

Multiphase flow through porous media

Liquid phase

Flux through a porous medium is defined as:

$$f_l = -K \frac{d\psi}{dz}$$

where f_l is liquid flux density [$\text{kg m}^{-2} \text{s}^{-1}$], ψ is the matric potential [J kg^{-1}] (note: $\psi = gh$), and K is the hydraulic conductivity [kg s m^{-3}].¹

When combined with the mass conservation equation, the result yields the transient vertical (1D) Richards (1931) equation:

$$\rho_l \frac{\partial \theta}{\partial t} = \frac{\partial}{\partial z} \left[K(\psi) \left(\frac{\partial \psi}{\partial z} + g \right) \right],$$

which is often expressed in its so-called ψ -based form (liquid flow only):

$$\rho_l C(\psi) \frac{\partial \psi}{\partial t} = \frac{\partial}{\partial z} \left[K(\psi) \left(\frac{\partial \psi}{\partial z} + g \right) \right],$$

where ρ_l is the density of the liquid [kg m^{-3}], g is the acceleration to to gravity ($\approx 9.80665 \text{ m s}^{-2}$ or $\text{J kg}^{-1} \text{ m}^{-1}$). The capacity [kg J] term is solved using the Campbell (1974) model:

$$C(\psi) = \frac{d\theta}{d\psi} = \frac{-\theta}{b\psi}$$

and conductivity:

$$K(\psi) = \begin{cases} K_s \left(\frac{\psi_e}{\psi} \right)^{2+3/b} & \text{if } \psi_m \leq \psi_e \\ K_s & \text{otherwise} \end{cases}$$

where K_s is the saturated conductivity [kg s m^{-3}], b is a shape parameter and ψ_e is the air-entry potential [J kg^{-1}]. Furthermore, water content (θ) [$\text{m}^3 \text{ m}^{-3}$] and ψ are assumed related by the power-law relationship:

$$\theta(\psi) = \begin{cases} \theta_s \left(\frac{\psi}{\psi_e} \right)^{-1/b} & \text{if } \psi_m \leq \psi_e \\ \theta_s & \text{otherwise} \end{cases}$$

or conversely,

$$\psi(\theta) = \begin{cases} \psi_e \left(\frac{\theta}{\theta_s} \right)^{-b} & \text{if } \theta \leq \theta_s \\ \psi_e & \text{otherwise} \end{cases}$$

where θ_s is the water content at saturation.

¹note: $K_{\text{kg} \cdot \text{s} \cdot \text{m}^{-3}} = \frac{\rho_l}{g} K_{\text{m} \cdot \text{s}^{-1}}$

Vapour phase within the soil matrix

Vapour flux density [$\text{kg m}^{-2} \text{s}^{-1}$] (i.e., molecular diffusion flux of water vapour in the air) is defined using Fick's Law:

$$f_v = -\rho_a D(\theta) \frac{dq}{dz},$$

where ρ_a is the density of air, [kg m^{-3}], $D(\theta)$ is the effective coefficient of turbulent diffusion of water vapour in pore-space air, often given the form attributed to Penman (1940):

$$D(\theta) = D_a \eta (\theta_s - \theta),$$

where D_a is coefficient of molecular diffusion of water vapour in air $\approx 2.12 \times 10^{-5} \text{ m}^2 \text{s}^{-1}$, and η is a dimensionless coefficient characterizing the turbulent diffusion in porous media, $\eta \approx 0.66$ (Penman, 1940). Note that the term $(\theta_s - \theta)$ is included correct for the volume of gas-filled porosity, where vapour flux is occurring. q is the specific humidity (moisture content) of air [kg kg^{-1}]:

$$q = \frac{\varepsilon e}{P} = w q^*(T, P),$$

where e and P are the vapour and total pressures, respectively [$\text{kg m}^{-1} \text{s}^{-2}$], ε is the ratio of the molecular weight of water to the molecular weight of air (≈ 0.622), w is the relative humidity [-], and q^* is the saturated specific humidity determined using the August-Roche-Magnus-Tetens relationship:

$$q^*(T, P) = \frac{0.38}{P} \exp \left(\frac{17.625T}{T + 243.04} \right),$$

where P is in [kPa] and T is in [$^{\circ}\text{C}$].

Within a porous medium, assuming isobaric (i.e., $P_{\text{pores}} \approx P_a$) and isothermal (constant temperature) conditions, the specific humidity gradient can be related to pore-space relative humidity (w_p) and soil temperature (T_s) by:

$$\frac{dq}{dz} = q^*(T_s, P) \frac{dw_p}{dz}.$$

By further assuming that the liquid and vapour phases are in equilibrium (i.e., have equal potential and the partial pressure of water has reached its vapour pressure, e), according to the ideal gas law, pore-space relative humidity can be determined from liquid potential by:

$$w_p = \frac{e}{e^*(T_s)} = \frac{q}{q^*(T_s, P)} = \exp \frac{\omega_w \psi}{RT_s},$$

where the gas constant $R = 8.3143 \text{ J mol}^{-1} \text{ K}^{-1}$, the molecular mass of water $\omega_w = 0.01802 \text{ kg mol}^{-1}$, and e^* is the saturated vapour pressure at the temperature of the soil surface T_s [K]. Next, using the chain rule,

$$\frac{dw_p}{dz} = \frac{dw_p}{d\psi} \frac{d\psi}{dz} = w_p \frac{\omega_w}{RT_s} \frac{d\psi}{dz},$$

Combining the above equations, results in:

$$f_v = -K_v \frac{d\psi}{dz},$$

where

$$K_v = \frac{\omega_w \eta \rho_a D_a q}{RT_s} (\theta_s - \theta),$$

By combining the flux of water vapour, the Richards equation can be re-stated as (Novák, 2012):

$$\rho_l \frac{\partial \theta}{\partial t} = \frac{\partial}{\partial z} \left[K(\psi) \left(\frac{\partial \psi}{\partial z} + g \right) + \rho_a D(\theta) \frac{\partial q}{\partial z} \right].$$

and further:

$$\rho_l \frac{\partial \theta}{\partial t} = \frac{\partial}{\partial z} \left[[K(\psi) + K_v(\theta, q)] \frac{\partial \psi}{\partial z} + gK(\psi) \right].$$

Atmospheric exchange

At the soil surface–atmosphere interface, the flux density to the atmosphere remains (Novák, 2012):

$$f_v = \rho_a D \frac{dq}{dz}$$

Assuming that vertical fluxes do not change in the vicinity of the evaporating soil surface, integrating the above equation from the soil surface (z_s) and some effective height above the surface (z_a) yields:

$$-\int_{q_s}^{q_a} dq = \frac{f_v}{\rho_a} \int_{z_s}^{z_a} \frac{dz}{D(z)},$$

re-arranging:

$$f_v = \frac{1}{\int_{z_s}^{z_a} \frac{dz}{D(z)}} \rho_a (q_s - q_a).$$

Next, by defining a water vapour turbulent transport coefficient (k_v) [m s^{-1}]:

$$k_v = \frac{1}{\int_{z_s}^{z_a} \frac{dz}{D(z)}},$$

vapour flux density to the atmosphere from the gas-filled pores can be written as:

$$\begin{aligned} f_{v,g} &= \rho_a k_v (q_s - q_a) \\ &= \rho_a k_v [q_s - w_a q^*(T_a, P_a)], \end{aligned}$$

and

$$f_{v,l} = \rho_a k_v (q^*(T_s, P_a) - q_a),$$

where w_a is the atmospheric relative humidity, T_a is air temperature [K], and P_a is air pressure [$\text{Pa} = \text{kg m}^{-1} \text{s}^{-2}$]. (Note that for water, the units for flux density [$\text{kg m}^{-2} \text{s}^{-1}$] is equivalent to [mm s^{-1}].) The water vapour turbulent transport coefficient can be related to wind speed ($u(z)$) [m s^{-1}] using boundary-layer theory:

$$k_v \approx \frac{\kappa^2 u(z)}{\left[\ln \left(\frac{z}{z_0} \right) + \frac{\beta}{L_*} z \right]^2} \approx \frac{\kappa^2 u(z)}{\ln \left(\frac{z-d_e}{z_0} \right)^2},$$

where κ is the von-Kármán constant (≈ 0.4), z_0 is the roughness length [m], β and L_* are the Monin-Obukhov (1954) coefficient of atmospheric stability and characteristic length, respectively, and d_e is the so-called zero-plane displacement height [m]. Alternatively, the Penman (1948)/Penman-Monteith (1965) relationship to aerodynamic resistance (r_a), soil surface resistance (r_s) and canopy resistance (r_c) [s m^{-1}] can be used, where:

$$k_v = \frac{1}{r_a + r_s + r_c}.$$

Lastly, the effective mass flux evaporating from the soil surface must account for the portion of the soil surface area exposed to gas-filled pores and the area of direct liquid exposure:

$$\begin{aligned} f_e &= (\theta_s - \theta) f_{v,g} + \theta f_{v,l} \\ &= \rho_a k_v [\theta_s (q_s - q_a) + \theta (q_s^* - q_s)] \end{aligned}$$

Numerical solution to the Richards equation

A finite-difference solution to the 1D Richards equation is solved using a cell-centered finite volume solution scheme following Bittelli et.al. (2015). The water balance in cell i is:

$$\frac{\rho_l (\theta_i^{j+1} - \theta_i^j) (z_{i+1} - z_{i-1})}{2\Delta t} = \frac{K_i (\psi_{i+1} - \psi_i)}{z_{i+1} - z_i} - \frac{K_{i-1} (\psi_i - \psi_{i-1})}{z_i - z_{i-1}} + u_i$$

or in its ψ -based form:

$$\frac{\rho_l C_i (\psi_i^{j+1} - \psi_i^j) (z_{i+1} - z_{i-1})}{2\Delta t} = \frac{K_i (\psi_{i+1} - \psi_i)}{z_{i+1} - z_i} - \frac{K_{i-1} (\psi_i - \psi_{i-1})}{z_i - z_{i-1}} + u_i$$

the source term u_i could include extraction by roots, evaporation, etc. In most cases, for vertical simulations, the source term represents gravitational flux, where:

$$u_i = g (K_{i-1} - K_i).$$

The solution to the above equation is done implicitly, by solving a system equations. Transforming the above to:

$$\begin{aligned}\varrho(i) &= \rho_l C_i \frac{(z_{i+1} - z_{i-1})}{2\Delta t} \\ k(i) &= \frac{K_i}{z_{i+1} - z_i} \\ c(i) &= a(i+1) = k(i) \\ b(i) &= k(i-1) + k(i) + \varrho(i) \\ d(i) &= \varrho(i)\psi_i^j + u(i)\end{aligned}$$

the finite-difference scheme for 4 layers is put in matrix form:

$$\begin{bmatrix} b(1) & c(1) & 0 & 0 \\ a(2) & b(2) & c(2) & 0 \\ 0 & a(3) & b(3) & c(3) \\ 0 & 0 & a(4) & b(4) \end{bmatrix} \begin{bmatrix} \psi^{j+1}(1) \\ \psi^{j+1}(2) \\ \psi^{j+1}(3) \\ \psi^{j+1}(4) \end{bmatrix} = \begin{bmatrix} d(1) \\ d(2) \\ d(3) \\ d(4) \end{bmatrix}$$

Newton-Raphson transformation

A finite-difference solution to the 1D Richards equation is solved using the Newton-Raphson solution scheme following Bittelli et.al. (2015). The water balance in cell i is:

$$\begin{aligned}F_i &= \frac{\overline{K}_i (\psi_{i+1} - \psi_i)}{z_{i+1} - z_i} - \frac{\overline{K}_{i-1} (\psi_i - \psi_{i-1})}{z_i - z_{i-1}} \\ &\quad - g (\overline{K}_i - \overline{K}_{i-1}) - \frac{\rho_l}{2\Delta t} (\theta_i^{k+1} - \theta_i^k) (z_{i+1} - z_{i-1}),\end{aligned}$$

where F is the mass balance residual and \overline{K} is the mean elemental hydraulic conductivity. Applying integral transform methods (Bittelli et.al., 2015), \overline{K} is defined by:

$$\overline{K}_i = \frac{K_{i+1}\psi_{i+1} - K_i\psi_i}{(1 + 3/b)(\psi_{i+1} - \psi_i)}.$$

The solution to the above equation is then considered a minimization problem, whereby Newton-Raphson:

$$\frac{\partial F}{\partial \psi} \cdot (\psi^{k+1} - \psi^k) = -F \rightarrow 0$$

or in a matrix-form for a 3-layer profile is:

$$\begin{bmatrix} \frac{\partial F_1}{\partial \psi_1} & \frac{\partial F_1}{\partial \psi_2} & \frac{\partial F_1}{\partial \psi_3} \\ \frac{\partial F_2}{\partial \psi_1} & \frac{\partial F_2}{\partial \psi_2} & \frac{\partial F_2}{\partial \psi_3} \\ \frac{\partial F_3}{\partial \psi_1} & \frac{\partial F_3}{\partial \psi_2} & \frac{\partial F_3}{\partial \psi_3} \end{bmatrix} \begin{bmatrix} \psi_1^{k+1} - \psi_1^k \\ \psi_2^{k+1} - \psi_2^k \\ \psi_3^{k+1} - \psi_3^k \end{bmatrix} = \begin{bmatrix} -F_1 \\ -F_2 \\ -F_3 \end{bmatrix}$$

where k indicates the k^{th} iteration of the minimization attempt. The mass-balance function F is solved by:

$$F_i = f_i - f_{i-1} + u_i - u_{i-1} - \frac{\rho_l}{2\Delta t} (\theta_i^{k+1} - \theta_i^k) (z_{i+1} - z_{i-1}),$$

where

$$f_i = -\frac{\bar{K}_i (\psi_{i+1} - \psi_i)}{z_{i+1} - z_i} = -\frac{K_{i+1}\psi_{i+1} - K_i\psi_i}{(1 + 3/b)(z_{i+1} - z_i)}$$

and letting

$$\begin{aligned} u_i &= -gK_i \\ \varrho_i &= -\rho_l \theta_i \frac{z_{i+1} - z_{i-1}}{2\Delta t}, \end{aligned}$$

then substituting yields:

$$\begin{aligned} F_i &= \frac{K_i\psi_i - K_{i-1}\psi_{i-1}}{(1 + 3/b)(z_i - z_{i-1})} - \frac{K_{i+1}\psi_{i+1} - K_i\psi_i}{(1 + 3/b)(z_{i+1} - z_i)} \\ &\quad - g(K_i - K_{i-1}) - \frac{\rho_l}{2\Delta t} (\theta_i^{k+1} - \theta_i^k) (z_{i+1} - z_{i-1}). \end{aligned}$$

Differentiating F_i with respect to ψ_i , ψ_{i-1} and ψ_{i+1} yields:

$$\begin{aligned} f'_i &= \frac{\partial f_i}{\partial \psi_i} = \frac{K_i}{z_{i+1} - z_i} \\ u'_i &= \frac{\partial u_i}{\partial \psi_i} = \frac{gK_i}{\psi_i} (2 + 3/b) \\ \varrho'_i &= \frac{\partial \varrho_i}{\partial \psi_i} = \rho_l \frac{\theta_i}{b\psi_i} \frac{z_{i+1} - z_{i-1}}{2\Delta t} \end{aligned}$$

and thus

$$\begin{aligned} \frac{\partial F_i}{\partial \psi_i} &= \frac{K_i}{z_{i+1} - z_i} + \frac{K_{i-1}}{z_i - z_{i-1}} + u'_i + \varrho'_i \\ \frac{\partial F_i}{\partial \psi_{i-1}} &= \frac{-K_{i-1}}{z_i - z_{i-1}} - u'_{i-1} \\ \frac{\partial F_i}{\partial \psi_{i+1}} &= \frac{-K_i}{z_{i+1} - z_i} \end{aligned}$$

Solving the Newton-Raphson system of equations is then accomplished like the linear methods described above.

Boundary conditions

Boundary conditions for the top of the vertical profile can either be a constant flux (e.g., evaporation) or constant potential:

- For constant flux, the source is added to F_1 ;
- For a constant potential, ψ_1^{j+1} is known and is set at the start of the time step (since the Newton-Raphson method computes changes in ψ to bring F to zero, F_1 and $\frac{\partial F_1}{\partial \psi_i}$ are set to zero before solving; therefore the value of ψ_1 will remain constant throughout the time step. For infiltration, set $\psi_1 = \psi_e$.

At the bottom of the profile, either a constant potential (i.e., water table) or a free drainage condition can be specified using a ghost below the profile that never becomes part of the solution. Free drainage is set by:

$$\begin{aligned}\psi_{n+1} &= \psi_n \\ \theta_{n+1} &= \theta_n \\ K_{n+1} &= K_n\end{aligned}$$

after each outer iteration, where n is the number of finite difference cells being solved.

References

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