## Community Atlas of Stellar Streams

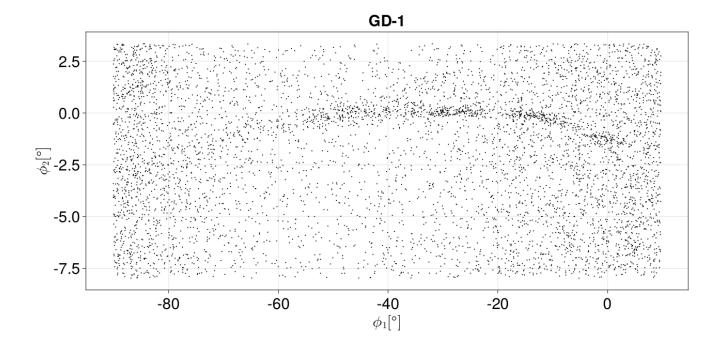
## Searching the streams at first order

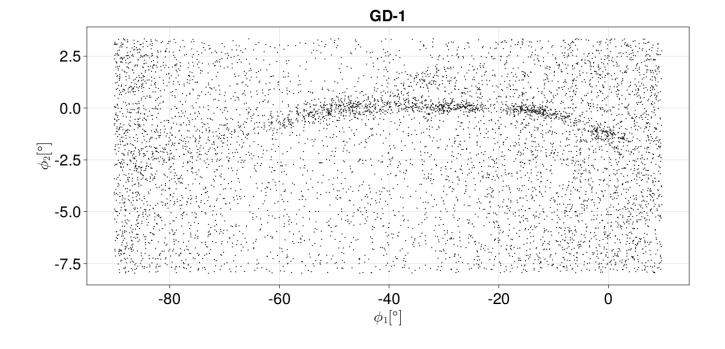
Here we show the result of making approximate cuts in proper motion (PM) and color-magnitude (CM) for the five streams using Gaia DR3. We also show the case of GD-1 with DR2 for comparison. The PM cut is guided by the Galstream track:  $\mu^{(t)}$ . The filtering is done by keeping only stars that satify:  $\forall \phi_1: |\mu_1^{(\star)} - \mu_1^{(t)}(\phi_1^{(\star)})| < 0.7$  and  $\forall \phi_1: |\mu_2^{(\star)} - \mu_2^{(t)}(\phi_1^{(\star)})| < 0.7$  where  $\mu_{1,2}$  are the PMs not corrected by the solar reflex-motion, and  $\star$  denotes the Gaia data. The process is performed by the function "example" given below. There you can see the input parameters used for age, metal,  $\sigma_c$  (width in color) and  $\sigma$  (width in PM).

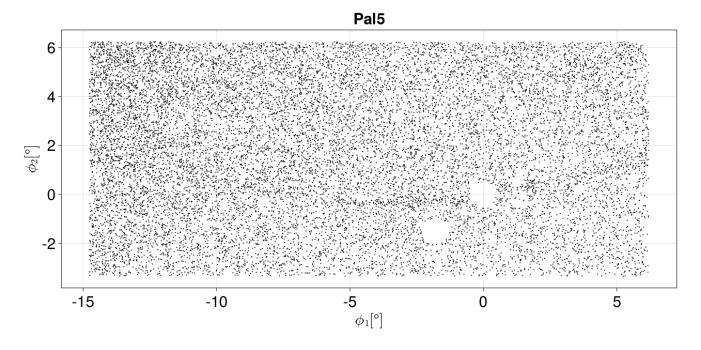
```
1 begin
        name_s = ["GD-1", "GD-1", "Pal5", "Jhelum", "Fjorm-M68","PS1-A"]
 2
 3
        name_t = ["GD-1-PB18", "GD-1-PB18", "Pal5-PW19", "Jhelum-b-B19", "M68-P19",
   "PS1-A-B16"]
               = [12.0, 12.0, 12.0, 12.0, 11.2, 12.0]*10^9 #yr
 4
        age
        metal = [-1.5, -1.5, -1.4, -1.2, -2.2, -1.7]
 5
        filters = fill("UBVRIplus",length(name_s))
 6
               = fill("DR3",length(name_s))
 7
        dr[1] = "DR2"
 8
 9
        tol_curation = [0.3, 0.3, 1.0] # tolerances in \mu_{\alpha} \times \cos \delta, \mu_{\alpha} \delta, \Pi.
        col_bounds = (-1.0, 4.0)
10
        box_{\mu} = [[-14, -10.], [-4., -2.]]
11
12
        \sigma_c = 1
13
        \sigma = 0.7
14
        file_orig, file_corr, file_phot, file_iso, file_filt, file_plot =
15
   name_files_all(dr, name_s, age, metal)
16
        nothing
17 end
```

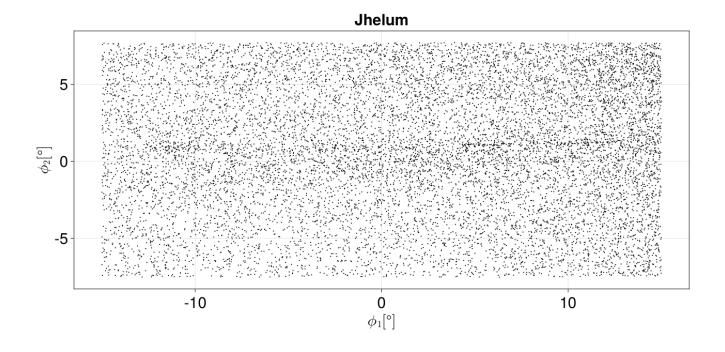
Now lets reopen the filtered files and display the plots without the tracks so as not to cheat the eye.

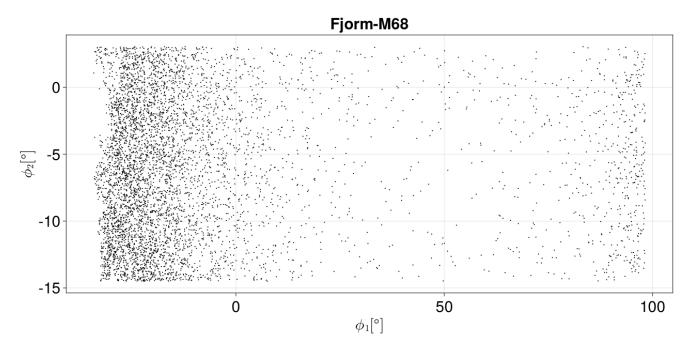
```
1 begin
2    figs = Vector{Figure}(undef, length(name_s))
3    for i ∈ eachindex(name_s)
4         df = DataFrame(CSV.File(file_filt[i], delim=",", ignorerepeated=true))
5         figs[i]=plot_scatter_on_sky_self_frame(name_s[i], df, file_plot[i])
6    end
7 end
```

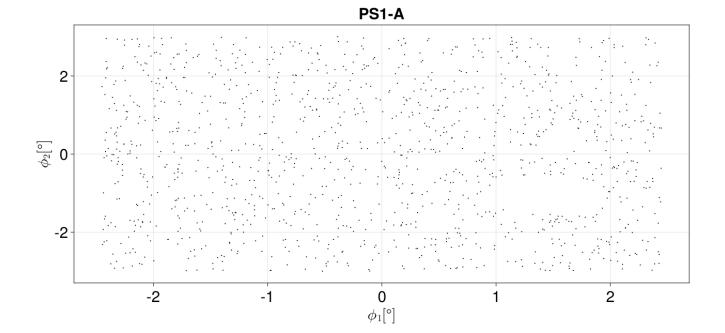








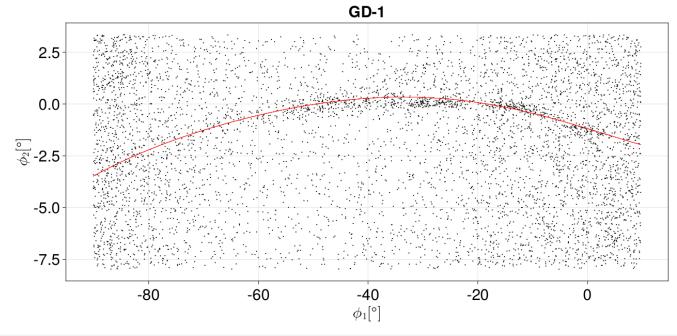




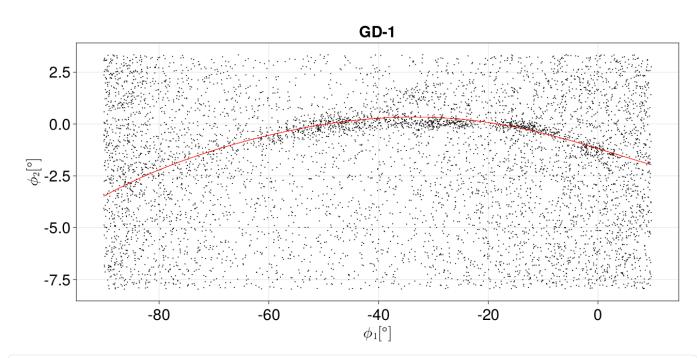
```
🖣 cats.jl — Pluto.jl
```

```
WARNING: AstropyDeprecationWarning: The matrix_product function is deprecat ②
ed and may be removed in a future version.
       Use @ instead. [gala.coordinates.greatcircle]
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```

Now lets plot together with the tracks.



## 1 <u>figs\_t</u>[1]



1 <u>figs\_t</u>[2]