

# Time limits for measurement of gravitational waves with dynamical Casimir effect in solid-state detectors

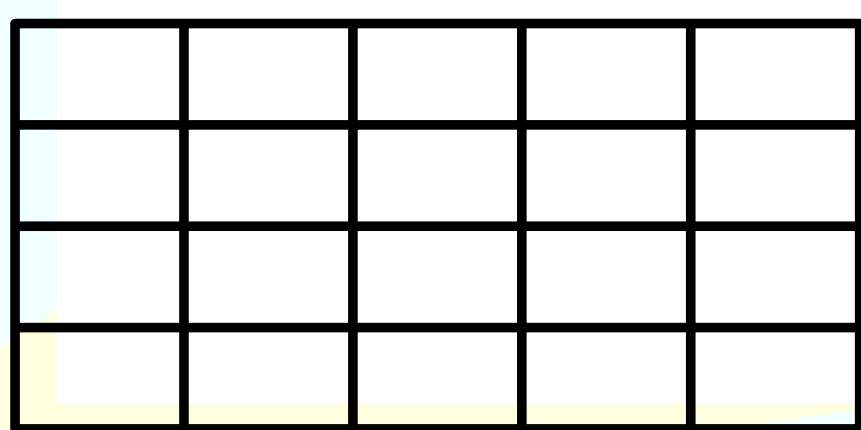
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A distributed **on-chip** resonator:  
a planar microwave resonator  
or a surface-acoustic wave resonator

## Can a gravitational wave change a resonator state?

Space-time metric periodically varies in two perpendicular directions

linearly-polarized gravitational wave



Some interesting sources

100 kHz

primordial black holes and exotic compact objects  
axion decay  
superradiance

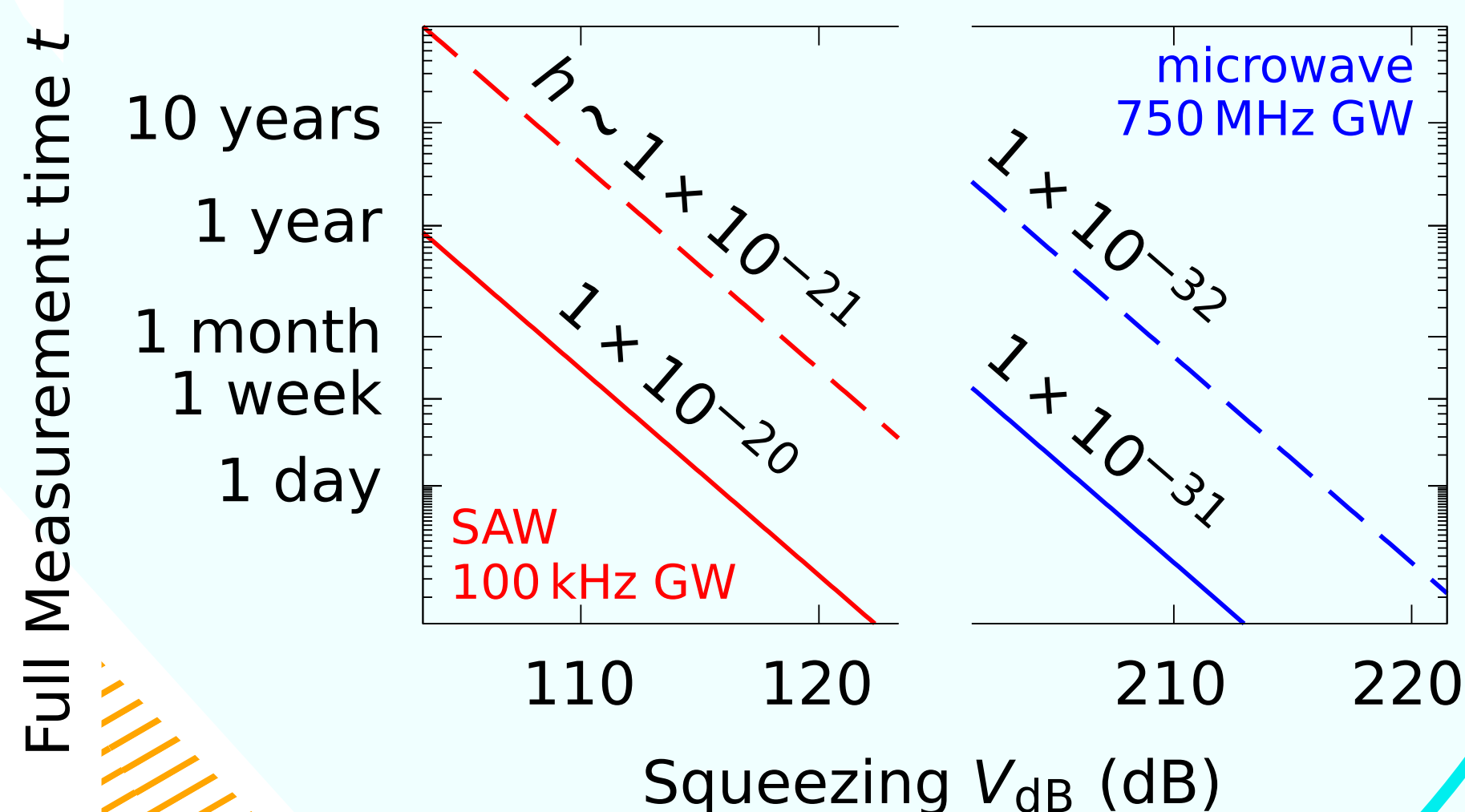
750 MHz

phase transitions in the early Universe  
oscillons

## Many mutually-incoherent measurements

$$t \sim \left( \frac{\omega_0/2\pi}{\text{MHz}} \right)^{-1} 10^{-13 - \lg(\Delta\epsilon^2) - \lg QD - \frac{1}{5} V_{\text{dB}}} \text{ years}$$

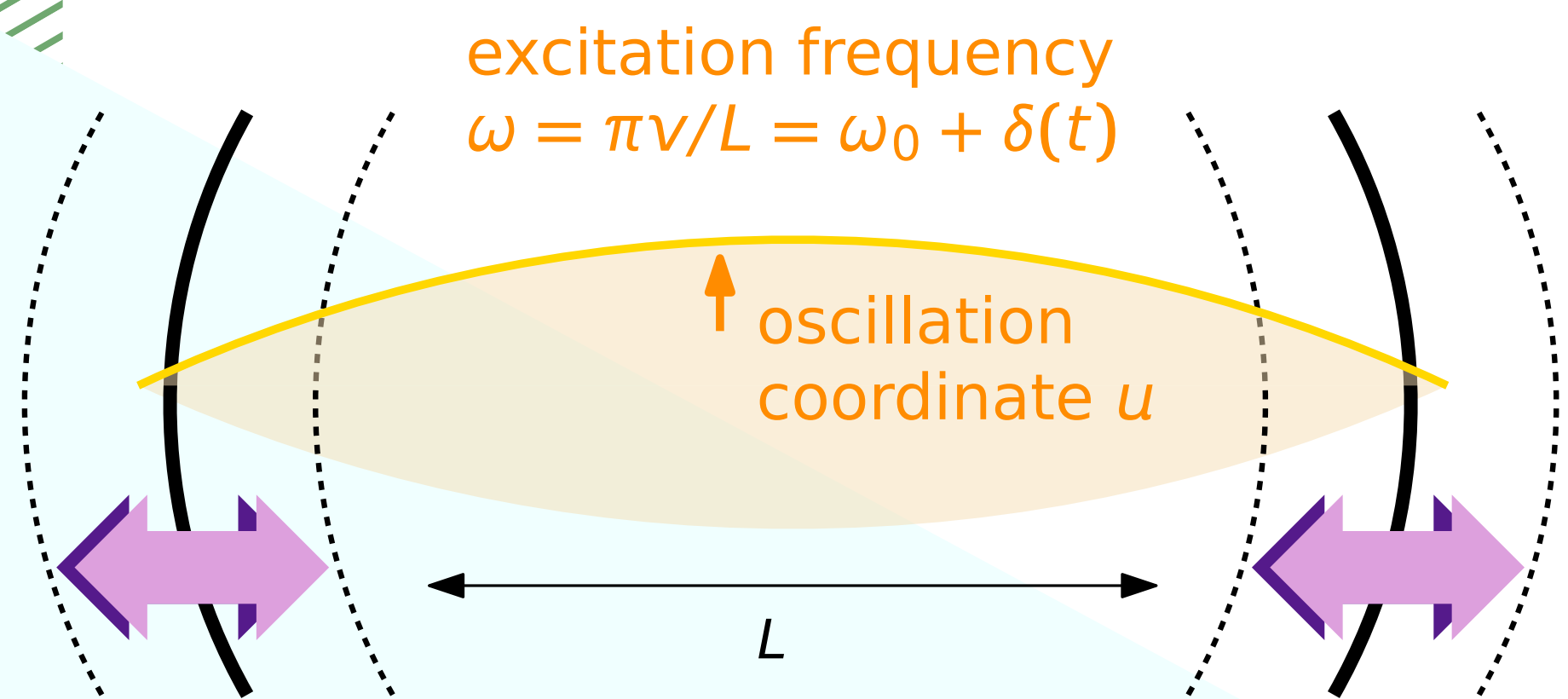
$D = 100$  devices operate simultaneously.  
Each measurement bin ten is times shorter than the coherence time  $Q/\omega$ , where  $Q = 10^6$ .



Such squeezing may be gained already this millenium. Currently,  $V_{\text{dB}} \lesssim 10$  dB.

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This can be done even sooner—if a robust quantum error-correction is ever reached.



Resonator length varies as  
 $L = L_0(1 + h_+)$

Metric perturbation  
 $h_+ = \epsilon \sin \Omega t$

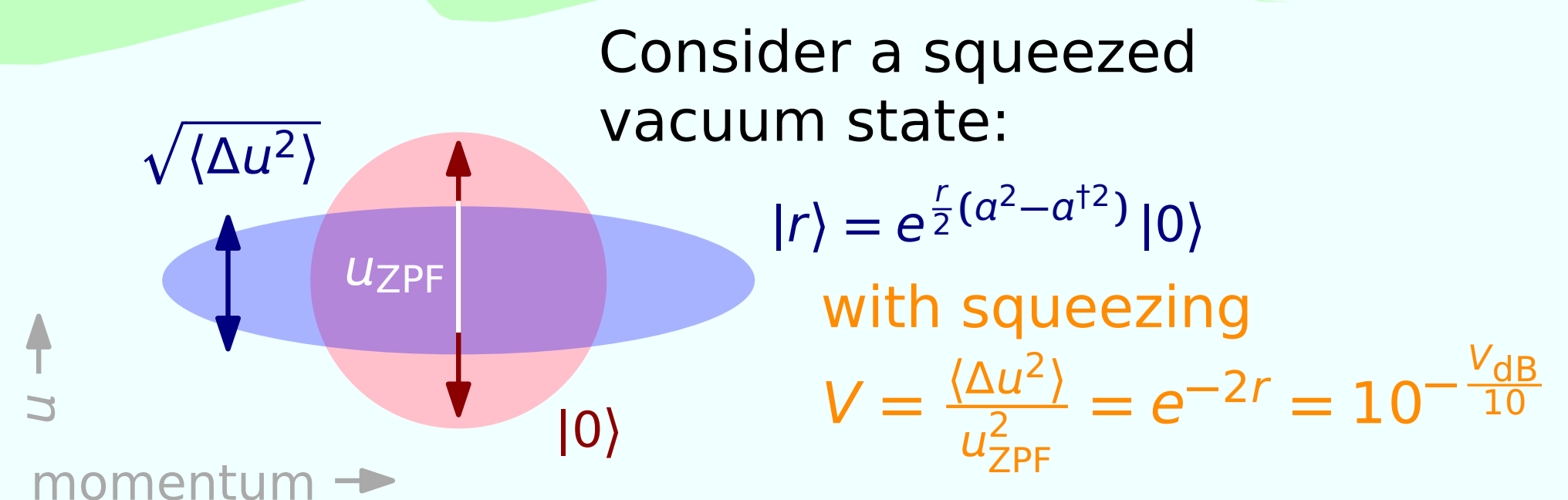
## Dynamical Casimir effect

At  $\Omega = \omega_0/2$ , fluctuations are parametrically amplified—if the resonator length varies vigorously

$\epsilon \sim 10^{-20}$   
@  $\Omega = 100$  kHz  
 $\rightarrow \langle n \rangle = 1$  phonon in  
 $t \sim 1/\Omega\epsilon^2 \sim 10^{27}$  years

Gravitational waves are far too weak near Earth to excite photons or phonons

## Still, they alter a quantum state. Can that be measured?



Consider a squeezed vacuum state:

$$|r\rangle = e^{\frac{r}{2}(a^2 - a^2)} |0\rangle$$

with squeezing

$$V = \frac{\langle \Delta u^2 \rangle}{u_{\text{ZPF}}^2} = e^{-2r} = 10^{-\frac{V_{\text{dB}}}{10}}$$

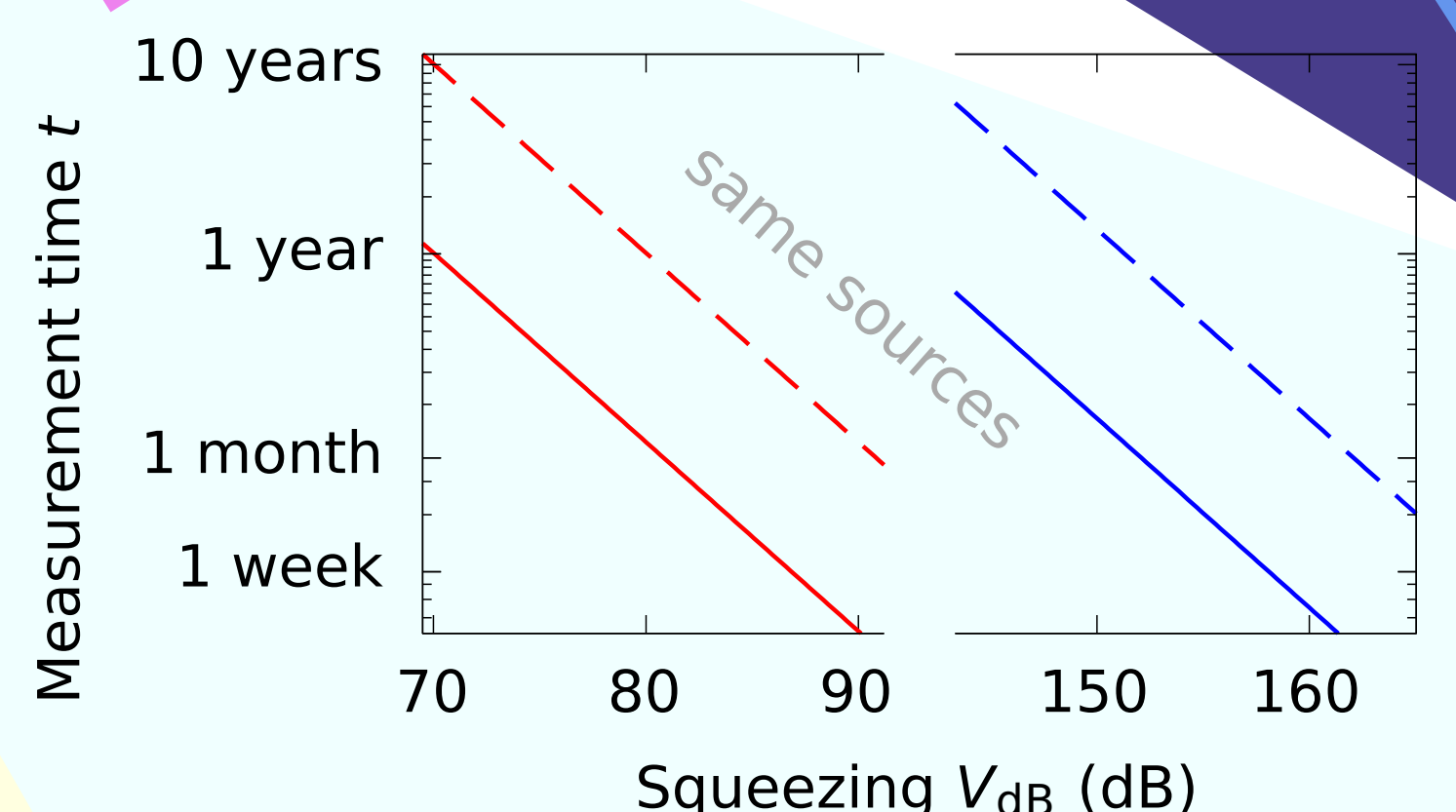
$H_\epsilon(t) = 2\omega_0^2 t^2 (1 + \sinh^4 r)$  information encoded by a wave

$N = \frac{1}{\langle \Delta\epsilon^2 \rangle H_\epsilon(t_{\text{bin}})}$  measurements—each  $t_{\text{bin}}$  long—

yield at least  $\sqrt{\langle \Delta\epsilon^2 \rangle}$  error in the metric magnitude  $\epsilon$

## Infinite coherence limit

$N = 1$  measurement with  $D = 1$  device



$$t \sim \left( \frac{\omega_0/2\pi}{\text{MHz}} \right)^{-1} 10^{-14 - \frac{1}{2} \lg(\Delta\epsilon^2) - \frac{1}{10} V_{\text{dB}}} \text{ years}$$