Martin P. King

**Computational Methods with MATLAB Examples and Exercises**

A notebook of basic computational methods in example m-files, MATLAB exercises, and author’s solutions



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# Foreword

The materials covered in this notebook are computational methods commonly taught to second or third year undergraduates (in the UK and Australia systems) in engineering, physical sciences, and applied mathematics. Students normally have learned programming, simple linear algebra and numerical methods before they take a course of this type. A course like this is often an essential component of certain accredited engineering degrees and is also a foundational preparation before advancing to more complex or modern methods.

This is not a textbook and it does not provides much explanation of the methods, theories, and equations. Rather, it is a collection of computational lab exercises and many example codes, all in MATLAB. Example solutions are also provided immediately after the exercises. It may also be useful for looking up ‘classical’ computational methods to solve simple physical problems. The contents are based on courses for mechanical engineering students I taught some years ago.

The exercises involve modifying codes that are provided or writing your own codes, as well as perform some investigations on questions given in the exercises. Some information and tips are also given to help you complete the exercises. Whether you have taken a similar course or not, it is likely that you need to refer to a textbook or search for relevant information on the Web. I provide a list of selected references in *Bibliography*.

Of course you are encouraged to complete the exercises before looking at my suggested solutions. Plan your methods first, then implement them in codes, fix bugs, check the solutions, and then look for ways to further improve your codes.

Most of the methods I touch on are highly standard, widely taught and shared. I have also learned certain MATLAB techniques from people who generously share their codes on the Web. I have not recorded the numerous and varied sources I learned from. I do not intentionally copy anyone’s works verbatim, but for the reasons I just mentioned, I do not claim originality in what are presented.

This notebook is provided free of charge with a Creative Commons CC BY-NC 4.0 International License (see copyright page). You are welcome to send me any comment on [martin.p.king@bath.edu](mailto:martin.p.king@bath.edu). Enjoy!

Martin P. King

Bergen, Norway, 2022.

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# 2D Laplace Equation (an elliptic PDE)

**2D Laplace Equation (an elliptic PDE):**

Methods:

*a)* Inverse Laplacian

**D**T=B → T=**D**\B

Example codes: laplace2d\_fdm\_direct\_eg1.m,

laplace2d\_fdm\_direct\_eg2.m

*b)* Iterative methods

- Jacobi, example code: laplace2d\_fdm\_jacobi.m

- Gauss-Seidel with SOR (Liebmann) (this exercise)

**Exercise**

This problem involves finding the steady solutions for the heat equation:

In other words, by setting the time derivative to 0, you are required to solve:

This is also referred to as the Laplace equation (if the RHS is non-zero; then it is called the Poisson equation).

The computational domain is a two-dimensional square plate measuring 2 units long on each side. At one edge T=6 and at all the other three edges T=0 are maintained. Initially, T=0 in the whole plate away from the edge.

**a)** Solve this problem by the iterative method of Gauss-Seidel with successive over-relaxation (SOR).

Write a MATLAB code to perform the computation. Your script should not need to be longer than 100 lines including the compulsory comments, although no marks will be deducted for exceeding this length.

**b)** Consider the following questions in your investigation: What is the value for T in the exact centre of the plate at steady state? How do you know if your answer is physically reasonable? What is the convergence criterion you use? What is the grid size you choose and why is that your choice? What is the SOR parameter (*α*), you choose, and why?

**N.B.** You may find referring to laplace2d\_fdm\_jacobi.m; laplace2d\_fdm\_direct\_eg1.m; and laplace2d\_fdm\_direct\_eg2.m helpful.

**Useful formulas**

The PDE here can be discretised as:

If *Δx2=Δy2*, then

If *ui,j* is the current Gauss-Seidel estimate at the point (*i,j*) and is the previous estimate, the following ‘updating’ procedure may be implemented:

where *-1<α<1*. If *α=0*, then this is just the Gauss-Seidel iteration. If *α<0*, then this is an under-relaxtion; if *α>0*, then this is an over-relaxation.

**Solutions**

I realised that with minimal changes, I can immediately solve the PDE using laplace2d\_fdm\_direct\_eg1.m. I set xmax=2, ymax=2, Thot=6 and found that the temperature at the centre of the plate at steady state is 1.5. I set dx=0.05 and rerun and obtain the same answer. My first instinct is that 1.5 is the exact answer. However, I am not yet solving this problem using Gauss-Seidel with SOR.

Since I know the Jacobi method is also an iterative method, I modify laplace2d\_fdm\_jacobi.m to obtain the following code (see laplace2d\_fdm\_gssor.m):

laplace2d\_fdm\_gssor.m

% sample solution for Lab 8, Q 1.

% by martin king, 31 Aug 2008

%------------------------------

% solves 2D Laplace Equation by SOR Gauss-Seidel iteration.

% 0=u\_xx+u\_yy in domain [0,2]x[0,2]

% bc's: one side u=6, 3 sides u=0.

% ic: u=0.

clear all; close all;

format long;

xmin=0.; xmax=2.0; ymin=0.; ymax=2.0; % defining domain

dx=0.05; dy=0.05; % grid spacing

x=[xmin:dx:xmax]; y=[ymin:dy:ymax]; % defining mesh

[X,Y]=meshgrid(x,y);

[nx,ny]=size(X); % setting number of grid points

Thot=6.0; Tcold=0.; % boundary values

wallhot=find(X==xmin); % grid point indices for hot wall

wallcold=find(X==xmax|Y==ymin|Y==ymax); % grid point indices for cold wall

u=zeros(nx,ny);

u(wallcold)=Tcold; u(wallhot)=Thot; %boundary conditions

uold=u;

alpha=0.9; %-1<alpha<1; negative for under-relaxation, positive for over-relaxation

err=1; %so that first iteration must run

iter=0;

while err>=0.5e-10

for ix=2:nx-1

for iy=2:ny-1

u(iy,ix)=(1+alpha)\*0.25\*(u(iy,ix-1)+u(iy,ix+1)+u(iy-1,ix)+u(iy+1,ix))-alpha\*u(iy,ix);

end

end

err=norm(u-uold);

uold=u;

iter=iter+1

end

%

%plotting

mesh(X,Y,u), axis([0 2 0 2 0. 6]), view(44,38);

xlabel x, ylabel y, zlabel T

colormap(1e-6\*[1 1 1]);

xmid=ceil(nx/2);ymid=ceil(ny/2);

title(['u(0,0)=',num2str(u(xmid,ymid))]);

umid=u(xmid,ymid)

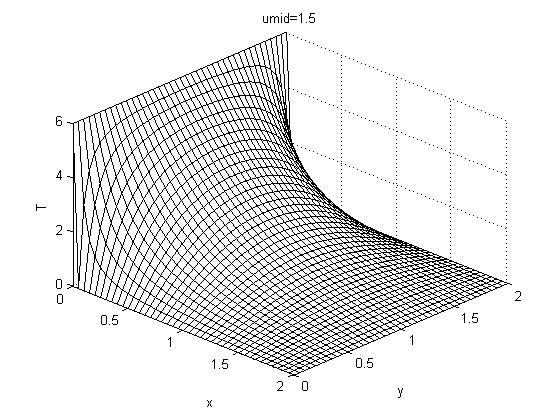
%

The two codes are essentially the same except what are inside the while loop. You can see that u(iy,ix) is immediately updated, and that I also use SOR.

I have set the stopping criterion as 0.5e-10, therefore my approximations have reached the 10th decimal place (although this doesn’t mean that my final approximation is correct to the 10th decimal place when compared with the analytical answer). I increased α from 0.2 to more than 0.9 and found that at around 0.9, the smallest number of iterations are required to satisfy the convergence criterion. I chose dx=dy=0.05, as this seems to be a good compromise between accuracy and number of iterations to reach convergence. The suitable values for α and the grid sizes may be problem-dependent.

The answer I obtained with the above code is: 1.500000000002050. Besides that the direct method also gives me an answer which is very close to this, I noticed that this is also the average temperature of the 4 edges, i.e. (6+0+0+0)/4. A few experiments suggest to me that as long as I have a square plate of any side length with any uniform temperature on an edge, this kind of prediction is always true! This is a consequence of diffusion, which can be thought of as a Brownian motion problem. A point in the middle of the plate is equi-distanced from all sides. Therefore, the concentration of the species diffused from all edges is the average concentration at the middle of the plate at steady state. Note also, that this must be correct because the concentration at any point is the average concentrations of the surrounding points.

Plotted figure:



# 2D Poisson Equation (an elliptic PDE)

**Exercise**

The governing equation for two-dimensional steady heat conduction with a heat source is

Here, we consider a heat source with a Gaussian-shaped spatial distribution

applied at the centre of a square plate measuring 2units long on each side. The temperatures T=5; T=3; T=2, T=0 are maintained at the 4 sides respectively. For *σ=0.2*, find the value of *A* necessary to maintain T=3.0 at the centre of the plate.

**a)** You are required to solve this problem by the Gauss-Seidel iterative method with successive over-relaxation (SOR). Write a MATLAB code to perform the computation.

**b)** Include also in your discussion on the following: What is the convergence criterion used? What is the grid size you choose and why? What is the SOR parameter (*α*)you choose and why?

**Solutions**

First I discretise the PDE with FDM.

If we use

,

the above discretised equation can be rearranged to

Even though the question asks us to do Gauss-Seidel with SOR, I want to solve this first in Jacobi method by modifying the provided m-file laplace2d\_fdm\_jacobi.m from the last exercise because this is a simpler task and will provide an answer to guide my further work. The resulted code I used is called poisson2d\_fdm\_jacobi.m.

By trial-and-error, I found that (using dx=dy=0.05)

A=1.539

will give T at the middle of the plate to be

T=3.000005008124929.

The relative error is about 1.7e-4% which is very small.

I then modified the m-file to perform Gauss-Seidel iteration (without SOR). The resulted m-file is named poisson2d\_fdm\_gs.m. Using the same value of A as above, I obtained

T=3.000005008528098.

The results are slightly different but have errors of same orders of magnitude. (Note the main changes from lines 40 to 48).

Next, I add the successive-over-relaxation part. The resulted m-file is called poisson2d\_fdm\_gssor.m (script shown below). Note lines 35 and 46. I also added niter in lines 38 and 51 to record the number of iterations required for convergence. From this, I found that if I used alpha=0.9, the calculation needed the least number of iterations.

For the Gauss-Seidel SOR method, I obtained

T=3.000005008931867

by using A=1.5329.

Note that the question has not specified how the 4 boundary conditions are arranged. I think there are 3 combinations. From

.

There is 4! because there are 4 different b.c.'s one for each side; there is the division by 4 because the square has rotational symmetry of in 90 degrees; there is the division by 2 because the square has mirror symmetry.

I have also coded in the three combinations of b.c's arrangements (see lines 19 to 26). Interestingly, as far as finding an A so that T in the middle is 3 is concerned, the answer is unaffected by which boundary case is used.

You should also notice that for f(x)=0, T in the middle is the average of the boundary values, i.e. T=(0+5+3+2)/4=2.5. If I set A=0., I find that this is indeed the result from the numerical iteration.

poisson2d\_fdm\_gssor.m

% sample solution m-file

% by martin king, 30 Sep 2009

%------------------------------

% solves 2D Poisson Equation by Gauss-Seidel iteration with SOR.

% 0=u\_xx+u\_yy+f(x,y) in domain [0,2]x[0,2]

% bc's: u=5, u=3, u=2, u=0

% ic: u=0.

clear all; close all; clc;

format long;

xmin=-1.; xmax=1.; ymin=-1.; ymax=1.; % defining domain

dx=0.05; dy=0.05; % grid spacing

x=[xmin:dx:xmax]; y=[ymin:dy:ymax]; % defining mesh

[X,Y]=meshgrid(x,y);

[nx,ny]=size(X); % setting number of grid points

T5=5.0; T3=3.0; T2=2.0; T0=0.0; % boundary values

switch 3

case 1

wall5=find(Y==ymax); wall3=find(X==xmax); wall2=find(Y==ymin); wall0=find(X==xmin);

case 2

wall5=find(X==xmax); wall3=find(Y==ymin); wall2=find(Y==ymax); wall0=find(X==xmin);

case 3

wall5=find(Y==ymax); wall3=find(Y==ymin); wall2=find(X==xmax); wall0=find(X==xmin);

end

u=zeros(nx,ny);

u(wall5)=T5; u(wall3)=T3; u(wall2)=T2; u(wall0)=T0; % boundary conditions

uprev=u;

f=@(x,y,A) A/0.2\*exp(-0.5\*(x^2+y^2)/0.2^2); % defining a function

A=1.5329;

alpha=0.9;

err=1; %so that 1st iteration must run

niter=0; %iteration counter

while err>=0.5e-10

%update=del2(u)+u;

%the above line does the same job as these nested loops:

for ix=2:nx-1

for iy=2:ny-1

u(ix,iy)=0.25\*(u(ix,iy+1)+u(ix,iy-1)+u(ix-1,iy)+u(ix+1,iy)+ ...

A/0.2\*exp(-0.5\*(x(ix)^2+y(iy)^2)/0.2^2)\*dx^2);

u(ix,iy)=(1+alpha)\*u(ix,iy)-alpha\*uprev(ix,iy);

end

end

err=norm(u(2:end-1,2:end-1)-uprev(2:end-1,2:end-1)); %calculating aprox error

uprev(2:end-1,2:end-1)=u(2:end-1,2:end-1); %updating internal points

niter=niter+1

end

%

%plotting

mesh(X,Y,u), axis([xmin xmax ymin ymax 0 6]), view(44,38);

xlabel x, ylabel y, zlabel T

colormap(1e-6\*[1 1 1]);

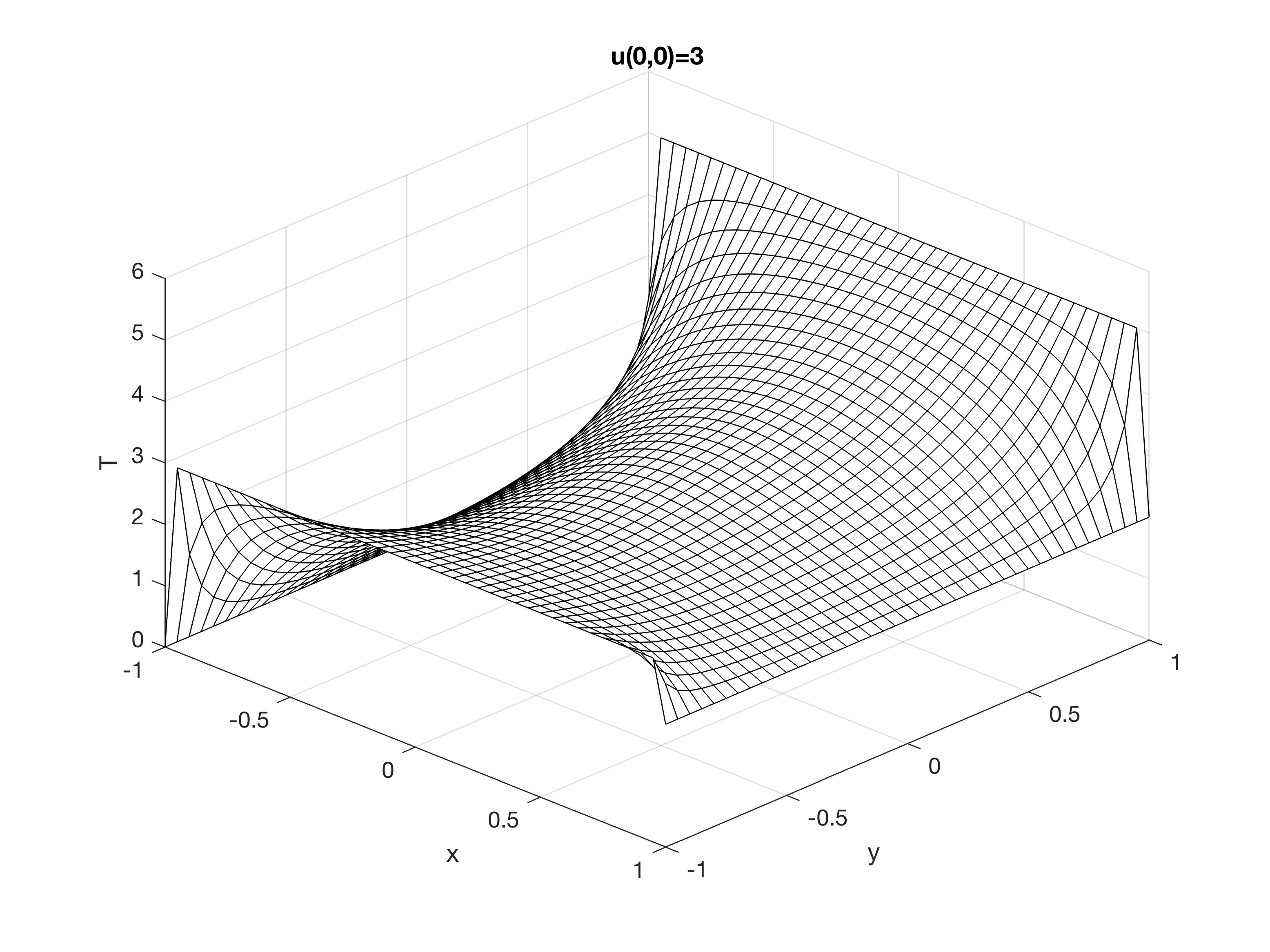
xmid=ceil(nx/2);ymid=ceil(ny/2);

title(['u(0,0)=',num2str(u(xmid,ymid))]);

umid=u(xmid,ymid)

%

Plotted figure for Case 3:



# 1D Time-dependent Heat Conduction Equation (a parabolic PDE)

**1D Time-dependent Heat Conduction Equation (a parabolic PDE):**

This is more generally known as the diffusion equation.

Methods:

*a)* Explicit time-marching and 2nd-order central finite differencing in space. Codes: heat1D\_fdm\_explicita.m, heat1D\_fdm\_explicitb.m.

*b)* Implicit time-marching and 2nd-order central finite differencing in space.

Code: heat1D\_fdm\_implicit.m.

*c)* Crank-Nicolson with the THETA schme. Code heat1D\_fdm\_CN.m and explanation below.

This chapter has no exercise. I provide four example m-files to solve this equation, and explanations for heat1D\_fdm\_CN.m below.

**Explanation for heat1D\_fdm\_CN.m**

This short note explains the solution strategy for the m-file. The aim is to solve the time-dependent 1D heat equation

We use the Crank-Nicolson method with the second-order central finite differencing. The THETA scheme is implemented too. After rearranging the discretised equation, we obtain the following (with *C=Δt/Δx2*):

**Up to line 31:**

This part sets up the computational domain, values for necessary parameters, boundary conditions, timestepping parameters. Together with the comments, they should be self-explanatory.

**Lines 33 and 34:**

In line 33, I create an ‘operator’ matrix for the LHS, and also one for the RHS for the discretised equation above.

**Lines 36 to 46:**

Here is the for loop which does the timestepping. I first need to create the RHS by applying TR=R\*TR’ in line 39. Then in lines 40 to 43, for TR(1) and TR(end), I add the boundary terms from both the LHS and RHS in the discretised equation above. Now I should have a complete RHS vector. I find T of the internal points in the next timestep by doing an inverse Laplacian line 45.

The default value for theta is set to 0 in line 15. Change it to -1 and then 0, and see what happen. The default C=5. For an explicit method to work, recall what is the stability condition you learned from the lectures.

heat1D\_fdm\_CN.m

%example m-file.

%finite differencing method for the 1D heat equation.

%solving dT/dt=d^2T/dx^2.

%crank-nicolson with 2nd order central differencing in space.

%with theta scheme.

%see also explanatory note.

%theta=1: implicit.

%theta=-1: explicit.

%theta=0: crank-nicolson.

%------------------------------------------

%

clear all, close all;

theta=0.; %see above

%defining the spatial domain

xmin=0; xmax=2; dx=0.1; x=xmin:dx:xmax; nx=length(x);

%where C=dt/dx^2

C=5; dt=C\*dx.^2;

Thot=5.0; Tcold=2.0; %Dirichlet boundary values

%grid point indices for hot and cold ends:

hotend=find(x==xmin); coldend=find(x==xmax);

T=zeros(1,nx); T(hotend)=Thot; T(coldend)=Tcold; Tnew=T; %initial T; and applying b.c. too

tlast=1.5; %final time

ntsteps=ceil(tlast/dt); %number of timesteps

dt=tlast/ntsteps; %re-correct dt

C=dt./dx.^2; %re-correct C

L=toeplitz([2+2\*C\*(1+theta) -C\*(1+theta) zeros(1,nx-4)]); %operator for LHS

R=toeplitz([2-2\*C\*(1-theta) C\*(1-theta) zeros(1,nx-4)]); %operator for RHS

for istep=1:ntsteps %time stepping

Tstore(istep,:)=T; %storing T to Tstore

TR=T(2:end-1); %setting rhs vector; only the internal points

TR=R\*TR'; %applying the RHS operator

TR(1)=TR(1)+C\*(1+theta)\*T(1); %add the b.c.'s for the implicit (LHS) terms

TR(end)=TR(end)+C\*(1+theta)\*T(end);

TR(1)=TR(1)+C\*(1-theta)\*T(1); %add the b.c.'s for the explicit (RHS) terms

TR(end)=TR(end)+C\*(1-theta)\*T(end);

T(2:end-1)=L\TR; %implicit euler's

end

Tstore(ntsteps+1,:)=T; %T at tlast

%for plotting

t=[0:dt:tlast];

[xx,tt]=meshgrid(x,t);

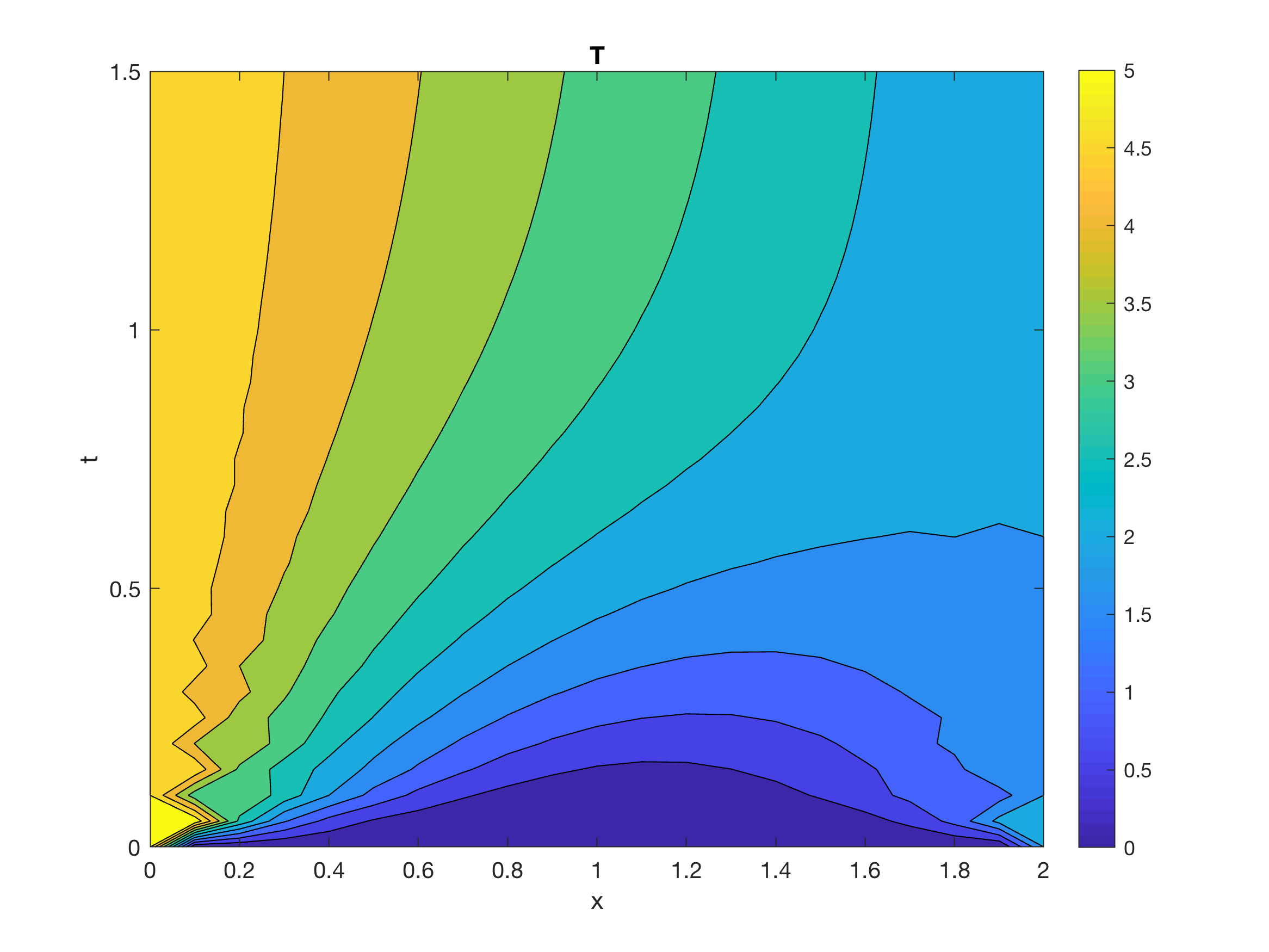
contourf(xx,tt,Tstore),colorbar;

xlabel('x'),ylabel('t');

title('T')

%

Plotted figure:



# 2D Time-dependent Heat Conduction Equation (a parabolic PDE)

**2D Time-dependent Heat Conduction Equation (a parabolic PDE):**

Methods:

*a)* Implicit (Backward) Euler time-marching with 2nd-order central finite differencing in space. Code diffusion2d\_fdm\_implicit.m and explanatory note below.

*b)* Explicit (Forward) Euler time-marching with 2nd-order central finite differencing in space (this exercise).

*c)* Crank-Nicolson method. Code: Not available; but once you know how to do *a)* and *b)* above, this should be straightforward too.

This chapter is longer than the others. I first provide explanation for an m-file solving the equation with the implicit Euler time-marching method. And then an exercise to program a code for explicit Euler time-marching is given, followed by suggested solutions.

**Explanation for diffusion2d\_fdm\_implicit.m**

This short note explains the solution strategy for the above m-file. The aim of the m-file is to solve the 2D diffusion equation

The method used is the Second-Order Central Differencing for the spatial derivatives with Implicit (Backward) Euler timestepping (SOCDBE).

I aim to use MATLAB functions to simplify the task whenever possible, instead of doing the coding explicitly. For example, I don’t create matrices and vectors manually using for loops, etc. The rules I give myself are that I don’t use the MATLAB toolboxes and I have only one self-contained m-file.

**Up to line 35:**

This part sets up the computational domain, values for necessary parameters, boundary conditions, timestepping parameters. Together with the comments, they should be self-explanatory.

**Lines 37 to 41:**

The above PDE can be discretised for SOCDBE and then rearranged to become:

Where (n+1) denotes values of *u* (unknown) for the next timestep, while n denotes the most current *u*’s. Lines 38 and 40 make a 2D Laplacian matrix (operating only on the internal points, since we don’t need to solve *u* at the boundaries) which essentially does the second term in the LHS above. In line 41, we complete making an ‘operating’ matrix for the LHS by adding the first term on the LHS. I call this the **B** matrix. Note that B is a sparse matrix for MATLAB. You can visualise it by spy(B).

For each timestep we need to solve for the vector **u** in **Bu**=**r**. We will do this in MATLAB by u=B\r. But first I need to make the **r** vector.

**Lines 43 to 53:**

Before I worry about the RHS in the discretised equation above, I need to think about terms which contain the boundary points in the second term in LHS. In MATLAB, 4\*del2(u,dx,dy) is a useful way for computing the 2D Laplacian if u is known. Now uinit is a matrix containing the known *u*’s on the boundaries and 0’s in all the internal terms. So I obtain rhsb in line 45 in such a way. The meaningful boundary terms for the Laplacian of the internal points are extracted in line 47. In lines 49 and 51, I reshape rhsb into a vector (first step in making **r**) and change its sign because I need to move these terms to the RHS. The **r** vector also contains the current *u’*s (RHS of the discretised equation above). So I add them in line 53.

**Lines 54, etc:**

Here is the real action! At each timestep, I solve B\rhs. I store the result to utemp. This is the solution for internal points of u in the next timestep. The **B** operator matrix will not change, but **r** will need to be updated with the current internal *u’s* for use in the next timestep, so I do what is done in line 56 (the same is done in line 53). The for loop does the timestepping, advancing the solution to the next timestep from rhs. The rest of the code contains some plotting facilities.

**N.B.** Here the MATLAB functions toeplitz, kron, del2 and reshape make our life easier. You may not have such conveniences if you use other programming languages. I am especially proud of how I construct the boundary terms in the Laplacian in line 45.

diffusion2d\_fdm\_implicit.m

% example m-file

% the note king\_2008\_diffusion2d\_fdm\_implicit.pdf accompanies this mfile

% martin king, martin.king@eng.monash.edu.my, 4 Sept 2008

%------------------------------

% solves 2D Diffusion Equation by Backward/Implicit Euler timestepping.

% du/dt=a(u\_xx+u\_yy) in domain [0,1]x[0,1]

% bc's: one side x=0, x=1: u=5. y=0, y=1: u=0.

% ic: u=0.

clear all; close all;

format long;

a=1.; %the coefficient of diffusion

xmin=0.; xmax=1.0; ymin=0.; ymax=1.0; % defining domain

dx=0.1; dy=0.1; % grid spacing; this code will only work for dx=dy

x=[xmin:dx:xmax]; y=[ymin:dy:ymax]; % defining mesh

[X,Y]=meshgrid(x,y); %X,Y are only used for plotting later

[nx,ny]=size(X); % setting number of grid points

%where C=dt/dx^2, dt is timestep size

C=1.0; dt=C\*dx.^2;

Thot=5.0; Tcold=0.; % boundary values

wallhot=find(X==xmin|X==xmax); % grid point indices for hot wall

wallcold=find(Y==ymin|Y==ymax); % grid point indices for cold wall

u=zeros(nx,ny); %declaring u

u(wallcold)=Tcold;u(wallhot)=Thot; %boundary conditions

uinit=u; %store initial u

tlast=2.; %final time

ntsteps=ceil(tlast/dt); %number of timesteps

dt=tlast/ntsteps; %re-correct dt

C=dt./dx.^2; %re-correct C

I = speye(nx-2); II=speye((nx-2)^2);% identity matrices

D = -a\*dt\*dx^(-2)\*toeplitz([-2 1 zeros(1,nx-4)]); % 1D Laplacian, see formula for Backward Euler

% 2D Laplacian; difficult to explain, but this is the MATLAB way to make this:

L = kron(I,D) + kron(D,I);

B = II+L;

%forming the rhs (here, it's the 2D Laplacian of the b.c.'s) by taking DEL^2,

%internal points must first be zero for this to work

rhsb=-a\*dt\*4\*del2(uinit,dx,dy);

%the border of internal points in rhsb nicely forms the 2D Laplacian of b.c's

rhsb=rhsb(2:end-1,2:end-1);

%reshape rhs into a vector

rhsb=reshape(rhsb,(nx-2)\*(ny-2),1);

%don't forget we have to change the sign, when we move terms from LHS to RHS

rhsb=-1.\*rhsb;

rhs=rhsb+reshape(uinit(2:end-1,2:end-1),(nx-2)\*(ny-2),1);

for istep=1:ntsteps %time stepping

utemp=B\rhs; %all the above was done just to do this!

rhs=utemp+rhsb;

%

% 'animation'

if(mod(istep,2)==0.)

figure(1)

utemp=reshape(utemp,nx-2,ny-2); %reshaping into a matrix

u(2:end-1,2:end-1)=utemp; %update internal points

surf(X,Y,u); shading interp; view(28,16); drawnow;

xlabel x, ylabel y, zlabel T;

end

%

end

%

%plotting the last solution

figure(2)

mesh(X,Y,u), axis([xmin xmax ymin ymax 0. 6]), view(28,16);

xlabel x, ylabel y, zlabel T;

colormap(1e-6\*[1 1 1]);

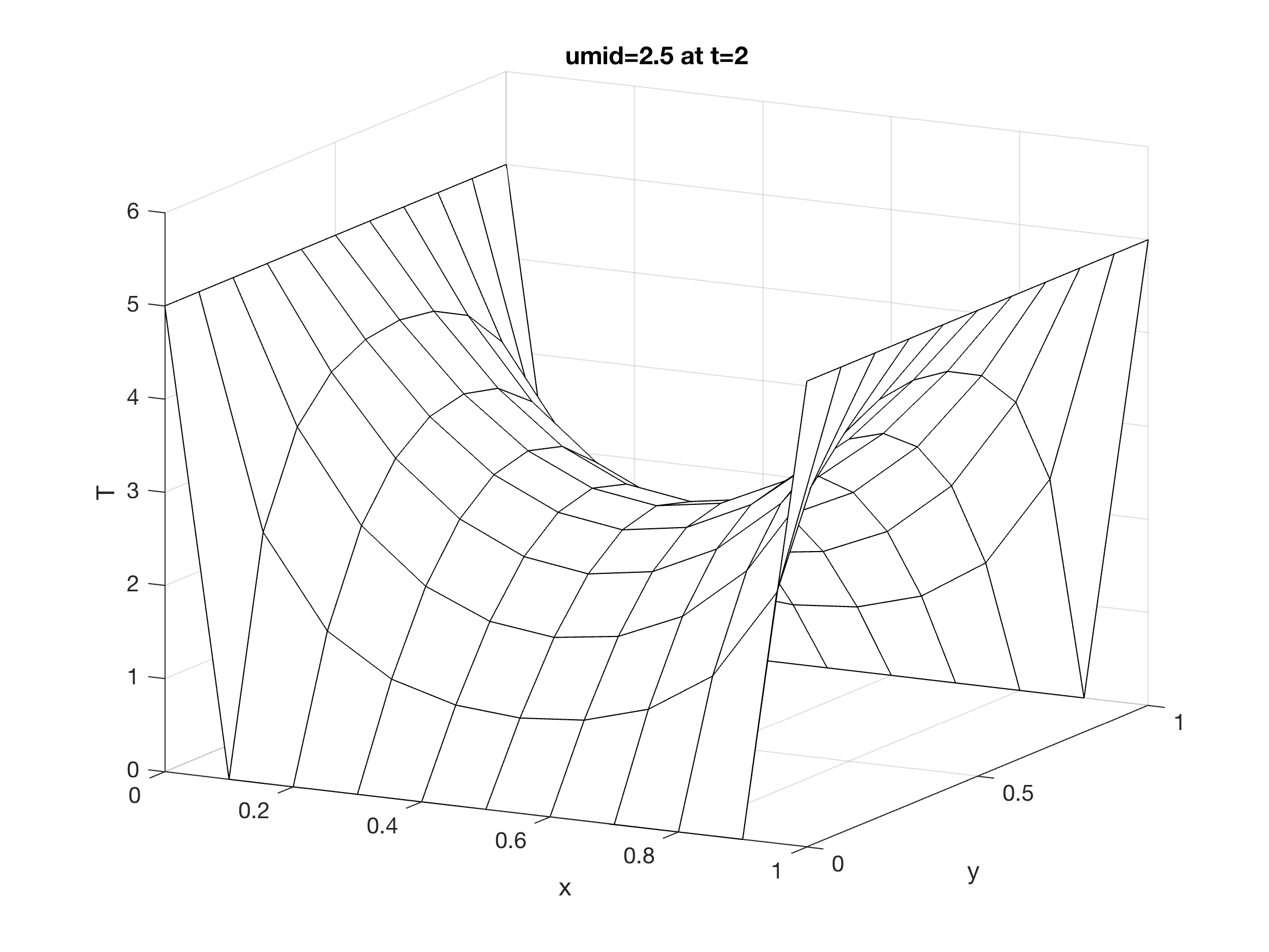
xmid=ceil(nx/2);ymid=ceil(ny/2);

title(['umid=',num2str(u(xmid,ymid)),' at t=',num2str(istep\*dt)]);

umid=u(xmid,ymid)

%

Plotted figure:



**Exercise**

This problem involves finding the solutions for the heat equation:

It is the time-dependent version of the problem from the earlier 2D Laplace Equation exercise.

Again, the computational domain is a two-dimensional square plate with edges measuring 2 units long. At one edge T=6 is maintained and at all the other three edges T=0. Initially, T=0 in the whole plate away from the edge.

**a)** Solve this problem by the explicit Euler time marching and Second-order central differencing for the spatial derivatives. Create a MATLAB code to perform the computation (you are not allowed to use any of the MATLAB toolboxes such as the PDE toolbox).

Investigate numerically to obtain the stability condition in terms of *dt/dx2*. In addition, include in your discussion the following.

**b)** What is the value for *t* (the time) in the exact centre of the plate when T reaches exactly 1 there? Explain your method and what step(s) have you taken to increase the accuracy in your answer.

**c)** Compute to a large enough final time. Do you find that the solution for T at the centre of the plate converges to a particular value? Is this value in agreement with the steady solution you found in the 2D Laplace Equation exercise?

**Suggestion:** You can consider reusing parts of diffusion2d\_fdm\_implicit.m as well as some of the techniques employed there. This m-file should be made available to you together with an explanatory note.

**Solutions**

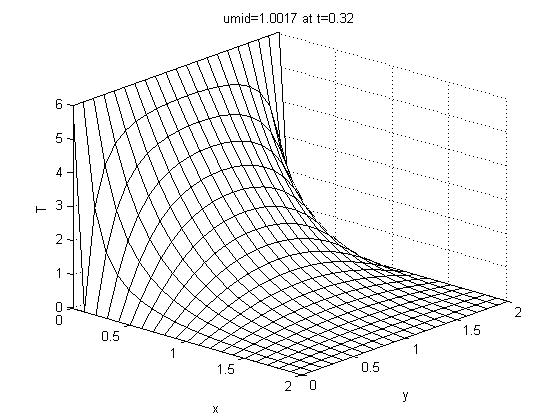
**a)** The PDE given in the question is discretised by Explicit Euler time marching and Second-Order Central Differencing for the spatial derivatives as

Like suggested in the question, I took diffusion2d\_fdm\_implicit.m and modified it. First, I set xmax=2 and ymax=2 in line 15 and Thot=6 in line 24.

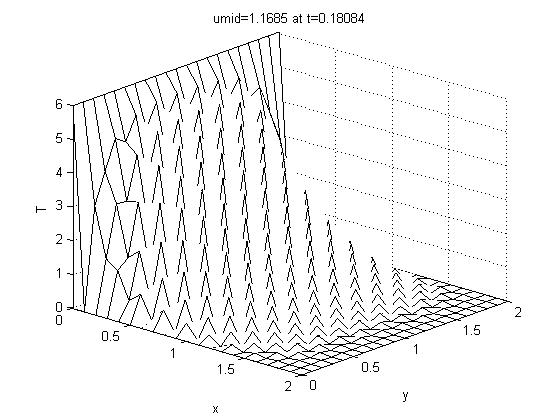
Then I deleted the part from line 37 to line 53 in diffusion2d\_fdm\_implicit.m as this is only necessary if we do a Backward Euler time-marching. The m-file I created to solve the current question is called diffusion2d\_fdm\_explicit.m. Obviously, there is usually more than one way to implement the coding. Here is just a suggested solution.

Lines 39 to 42 complete the timestepping like the discretised equation above. See also the comments in the mfile.

I first ran the code with C=0.1 (line 22). I have an if loop in line 44 to break the timestepping as soon as the solution at the middle (umid) of the plate exceeds 1. The following figure is obtained:



I then reran with C=0.2 and C=0.3. I noticed that the method is unstable at C=0.3. Therefore, I ran a case with C=0.25. And it was successful. ‘Funny’ solutions start to happen when I used C greater than 0.25. For example at C=0.27:



So I conclude that for numerical stability C=dt/dx2<=0.25 (Which, by the way, agrees exactly with the theory).

**b)** To accomplish this part. I first obtained the result using C=0.1 and dx=0.1. I get istep\*dt= 0.320000000000000 and (istep-1)\*dt= 0.319000000000000. At this stage, I am probably only confident that the answer is *t=0.3*...

I then ran all these cases with dx=dy=0.05:

|  |  |  |
| --- | --- | --- |
| C | (istep-1)\*dt | (istep)\*dt |
| 0.1 | 0.32 | 0.32025 |
| 0.05 | 0.320125 | 0.32025 |
| 0.02 | 0.32015 | 0.3202 |
| 0.01 | 0.3202 | 0.320225 |

At this stage, my confidence could be t=0.3202 when umid=1.0.

Just for the sake of experimenting I do the following two cases with *dx=dy=0.02*:

|  |  |  |
| --- | --- | --- |
| C | (istep-1)\*dt | (istep)\*dt |
| 0.01 | 0.320272 | 0.320276 |
| 0.005 | 0.320274 | 0.320276 |

The above two cases took a while to run. Now I may be confident that t=0.32027... At this point I stopped.

**c)** This should be easy. I set tlast=5 in line 31. And found that umid≈1.5 at the end. This agrees with the steady-state solution obtained in the 2D Laplace Equation exercise.

diffusion2d\_fdm\_explicit.m

% sample solution m-file.

% sample solution m-file.

% by martin king, 4 Sept 2008

%------------------------------

% solves 2D Diffusion Equation by Explicit Euler timestepping.

% du/dt=a(u\_xx+u\_yy) in domain [0,2]x[0,2]

% bc's: one side u=6, 3 sides u=0.

% ic: u=0.

clear all; close all;

format long;

a=1.; %the coefficient of diffusion

xmin=0.; xmax=2.0; ymin=0.; ymax=2.0; % defining domain

dx=0.1; dy=0.1; % grid spacing; this code will only work for dx=dy

x=[xmin:dx:xmax]; y=[ymin:dy:ymax]; % defining mesh

[X,Y]=meshgrid(x,y); %X,Y are only used for plotting later

[nx,ny]=size(X); % setting number of grid points

%where C=dt/dx^2, dt is timestep size

C=0.1; dt=C\*dx.^2;

Thot=6.0; Tcold=0.; % boundary values

wallhot=find(X==xmin); % grid point indices for hot wall

wallcold=find(X==xmax|Y==ymin|Y==ymax); % grid point indices for cold wall

u=zeros(nx,ny); %declaring u

u(wallcold)=Tcold;u(wallhot)=Thot; %boundary conditions

tlast=5.; %final time

ntsteps=ceil(tlast/dt); %number of timesteps

dt=tlast/ntsteps; %re-correct dt

C=dt./dx.^2; %re-correct C

xmid=ceil(nx/2);ymid=ceil(ny/2);

for istep=1:ntsteps %time stepping

rhstemp=a\*dt\*4\*del2(u,dx,dy); %forming the rhs by taking DEL^2 of known matrix u.

rhs=rhstemp(2:end-1,2:end-1); %the internal points in rhstemp nicely forms L[u]

rhs=rhs+u(2:end-1,2:end-1); %add the most current u to complete making rhs for the internal points

utemp=rhs; %internal points at next timestep are now just rhs

u(2:end-1,2:end-1)=utemp; %update internal points

% if u(xmid,ymid)>=1.

% break;

% end

%{

% 'animation'

if(mod(istep,2)==0.)

figure(2)

utemp=reshape(utemp,nx-2,ny-2); %reshaping into a matrix

u(2:end-1,2:end-1)=utemp; %update internal points

surf(X,Y,u); shading interp; view(35,25); drawnow;

xlabel x, ylabel y, zlabel T;

end

%}

end

%

%plotting the last solution

utemp=reshape(utemp,nx-2,ny-2); %reshaping into a matrix

u(2:end-1,2:end-1)=utemp; %update internal points

figure(1)

mesh(X,Y,u), axis([xmin xmax ymin ymax 0. 6]), view(42,26);

xlabel x, ylabel y, zlabel T;

colormap(1e-6\*[1 1 1]);

title(['umid=',num2str(u(xmid,ymid)),' at t=',num2str(istep\*dt)]);

umid=u(xmid,ymid)

%

# ODE – Numerical Order of Accuracy

**Exercise**

The radioactivity of an unstable isotope may be modelled by the following equation:

The function *u(t)* represents the concentration of the isotope. This concentration decays by a factor of two during a time interval *T* (the half-life). Find an analytical solution to the above ODE and show that

The objective of this question is to investigate numerically the order of accuracy for the Explicit Euler’s method, the Heun’s method and the Midpoint method which are covered in the lectures.

Use the m-file ode\_accuracy\_lab5\_student.m as a template. As an example, the Euler’s method has been coded at line 30. You are required to code in the Heun’s method (near line 33) and the Midpoint method (near line 37). If you find it necessary to, you may add to or modify other part of the code.

For all the methods, report and explain your findings. Do the orders of accuracy obtained agree with the theoretical expressions (usually derived using Taylor series) given during the lectures? Can you suggest a reason for the order of accuracy you observe for the Midpoint method?

**Solutions**

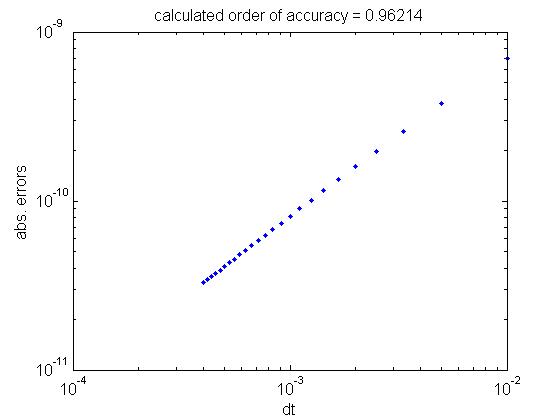
The solution to the ODE is

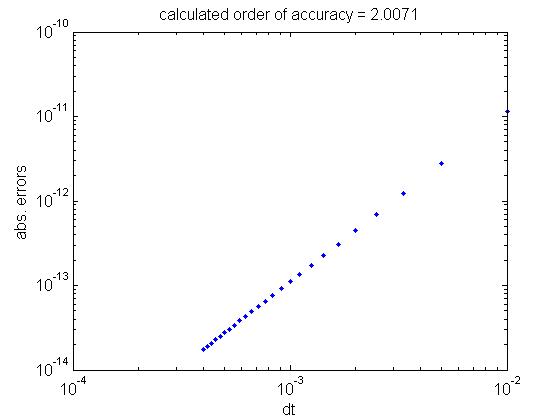
Since *T* is the half-life,

from here, it is easily shown that

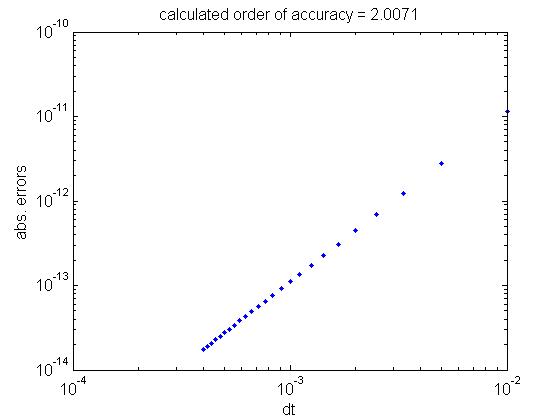
The suggested coding for case 2 (Heun) and case 3 (Midpoint) can be found in the provided m-file (ode\_accuracy\_lab5\_solution.m).

The following figures of absolute errors against dt (in log-log plots) are plotted for the Explicit, Heun and Midpoint methods respectively:

(Explicit)

****

(Heun)



(Midpoint)

These are exact errors at the end of the time integration. For the explicit Euler method, the order of error, Ea=O(dt2)*,* is for a local truncation error, i.e. for advancing a single *dt* step. Since the total number of timesteps integrated is proportional to 1/dt , the error at the end of the time integration (i.e. the total error or the global truncation error) is ~O(dt).

Both the Heun’s and the Midpoint methods have local truncation error of ~O(dt3) and therefore, global truncation error ~O(dt2)*.*

The results from our little numerical experiments confirm the theoretical analysis. Since Heun’s and the Midpoint methods are higher order methods, the absolute errors are also smaller than those for the Euler’s method for a particular *dt* used. This fact can be seen clearly in the figures above.

ode\_accuracy\_lab5\_solution.m

% Sample solution m-file.

% Suggested solutions to guide tutors.

%--------------------------------

% Investigating the order of accuracy for some common

% numerical schemes for ODES.

% Numerically solving the ODE

% du/dt=alpha\*u

clear all; % clearing variables from workspace

close all; % closing figure windows

format long; % display all the decimal places

alpha=-4.; % suggested alpha

finaltime=5.; % suggested final time

dtvect=[]; % declaring two empty matrices to store dt's and errors

errorvect=[];

% loop for different timestep sizes

for ndt=2:2:50

u0=1.; % initial condition

u=u0; % set first value of f to f0

dt=0.02/ndt; % timestep sizes

itimelast=round(finaltime/dt); % total number of timesteps required

for itime=0:itimelast-1 % timestepping loop

switch 3 % use switch to choose which case to run

case 1 %Explicit Euler

%BEGIN: WRITE YOUR CODE HERE, time-marching in u

u=u+dt\*(alpha\*u);

%END

case 2 %Heun

%BEGIN: WRITE YOUR CODE HERE, time-marching in u

k1=dt\*(alpha\*u);

k2=dt\*(alpha\*(u+k1));

u=u+0.5\*(k1+k2);

%END

case 3 %Midpoint method

%BEGIN: WRITE YOUR CODE HERE, time-marching in u

umid=u+0.5\*dt\*(alpha\*u);

u=u+dt\*(alpha\*umid);

%END

end

t=dt\*(itime+1); %advance to next time instant

exact=exp(alpha\*t)\*u0; %the exact solution to the ODE

error=abs(u-exact);

end

figure(1)

loglog(dt,error,'.','MarkerSize',16) %plotting errors against dt

hold on;

% store dt and final error to dt\_vect and error\_vect respectively

dtvect=[dtvect; dt]; errorvect=[errorvect; error];

end

hold off;

p=polyfit(log10(dtvect),log10(errorvect),1); %finding gradient of the points in figure 1

title(['calculated order of accuracy = ',num2str(p(1))]); %print it out as title

xlabel('dt');

ylabel('abs. Errors');

%

# ODE – Numerical Order of Accuracy and Stability

**Exercise**

We have encountered the following ODE in the previous exercise:

The function *u(t)* may represent the concentration of a radioactive isotope.

**a)** Use the m-file ode\_accuracy\_lab6\_student.m as a template. As before, the Euler’s method has been coded near line 30. This m-file is essentially identical to ode\_accuracy\_lab5\_student.m which you have used in the previous exercise. However, here you are required to code in the Runge-Kutta 4th-Order method for case 2. Investigate the order of accuracy for this method.

**b)** Obtain a stability condition (analytical) for the Euler’s method for this ODE, in terms of *dt*, the size of the timestep. Then, investigate this stability condition numerically on the above ODE. Does your finding agree with the theoretical derivation?

**Hint**: For part b, you can make some minimal changes to either one of the two m-files above.

**Solutions**

**a)** Please asee case 2 in ode\_accuracy\_lab6\_solution.m. The following snippet of MATLAB coding is used for RK4.

k1=dt\*(alpha\*u);

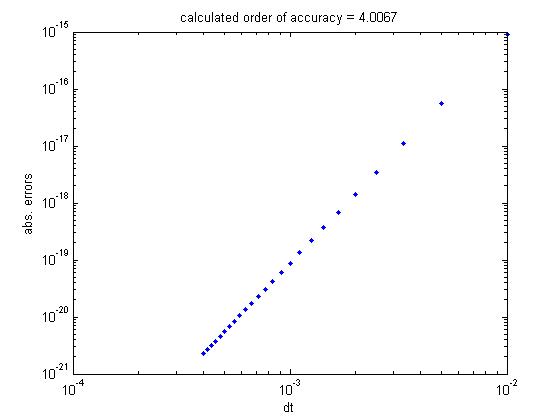
k2=dt\*(alpha\*(u+k1/2));

k3=dt\*(alpha\*(u+k2/2));

k4=dt\*(alpha\*(u+k3));

u=u+1/6\*(k1+2\*k2+2\*k3+k4);

If the RK4 is coded correctly (without a change to other default values) the following figure will result:

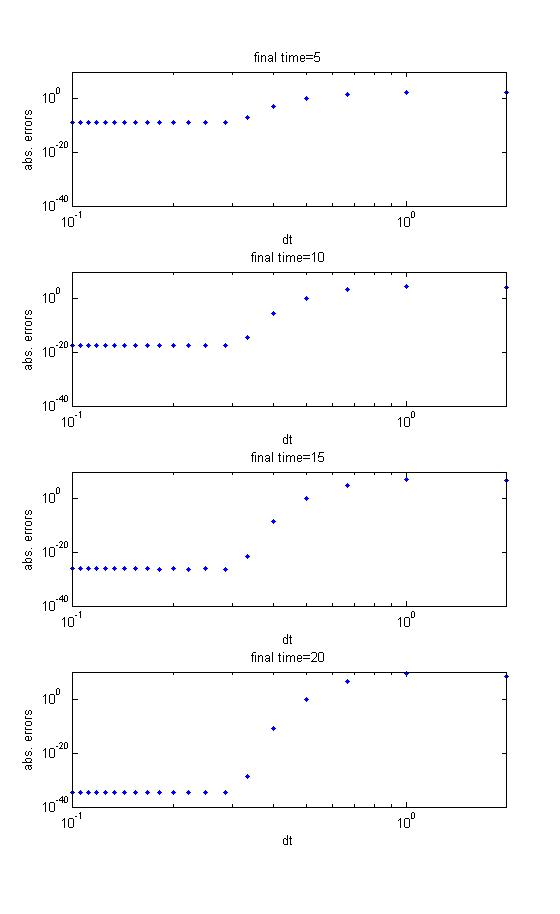


Thus, the RK4 method applied to the present ODE has an order of accuracy which agrees with the theoretical accuracy for RK4.

**b)** From an example in the lectures, the stability condition for an ODE of the present form is:

Here k=-4. Therefore:

You should notice that in part a, the largest dt used is 0.01, thus all the cases carried out in part a are numerically stable. To carry out a numerical investigation for the stability condition, I perform computations with a larger values of *dt* (see line 25 in ode\_stability\_lab6\_solution.m), i.e from dt=0.1 to 2. I also investigate cases for 4 different final times (see the outermost for loop at line 16). If you just go ahead and run this code, you should obtain this figure:



From these figures, I am able to conclude that the method starts to deteriorate at dt=0.3. The errors appear to reach an unacceptable O(1) at dt=0.5. Also, from the figures, we can see that above dt=0.5, as longer durations are being integrated (going down the panel of figures), the errors are larger. Whereas the reverse is true for dt<0.3, i.e. errors become smaller as larger final time are being integrated. This is consistent with the concept of numerical stability in the current context. Thus I may conclude that numerically, the Euler’s method applying to this ODE obeys the theoretical stability condition, although the errors may already become unacceptable at around dt=0.4, especially in cases in which the final times integrate to are short (see top figure).

**N.B.** There is no one fixed way to carry out the investigation for part b. What is shown here is only a guide or a suggestion. As long as you can perform an investigation numerically with a range of dt, and arriving at the same findings and conclusions, then regardless of methods or coding used, the solutions may be regarded as successful and complete.

ode\_stability\_lab6\_solution.m

% Sample solution m-file.

% Suggested solutions.

%--------------------------------

% Investigating the order of accuracy for some common

% numerical schemes for ODES.

% Numerically solving the ODE

% du/dt=alpha\*u

clear all; % clearing variables from workspace

close all; % closing figure windows

format long; % display all the decimal places

alpha=-4.; % suggested alpha

finaltime=5.; % suggested final time

dtvect=[]; % declaring two empty matrices to store dt's and errors

errorvect=[];

% loop for different timestep sizes

for ndt=2:2:50

u0=1.; % initial condition

u=u0; % set first value of f to f0

dt=0.02/ndt; % timestep sizes

itimelast=round(finaltime/dt); % total number of timesteps required

for itime=0:itimelast-1 % timestepping loop

switch 2 % use switch to choose which case to run

case 1 %Explicit Euler

%BEGIN: WRITE YOUR CODE HERE, time-marching in f

u=u+dt\*(alpha\*u);

%END

case 2 %Runge-Kutta 4-th Order

%BEGIN: WRITE YOUR CODE HERE, time-marching in f

k1=dt\*(alpha\*u);

k2=dt\*(alpha\*(u+k1/2));

k3=dt\*(alpha\*(u+k2/2));

k4=dt\*(alpha\*(u+k3));

u=u+1/6\*(k1+2\*k2+2\*k3+k4);

%END

end

t=dt\*(itime+1); %advance to next time instant

exact=exp(alpha\*t)\*u0; %the exact solution to the ODE

error=abs(u-exact);

end

figure(1)

loglog(dt,error,'.','MarkerSize',6) %plotting errors against dt

hold on;

% store dt and final error to dt\_vect and error\_vect respectively

dtvect=[dtvect; dt]; errorvect=[errorvect; error];

end

hold off;

p=polyfit(log10(dtvect),log10(errorvect),1); %finding gradient of the points in figure 1

title(['calculated order of accuracy = ',num2str(p(1))]); %print it out as title

xlabel('dt');

ylabel('abs. errors');

%

# ODE – Lorenz Equations

**Exercise a**

The most famous equations displaying chaos behaviour are arguably the Lorenz equations:

E.N. Lorenz (1917-2008) derived them as a simplified model for convective cells in the atmosphere. It turned out that Lorenz’s work on this simple set of equations would revolutionalise the whole field of dynamical systems.

For this question, you will investigate one important property of a chaotic system. Obtain the mfiles lorenz\_lab7q1\_student.m and lorenzrhs.m. Run lorenz\_lab7q1\_student.m, using *X0*=10 (the initial condition for *X*), and final time =100. This m-file is coded with the explicit Euler’s time-marching method. It will take some time to finish running, be patient.

Make sure that you can obtain the famous butterfly-shaped Lorenz attractor. The solutions, which have a distinctive structure, are attracted (hence the name) to a particular part of the *X-Y-Z* space. But you should recognise that the solutions are non-periodic with complicated fluctuations (see figure 2). You might also see from figure 2 that the solutions also jump from one ‘wing’ of the butterfly to another.

**a)** The time history of v=[*X Y Z*] is stored in a matrix called vsave (near line 51). Save vsave to a MATLAB data file (MAT-File). Run the code again, but decrease *X0* by 0.1%. Also save vsave for this second case to another MAT-File. Therefore now you have the time history of two vectors. Clear your MATLAB workspace of previous variables and load these two vectors. Plot the separation (i.e. the distance between two three-dimensional vectors) for these two vectors against time (you should scale the separation by the magnitude of one of the two vectors). In other words, plot:

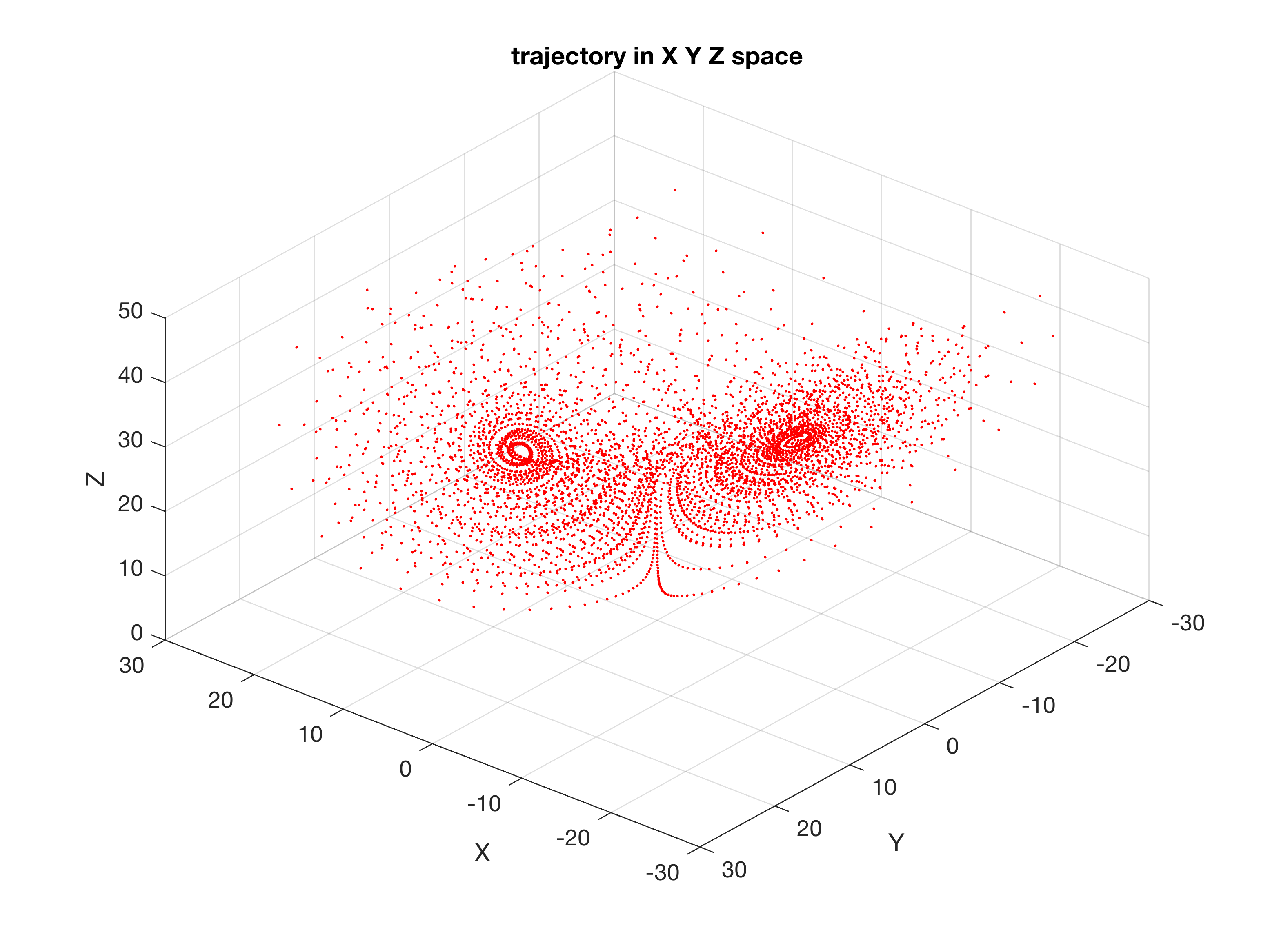
against *t*. Write another simple mfile to carry out this task if you need to. What can you say about the separations of the two solutions? For how long do the vectors stay within 10% of each other? **Hint:** you may find the MATLAB function norm()useful.

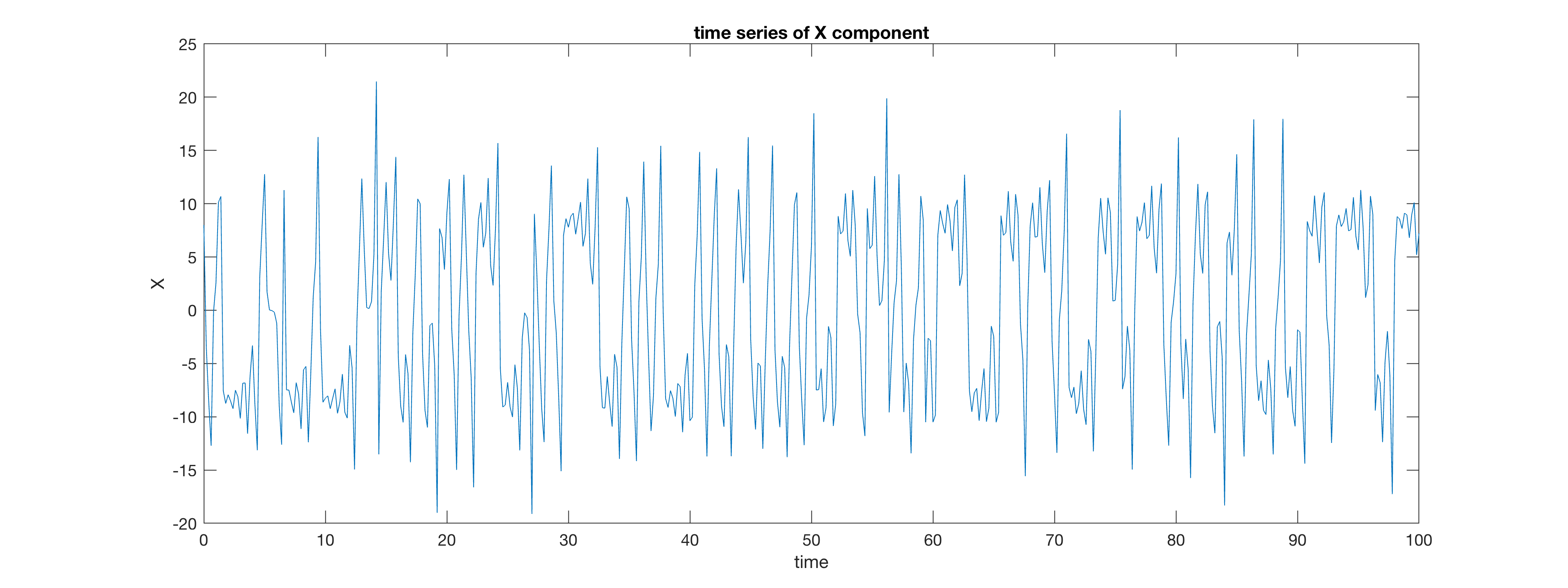
**b)** Repeat the investigation in a) by using a smaller timestep, say dt=0.002(note the default is set to dt=0.02). What do you observe?

**c)** With, again, dt=0.02 but using an extremely small difference in initial conditions (e.g. 1e-13), repeat the investigation. What do you observe?

**Solutions a**

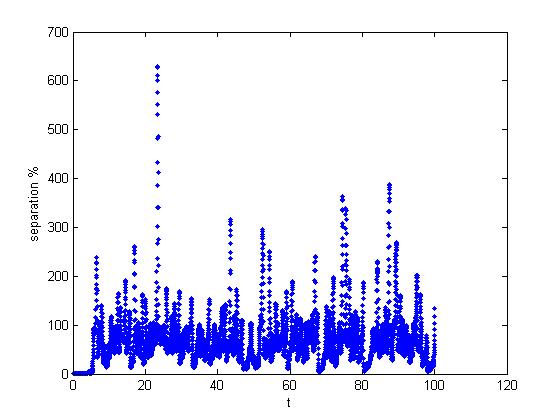
**a)** I first run lorenz\_lab7q1\_student.m as suggested. The following figures about the Lorenz attractor is obtained.





Next, I uncomment line 69 in lorenz\_lab7q1\_student.m so I can save the matrix vsave. I then run two cases with *X0*=*10*. and *X0*=*9.99* (which is *0.1%* from *10*). I then use lorenz\_loadplot\_solution.m to load and plot the normalized separations of the two solutions against integration time.

It is found that the solutions stay within *10%* of each other until t=5.4. From then on, the solutions are observed to have large but bounded differences. The differences are so large that they can be regarded as two different solutions.

****

**b)** I repeat the two cases with dt=0.002. And find that the solutions stay within 10% for t=5.5. After that the differences in the solutions quickly grow to be as large as those in case **a)**. It seems that a more accurate time integration may not increase the lead time in prediction.

**c)** I run a case with dt=0.02 and *X0*=10-(1e-13). Comparing with the solution for the case dt=0.02 and *X0*=10, the solutions stay within 10% until t=21.8. Thereafter, the differences of two solutions also become very large.

It is concluded that Lorenz equations are very sensitive to the initial condition used. Using a smaller timestep does not delay the time to reach large separation. It seems that obtaining two different solutions are ultimately inevitable, even in cases which very similar initial conditions are used. Note importantly that you may not get identical numerical results to mine because tiny differences between our softwares and computers can cause the divergence.

**Exercise b**

This question is a continuation of Exercise a.

**a)** A valid assumption is that a high-order method may help delay or restrict the separation of the two solutions which are started with a small difference in initial conditions. Modify lorenz\_lab7q1\_student.m to employ the Runge-Kutta 4th-order method. Repeat the investigations you carried out in part a of Exercise a) using the Runge-Kutta 4th-order method (with a difference in *X0* of 0.1% and dt=0.02). What do you observe?

**b)** What if now dt is being decreased to dt=0.002 (i.e. repeating as part b of Exercise a)?

**c)** Repeat the computation as in part c of Exercise a.

**d)** On many occasions in this course, you have encountered and applied the concepts of accuracy and stability, especially during the computing sessions. A third pillar of numerical methods is *consistency*. For a method to be *consistent*, the solutions must converge when smaller and smaller *dt* is used while other parameters remain the same. Using *X0*=10, and a suggested final time=50, verify that neither the explicit Euler’s method nor the Runge-Kutta 4th-Order is *consistent* for the Lorenz equations.

**Hint:** You may want to cancel plotting and cancel storing vsave to speed up the computations. You can check on the solutions to the equations at final times (just v) to see if they converge. You can try dt down to 2d-6. Be patient, cases with small timesteps may take quite a while to run.

**N.B.** Regarding part **c)**: In fact, it’s been recently suggested that no consistent numerical method has been found for the integrations of ODEs that exhibit chaotic behaviours. The computed numbers are not solutions of the continuous differential equations (in the limit dt→0), but are merely numerical noise which (luckily?) also resides in the attractor. If interested see Yao, L.S. and D. Hughes, 2008, Comment on “Computational periodicity as observed in a simple system” by Edward N. Lorenz, Tellus A, 60A, pp. 803-805.

**Solutions b**

**a)** I code in the R-K4 method in lorenz\_lab7q2\_solution.m.

k1=dt.\*lorenzrhs(v,r,sig,b);

k2=dt.\*lorenzrhs(v+k1/2,r,sig,b);

k3=dt.\*lorenzrhs(v+k2/2,r,sig,b);

k4=dt.\*lorenzrhs(v+k3,r,sig,b);

v=v+(k1+2.\*k2+2.\*k3+k4)/6.;

And redo the investigation as I do for part a of Exercise a. I find that the two solutions stay within 10% of each other until t=5.06. As always, the differences quickly grow to be large after that. It appears that a higher order method may not provide better predictability.

**b)** Now I redo part a) but with dt=0.002. It is found that the solutions stay close until t=5.06.

**c)** I find t=33.0. From the simple exercise performed so far, it may be commented that the only factor that can help in delaying the time when large separations start to occur is in having initial conditions which are closer to each other.

**d)** I run both the Euler’s and R-K 4 methods with the suggested *X0*=10 and final time=50. First using dt=0.02 and subsequently a number of cases with further reductions of dt until dt=0.000002. There is no indication that the final solution for **v** at t=50 is reaching some converging values. This may point to the serious problem in consistency for applying Euler’s and R-K 4 methods on systems of equations that exhibit chaos.

lorenz\_lab7q2\_solution.m

% Lab7 Q2.

% Sample solution m-file.

%---------------------------------------------------

% time integration of "chaotic" Lorenz 3-component system.

% employs 4th-Order Runge-Kutta. this code needs lorenzrhs.m to run

% Martin King, Aug 2008.

clear all;

close all;

%setting the initial conditions

disp('initial Y and Z are zero. Specify initial 0 < X <= 10 when asked ')

X0 = input('give an initial X value ');

Y0 = 0;

Z0 = 40;

time = input('final integration time units (at least 100) ');

% define constants of the system

r = 28;

sig = 10;

b = 8/3;

%timestep size

dt = 0.02;

v = [X0 Y0 Z0]; % vector of initial conditions

vprime = zeros(1,3);

% setting up figure 1 for plotting

figure(1)

axis([-30 30 -30 30 0 50]), view(-140,45)

xlabel('X'), ylabel('Y'), zlabel('Z')

title('trajectory in X Y Z space')

hold on

n = 0; % counter

vsave=[];

for t = 0:dt:time

%R-K 4

% lorenzrhs is a function written in lorenzrhs.m and returns the

% vprime vector

k1=dt.\*lorenzrhs(v,r,sig,b);

k2=dt.\*lorenzrhs(v+k1/2,r,sig,b);

k3=dt.\*lorenzrhs(v+k2/2,r,sig,b);

k4=dt.\*lorenzrhs(v+k3,r,sig,b);

v=v+(k1+2.\*k2+2.\*k3+k4)/6.;

p = plot3(v(1),v(2),v(3),'r.','MarkerSize',4);

drawnow

vsave=[vsave; v];

if mod(t,.2) == 0

n = n+1;

Xhist(n) = v(1);

thist(n) = t;

end

end

grid

hold off

scrsz = get(0,'ScreenSize');

%[left, bottom, width, height]

figure('Position',[2 2 scrsz(3)/2. scrsz(4)/3.]);

figure(2)

plot(thist(1:n),Xhist(1:n))

xlabel('time'); ylabel('X')

title('time series of X component')

%uncomment the following line to save vsave to file1.mat

%save file1 vsave;

%

# ODE – Eigenvalues and Eigenvectors

**Exercise**

This question originates from the physical example of a slender column under compression given in the lectures. It concerns the eigenvalue problem for the ODE:

If this equation is discretised by the Second-order Central Differencing, the matrix equation in the following form will result:

**D*y*=**λ***y***.

where **D** is a tridiagonal matrix, ***y***a column vector containing the unknown values of *y* in the internal points and λ a scalar. Therefore this is an eigenvalue problem for the matrix **D**.

**a)** Take the length of the column as L=4, and 19 internal nodes for the discretisation. Then use MATLAB to calculate and plot the first 4 eigenmodes for the column fixed at both ends under compression. Show also their corresponding eigenvalues (the 4 smallest ones).

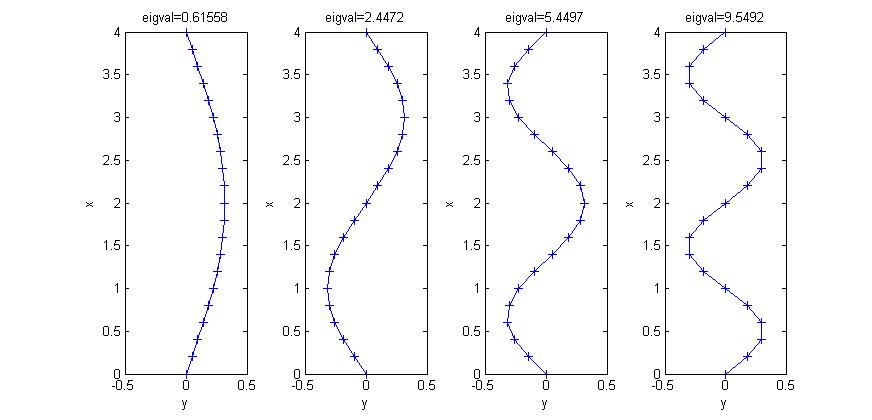
**b)** Since p2=P/EI, we can rewrite the above ODE as

Now, a concrete column with modulus *E* which changes linearly (along the column) from 30 to 100 is installed. Repeat part a).

**Hint:** You can use the MATLAB built-in function eig()to find the eigenvectors and eigenvalues. You are reminded that the results returned by eig()are not necessarily sorted in order in the eigenvalues.

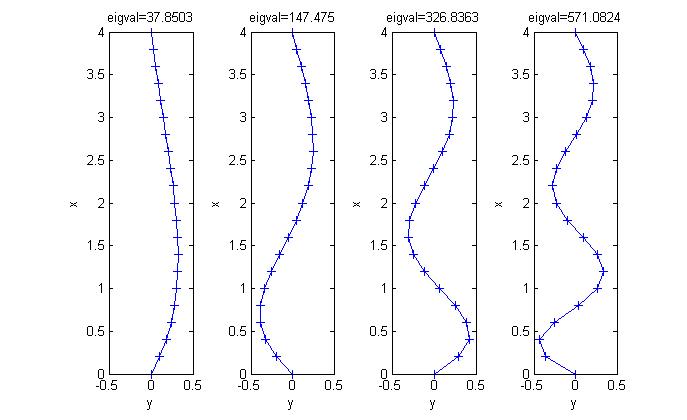
**Solutions**

**a)** The **D** is a tridiagonal matrix of 19 rows x 19 columns containing -dx^(-2)\*(1, -2, 1) in its three main diagonals respectively. I create **D** in line 11 in eigenvector\_solution.m. Although I call it A there. I used eig(), to find the eigenvectors and eigenvalues in line 22. In line 23, I sorted the eigenvalues. At the same time I obtained the order that it should be sorted and stored the indices in isort.

In the for loop from line 25, I plotted the first 4 eigenvectors, with the titles as the eigenvalues (p2) . The following is the figure I obtained:

**b)** This part is slightly tricky because we need to incorporate the linearly varying *E* in **D**. The part from lines 20 to 23 do that. I first create EE, which is a tridiagonal matrix containing the internal E’s in the principal diagonal (diag(E(2:end-1),0)). And the diagonal entries below the principal diagonal end with the second last internal point of E’s (diag(E(2:end-2),-1)), while the diagonal above starts with the second internal point of E’s diag(E(3:end-1),1). To complete, I multiply EE by the respective entries in A in line 23.

I obtained:



This is interesting because we now find that the ‘buckling’ (the wavy parts) shifted to the weaker part of the column near the bottom of the column.

# ODE – Boundary Value Problems

**Exercise**

(This exercise originated from Ravi Jagadeeshan, Monash University. But the sample solutions are mine.)

Consider the flow of a Newtonian fluid between two flat wide plates which are distance 2Hapart, and which is governed by the equation:

subject to the boundary conditions:

**(a)** Use MATLAB to solve this two-point boundary value problem for Δp=2.8⨯105 Pa*,* μ=0.492 Pa s*,* L=4.88 mand H=0.0025 m*.*

**(b)** Plot the dependence of velocity on position.

**(c)** Compare your numerical prdiction with the analytical solution by plotting the two solutions on the same plot.

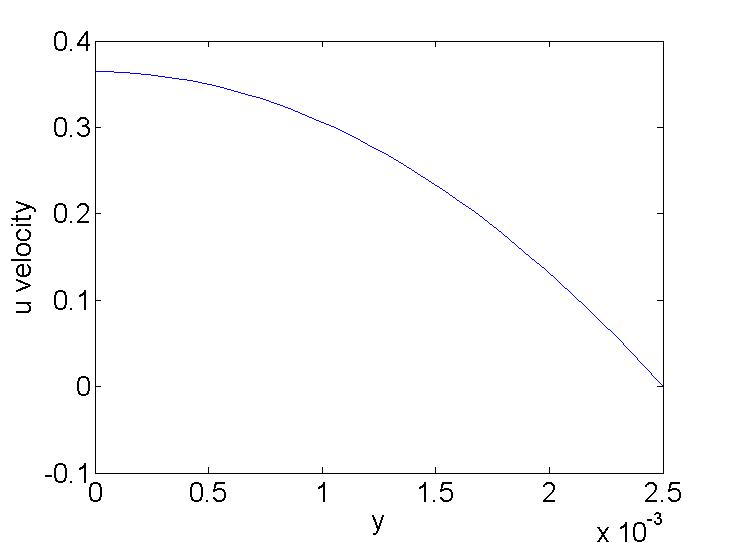
**(d)** Compute the average velocity of the fluid between the plates and compare your result with the analytical prediction.

**Solutions**

I ‘cheated’ with using MATLAB’s ODE45 function. But you gain from how to use it. If you have written a code from scratch to solve this, bravo! We can compare our answers below.

**(a) and (b)**

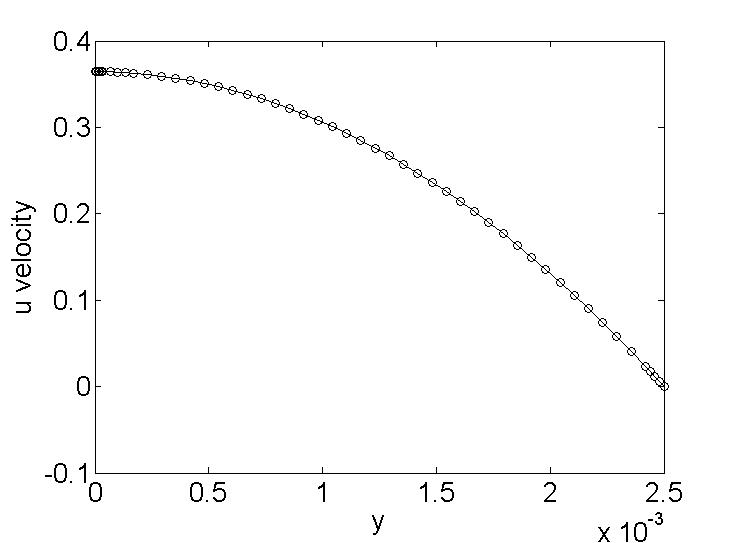
The codes (main.m, twoode.m, and res.m) I used are provided. Since the boundary condition is also given in the mid-plane (at y=0), I only solved half of the flow field. from y=0 to y=0.0025. Here’s the solution I obtained:



The solution for the complete flow field is symmetrical about y=0.

**(c)** You should be able to integrate the 2nd-order ODE twice (and using the appropriate boundary conditions given) to obtain the following analytical solution:

Now, I plotted the analytical solution in a solid line and the numerical solution in open circles. As can be seen from the figure, the numerical and analytical solutions agree very well.



**(d)** The average velocity is the velocity that would give the same flow rate as the parabolic velocity.

From this, it is found that Uave=0.2429... The average velocity from the numerical solution is found by

2\*trapz(y,u(:,1))/(2\*H)

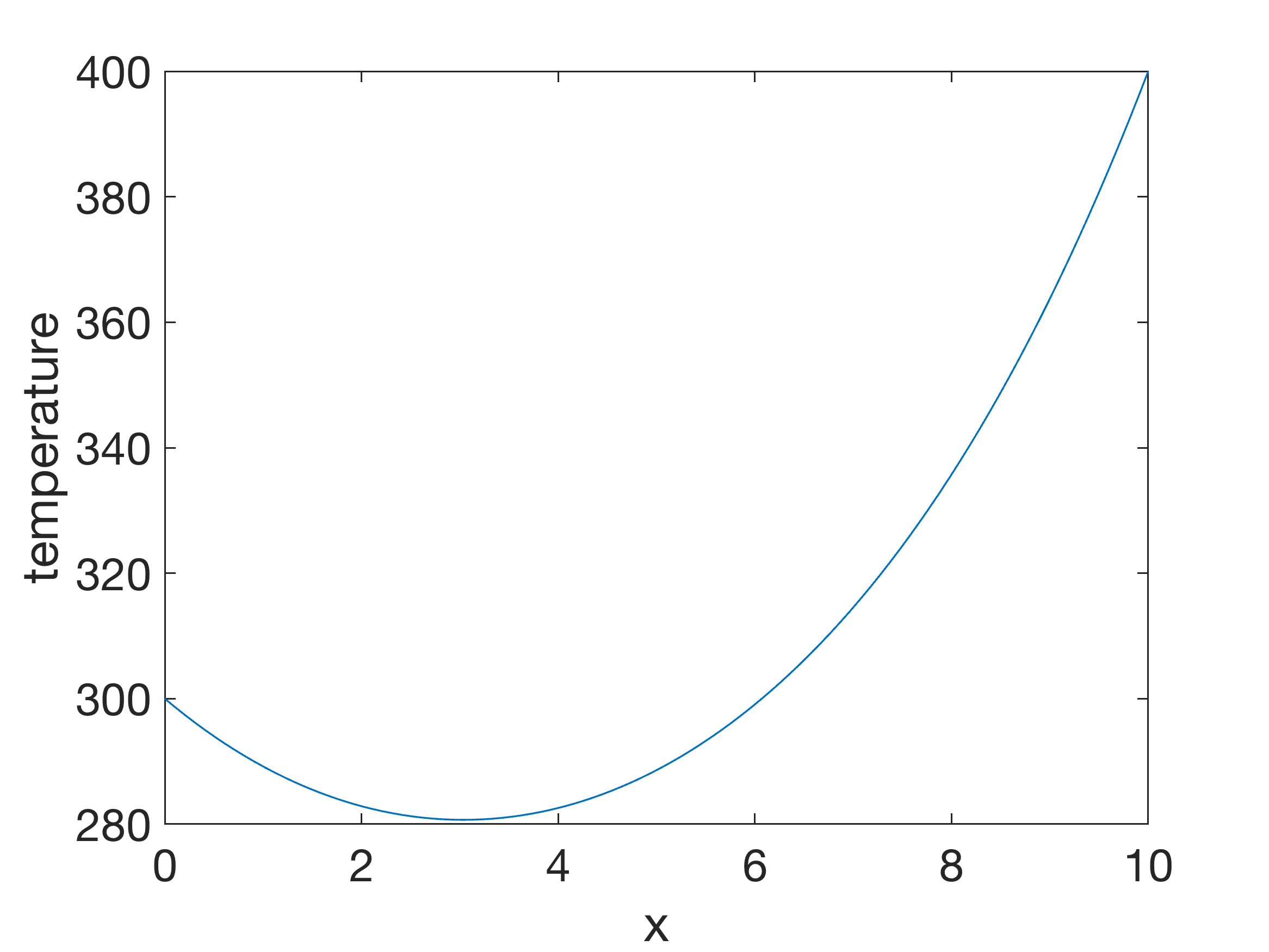
which gives also identical solution to the above. This is not too surprising as the numerical solution for velocity is so close to the analytical solution.

**Bonus example**

Here I will show a slightly different MATLAB technique of solving same type of equation. This ODE can model heat conduction in a one-dimensional solid with a temperature-dependent heat source/sink applied along it.

x=[0, 10], T(0)=300, T(10)=400, Ta=200, and h=0.05.

You can read the m-files main\_heat.m, twoode\_heat.m, and twobc\_heat.m provided. The following figure shows the numerical solution:



# Interpolation

**Exercise a**

**(a)** Interpolate the following data by each of the four MATLAB functions: polyfit, interp1, pchip and spline. Plot the results for -1 < *x* < 1.

|  |  |
| --- | --- |
| *x* | *y* |
| -1.00 | -1.0000 |
| -0.96 | -0.1512 |
| -0.65 | 0.3860 |
| 0.10 | 0.4802 |
| 0.40 | 0.8838 |
| 1.00 | 1.0000 |

**(b)**What are the values of each of the four interpolants at x=-0.3? Which of these values do you prefer? Why?

**(c)** The data were actually generated from a low-degree polynomial with integer coefficients. What is that polynomial?

**Solutions a**

Sorry, I have not written any solution for this exercise. It is a very straightforward one though.

**Exercise b**

The following data points are generated for 10 equally spaced points in the range -1 < *x* < 1 by the function:

|  |  |
| --- | --- |
| *x* | *f(x)* |
| -1.0000 | 0.0385 |
| -0.7778 | 0.0620 |
| -0.5556 | 0.1147 |
| -0.3333 | 0.2647 |
| -0.1111 | 0.7642 |
| 0.1111 | 0.7642 |
| 0.3333 | 0.2647 |
| 0.5556 | 0.1147 |
| 0.7778 | 0.0620 |
| 1.0000 | 0.0385 |

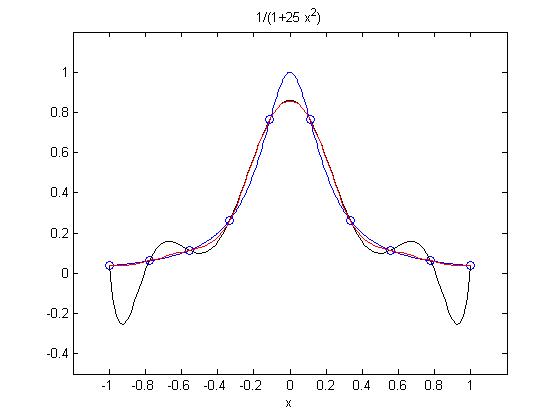
**(a)** Let *Pn*(*x*) denote the polynomial of degree *n*-1 that interpolates *f*(*x*) at *n* equally spaced points. Calculate and plot *Pn*(*x*) for the above data points. Investigate whether as *n* increases, *Pn*(*x*) converges to *f*(*x*).

**(b)** Change the distribution of the interpolation points so that they are not equally spaced. How does this affect convergence? Can you find a distribution so that *Pn*(*x*) → *f*(*x*) for all *x* in the interval?

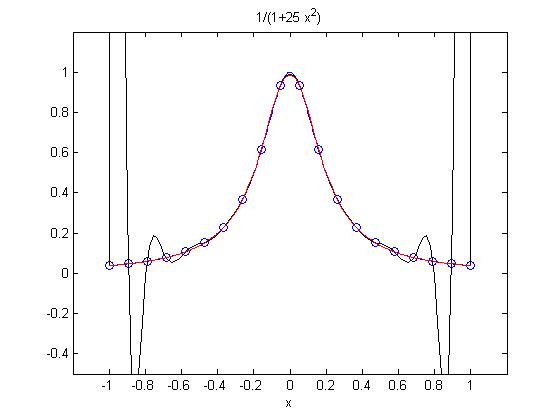
**Solutions b**

The sample solution m-file is called rungephe.m.

When you run it, the program pauses at each of the line with pause. You press any key to continue. I ran it and obtained the figure below:

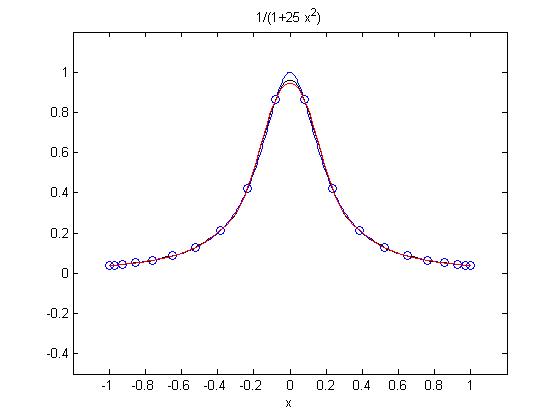


It is found that the degree-9 interpolating polynomial is oscillatory in regions near the sides (black curve above). Note that the polynomial has to pass through all the points (open circles) used in deriving it. The interpolation is not satisfactory for these ‘innocent looking’ points If I naively increase the number of points I use for obtaining a higher-degree polynomial, I obtain the following:



Here, I have used 20 points and so the interpolating is of degree-19. The problem at the sides has worsened. In fact, it is not getting better if I used more points. The large oscillations is an expected property of high-degree interpolating polynomials.

Like given in the lecture, the problem is absent for unevenly distributed points used for the interpolating. The optimal points are known as the Chebyshev points. I uncomment the part of the code which says ‘chebyshev points’ in the code above and rerun the code to obtain the following figure:



So now I am pretty happy because the 20 Chebyshev points (which you can see have higher concentration near the sides) allow me to get a good interpolating polynomial. When more Chebyshev points are used, I get convergence to the Runge function (figure is not shown here).

This exercise highlights the problem called Runge phenomenon where polynomial interpolation of equi-distanced points produces oscillatory behaviour on the edges of the interval. Splines (piece-wise) interpolation can resolve the problem. And if the polynomials are used for solving discretised equations, Chebyshev points could be used.

rungephe.m

% just an mfile to show runge phenomenon.

% martin king, 30 july 2008.

clear all;

close all;

ezplot('1./(1+25\*x.^2)',[-1 1]) %plotting the runge function

hold on;

axis([-1.2 1.2 -0.5 1.2])

pause;

x=linspace(-1,1,10); %setting equi-distanced x points and the corresponding y points

%chebyshev points

%i=[1:1:20];

%x=cos((2.\*i-1)/(2.\*length(i))\*pi);

%

y=1./(1+25\*x.^2);

plot(x,y,'bo')

pause;

p=polyfit(x,y,length(x)-1); %fit a polynomial to these points

xi=linspace(-1,1,100); %setting many x points for interpolation

yi=polyval(p,xi);

plot(xi,yi,'k-') %plotting the interpolated curve

pause;

% interpolation with splines

yi3=interp1(x,y,xi,'spline');

plot(xi,yi3,'r-')

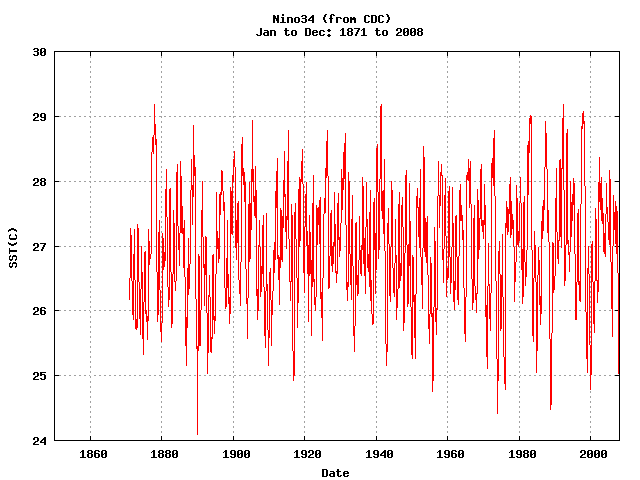
%

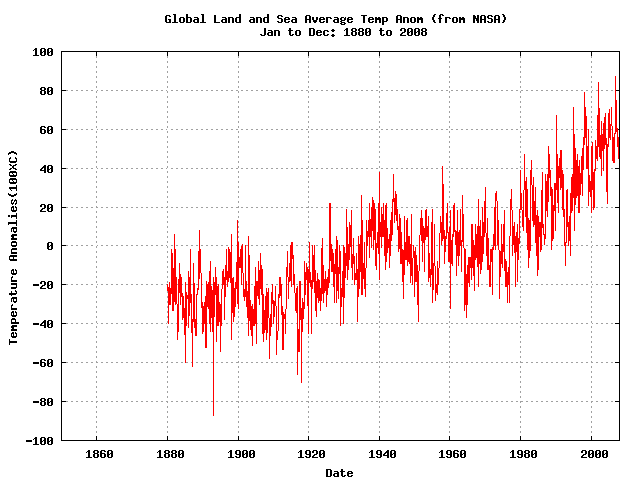
# Filters

This exercise is not well constructed. You are welcome to regard it as an example and read the suggested solutions to find out how I do it without attempting it yourself.

**Exercise**

This question requires you to perform some analysis on two climate time series. The first one is the Nino3.4 timeseries (monthly, from Jan 1871 to Dec 2007). Nino3.4 is the sea-surface temperature averaged over the area 5N-5S 170W-120W in the central-to-eastern Equatorial Pacific Ocean.

Nino3.4 is commonly used in the climate science as an indicator for the El Nino-Southern Oscillation phenomenon. The second time series is the global average land and sea surface temperature anomalies (monthly, from Jan 1880 to Dec 2007).



The two data files are provided for you. (They, and other climate indices, are also available from http://www.cdc.noaa.gov/gcos\_wgsp/Timeseries/).

You can carry out the following suggested analyses:

**(i)** Use the mydft.m and lowpass.m m-files provided to carry out low-pass filtering with a cut-off period of 12 months for both the nino3.4 and global surface temperature time series (i.e., we want to remove all fluctuations faster than, and including, the annual cycles).

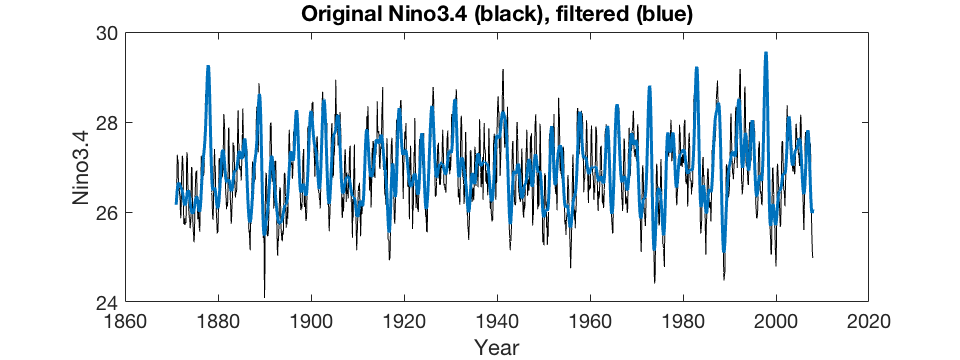
**(ii)** Use the mydft.m m-file provided to calculate the power spectra for the filtered Nino3.4 timeseries from (i). Is there any noticeable difference in the power spectra between calculations performed only on first half and the second half of the time series?

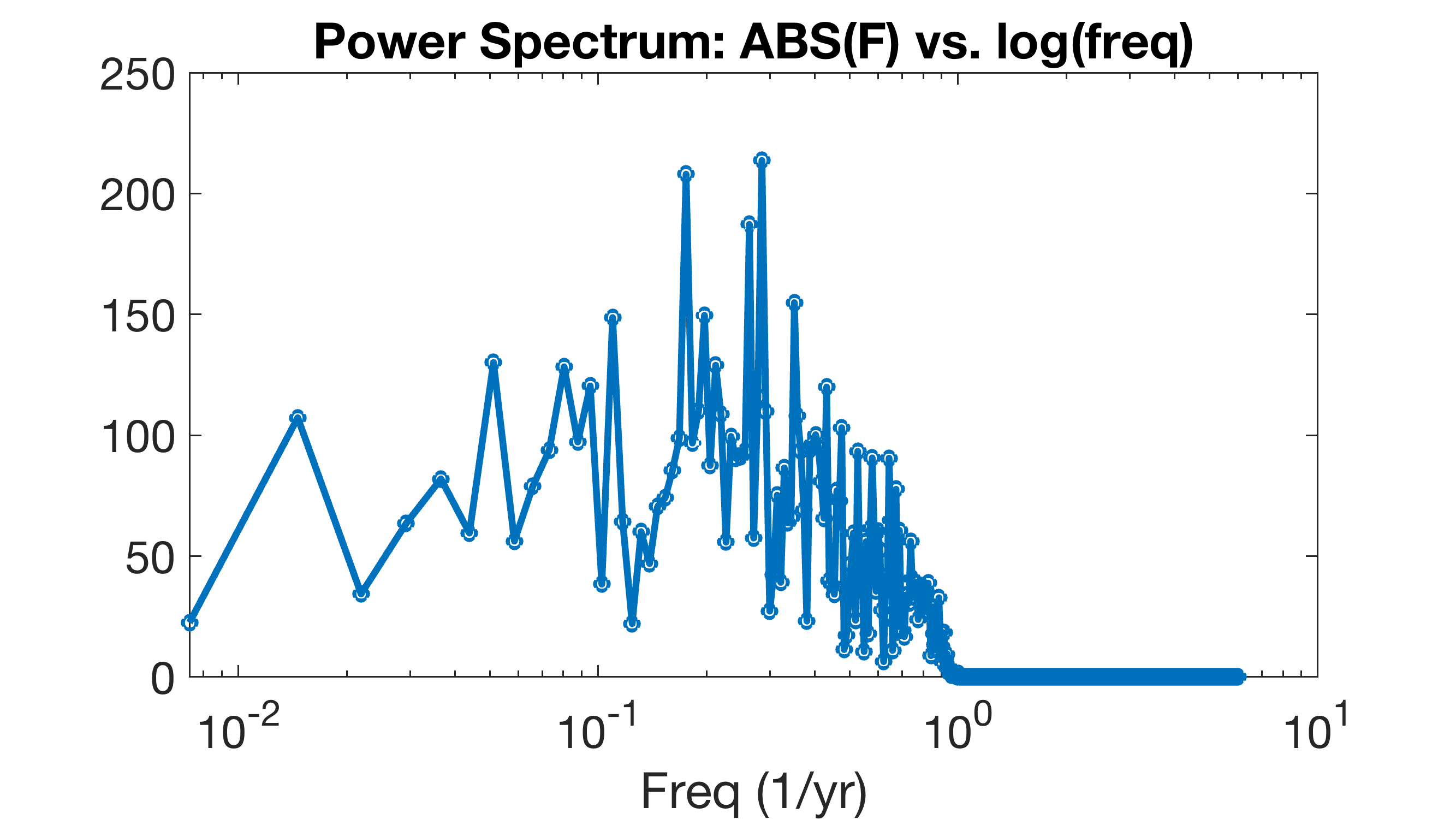
**(iii)** Calculate the correlation coefficient between the two low-pass filtered timeseries. Is there any noticeable change in the correlation coefficients from the first to the second half of the two time series?

Write down your observations from the analyses above and make some reasonable conclusions from your results.

**Solutions**

**(i) and (ii)** I coded filter\_solns.m to perform the task. The lowpass filter is applied in Lines 30 and 34. They call the mydft.m and lowpass.m provided. I used the cutoff frequency of 11.5/12 per year after the time interval of the time series (dt) is already defined as 1/12. The filtered time series Nino34f is 'fed' to mydft.m again for its Fourier Transform (Line 36) and then the power spectra for the filtered time series is plotted in 'figure(2)'. The following figures are produced:

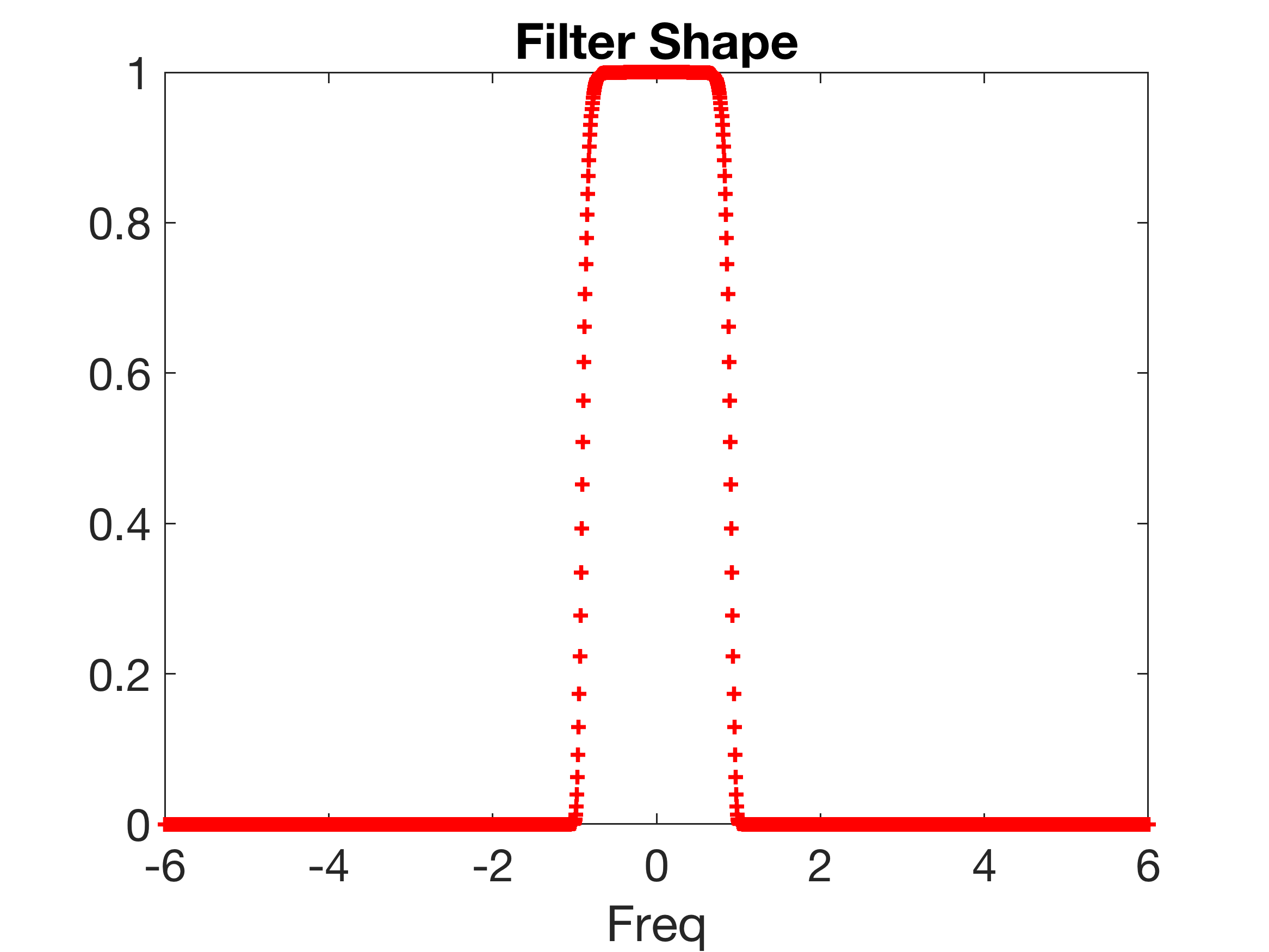




1/0.1754=5.7 yrs

1/0.285=3.5 yrs

1/0.2631=3.8 yrs



The power spectra is eliminated for frequencies greater than 1 per year, so we have probably done this correctly. We can also see that the Nino3.4 time series has cycles at around one per 3.5, 3.8, and 5.7 years. When I checked the power spectra calculated using the first and second halves of the available data, I found that the earlier part has peaked periods between 2.8 to 3.8 years, and the later part has peaked periods between 2.4 to 5.6 years. So it seems that the Nino3.4 spectra have changed, with more recent decades covering a broader range of periods as well as longer periods.

I modified the m-file slightly to also do the lowpass filtering of the global surface temperature (see filter\_lab9\_solns\_2.m). I used the last two lines of the two m-files to store the filtered time series data to nino34lowpass.txt and glbtsstlowpass.txt.

**(iii)** I cleared the workspace and loaded the filtered time series.

load nino34lowpass.txt

load glbtsstlowpass.txt

Since the nino3.4 index is shorter by 9 years, the correlation between them are calculated as:

[r,p,rlo,rup]=corrcoef(glbtsstlowpass(1:end),nino34lowpass(109:end)),

which results in r=[0.2, 0.29]. Therefore, it is found that the two time series have weak but statistically significant correlation. Calculated in similar way, the first halve of the available data results in r=[0.35, 0.47], while in the second halve r=[0.23, 0.36]. Therefore, it is possible that the two time series have stronger monthly correlation in the earlier part of the available data.

**Bonus codes**

I also provide highpass.m and bandpass.m. Here are examples of code snippets for how to use them from the main m-file.

For highpass filter:

%highpass, cutoff freq=11.5/12 per year:

% ie faster than 1 year period are included

nino34f=highpass(F,Fc,11.5/12);

and for bandpass filter:

%bandpass, freq\_low = 1/10 per year to freq\_high=11.5/12 per year:

% ie only between 10 years and 1 year periods are included.

nino34f=bandpass(F,Fc,1/10,11.5/12);

# Bibliography

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Kiusalaas, 2015, Numerical Methods in Engineering with MATLAB, 3rd Ed., Cambridge University Press.

Faires, Burden, 2003, Numerical Methods, 3rd Ed., Thomson, Brooks/Cole.

Moler, 2004, Numerical Computing with MATLAB, SIAM. (*the free electronic version and m-files are available at* [www.mathworks.com/moler](http://www.mathworks.com/moler))

Downey, 2021, Physical modeling in MATLAB, Green Tea Press, 4th Ed., (*a free textbook available at* <http://greenteapress.com/matlab>)

Evans, Blackledge, Yardley, 2000, Numerical Methods for Partial Differential Equations, Springer.

Griffiths, Dold, Silvester, 2015, Essential Partial Differential Equations: Analytical and Computational Aspects, Springer.

Larsson, Thomee, 2009, Partial Differential Equations with Numerical Methods, Springer.

Morton, Mayers, 2005, Numerical Solution of Partial Differential Equations, 2nd Ed., Cambridge University Press.

Smith, 1986, Numerical Solution of Partial Differential Equations: Finite Difference Methods, Oxford University Press.

**Additional Resources on the WWW**

**MATLAB Documentation:**

https://se.mathworks.com/help/matlab/

**Self-paced MATLAB Online Courses:**

<https://matlabacademy.mathworks.com/>

**MATLAB Central Repository:**

<http://www.mathworks.com/matlabcentral/>

**MITOCW Introduction to MATLAB:**

https://ocw.mit.edu/courses/electrical-engineering-and-computer-science/6-057-introduction-to-matlab-january-iap-2019/index.htm

**Mathworld's entry on 'Numerical Methods':**

http://mathworld.wolfram.com/topics/NumericalMethods.html

**Numerical Recipes:**

<http://numerical.recipes/>

**OCTAVE: A Free MATLAB-like Software:**

http://www.gnu.org/software/octave/

**OCTAVE Online:**

https://octave-online.net/

**Scilab: Another Open Source Numerical Computing Software**:

http://www.scilab.org/

THE END