

NUR - Hand-in Exercise 2

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Abstract

In this document, I present the results for the Hand-in Exercise 2 for the course Numerical Recipes for Astrophysics.

1 Satellite galaxies around a massive central

The exercise is done in the script `satellite.py`. The necessary explanations of the methods used are in the comments of the code. For the purpose of this task, we implement a Random Number Generator (RNG) as follows:

```
1 class CombinedRNG:
2     '''
3     Pseudo-random number generator combining the XOR-shift and Multiply-With-Carry (MWC)
4     methods.
5     We set the seed to the current time in milliseconds, as a source of variability
6     '''
7     def __init__(self, seed = None):
8         # If no seed is provided, use the current time in milliseconds
9         if seed is None:
10             seed = int(time.time() * 1000)
11         # Initialize the states for XOR-shift and MWC methods
12         self.xor_state = seed
13         self.mwc_state = seed
14
15     def xor_shift(self):
16         # XOR-shift method
17         # The seed is converted to a 64-bit unsigned integer
18         # By performing a shift of 21 bits to the left and XORing the result with the
19         # original state
20         self.xor_state ^= (self.xor_state << 21) & 0xFFFFFFFFFFFFFFFF
21         # Then, a shift of 35 bits to the right and again XOR with the previous result
22         self.xor_state ^= (self.xor_state >> 35)
23         # Finally, a shift of 4 bits to the left and XOR with the previous result
24         self.xor_state ^= (self.xor_state << 4) & 0xFFFFFFFFFFFFFFFF
25         return self.xor_state
26
27     def mwc(self):
28         # Multiply-With-Carry (MWC) method
29         # Set the multiplier
30         a = 4294957665
31         # Extract the upper and lower 32 bits of the state
32         mwc_upper = (self.mwc_state & 0xFFFFFFFF00000000) >> 32
33         mwc_lower = self.mwc_state & 0x00000000FFFFFFFF
34         # Calculate the next state by multiplying the lower 32 bits with the multiplier
35         # and adding the upper 32 bits
36         x_next = a * mwc_lower + mwc_upper
37         # Update the state using AND operation with a 64-bit mask
38         # Only the lower 64 bits are kept, the rest are set to zero
39         self.mwc_state = x_next & 0xFFFFFFFFFFFFFFFF
40         return x_next >> 32
41
42     def combined_rng(self):
43         # Combine the XOR-shift and MWC methods by XORing their outputs
44         return (self.xor_shift() ^ self.mwc()) & 0xFFFFFFFFFFFFFFFF
```

```

43
44     def uniform(self, low, high, num_samples):
45         # Generate a list of uniformly distributed random numbers
46         # By scaling the output of the combined RNG to the desired range
47         return [low + (high - low) * self.combined_rng() / 0xFFFFFFFFFFFFFFFF for _ in
48                 range(int(num_samples))]
49
50 # Create an instance of the combined RNG
51 rng = CombinedRNG()

```

satellite.py

For question (a), we do the following:

```

1  ## 1a)
2
3  # Define the parameters
4  a = 2.4
5  b = 0.25
6  c = 1.6
7  xmax = 5
8  Nsat = 100
9
10 def integrand(x):
11     # Take into account singularity at x=0
12     # Consider the volume element is 4*pi*x^2 (4*pi is taken outside)
13     if x == 0:
14         return 0
15     return Nsat * (x/b)**(a-3) * exp(-(x/b)**c) * x**2
16
17 # Implement a numerical integration
18 # Using the trapezoidal rule
19 def trapezoidal_rule(f, a, b, n):
20     '''
21     Inputs:
22     f: Function to integrate
23     a, b: Integration limits
24     n: Number of intervals
25     It works by approximating the integral of f(x) between a and b
26     by the sum of the areas of trapezoids formed by the function and the x-axis
27     '''
28     # Width of each trapezoid
29     h = (b - a) / n
30     # Initialize the result with the average of the function at the limits
31     result = 0.5 * (f(a) + f(b))
32     for i in range(1, n):
33         # Add the value of the function at each interior point
34         # Each of the function values corresponds to the height of a trapezoid
35         result += f(a + i * h)
36     # Multiply the sum by the width of the trapezoids
37     # This gives the total area under the curve
38     result *= h
39     return result
40
41
42 # Romberg integration
43 def romberg(f, a, b, n):
44     '''
45     Inputs:
46     f: Function to integrate
47     a, b: Integration limits
48     n: Number of rows in the Romberg table
49     '''
50     R = np.zeros((n, n))
51     # At each iteration, calculates an approximation of the integral
52     # using the trapezoidal rule with 2 ** i intervals
53     # This forms the first column of the Romberg table
54     for i in range(n):
55         R[i, 0] = trapezoidal_rule(f, a, b, 2 ** i)
56     # At each iteration, calculates an improved approximation of the integral
57     # using Richardson extrapolation

```

```

58 # Takes a weighted average of the current approximation and the previous one
59 # The weights are chosen to cancel out as much of the error as possible
60 # This fills the rest of the Romberg table
61 for j in range(1, n):
62     for k in range(j, n):
63         R[k, j] = R[k, j - 1] + (R[k, j - 1] - R[k - 1, j - 1]) / (4 ** j - 1)
64     return R[-1, -1] # Return the last element of the last row of the Romberg table
65
66
67 # Apply the method above to the integral
68 n = 10
69 m = 6
70 R = romberg(integrand, 0, xmax, m)
71 result = R * 4 * pi
72 A = Nsat / result
73 with open("normalization.txt", "w") as file:
74     file.write("The result of the numerical integration is: " + str(A))

```

satellite.py

The normalization factor A we obtain is the following:

```

1 The result of the numerical integration is: 9.05216643171278

```

normalization.txt

. For question (b), we generate 3D satellite positions that statistically follow the satellite profile $n(x)$. We sample them from the probability density function $p(x)$, using the inverse transform sampling. We do this with the following code:

```

1 ## 1b)
2
3 # Define the number density of galaxies
4 # This represents the number of galaxies per unit volume at a given radius x
5 def n(x):
6     return A * Nsat * (x/b)**(a-3) * exp(-(x/b)**c)
7
8 # Define the number of galaxies N(x) in a shell of radius x and thickness dx
9 # We compute N(x) by multiplying the number density n(x) by the volume element 4*pi*x^2
10 def N(x):
11     return 4 * np.pi * x**2 * n(x)
12
13 # Define the probability distribution function p(x) given that p(x)dx = N(x)dx / Nsat
14 # It gives the probability of finding a galaxy at a given radius x
15 def p_x(x):
16     return N(x) / Nsat
17
18
19 # Implement inverse transform sampling to sample from the distribution
20 def inverse_transform_sampling(pdf, n_samples, x_min, x_max):
21     x_values = np.linspace(x_min, x_max, 10000)
22     cdf_values = np.zeros_like(x_values)
23     # Calculate the cumulative distribution function using numerical integration (
24     # trapezoidal rule)
25     # The cdf at a point x is the integral of the pdf from x_min to x
26     for i in range(1, len(x_values)):
27         cdf_values[i] = cdf_values[i-1] + trapezoidal_rule(pdf, x_values[i-1], x_values[i], 1)
28     cdf_values /= cdf_values[-1] # Normalize to ensure CDF ranges from 0 to 1
29
30     # Generate random numbers uniformly distributed between 0 and 1
31     random_numbers = rng.uniform(0, 1, num_samples=n_samples)
32
33     # Apply the inverse CDF to the random numbers
34     # For each random number, find the corresponding value of x
35     # such that the CDF of x is equal to the random number
36     # This is done by finding the first x value for which the CDF is greater than the
37     # random number
38     # The corresponding x value is then stored as a sampled point
39     sampled_points = np.zeros(n_samples)

```

```

38     for i in range(n_samples):
39         for j in range(len(cdf_values)):
40             if random_numbers[i] < cdf_values[j]:
41                 sampled_points[i] = x_values[j]
42                 break
43
44     return sampled_points
45
46 # Generate 10,000 sampled points
47 n_samples = 10000
48 x_min, x_max = 10**-4, 5
49 h = (x_max - x_min) / n_samples
50 sampled_points = inverse_transform_sampling(p_x, n_samples, x_min, x_max)
51
52 # Calculate the bin edges
53 bin_edges = np.logspace(np.log10(x_min), np.log10(x_max), 21)
54
55 # Calculate histogram and divide each bin by its width
56 hist, _ = np.histogram(sampled_points, bins=bin_edges, density=False)
57 bin_widths = np.diff(bin_edges)
58 hist_density = hist / (bin_widths * n_samples / Nsat) # Correcting for the
59               normalization offset
60
61 # Plot the analytical function N(x) and the histogram of the sampled points on a log-log
62   scale
63 x_values = np.logspace(-4, np.log10(x_max), 20)
64 N_x_values = [N(x) for x in x_values]
65
66 plt.figure(figsize=(10, 6))
67 plt.loglog(x_values, N_x_values, label='Analytical N(x)', color='r')
68 plt.bar(bin_edges[:-1], hist_density, width=bin_widths, label='Sampled Points Histogram'
69         )
70 plt.xlabel('Relative Radius x')
71 plt.ylabel('Number of Galaxies N(x)')
72 plt.ylim(10**-3, 10**5)
73 plt.legend()
74 plt.title('Analytical N(x) vs Sampled Points Histogram')
75 plt.grid(True)
76 plt.savefig('my_solution_1b.png', dpi=600)
77
78 ## 1c)
79
80 # Define a selection method following these rules:
81 # Select each galaxy with same probability
82 # Not draw the same galaxy twice
83 # Not reject any drawn galaxy
84 def reservoir_selection(sampled_points, n_galaxies):
85     selected_galaxies = []
86     num_samples = len(sampled_points)
87     for i in range(num_samples):
88         # If we haven't selected n_galaxies yet, add the current galaxy
89         if i < n_galaxies:
90             selected_galaxies.append(sampled_points[i])
91         else:
92             # If we have already selected n_galaxies, decide whether to replace one of
93             them
94             # Generate a random index j between 0 and i
95             j = rng.combined_rng() % (i + 1)
96             # If j is less than n_galaxies, replace the galaxy at index j with the
97             current galaxy
98             if j < n_galaxies:
99                 selected_galaxies[j] = sampled_points[i]
100     return selected_galaxies
101
102 # Define a sorting method (quicksort)
103 # Takes an array and two indices: low (starting index of the array) and high (last index
104   )
105 def quick_sort(arr, low, high):
106     # If the low index is less than the high index, there are elements in the array to

```

```

102     be sorted
103     if low < high:
104         # partition function called to partition the array around a pivot.
105         # The pivot's final position in the sorted array is returned as pi
106         pi = partition(arr, low, high)
107
108         # quick_sort function recursively called for the parts of the array before pi
109         and after pi
110         quick_sort(arr, low, pi - 1)
111         quick_sort(arr, pi + 1, high)
112     return arr
113
114 # This function partitions the array around a pivot
115 def partition(arr, low, high):
116     # Choose the pivot as the median of the first, middle, and last elements
117     mid = low + (high - low) // 2
118     pivot = arr[mid]
119
120     # If the first element is greater than the middle element, swap them
121     if arr[low] > arr[mid]:
122         arr[low], arr[mid] = arr[mid], arr[low]
123
124     # If the middle element is greater than the last element, swap them
125     if arr[mid] > arr[high]:
126         arr[mid], arr[high] = arr[high], arr[mid]
127
128     # If the first element is greater than the middle element, swap them
129     if arr[low] > arr[mid]:
130         arr[low], arr[mid] = arr[mid], arr[low]
131
132     # Now the middle element is the median of the first, middle, and last elements
133     pivot = arr[mid]
134
135     # Swap the pivot with the high element
136     arr[mid], arr[high] = arr[high], arr[mid]
137
138     # i is set to one index less than the low index
139     i = (low - 1)
140     # The array is iterated from the low to the high index
141     for j in range(low, high):
142         # If the element is less than or equal to the pivot, the element is swapped with
143         the one at index i + 1
144         if arr[j] <= pivot:
145             i = i + 1
146             arr[i], arr[j] = arr[j], arr[i]
147
148     # After all elements have been checked, the pivot is swapped with the element at
149     index i + 1
150     # placing the pivot in its correct position in the sorted array
151     arr[i + 1], arr[high] = arr[high], arr[i + 1]

```

satellite.py

We compare the distribution of the sampled points with $N(x)$, the analytical function describing the number of galaxies in a shell of thickness dx at a given distance x . From Fig.1, we can see that the sampled points match the expected distribution. Nevertheless, we can see how the samples are not present in the first interval of x values. Inverse sampling works by transforming uniform random numbers into samples from the desired distribution, using the cumulative distribution function (CDF) of the PDF. The CDF maps values from the sample space to the interval $[0, 1]$, and its inverse can be used to map values from $[0, 1]$ back to the sample space. In the case of smaller values, a small range of uniform random numbers can map to a large range of small values in the sample space, leading to undersampling in these regions. This might explain why there is a discrepancy at smaller values in our histogram.

For question (c), we select 100 random galaxies from the ones sampled in point (b), using Reservoir method. It is a sampling algorithm that chooses a random sample, without replacement, of k items from a population of unknown size n in a single pass over the items. The 100 drawn galaxies are ordered from smallest to largest radius. We do this as follow:

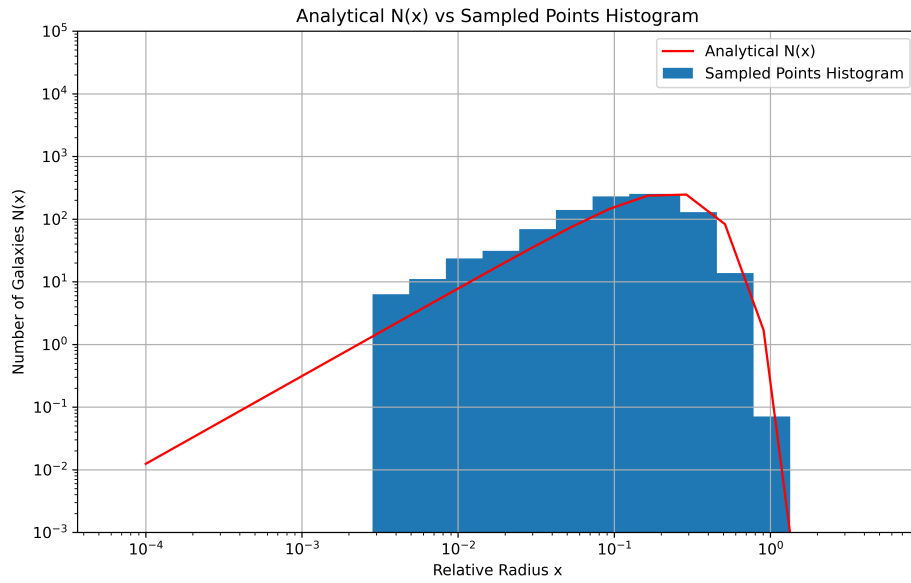


Figure 1: Plot in logarithmic scale showing $N(x)$, function of the number of galaxies at given distance x , and the histogram of the 10000 sampled points. The sampled points match the analytical distribution, but there is a problem of undersampling in the region of smaller radii.

```

1     return (i + 1)
2
3     # Select 100 galaxies
4     n_galaxies = 100
5     selected_galaxies = reservoir_selection(sampled_points, n_galaxies)
6
7     # Sort the galaxies using quicksort from smallest to higher radius
8     sorted_galaxies = quick_sort(selected_galaxies)
9
10    # Plot the cumulative number of the chosen galaxies
11    fig1c, ax = plt.subplots()
12    ax.plot(sorted_galaxies, np.arange(n_galaxies))
13    ax.set(xscale='log', xlabel='Relative radius',
14           ylabel='Cumulative number of galaxies',
15           xlim=(x_min, x_max), ylim=(0, n_galaxies))
16    plt.savefig('my_solution_1c.png', dpi=600)
17
18    ## 1d)
19
20    # Define Ridders' method for numerical differentiation
21    def ridders_method(f, x, h, m, tol=1e-10):
22        # Compute first approx with central difference for high h
23        D = [(f(x + h) - f(x - h)) / (2 * h)]
24
25        # Decrease h by a factor of 2 and calculate a new approximation. Repeat until you
26        # have m approximations
27        for i in range(1, m):
28            h /= 2
29            D.append((f(x + h) - f(x - h)) / (2 * h))
30            for j in range(i):
31                D[j] = (4**(j+1) * D[j+1] - D[j]) / (4**(j+1) - 1)
32
33        # Terminate when the improvement over previous best approximation is smaller than
34        # the target error or if error grows
35        best_approximation = None
36        for i in range(1, len(D)):
37            if abs(D[i] - D[i-1]) < tol:

```

```

36         best_approximation = D[i]
37         break
38
39     # Terminate early if the error grows and return the best approximation from
40     before that point.
41     if abs(D[i] - D[i-1]) > abs(D[i-1]):
42         best_approximation = D[i-1]
43         break
44
45     if not best_approximation:
46         print("Tolerance not reached or error grew significantly. Consider increasing m.
47 ")

```

satellite.py

The number of galaxies within a radius x are shown in Fig. 2. The radii selected are starting from higher values because of the previous undersampling. The increasing trend makes sense as the number of galaxies within a given radius should increase as the radius increases.

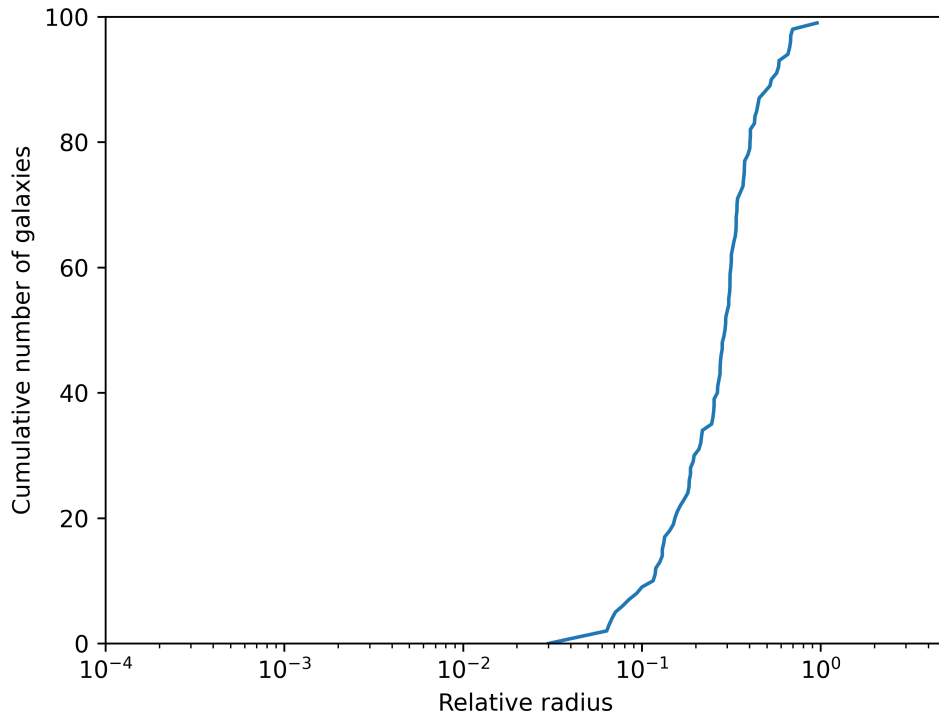


Figure 2: Plot showing the cumulative number of galaxies randomly selected with respect to the relative radius x . The number of galaxies progressively increases with larger distances.

For question (d), we compute the derivative of the function $n(x)$ using both the Ridders' method and the analytical procedure. The code is below:

```

1
2     # Now the middle element is the median of the first , middle, and last elements
3     pivot = arr[mid]
4
5     # Swap the pivot with the high element
6     arr[mid], arr[high] = arr[high], arr[mid]
7
8     # i is set to one index less than the low index
9     i = (low - 1)
10    # The array is iterated from the low to the high index

```

```

11     for j in range(low , high):
12         # If the element is less than or equal to the pivot , the element is swapped with
13         the one at index i + 1
14         if arr[j] <= pivot:
15             i = i + 1
16             arr[i], arr[j] = arr[j], arr[i]
17
18     # After all elements have been checked, the pivot is swapped with the element at
19     index i + 1
20     # placing the pivot in its correct position in the sorted array
21     arr[i + 1], arr[high] = arr[high], arr[i + 1]
22     # Return the final position of the pivot
23     return (i + 1)
24
25 # Select 100 galaxies
26 n_galaxies = 100
27 selected_galaxies = reservoir_selection(sampled_points , n_galaxies)
28
29 # Sort the galaxies using quicksort from smallest to higher radius
30 sorted_galaxies = quick_sort(selected_galaxies)
31
32 # Plot the cumulative number of the chosen galaxies
33 fig1c , ax = plt.subplots()
34 ax.plot(sorted_galaxies , np.arange(n_galaxies))
35 ax.set(xscale='log' , xlabel='Relative radius' ,
36        ylabel='Cumulative number of galaxies' ,
37        xlim=(x_min , x_max) , ylim=(0 , n_galaxies))
38 plt.savefig('my_solution_1c.png' , dpi=600)
39
40 ## 1d)
41
42 # Define Ridders' method for numerical differentiation
43 def ridders_method(f , x , h , m , tol=1e-10):
44     # Compute first approx with central difference for high h
45     D = [(f(x + h) - f(x - h)) / (2 * h)]
46
47     # Decrease h by a factor of 2 and calculate a new approximation. Repeat until you
48     have m approximations

```

satellite.py

The obtained results can be seen in:

```

1 Analytical result: -0.615624912159
2 Numerical result: -0.615624525721

```

derivative.txt

2 Heating and cooling in HII regions

The code relative to this exercise is in the script

```

1 import time
2 import numpy as np
3
4 '''
5 This script contains the code for Exercise 2.
6 '''
7
8 ## 2a)
9
10 # Define the Newton-Raphson method
11 def newton_raphson(f , df , x0 , tol=1e-6 , max_iter=100):
12     '''
13     Find the root of the function f(x) = 0
14     Takes the function f , its derivative df , an initial guess x0 , and optional arguments
15     tol and max_iter
16     Improve iteratively the estimate of the root

```



```

16     until the change in x is less than the specified tolerance
17     or the maximum number of iterations is reached
18     '''
19     x = x0
20     for i in range(max_iter):
21         x_new = x - f(x) / df(x)
22         if abs(x - x_new) < tol:
23             return x_new, i+1
24         x = x_new
25     return None, i+1 # Ensure that a tuple is returned
26
27 # Define the constants
28 k = 1.38e-16 # erg/K
29 aB = 2e-13 # cm^3 / s
30 Z = 0.015 # metallicity
31 Tc = 1e4 * Z**2 # stellar temperature in K
32 psi = 0.929
33 A = 5e-10 # erg
34 epsilon_CR = 1e-15 # s^-1
35
36 # Define the equilibrium function
37 def equilibrium_1(T):
38     term1 = psi*Tc*k
39     term2 = (0.684 - 0.0416 * np.log(T/(1e4 * Z*Z)))*T*k
40     return term1 - term2
41
42 # Define the derivative of the equilibrium function
43 def derivative_1(T):
44     term1 = -psi*Tc*k / (T*T) # wrong??? why does it work
45     term2 = (0.684 - 0.0416 * np.log(T/(1e4 * Z*Z))) + 0.0416 * T / (T * np.log(10))
46     return term1 - term2*k
47
48 # Measure the time taken and find the root of the equilibrium function
49 # The initial guess is the midpoint of the bracket [1, 10^7]
50 start_time = time.time()
51 T_eq, num_steps = newton_raphson(equilibrium_1, derivative_1, (1 + 10**7) / 2)
52 end_time = time.time()
53
54 # Print the result
55 with open('2a.txt', 'w') as f:
56     if T_eq is not None:
57         f.write(f"The equilibrium temperature is {T_eq:.2f} K.\n")
58         f.write(f"The Newton-Raphson method found the root in {num_steps} steps.\n")
59         f.write(f"The time taken was {end_time - start_time:.10f} seconds.\n")
60     else:
61         f.write("The Newton-Raphson method did not converge.\n")
62
63
64 ## 2b)
65
66 # Define the equilibrium function
67 def equilibrium_2(T, n_e):
68     T4 = T / 1e4
69     term1 = 0.54 * T4**0.37 * aB * n_e * n_e * k * T
70     term2 = A * n_e * epsilon_CR
71     term3 = 8.9e-26 * n_e * T4
72     return term1 - term2 - term3
73
74 # Bisection method
75 def bisection_method(f, a, b, n_e, tol=1e-10, max_iter=100):
76     if f(a, n_e) * f(b, n_e) >= 0:
77         print("Bisection method fails.")
78         return None, None, None
79
80     num_steps = 0
81     start_time = time.time()
82     while (b - a) / 2 > tol and num_steps < max_iter:
83         c = (a + b) / 2
84         if f(c, n_e) == 0:
85             end_time = time.time()

```

```

86         return c, num_steps + 1, end_time - start_time
87
88     if f(c, n_e) * f(a, n_e) < 0:
89         b = c
90     else:
91         a = c
92     num_steps += 1
93
94     end_time = time.time()
95     return (a + b) / 2, num_steps, end_time - start_time
96
97 # Secant method
98 def secant_method(f, x0, x1, n_e, tol=1e-10, max_iter=100):
99     num_steps = 0
100     start_time = time.time()
101     while num_steps < max_iter:
102         x_new = x1 - f(x1, n_e) * (x1 - x0) / (f(x1, n_e) - f(x0, n_e))
103         if abs(x_new - x1) < tol:
104             end_time = time.time()
105             return x_new, num_steps + 1, end_time - start_time
106
107         x0, x1 = x1, x_new
108         num_steps += 1
109
110     end_time = time.time()
111     return x_new, num_steps, end_time - start_time
112
113 # Densities to consider
114 densities = [1e-4, 1, 1e4]
115
116 for n_e in densities:
117     # Initial interval for bisection method
118     a_bisection = 1
119     b_bisection = 1e7
120     T_eq_bisection, num_steps_bisection, time_taken_bisection = bisection_method(
121         equilibrium_2, a_bisection, b_bisection, n_e)
122
123     if T_eq_bisection is None:
124         # If bisection method fails, use brent method
125         a_brent = 1
126         b_brent = 1e7
127         T_eq_secant, num_steps_brent, time_taken_brent = secant_method(equilibrium_2,
128             a_brent, b_brent, n_e)
129
130         # Print the result
131         with open('2b.txt', 'a') as f: # open in append mode
132             if T_eq_secant is not None:
133                 f.write(f"For n_e = {n_e} cm^-3, the equilibrium temperature is {
134                     T_eq_secant:.2f} K (using secant method).\n")
135                 f.write(f"The secant method found the root in {num_steps_brent} steps.\n
136                     ")
137                 f.write(f"The time taken was {time_taken_brent:.6f} seconds.\n\n")
138             else:
139                 if T_eq_bisection is not None:
140                     f.write(f"For n_e = {n_e} cm^-3, the equilibrium temperature is {
141                         T_eq_bisection:.2f} K (using bisection method).\n")
142                     f.write(f"The bisection method found the root in {
143                         num_steps_bisection} steps.\n")
144                     f.write(f"The time taken was {time_taken_bisection:.6f} seconds.\n\n")
145                 else:
146                     f.write(f"For n_e = {n_e} cm^-3, both bisection and secant methods
147                         failed to converge.\n\n")

```

heating.py

The output of point (a) is in

```

1 The equilibrium temperature is 3.12 K.
2 The Newton-Raphson method found the root in 18 steps.
3 The time taken was 0.0000500679 seconds.

```

2a.txt

The output of (b) is in

```
1 For n_e = 0.0001 cm^-3, the equilibrium temperature is -56186.85-3.86j K (using secant
  method).
2 The secant method found the root in 100 steps.
3 The time taken was 0.000326 seconds.
4 For n_e = 0.0001 cm^-3, the equilibrium temperature is -56186.85-3.86j K (using secant
  method).
5 The secant method found the root in 100 steps.
6 The time taken was 0.000111 seconds.
7 For n_e = 0.0001 cm^-3, the equilibrium temperature is -56186.85-3.86j K (using secant
  method).
8 The secant method found the root in 100 steps.
9 The time taken was 0.000110 seconds.
```

2b.txt