.1.14)}} = \\
 &= \frac{g^2 \beta}{8N} a^4 \sum_{x, \mu, \nu} F_{\mu\nu}^i F^{i\mu\nu} \xrightarrow{a \rightarrow 0} \frac{g^2 \beta}{8N} \int d^4 x F_{\mu\nu}^i F^{i\mu\nu}
 \end{aligned} \tag{1.2.11}$$

where the Yang-Mills action (1.1.26) can be recognized if  $\beta = \frac{2N}{g^2}$ .



### 1.2.4.1 Wilson Loops

Now that the lattice action has been defined, observables need to be taken into account. As mentioned below equation (1.2.6), the trace of the product of link variables along any closed path is a gauge-invariant quantity. This means that any quantity of the sort is a good candidate for an observable and, in fact, given any closed path  $\gamma$  on the lattice, the Wilson loop is defined as follows:

$$W[\gamma] \equiv \text{Tr} \left[ \prod_{(x,\mu) \in \gamma} U_\mu(x) \right] \quad (1.2.12)$$

It is worthwhile to mention that this definition does not depend on the type of lattice and of boundary conditions chosen for such lattice. If the path  $\gamma$  does not include any temporal link, the loop is called *spacelike Wilson loop*, otherwise it is called *timelike Wilson loop*. If all the links in  $\gamma$  lie on the same plane, the Wilson loop is *planar*, otherwise it is *nonplanar*.

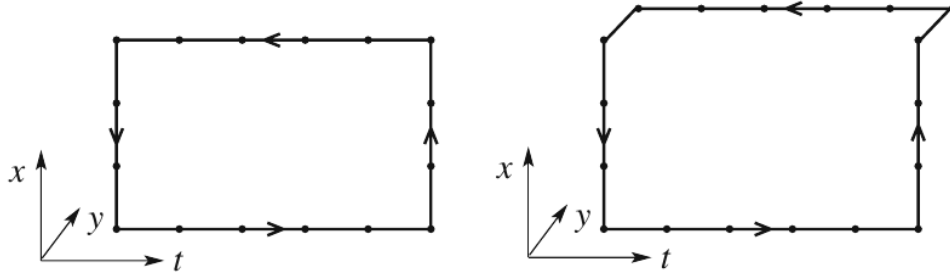


Figure 1.2.4: Example of a planar (left) and nonplanar (right) Wilson loop.

Wilson loops are important for a variety of reasons, that will not be elaborated further as this is not the focus of this project, however it is worth knowing that they can be used as operators for purely gluonic bound states, the so-called *glueballs*: the mass spectrum of these states can be obtained through the exponential decay of Wilson loops' correlation functions.

### 1.2.4.2 Polyakov Loops

Another important observable can be defined in lattices where periodic boundary conditions in the time direction are assumed. Thanks to periodicity, a whole new class of closed paths arises: all paths that wind around the time direction. A Polyakov loop is thus defined as a line at a fixed spatial position  $\vec{x} = (x^1, x^2, x^3)$  that runs along the whole time direction:

$$P(\vec{x}) \equiv \text{Tr} \left[ \prod_{t=0}^{T-1} U_0(t, \vec{x}) \right] \quad (1.2.13)$$

The Polyakov loop has a lot of interesting properties: its physical interpretation is the world-line of a single static quark, therefore its expectation value is an order parameter of the deconfinement transition, as well as of a new global symmetry, called *center symmetry*, arising from the periodicity of the time direction.

This means that, if the system is confined (quarks and gluons cannot move around freely), its expectation value is 0, while if the system is deconfined,  $\langle P(\vec{x}) \rangle \neq 0$ .

Another property, that will be used in the simulations, is that the expectation value of the product of two Polyakov loops at distance  $r = a|\vec{x} - \vec{y}|$  is a function of the static quark-antiquark potential  $V(r)$ , as indicated below:

$$\langle P(\vec{x})P^\dagger(\vec{y}) \rangle \propto e^{-TaV(r)} (1 + O(e^{-Ta\Delta E})) \quad (1.2.14)$$

where  $\Delta E$  is the difference between  $V(r)$  and the first excited energy level of the quark-antiquark pair.

The value of the potential  $V(r)$  will be used, in this project, to study the recovery of rotational invariance when approaching the continuum limit.

Various methods of performing computer simulations of lattice field theory and how to obtain these quantities, such as plauette values, Wilson and Polyakov loops, is shown in the following chapter.

# Computer Simulation of Pure Gauge Theories

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## 2.1 Introduction

Monte Carlo simulations are a powerful tool that can be used to evaluate observables in lattice field theory. Given an observable  $O$  (such as a plaquette, a Wilson or Polyakov loop, etc.), its vacuum expectation value is formally given by the functional integral

$$\langle O \rangle = \frac{1}{Z} \int \mathcal{D}[U] e^{-S[U]} O[U] \quad \text{with} \quad Z = \int \mathcal{D}[U] e^{-S[U]} \quad (2.1.1)$$

where  $\mathcal{D}[U]$  is to be intended as a Haar measure.

This expression, however, cannot be evaluated analytically except for very small lattices, therefore (2.1.1) is approximated by an average of the observable evaluated on  $N$  sample gauge field configurations  $\{U_n\}$ <sup>1</sup>, distributed according a probability density  $\propto e^{-S[U_n]}$ .

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<sup>1</sup>Here the subscript  $n$  distinguishes the different configurations.

The expectation value is then obtained by computing the following sum for a sufficient number of configurations generated by the proper Monte Carlo algorithm(s):

$$\langle O \rangle \simeq \frac{1}{N} \sum_{\{U_n\}} O[U_n] \quad (2.1.2)$$

Because of the probability density  $\propto e^{-S[U_n]}$ , only configurations  $\{U_n\}$  that minimize the action are *good configurations*, as all other configurations are exponentially suppressed. For this reason, generating totally random gauge fields on the lattice links is not an efficient way to evaluate (2.1.2), as most of the  $\{U_n\}$  will have a very small Boltzmann factor (the  $e^{-S[U_n]}$ ) and expression (2.1.2) will give incorrect results unless a huge number (orders of magnitude higher than what is reasonably possible) of different configurations is tried.

In order to avoid the evaluation of a great number of configurations that would contribute little-to-nothing to the observables' values, a sequence of configurations  $\{U_n\}$  is generated through a Markov chain process, built such that its stationary distribution minimizes the action  $S[U]$ . This process is usually called *thermalization*.

Every process of this type must begin from a starting configuration, usually chosen by the user. The two mostly used starting configurations are: a *cold start*, if the simulation begins with the fields in an ordered way (for example, all the gauge links set to the identity), or a *hot start*, if the simulation begins with random gauge fields in every link. Of course, the Markov chain must have the same stationary distribution if the hot or cold start is chosen.

In this chapter the main Monte Carlo algorithms used to thermalize the starting configuration are presented, then the methods used for evaluating some observables of interest, such as plaquettes and Polyakov loops, are explained.

## 2.2 Markov Processes

As anticipated before, a Markov chain is needed to evolve an arbitrary starting configuration  $U_0$  up to a region of the configuration space with a relatively large Boltzmann factor  $e^{-S}$ , therefore with high probability. In Figure (2.2.1) a scheme of a Markov chain is shown.

A Markov process is characterized by a probability of transitioning from a configuration  $\alpha$  to another configuration  $\beta$  depending only on  $\alpha$  and  $\beta$  and not on the history of the process, namely the previous transitions that already occurred. In formulae, this is expressed as

$$P(U_n = \beta | U_{n-1} = \alpha) = T(\beta | \alpha) \quad (2.2.1)$$

That is to say that the transition matrix  $T$  does not depend on the index  $n$ , representing the computer time.

Being a transition probability, the matrix  $T$  must obey the following equations

$$0 \leq T(\beta | \alpha) \leq 1 \quad \forall \alpha, \beta \quad (2.2.2)$$

$$\sum_{\beta} T(\beta | \alpha) = 1 \quad \forall \alpha \quad (2.2.3)$$

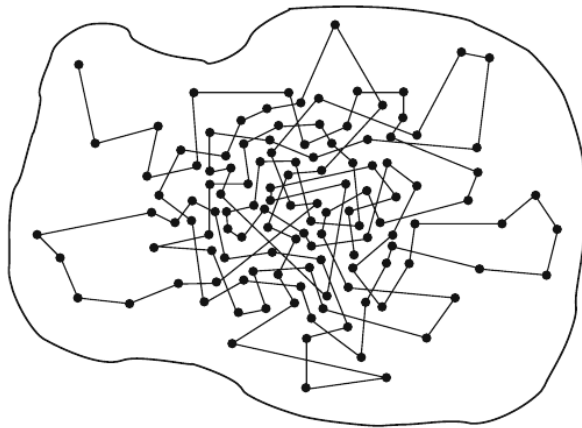


Figure 2.2.1: Schematic representation of a motion through the space of configurations through a Markov process.

where (2.2.2) is a consequence of  $T(\beta|\alpha)$  representing a probability, and (2.2.3) means that the probability of transitioning to any configuration must be 1 (of course, the case when  $\alpha = \beta$  is included as well).

In order for the stochastic process not to have any sink or source of probability, the following balance equation must be satisfied:

$$\sum_{\alpha} T(\beta|\alpha)P(\alpha) = \sum_{\alpha} T(\alpha|\beta)P(\beta) \quad (2.2.4)$$

This equation states that the probability of transitioning to the configuration  $\beta$ , written in the l.h.s. as the sum of the transition probability from the configuration  $\alpha$  weighted by the probability  $P(\alpha)$  that the system is actually in that configuration, must be equal to the probability of transitioning out of the configuration  $\beta$ , given by the probability of finding the system in the configuration  $\beta$  times the transition probability  $T(\alpha|\beta)$  over all the final configurations, in the right-hand side. Thanks to (2.2.3), the r.h.s. is easily proven to be equal to  $P(\beta)$ .

A sufficient (but not necessary) condition to obey the balance equation (2.2.4) is obtained by requiring that it holds true term-by-term, thus obtaining the detailed balance condition:

$$T(\beta|\alpha)P(\alpha) = T(\alpha|\beta)P(\beta) \quad (2.2.5)$$

Although it is not a necessary condition, most algorithms, including the ones discussed in the following section, satisfy it.

## 2.3 Monte Carlo Algorithms

Monte Carlo algorithms are a class of algorithms that, singularly or combined together, allow to advance the Markov chain, while satisfying the balance condition (2.2.4).

Each algorithm, if applied once, allows the transition from a configuration  $U_{n-1}$  to a configuration  $U_n$  (eventually the same as  $U_{n-1}$ ). The repeated application of the algorithm allows to advance through the Markov chain.

### 2.3.1 Metropolis Algorithm

The first algorithm presented is the Metropolis algorithm. It is not very much efficient and it is not often used in simulations unless for didactical purposes, however it contains the fundamental steps that are present in some of the more advanced algorithms and it is quite easy to understand. For this reasons it is usually viewed as the “*ancestor*” of all Monte Carlo algorithms and it is present in every textbook on the subject.

This algorithm consists in two steps that implement in one of the most simple ways the detailed balance condition (2.2.5):

- 1) A candidate configuration  $\beta$  is chosen, according to some *a priori* selection probability  $T_0(\beta|\alpha)$ , where  $\alpha = U_{n-1}$ .
- 2) The candidate configuration  $\beta$  is accepted as the new configuration  $U_n$  with the acceptance probability

$$T_A(\beta|\alpha) = \min \left( 1, \frac{T_0(\alpha|\beta) \exp(-S[\beta])}{T_0(\beta|\alpha) \exp(-S[\alpha])} \right) \quad (2.3.1)$$

If it is not accepted, the unchanged configuration is considered again ( $U_n = \alpha$ ) and the measurements are eventually made again.

These two steps are repeated a sufficient amount of times up until the needed measurements are taken.

Note that  $P(\alpha) = \frac{e^{-S[\alpha]}}{Z} \propto e^{-S[\alpha]}$  has been used.

The total transition probability  $T$  is obtained through the product  $T = T_0 T_A$ , as the two steps are independant from each other, and it is straightforward to see that it satisfies the detailed balance condition (2.2.5):

$$\begin{aligned} T(\beta|\alpha)P(\alpha) &= \frac{1}{Z} T_0(\beta|\alpha) T_A(\beta|\alpha) \exp(-S[\alpha]) = \\ &= \frac{1}{Z} T_0(\beta|\alpha) \min \left( 1, \frac{T_0(\alpha|\beta) \exp(-S[\beta])}{T_0(\beta|\alpha) \exp(-S[\alpha])} \right) \exp(-S[\alpha]) = \\ &= \frac{1}{Z} \min (T_0(\beta|\alpha) \exp(-S[\alpha]), T_0(\alpha|\beta) \exp(-S[\beta])) = \\ &= \frac{1}{Z} T(\alpha|\beta) \exp(-S[\beta]) = \\ &= T(\alpha|\beta)P(\beta) \end{aligned}$$

□

In many cases a symmetric selection probability is used  $T_0(\alpha|\beta) = T_0(\beta|\alpha)$ , thus (2.3.1) simplifies to:

$$T_A(\beta|\alpha) = \min (1, e^{-\Delta S}) \quad \text{with} \quad \Delta S = S[\beta] - S[\alpha] \quad (2.3.2)$$

That means that if the new configuration lowers the action, the change is accepted with probability 1 (as  $e^{-\Delta S} > 1$ ), otherwise it is accepted with a certain probability that decays exponentially as the difference in the action of the two configurations becomes

greater. This ensures that the algorithm *moves across* the configuration space towards the minimums of the action, while allowing quantum fluctuations in order to not “*get stuck*” on a local minimum.

If the change in the action is local (it involves a single link variable),  $\Delta S$  can be computed using only the field values in the local neighbour. This will be the case for  $SU(N)$  gauge theories.

### 2.3.1.1 Application to $SU(N)$ Gauge Theories

For a  $SU(N)$  gauge theory, the algorithm is implemented in the following way. During each iteration, a single link is changed, then the acceptance probability is computed, a random number is extracted and, if it is less than the acceptance probability the change is accepted, otherwise it is rejected. The algorithm is then iterated a certain number of times and measures are taken.

The candidate link  $U'_\mu(x)$  for step 1 is generated in the vicinity of the old value  $U_\mu(x)$ , in order not to have a too great  $\Delta S$  that would lead to too low acceptance rates. This can be done by exploiting the property that the product of any two elements of  $SU(N)$  is still an element of  $SU(N)$ , therefore extracting a matrix  $X \in SU(N)$  near the identity allows to write the candidate link as:

$$U'_\mu(x) = XU_\mu(x) \quad (2.3.3)$$

The matrix  $X$  is chosen such that it has the same probability as  $X^{-1}$ , this way the selection probability  $T_0$  is symmetric and the computation of the acceptance probability  $T_A$  becomes easier.

For this reason, only the variation of the action  $\Delta S$  must be computed, where of course the action is the Wilson action (1.2.8). In particular, only the plaquettes containing the candidate link must be evaluated: the change in the action is local, so all the other plaquettes will have the same value both before and after the change of the link. Hence  $\Delta S = S[U'_\mu(x)]_{loc} - S[U_\mu(x)]_{loc}$ .

In a SH lattice, each link is shared between 6 plaquettes, in 4 spacetime dimensions. For each plaquette the change of its value is given by the change of the link, while the product of the other three gauge links, that is called *staple* and will be indicated as  $P_i$ , remains unchanged.

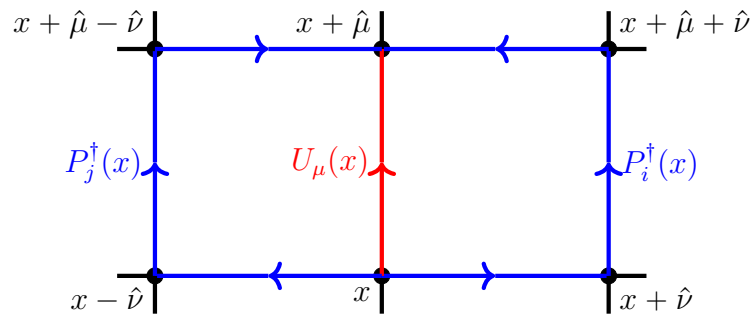


Figure 2.3.2: Example of a link (red) with two of its staples (blue).

Therefore, the local contribution to the action can be computed as:

$$S[U_\mu(x)]_{loc} = \frac{\beta}{2N} \sum_{i=1}^6 \text{Re Tr} [\mathbb{1} - U_\mu(x) P_i] = \frac{\beta}{2N} \text{Re Tr} \left[ 6\mathbb{1} - U_\mu(x) \sum_{i=1}^6 P_i \right]$$

where the sum over all the staples is:

$$A = \sum_{i=1}^6 P_i = \sum_{\nu \neq \mu} (U_\nu(x + \hat{\mu}) U_\mu^\dagger(x + \hat{\nu}) U_\nu^\dagger(x) + U_\nu^\dagger(x + \hat{\mu}) U_\mu^\dagger(x - \hat{\nu}) U_\nu(x - \hat{\nu})) \quad (2.3.4)$$

The change of the action can now be computed as

$$\Delta S = S[U'_\mu(x)]_{loc} - S[U_\mu(x)]_{loc} = \frac{\beta}{2N} \text{Re Tr} [(U_\mu(x) - U'_\mu(x)) A] \quad (2.3.5)$$

where  $A$  is not affected by the change of  $U_\mu(x)$ .

### 2.3.2 Heat Bath Algorithm

The heat bath algorithm is an *enhanced* version of the Metropolis algorithm, that combines the two steps into a single one and chooses the new candidate according to the probability distribution obtained by the computing the surrounding staples:

$$dP(U) = dU \exp\left(\frac{\beta}{2N} \text{Re Tr}[UA]\right) \quad (2.3.6)$$

where  $dU$  denotes the Haar integration measure of the gauge group and  $A$  is computed according to (2.3.4). This probability distribution can be computationally quite demanding, but has the advantage that, unlike the Metropolis algorithm, the link variable always changes.

The implementation details depend on the gauge group, for this reason, the heat bath method for the gauge group  $SU(2)$  will be now explained<sup>2</sup>.

Since the sum of any two  $SU(2)$  elements is proportional to another  $SU(2)$  element, the sum of all staples  $A$  (2.3.4) can be written in the form

$$A = V \sqrt{\det(A)} \quad \text{with} \quad V \in SU(2) \quad (2.3.7)$$

where it can be proven that  $\det(A) \geq 0$ . If  $\det(A) = 0$ , a random link variable is generated, otherwise the matrix  $V$  is well-defined. Plugging into (2.3.6) and using the invariance of the Haar measure under transformation of the origin of the group space ( $dU = d(UV) = dX$ ), the distribution probability of the matrix  $X = UV$  is

$$dP(X) = dX \exp\left(\frac{\beta}{2N} \sqrt{\det(A)} \text{Re Tr}[X]\right) \quad (2.3.8)$$

If a matrix  $X$ , distributed according to (2.3.8), is generated, then the new candidate link, distributed according to (2.3.6), is obtained as  $U'_\mu(x) = XV^\dagger = \frac{1}{\sqrt{\det(A)}} X A^\dagger$ .

---

<sup>2</sup>As matrixes of  $SU(N)$  can be “built” combining  $SU(2)$  matrixes, the general case is just a little more complicated, but it follows the same principles.



Any  $U \in SU(2)$  matrix can be written in the following representation, using 4 real numbers:

$$U = x_0 \mathbf{1} + i \mathbf{x} \cdot \boldsymbol{\sigma} \quad \text{with} \quad \det(U) = |x|^2 = \sum_{i=0}^3 x_i^2 = 1 \quad (2.3.9)$$

where  $\boldsymbol{\sigma} = (\sigma_1, \sigma_2, \sigma_3)$  is a vector built using the Pauli matrixes and  $x = (x_0, \mathbf{x})$  can be seen as a normalized 4-components vector. Using this representation, the Haar measure in (2.3.8) can be written as:

$$\begin{aligned} dX &= \frac{1}{\pi^2} d^4 x \delta(x_0^2 + \mathbf{x}^2 - 1) = \\ &= \frac{1}{\pi^2} d^4 x \frac{1}{2\sqrt{1-x_0^2}} \left( \delta(|\mathbf{x}| - \sqrt{1-x_0^2}) + \delta(|\mathbf{x}| + \sqrt{1-x_0^2}) \right) \end{aligned} \quad (2.3.10)$$

where a well known property of the Dirac delta function has been used.

The volume element can be rewritten in terms of the components of the vector  $x$ :

$$d^4 x = dx_0 d|\mathbf{x}| |\mathbf{x}|^2 \underbrace{d(\cos \theta) d\varphi}_{d^2 \Omega} \quad (2.3.11)$$

Plugging back into (2.3.10) and integrating out the  $|\mathbf{x}|$  thanks to the delta functions (actually, only the first one contributes, as  $x_i^2 \leq 1 \forall i$ ), the Haar measure takes the form:

$$dX = \frac{1}{\pi^2} dx_0 d^2 \Omega \frac{1-x_0^2}{2\sqrt{1-x_0^2}} = \frac{1}{2\pi^2} dx_0 d^2 \Omega \sqrt{1-x_0^2} \quad (2.3.12)$$

Then, in terms of the variables, the probability distribution becomes:

$$dP(X) = \frac{1}{2\pi^2} dx_0 d(\cos \theta) d\varphi \sqrt{1-x_0^2} \exp\left(\frac{\beta}{2} \sqrt{\det(A)} x_0\right) \quad (2.3.13)$$

with  $x_0 \in [-1, 1]$ ,  $\cos \theta \in [-1, 1]$ ,  $\varphi \in [0, 2\pi)$ , where  $\text{Tr}(X) = 2x_0$  has been used.

Thus, the problem of generating a matrix  $X$  distributed according to (2.3.8) has been reduced to the determination of three random variables  $x_0$ ,  $\theta$ ,  $\varphi$ , whose distributions factorize, so they can be determined independetly from each other.

The random variable  $x_0$ , being distributed according to  $\sqrt{1-x_0^2} \exp\left(\frac{\beta}{2} \sqrt{\det(A)} x_0\right)$ , is determined through the auxiliary variable  $\lambda \in [0, 1]$  such that  $x_0 = 1 - 2\lambda^2$ , therefore

$$dx_0 \sqrt{1-x_0^2} \exp\left(\frac{\beta}{2} \sqrt{\det(A)} x_0\right) \propto d\lambda \lambda^2 \sqrt{1-\lambda^2} \exp\left(-\beta \sqrt{\det(A)} \lambda^2\right) \quad (2.3.14)$$

The variable  $\lambda$  is generated with the polynomially modified Gaussian distribution density

$$p_1(\lambda) = \lambda^2 e^{-\beta \sqrt{\det(A)} \lambda^2} \quad (2.3.15)$$

and accepted with an accept/reject step using the square root function

$$p_2(\lambda) = \sqrt{1-\lambda^2} \quad (2.3.16)$$

There are several algorithms that can perform this computation in an efficient way [11, 12].

After this, the length  $|x| = \sqrt{1 - x_0^2}$  is computed, in order to determine the remaining variables.

The variables  $\cos \theta$  and  $\varphi$  are uniformly distributed, therefore a possible way of proceeding is by generating three random uniformly distributed numbers  $r_1$ ,  $r_2$  and  $r_3$  in the interval  $[-1, 1)$  and accepting them only if  $r_1^2 + r_2^2 + r_3^2 \leq 1$ . Then, the 3-vector  $(r_1, r_2, r_3)$  is normalized to length  $|x|$  computed before, obtaining  $\mathbf{x} = (x_1, x_2, x_3)$ .

After these steps, the vector  $x = (x_0, \mathbf{x})$  can be used to generate the matrix  $X$  according to (2.3.8), using representation (2.3.9), thus the new link variable can be obtained.

This algorithm rapidly leads the Markov process to a minimum of the action, however the risk of “getting stuck” on a local minimum is present. For this reason, it is usually combined with other algorithms, like the overrelaxation algorithm.

### 2.3.3 Overrelaxation Algorithm

The overrelaxation algorithm tries to change the link variables as much as possible, exploiting the property that new configurations are always accepted if they do not change the action, as can be seen in (2.3.2).

The case with gauge group  $U(1)$  is the simplest one: the group element can be written as  $U = e^{i\varphi}$ , the sum of staples becomes  $A = \rho e^{i\alpha}$  and the local action can be written as:

$$S[U]_{loc} = \frac{\beta}{2} \text{Re}(UA) = \frac{\beta}{2} \rho \text{Re}(e^{i\varphi} e^{i\alpha}) = \frac{\beta}{2} \rho \cos(\varphi + \alpha) \quad (2.3.17)$$

The reflection  $\varphi + \alpha \rightarrow -\varphi - \alpha$  or the change  $\varphi + \alpha \rightarrow 2\pi - \varphi - 2\alpha$  leaves the local action invariant, thus a change of this type is always accepted. For a non-Abelian group, this change is performed through the ansatz:

$$U \rightarrow U' = V^\dagger U^\dagger V^\dagger \quad (2.3.18)$$

where  $V$  is a gauge group element chosen such that the local action is invariant. The selection probability for a transformation of this kind is symmetric, as can be easily proven by inverting (2.3.18), obtaining  $U = V^\dagger U'^\dagger V^\dagger$ .

For the gauge group  $SU(2)$ , the matrix  $V$  is chosen proportional to the sum of staples:  $V = \frac{A}{\sqrt{\det(A)}}$ , like in (2.3.7). Because of this,

$$\text{Tr}(U'A) = \text{Tr}\left(V^\dagger U^\dagger V^\dagger \sqrt{\det(A)} V\right) = \sqrt{\det(A)} \text{Tr}(V^\dagger U^\dagger) = \text{Tr}(A^\dagger U^\dagger) = \text{Tr}(UA) \quad (2.3.19)$$

therefore the choice for  $U'$  leaves the action invariant. In case  $\det(A) = 0$ , any random link variable is accepted.

As the overrelaxation algorithm leaves the action invariant, it moves the Markov chain in the subspace of constant action, called *micro-canonical* ensemble, thus it is not ergodic: it does not visit every possible point in the configuration space if given enough time. For this reason, it needs to be used in combination with other updating algorithms, such as some steps of the Metropolis or heat bath algorithms.

## 2.4 Measurements

In this section, the techniques used to compute the value of some observables are explained. After the proper Monte Carlo algorithms have been iterated a sufficient number of times, the lattice is *thermalized* and measures can be taken. The observables presented below are the plaquettes and Polyakov loops, because they are the ones that will be evaluated in the simulation in the fourth chapter.

### 2.4.1 Plaquettes

The mean value of the plaquette, indicated as  $\langle \square \rangle$ , is obtained by evaluating expression (1.2.7) for all possible plaquette in the lattice, summing all the results and taking the trace of the result, divided by the total number of plaquettes present in the lattice. In formulae:

$$\langle \square \rangle = \frac{1}{12Nn_s} \sum_{x \in \Lambda} \sum_{\mu=0}^3 \sum_{\substack{\nu=0 \\ \nu \neq \mu}}^3 \text{Re Tr } U_{\mu\nu}(x) \quad (2.4.1)$$

where  $N$  is the number of color charges (the same  $N$  in  $SU(N)$ ),  $12 = 4 \cdot 3$  is the total number of ways of choosing the first direction  $\mu$  times the number of possible ways of choosing the second direction  $\nu$ , and  $n_s = TL_1L_2L_3$  is the total number of sites of the lattice.

The expectation value of the plaquette is important because it is directly linked to the Wilson action, thus it gives an intuition of the thermalization of the lattice.

### 2.4.2 Polyakov Loops

Polyakov loops are another important observable that will be evaluated in the simulations. In particular, the product of two Polyakov loops will be considered, in order to obtain the static quark potential, as explained in (1.2.14).

In order to achieve better statistics, the discrete translational invariance of the lattice is exploited, by computing the so-called *nonzero momentum Polyakov correlators*: because of this invariance, the quantity  $\langle P(0)P^\dagger(x) \rangle$ , with  $x \in \Lambda$  is the same as  $\langle P(y)P^\dagger(x+y) \rangle \forall y \in \Lambda$ , therefore the expectation value is obtained as

$$\langle P(0)P^\dagger(x) \rangle = \frac{1}{n_s} \sum_{y \in \Lambda} \langle P(y)P^\dagger(x+y) \rangle \quad (2.4.2)$$

In this way, from the same configuration, more samples of the same quantity can be obtained, making the average more statistically accurate.

## 2.5 Summary of a Simulation

Now that the Monte Carlo algorithms and the way observables are “read” from the lattice have been described, a brief explanation on how a full simulation works is presented.

### 2.5.1 Initialization

The first step is the initialization: the program reads from a file the size of the lattice (the number of sites in each direction:  $t, x, y, z$ ), the parameter  $\beta$  (that is the same in the Wilson action (1.2.8)), the thermalization time  $t_{TH}$ , the reunitarization period  $t_U$ , the number of measurements to be made  $n_M$ , the number of updates to be done between each measurement  $n_{BM}$  and the start type, either fully ordered (cold start) or totally random (hot start). The number of color charges and other parameters, like the type of measurements that have to be made are set in the source code of the program. Then, according to these parameters, the variables are initialized to the proper value and the simulation can begin.

### 2.5.2 Thermalization

The thermalization step is needed, as anticipated before, to advance the Markov process up to equilibrium configurations. This is done by updating the configuration with the Monte Carlo algorithms until the number of updates equals the thermalization time  $t_{TH}$  indicated in the input file. Every update step consists of a certain number of heat bath steps followed by another number of overrelaxation iterations. Both these numbers are set in the source code and for these simulations 1 iteration of heat bath and 3 of overrelaxation have been used.

The thermalization time is determined by looking, in other simulations, at the behaviour of some observables in different types of start: for instance, when some observable from a cold start assumes a mean value compatible to the one obtained from a hot start, the system is assumed as thermalized.

The thermalization time  $t_{TH}$  is then set as 2 or 3 times the number of steps needed to achieve thermalization, in order to keep a certain margin of confidence.

Additionally, every  $t_U$  iterations, all the lattice link variables are *reunitarized*, that means that each one matrix is checked, and eventually corrected, that it still is unitary (i.e.,  $U^\dagger = U^{-1}$ ), because unitarity can be lost due to computer roundig errors. After each update step, the mean value of the plaquette is written to a file.

### 2.5.3 Measurements

After the lattice has achieved thermalization, observables can finally be measured. Each observable is computed once, according to the methods illustrated in the previous section, and written to a file, then the system is updated a number of times equal to  $n_{BM}$ , with a reunitarization process every  $t_U$  updates, in the same way as the thermalization step. This process is iterated until the observables are measured  $n_M$  times.

After that, the simulation is concluded.

## 2.6 Statistical Analysis

In this final section, an overview on the analysis of the data collected from the simulations is given. After a run of the program is finished, one usually has a few thousands expectation values of each observable, obtained in different configurations of the fields. The

average of these values is used as an indicator for each observable, with its own statistical error.

Let us assume that the values  $(o_1, \dots, o_N)$  of some observable  $O$  have been computed from a Markov sequence of Monte Carlo generated configurations in equilibrium. The expectation value and variance of these variables are the same:

$$\langle o_i \rangle = \langle O \rangle \quad \sigma_{o_i}^2 = \langle (o_i - \langle o_i \rangle)^2 \rangle = \sigma_O^2 \quad (2.6.1)$$

whose estimators can be:

$$\hat{O} = \frac{1}{N} \sum_{i=1}^N o_i \quad (2.6.2)$$

$$\hat{\sigma}_O^2 = \frac{1}{N-1} \sum_{i=1}^N (o_i - \hat{O})^2 \quad (2.6.3)$$

These results hold true for both correlated and uncorrelated datasets.

### 2.6.1 Uncorrelated Data

If the values are uncorrelated, the product  $\langle o_i o_j \rangle$  factorizes

$$\langle o_i o_j \rangle = \langle o_i \rangle \langle o_j \rangle = \langle O \rangle^2 \quad \forall i \neq j \quad (2.6.4)$$

and the variance  $\hat{\sigma}_O^2$  allows to determine the statistical error of  $\hat{O}$ . In fact, the sample mean value (2.6.2) is itself a random variable, since its value may change from one set of  $N$  configurations to another. Its variance is:

$$\begin{aligned} \sigma_O^2 &= \left\langle \left( \hat{O} - \langle O \rangle \right)^2 \right\rangle = \left\langle \left( \frac{1}{N} \sum_{i=1}^N (o_i - \langle O \rangle) \right)^2 \right\rangle = \\ &= \frac{1}{N^2} \left\langle \sum_{i,j=1}^N (o_i - \langle O \rangle) (o_j - \langle O \rangle) \right\rangle = \\ &= \frac{1}{N^2} \left\langle \sum_{i=1}^N o_i^2 \right\rangle + \frac{1}{N^2} \left\langle \sum_{i \neq j} o_i o_j \right\rangle - \frac{2}{N} \langle O \rangle \sum_{i=1}^N o_i + \frac{1}{N^2} \left\langle \sum_{i,j=1}^N \langle O \rangle^2 \right\rangle = \\ &= \frac{1}{N} \langle O^2 \rangle + \frac{1}{N^2} \sum_{i \neq j} \langle o_i o_j \rangle - 2 \langle O \rangle^2 + \langle O \rangle^2 = \\ \sigma_O^2 &= \frac{1}{N} \langle O^2 \rangle - \langle O \rangle^2 + \frac{1}{N^2} \sum_{i \neq j} \langle o_i o_j \rangle \end{aligned} \quad (2.6.5)$$

For uncorrelated  $o_i$ , the last two terms cancel each other because of (2.6.4) and

$$\sigma_O^2 = \frac{1}{N} \sigma_O^2 \quad (2.6.6)$$

Therefore, for  $N$  uncorrelated measurements, the standard deviation  $\sigma$ , i.e., the statistical error, is  $\sigma_{\hat{O}}$  and the final result is

$$\hat{O} \pm \sigma \quad \text{with} \quad \sigma = \frac{\hat{\sigma}_O}{\sqrt{N}} \quad (2.6.7)$$

As can be easily seen, the statistical error decreases  $\propto 1/\sqrt{N}$  with the number of uncorrelated configurations  $N$ .

### 2.6.2 Correlated Data

However, not every data sample is statistically uncorrelated and, since in this case it is taken from a computer-time series in the Monte Carlo simulation, the correlation between the data depends on the Monte Carlo algorithms and how many times they have been iterated between each measurement. If the data sample is correlated, the *autocorrelation function*

$$C_O(t) = C_O(o_i, o_{i+t}) = \langle (o_i - \langle o_i \rangle) (o_{i+t} - \langle o_{i+t} \rangle) \rangle = \langle o_i o_{i+t} \rangle - \langle o_i \rangle \langle o_{i+t} \rangle \quad (2.6.8)$$

is nonvanishing. Note that  $C_O(0) = \sigma_O^2$ .

Typically, the normalized correlation function  $\Gamma_O(t)$  exhibits an exponential behaviour for large  $t$ :

$$\Gamma_O(t) \equiv \frac{C_O(t)}{C_O(0)} \underset{t \rightarrow +\infty}{\approx} e^{-t/\tau_O} \quad (2.6.9)$$

where  $\tau_O$  is called the exponential autocorrelation time for the observable  $O$ . The exponential autocorrelation time  $\tau$  is the supremum of the values  $\tau_O$  for all possible observables:

$$\tau = \sup_O \tau_O \quad (2.6.10)$$

Hence, for the correlated case, the second line of (2.6.5) can be continued as follows, using (2.6.8):

$$\begin{aligned} \sigma_O^2 &= \frac{1}{N^2} \left\langle \sum_{i,j=1}^N (o_i - \langle O \rangle) (o_j - \langle O \rangle) \right\rangle = \\ &= \frac{1}{N^2} \sum_{i,j=1}^N C_O(|i-j|) = \frac{1}{N^2} \sum_{t=-(N-1)}^{N-1} \sum_{k=1}^{N-|t|} C_O(|t|) = \\ &= \sum_{t=-N}^N \frac{N-|t|}{N^2} C_O(|t|) = \frac{C_O(0)}{N} \sum_{t=-N}^N \Gamma_O(|t|) \underbrace{\left(1 - \frac{|t|}{N}\right)}_{\rightarrow 1} \simeq \\ &\simeq \frac{\sigma_O^2}{N} 2 \left( \frac{1}{2} + \sum_{t=1}^N \Gamma_O(t) \right) \equiv \frac{\sigma_O^2}{N} 2\tau_{O,int} \end{aligned} \quad (2.6.11)$$

where the *integrated autocorrelation time* has been introduced as:

$$\tau_{O,int} \equiv \frac{1}{2} + \sum_{t=1}^N \Gamma_O(t) \underset{N \rightarrow \infty}{\simeq} \int_0^\infty e^{-t/\tau} dt = \tau \quad (2.6.12)$$

Usually  $\tau_{O,int}$  is estimated taking at least  $1000\tau_{O,int}$  data values starting with small size lattices and proceeding to larger sizes while evaluating  $\Gamma(t)$ .

The variance computed in this way is larger than the one computed using (2.6.6), therefore for the correlated case the value of an observable  $O$  is given by

$$\hat{O} \pm \sigma_O \quad \text{with} \quad \sigma_O = \sqrt{\frac{2}{N} \tau_{O,int} \hat{\sigma}_O^2} \quad (2.6.13)$$

# Symmetries and Non-Hypercubic Lattices

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## 3.1 Spacetime Symmetries Restoration

The restoration of spacetime symmetries in the continuum limit is an important aspect of lattice field theory. Minkowskian spacetime is invariant under the action of the Poincaré group, that includes translations, rotations in the 3-dimensional space, and boosts. When performing the Wick rotation, boosts become rotations in the planes formed by each one of the *spacelike* axis and the euclidean time. The euclidean spacetime is therefore invariant, in the continuum, under translations:

$$x^\mu \rightarrow x^\mu + \varepsilon^\mu \tag{3.1.1}$$

and under rotations in 4 dimensions:

$$x^\mu \rightarrow R^\mu_\nu x^\nu \quad \text{with} \quad R \in O(4) \tag{3.1.2}$$

These invariances do not hold true anymore if the spacetime is discretized: a generic lattice is, in fact, invariant only under translations of multiples of the lattice spacing  $a$ :

$$x \rightarrow x + a\mu \quad \text{with} \quad x \in \Lambda \quad (3.1.3)$$

and under certain rotations:

$$x \rightarrow \Gamma x \quad \text{with} \quad x \in \Lambda, \Gamma \in G_\Lambda \quad (3.1.4)$$

where  $G_\Lambda \subset O(4)$  is a discrete subset of the full rotational group  $O(4)$  that depends on the lattice  $\Lambda$ . For example,  $\Lambda = \Lambda_{SH}$  is a SH lattice, then  $G_{\Lambda_{SH}}$  is the 384-elements group with all possible reflections on every coordinate plane and rotations of multiples of  $90^\circ$  around each one of the 4 axis.

The fact that the symmetries between the discrete and the continuum cases are different is not a problem itself, as long as when approaching the continuum limit the lattice symmetries tend to the continuum ones. This is the case for the translational symmetry: if  $a \rightarrow 0$ , (3.1.3) becomes (3.1.1). However, the rotational symmetry is not restored when the continuum limit is taken ((3.1.4) does not become (3.1.2) when  $a \rightarrow 0$ ), as the Simple Hypercubic lattice is invariant under rotations of multiples of  $90^\circ$  independently from the value of the lattice spacing.

For this reason, studies on the restoration of rotational symmetry have been made, through the computation of rotational invariant quantities.

### 3.1.1 Rotational Invariance of the Static Quark Potential

An example of these studies is the 1982 article of Lang and Rebbi [13], where the static quark potential  $V(r)$ , obtained from the correlator of two Polyakov loops (see (1.2.14)), has been studied at different lattice spacings. Through simulations of  $SU(2)$  lattice gauge theories the potential

$$V(r) = -T \ln \langle P^\dagger(r) P(0) \rangle \quad (3.1.5)$$

has been obtained for different values of  $r = (r_x, r_y, r_z)$  on lattices extending for  $n_s$  lattice spacings in each of the space directions and  $n_t$  lattice sites in the time direction. Furthermore, the full gauge group  $SU(2)$  has been approximated by its discrete icosahedral subgroup  $\tilde{Y}$  in order to improve the efficiency of the computation. The various (about 1000) potentials obtained for each site  $r$  have then been fitted according to the expression

$$V(r) = c_0 + c_1/r + c_2 r \quad (3.1.6)$$

Finally, a representation of the equipotential surfaces of expression (3.1.6) has been plotted in Figure (3.1.1).

In Figure (3.1.1a) a lattice with  $n_s = 8$ ,  $n_t = 4$  and  $\beta = \frac{4}{g^2} = 2$  has been used, while Figure (3.1.1b) has been obtained with data from a lattice with  $n_s = 16$ ,  $n_t = 6$  and  $\beta = 2.25$ . As the lattice spacing depends, in first approximation, from  $\beta$  in the following way:

$$a(\beta) \approx e^{-\frac{12\pi^2}{11N^2}\beta} = e^{-\frac{3\pi^2}{11}\beta} \quad (3.1.7)$$



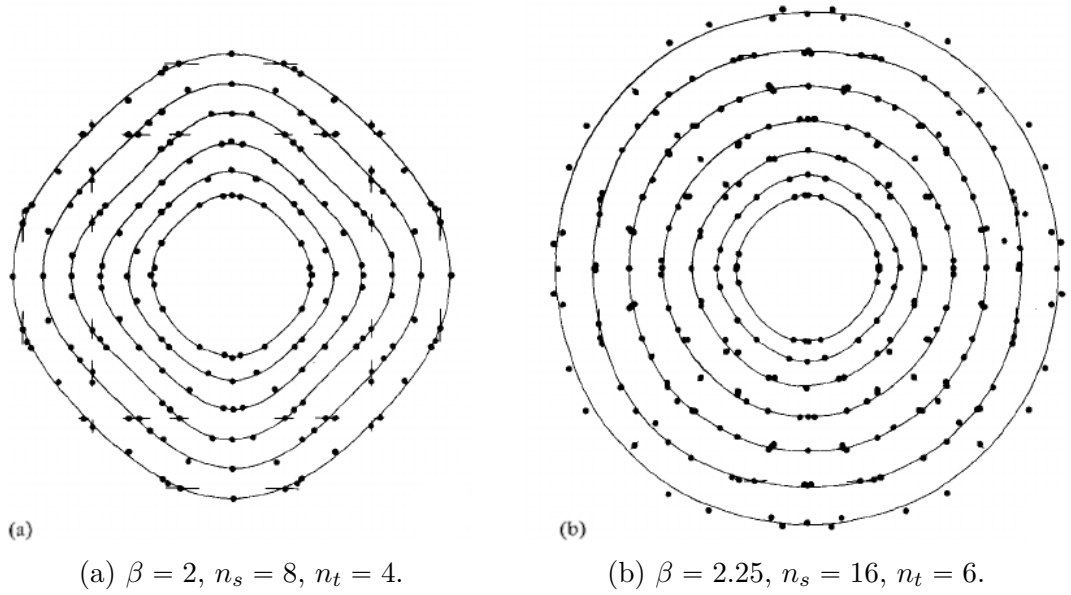


Figure 3.1.1: Representation of equipotential surfaces for larger (3.1.1a) and smaller (3.1.1b) lattice spacing.

the first plot corresponds to a higher value of  $a$  than the second one<sup>1</sup>.

As can be easily seen, by lowering the lattice spacing equipotential surfaces tend to become circles, therefore the rotational invariance, that is broken in the lattice for any value of the lattice spacing, gets restored in the expectation value of the observables, making lattice field theory a “good” theory capable of making meaningful predictions.

## 3.2 Other Types of Lattice

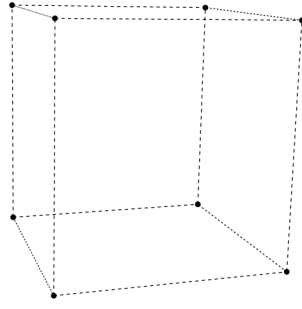
In section Section 1.2.2.1 the Simple Hypercubic lattice has been defined. Of course, it is not the only possible choice, although it is the simplest. In fact, in order to further investigate the restoration of rotational symmetry and to make better predictions on rotational invariant quantities, other types of lattices have been used to simulate lattice field theories.

### 3.2.1 Body-Centered Tesseract

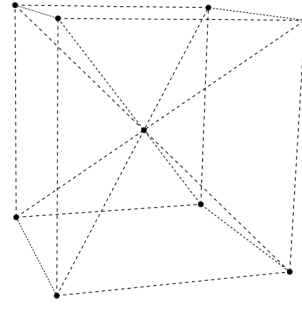
For example, the Body-Centered Tesseract (BCT) has been used for simulation of Yang-Mills theories. It consists of packing the spacetime with tesseracts, as the name suggests, but considering both the corners and the centers of every hypercube as lattice sites (see Figure (3.2.2) for a tridimensional representation of a cubic cell (3.2.2a) and a body-centered cubic cell (3.2.2b)).

Every site of the BCT lattice has, therefore, 24 nearest neighbours: 16 are identified by

<sup>1</sup>The rigorous determination of the lattice spacing is a rather complicated matter that will not be explained further, as it is not the purpose of this project.



(a) Simple cube.



(b) Body-centered cube.

Figure 3.2.2: Tridimensional representation of a simple cubic cell (3.2.2a) and a body-centered one (3.2.2b).

all possible sign permutations of  $(\pm\frac{1}{2}, \pm\frac{1}{2}, \pm\frac{1}{2}, \pm\frac{1}{2})$ , the 8 remaining are the ones of the SH lattice. The cell of this lattice is known as 24-cell, shown in Figure (3.2.3), and the plaquettes are triangular.

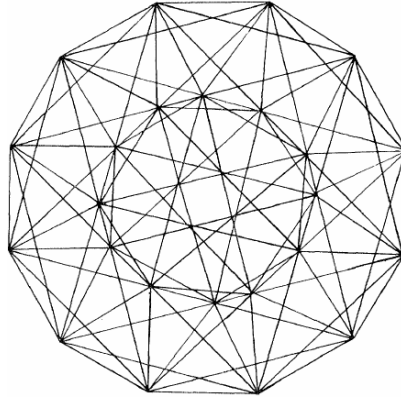


Figure 3.2.3: Bidimensional projection of the 24-cell.

This lattice has a 1152-elements symmetry group  $G_{\Lambda_{BCT}}$ , allowing a lot more rotations and reflections than the 384 of the SH lattice.

### 3.2.2 $F_4$ Coroots Lattice

$F_4$  is one of the five exceptional simple Lie groups, with Dynkin diagram  $\bullet \rightarrow \bullet \rightarrow \bullet \rightarrow \bullet$ . A more detailed explanation of exceptional Lie groups and algebras can be found in [14]. Its root lattice is a 4-dimensional body-centered hypercubic lattice, a BCT, while its dual, that is called the  $F_4$  coroots lattice, is the 4-dimensional lattice with the symmetry group of highest order. Each site of this lattice has 48 nearest neighbours:

- 24 corresponding to the roots of  $F_4$ , individuated by all possible sign and position permutations of  $(\pm 1, \pm 1, 0, 0)$ ;
- 24 corresponding to the coroots, the roots' dual vectors, individuated by

- the 8 possible sign and coordinate permutations of  $(\pm 1, 0, 0, 0)$
- the 16 possible sign permutations of  $(\pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2}, \pm \frac{1}{2})$

This lattice is made up of two BCT lattices: the first one is the dual lattice, that is the same as Section 3.2.1, the other is the root lattice, that is a BCT with lattice spacing  $\sqrt{2}$  as can be seen as represented in Figure (3.2.4).

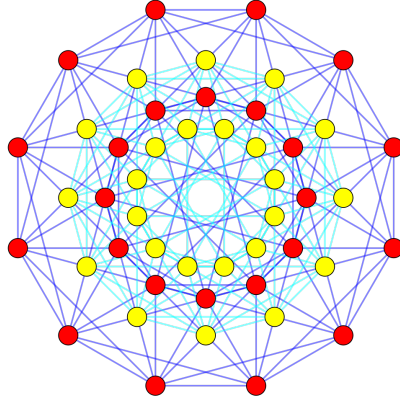


Figure 3.2.4: Bidimensional projection of the elementary cell of the  $F_4$  coroots lattice. The roots are represented in red and the coroots in yellow.

This lattice has a 2304-elements symmetry group  $G_{\Lambda_{F_4}}$ , twice as the BCT symmetry group.

### 3.3 Simulations on Higher Symmetric Lattices

These lattices have been used, in literature, to perform Monte Carlo simulations of quantum field theories.

#### 3.3.1 Scalar Fields on $F_4$ Lattice

In [15], Neuberger presented the formulation of scalar fields in terms of the  $F_4$  lattice, where he discretized the action (1.1.24) assuming the simplest nearest neighbours interaction:

$$S = \frac{1}{6} \sum_{\langle xy \rangle \in \Lambda} (\phi(x) - \phi(y))^2 + \sum_{x \in \Lambda} \left( \frac{1}{2} m^2 \phi^2(x) + \frac{\lambda}{4} \phi^4(x) \right) \quad (3.3.1)$$

where  $\langle xy \rangle$  indicates sum over nearest neighbours.

The Fourier transformation of the field  $\phi(x)$  in the limit of infinite lattice volume is given by:

$$\begin{aligned} \tilde{\phi}(k) &= \int d^4x e^{ik \cdot x} \phi(x) \\ \phi(x) &= \frac{1}{2(2\pi)^4} \int d^4k e^{-ik \cdot x} \tilde{\phi}(k) \end{aligned}$$

and the kinetic energy of (3.3.1) at infinite volume is:

$$\begin{aligned} \text{K.E.} &= \frac{1}{6(2\pi)^4} \int d^4k |\tilde{\phi}(k)|^2 \left( \sum_{\mu} (1 - \cos(k_{\mu})) + \sum_{\pm} \left( 1 - \cos\left(\frac{k_1 \pm k_2 \pm k_3 \pm k_4}{2}\right) \right) \right) = \\ &= \dots = \frac{1}{12(2\pi)^4} \int d^4k |\tilde{\phi}(k)|^2 \left( k^2 - \frac{1}{72} k^4 + O(k^6) \right) \end{aligned} \quad (3.3.2)$$

The fact that the fourth order term is proportional to the square of the second order one is a feature of all symmetry-preserving lattice actions.

This work lead to some analytical [16] and numerical [17] results and to an upper bound prediction for the mass of the Higgs boson [18] more than 20 years before its actual discovery.

### 3.3.2 Gauge Theories on BCT Lattice

In [19], Celmaster presented the first computations for an  $SU(2)$  theory on a BCT. Being a different lattice, where the plaquette is a triangle and not a square, a new action has to be defined: one that is invariant under the BCT group, with the correct classical continuum limit is:

$$S_{BCT} = \frac{1}{2g^2} \sum_{\Delta} \text{Re Tr } U_{\Delta} = \frac{\beta}{8} \sum_{\Delta} \text{Re Tr } U_{\Delta} \quad (3.3.3)$$

where  $U_{\Delta}$  is the triangular plaquette, defined as

$$U_{\Delta} = U_v(x)U_w(x + \hat{v})U_{-v-w}(x + \hat{v} + \hat{w}) = U_v(x)U_w(x + \hat{v})U_{v+w}^{\dagger}(x) \quad (3.3.4)$$

where  $v$  and  $w$  are labels indicating the possible nearest-neighbours directions, described in Section 3.2.1, such that  $v + w$  is a valid direction.

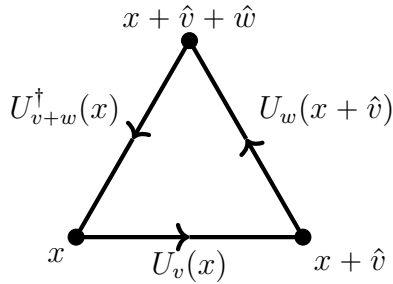


Figure 3.3.5: Schematization of an elementary triangular plaquette.

For example, if  $v = (+1, 0, 0, 0)$  and  $w = (0, +1, 0, 0)$ , then they add up to  $v + w = (+1, +1, 0, 0)$ , that is not one of the nearest-neighbours vectors, whereas  $v = (+1, 0, 0, 0)$  and  $w = (-\frac{1}{2}, +\frac{1}{2}, -\frac{1}{2}, +\frac{1}{2})$  add up to  $v + w = (+\frac{1}{2}, +\frac{1}{2}, -\frac{1}{2}, +\frac{1}{2})$ , that is a valid direction. In Figure (3.3.5) an example of an elementary triangular plaquette is shown.

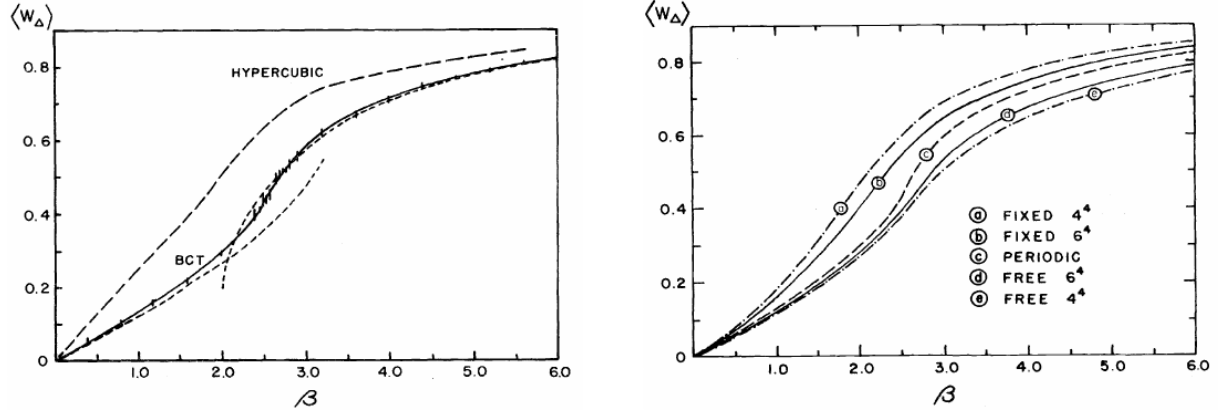
There are 96 such triangles touching each site and each edge is shared between 8 triangles. By comparison, on the SH lattice there are 24 elementary square plaquettes touching each site and each edge is contiguous to 6 squares.

In [20] and [21] Celmaster presented the first simulation results for the  $SU(2)$  gauge theory on the BCT lattice, explaining the algorithm used for the simulations in [22].

### 3.3.2.1 Average Plaquette Value on BCT

In [21] a study on the average value of the plaquette is presented: it is plotted w.r.t.  $\beta$  in Figure (3.3.6a).

As can be seen, the BCT average plaquette better follows the strong and weak coupling expansions, that are analytical approximations valid, respectively, for  $\beta \rightarrow 0$  ( $\Leftrightarrow g \rightarrow \infty$ ) and for  $\beta \rightarrow \infty$  ( $\Leftrightarrow g \rightarrow 0$ ).



(a) Comparison between  $\langle W_\Delta \rangle$  on BCT and SH lattices. The short-dashed lines represent weak and strong coupling expansions.

(b) Dependence of  $\langle W_\Delta \rangle$  on boundary conditions in different BCT lattices, with different boundary conditions.

Figure 3.3.6: Average value of the plaquette  $\langle W_\Delta \rangle$  on a BCT lattice as a function of  $\beta$ .

In the same article the influence of the boundary conditions on the average value of the plaquette is also investigated. Three different boundary conditions are simulated: periodic, that means that lattice has the topology of a 4-dimensional torus, free, that means that links on the boundary are free to assume any  $SU(2)$  value, and fixed, that means that links on the boundary are constrained to assume a certain  $SU(2)$  value. The results are plotted in Figure (3.3.6b).

From the plot, it can be seen that  $\langle W_\Delta \rangle_{free} \leq \langle W_\Delta \rangle_{periodic} \leq \langle W_\Delta \rangle_{fixed}$ , which is an inequality discussed by Mütter and Schilling in [23].

It is also pointed out that computer runtime of BCT simulations was slower by a factor of approximately 2.5 than SH simulations, but this is compensated by having 3 times as many degrees of freedom, that become  $\frac{11}{3}$  as many if a gauge fixing is done, and  $\frac{16}{3}$  as many plaquettes for each site. Furthermore, having 12 symmetry axes instead of 4 provides a higher symmetry of the observables allowing for better checks on computations.

### 3.3.2.2 String Tension

In [20] the  $SU(2)$  string tension is obtained from measurements of ratios of Wilson loops on a  $6^4$  BCT lattice.

On a BCT lattice, there are two types of Wilson loops:

- Rectangular Wilson loops  $W_R(n, m) = \frac{1}{2} \langle \text{Tr } U_R(n, m) \rangle$
- Triangular Wilson loops  $W_T(n) = \frac{1}{2} \langle \text{Tr } U_T(n) \rangle$

where  $U_R(n, m)$  is the product of link variables over an  $n \times m$  rectangle and  $U_T(n)$  is the product of link variables over an equilateral triangle of side  $n$ .

Assuming that Wilson loops are fitted by:

$$W = e^{-\sigma A - b(\text{perimeter}) + \text{const.}} \quad (3.3.5)$$

where  $A$  is the area of the loop, then the string tension  $\sigma$  can be obtained through logarithmic ratios:

$$\chi(2) = -\ln\left(\frac{W_R(1, 1)W_R(2, 2)}{W_R^2(1, 2)}\right) \quad (3.3.6)$$

$$R_T(3) = -\frac{2}{\sqrt{3}}\ln\left(\frac{W_T(1)W_T(3)}{W_T^2(2)}\right) \quad (3.3.7)$$

Results of the simulations are plotted in Figure (3.3.7), assuming no correlations between loop fluctuations in the computation of the error bars.

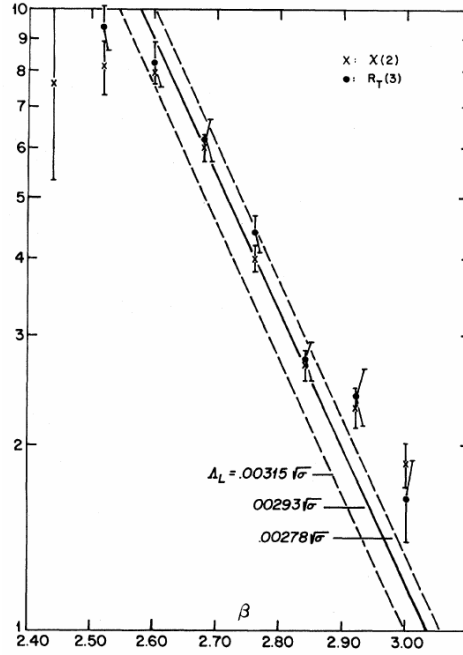


Figure 3.3.7:  $SU(2)$  logarithmic loop ratios as a function of  $\beta$  on a  $6^4$  BCT lattice.

The logarithmic loop ratio  $\chi(2)$  agrees with the asymptotic freedom curve in the region  $2.68 \leq \beta \leq 2.84$  far better than any other result obtained with simulations on SH lattices. This is a consequence of the fact that the BCT action has smaller  $O(a^2)$  corrections than usual at small  $\beta$ , therefore finite-spacing corrections are smaller for this action than for others previously examined. Therefore, BCT lattices with less sites than SH lattices could be used to obtain data with the same accuracy, at a given lattice spacing  $a$ . In the article is stated that, if the lattice spacing is sent  $a \rightarrow a/2$ , a SH lattice would require 16 times as more sites in order to keep the physical size of the Wilson loop constant, while if the same operation is done on a BCT lattice, the increment in the number of sites would be much less, using about 16 times less CPU time than the Wilson action.

Furthermore,  $R_T(3)$  is compatible with  $\chi(2)$  within error bars: this is a confirmation of the area-law dependence of equation (3.3.5).

In conclusion, simulations on BCT lattices provide better results, with more rotational invariance than those on other lattices, at a cost of a slightly higher computational time.





# Simulations Results

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4.1 Rotational Invariance Restoration

4.2 BCT Lattice Code



# Conclusions

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