

Implementation of a Total Power Radiometer in Software Defined Radios

by

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DEDICATION

I would like to dedicate this thesis to my wife Jennifer who has stood by me through this long journey towards my Masters. I would also like to thank my parents without whose support I would not have been able to complete this work.

I would also like to thank my friends and family for their continued support and constant encouragement.

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ABSTRACT

A software defined radio is defined as a communication system where components of a communication system that are typically done in hardware are now done in software. The result is a highly flexible communication system that can adapt to changes to the system based on requests by the user or due to conditions in the radio frequency channel. Software Defined Radios (SDRs) have been used for a variety of applications, mostly in the area of communications. However, they have not been widely applied to remote sensing applications such as a radiometer. A SDR based radiometer offers a very flexible and robust system more so than a more traditional radiometer which has fixed hardware and usually little room for flexibility and adaptability based on the needs of the radiometer application or due to changes in the radiometers environment. The price of SDRs have also been steadily dropping making implementing them into radiometers a more attractive option compared to using traditional RF hardware. In this thesis we will look into the feasibility and theory of using an off the shelf SDR hardware platform for a radiometer application. In addition, we will look into how we can use the GNURadio software to create a total power radiometer within software and how this software can be used to make easy changes to the functionality of the radiometer. Finally we will look at the preliminary results obtained from laboratory tests of the SDR radiometer system.

CHAPTER 1. INTRODUCTION

This thesis explores using software defined radio (SDR) technology to develop a radiometer that has performance on par with that of traditional radiometer. In addition to demonstrating a SDR-based radiometer can achieve similar performance, in terms of sensitivity and stability, to a traditional radiometer, it is shown that the inherent flexibility of SDR technology allows implementing functionality beyond what a traditional radiometer typically provides. Furthermore, by reducing cost and increasing flexibility, SDR technology may become an attractive path for broadening the accessibility of radiometry to the general research community.

Motivation. Remote sensing refers to recording, observing and perceiving objects or events that are far away (i.e. remote)[Weng (2012)]. Since the object of interest is remote, we cannot physically interact with it using local measurement methods such as placing sensors or probes on the object. Remote sensing can be accomplished in a variety of ways and the following lists the basic approaches:

1. Visible light: Photography and Photogrammetry,
2. Thermal, Far infrared: Thermal Infrared Radiometry
3. Laser distance measurement: Lidar,
4. Radio Frequency (RF): Radiometry.

Each of these methods has an associated set of pros and cons, thus remote sensing often uses a combination of methods to paint a complete picture of an object. This thesis focuses on the apparatus used in radiometry to capture RF signals, called a radiometer.

Radiometry can be broken in to two methods; active or passive. An active system is one in which a radio frequency signal or pulse is generated and transmitted to the object of interest.

The reflection of this RF signal, or in some cases lack of reflection, gives us information about the object. A passive system is one in which no RF signal or pulse is generated. Instead, this type of radiometer simply listens to the RF energy that is naturally generated by the object or that may be reflected from another source, such as the Sun. Radiometers have been focused on Earth for such purposes as helping scientists better understand its water cycle by monitoring ocean salinity [Hardy et al. (1974)] and soil moisture [Liu et al. (2013)]. Radiometers such as these are already in service on satellites such as the Soil Moisture and Ocean Salinity (SMOS) satellite [McMullan et al. (2008)] launched by the European Space Agency (ESA). Additional applications of radiometry include: assessing vegetation health and observing celestial objects[Ulaby and Long (2014)].

Problem Statement. While radiometers have proved to be an excellent tool for remote sensing and have been used in research applications for over fifty years, they have not made it into wide spread use. This is due to the fact that many traditional radiometers have the following hurdles:

1. They are expensive,
2. They require advanced knowledge to implement and use,
3. They are typically built and designed for a custom application.

This thesis aims to help address each of these hurdles as follows. The use of commercial off the shelf (COTS) parts and solutions will help reduce cost and the need for a custom design. Software will be used to define key components of the radiometer that have traditionally been implemented in hardware and to ease adaptation to multiple applications. A user friendly graphical user interface (GUI) will help reduce the advanced knowledge required to implement and use the radiometer.

Contributions. The primary contributions of this thesis are:

1. Development of a software defined radio-based radiometer
2. Python based scripts for analyzing data generated by the radiometer

3. A Radio Frequency Interference (RFI) mitigation technique implemented using software defined radio technology

Organization. The remainder of this thesis is organized as follows. Chapter 2 gives a discussion of related works. Chapter 3 gives background on traditional radiometers and on software defined radios. Chapter 4 presents details on how software defined radio technology was used to implement a radiometer. Chapter 5 describes the experimentation setup used to verify and evaluate the operation of the implemented radiometer. Chapter 6 examines the results obtained from our performance evaluation experiments. Chapter 7 concludes this thesis and outlines avenues of future work.

CHAPTER 2. RELATED WORKS

Three areas closely related to the work in this thesis are:

1. Digital radiometers,
2. Software defined radio based radiometers,
3. Radio frequency interference mitigation.

Digital radiometers digitize some part of the signal or the output from the power detection stage of the radiometer. This is similar to a software defined radio radiometer except a software defined radio radiometer digitizes the complete radio frequency (RF) signal. The software defined radio based radiometers discussed is close to the work in this thesis, but do not use commercial off the shelf components or readily available software such as GNURadio to develop the radiometer in software. The radio frequency interference mitigation discussed uses radiometers or digital radiometers that do not collect the entire signal information and therefore use different methods to detect and mitigate the offending signal. The work in this thesis will cover a different method for RFI due to both frequency and amplitude information being preserved when the signal is digitized.

2.1 Digital Radiometers

A digital radiometer replaces a portion or all of the radiometer components with digital components[Ruf and Gross (2010)]. The key component in a digital radiometer is the analog to digital converters that convert the analog information to a digital signal that can be processed. There are two types of digital radiometers:

1. Hybrid radiometer,

2. Direct digital sampled radiometer.

A hybrid radiometer is a radiometer that uses some traditional analog components but digitizes a certain portion of the radiometer information [Skou and Vine (2006)]. This is often the analog voltage output from the diode of a square law detector. However, it can be other stages such as a digital correlator [Villarino et al. (2002)][Wang et al. (2012)] or a different method for calibration such as digital reference averaging [Bremer (1979)].

Both types of digital radiometers are designed to retain only a portion of the information contained in the RF signal. This information is the total power contained within the signal which is the primary purpose of a radiometer. As it will be discussed later though, retaining the full signal allows for more in-depth digital signal analysis.

2.1.1 Direct Digital Sampled Radiometer

A direct digital sampled radiometer samples the incoming RF signal directly and then uses digital signal techniques to extract the total power from the signal.

Iowa State University owns a 1.4 GHz, dual polarization, correlating radiometer. This radiometer was built at the University of Michigan and put into service at ISU in 2006. The ISU radiometer is unique in that it is one of the few direct sampling radiometers in useErbas et al. (2006). This radiometer takes the RF signal and using analog components amplifies and filters the signal, and then sends it directly to an analog to digital converter. At the time the ISU radiometer was built an analog to digital converter (ADC) that is able to sample accurately at 1.4 GHz was expensive and hard to come. However, the ISU radiometer does not sample at 2.8 GHz or above, instead it samples at 1.4 GHz. The ISU radiometer is able to do this because it is only interested in power information which we can extract with an under-sampled signal. The I/Q information from the (ADC) is then sent to a Field Programmable Gate Array (FPGA) which then processes the data. The ISU radiometer is also a correlating radiometer that looks at both the vertical polarization (V-Pol) and the horizontal polarization (H-Pol) and then correlates this information[Fischman (2001)].

2.2 Software Defined Radio Based Radiometers

With the advent of software defined radios becoming widely available, there has been other works with using software defined radios for radiometry. The related works presented here outline work that has contributed to using software defined radios in developing radiometers.

Shirleys Bay Radio Astronomy Consortium. Radiometers used in radio astronomy [Tiuri (1964)] is nothing new and has been postulated for some time [Ohm and Snell (1963)]. The Shirleys Bay Radio Astronomy Consortium (SBRAC) located in Smiths Falls, Ontario is an organization that is restoring a radio telescope used for radiometry in astronomy. Recently the SBRAC is using software defined radios attached to their eighteen meter radius dish to obtain information by looking at the hydrogen line at 1420.4058 MHz[Leech and Ocame (2007)]. The person in charge of this facility, Marcus Leech, contributed software to the GNURadio specifically for radio astronomy applications. It was this software branch that was used as the base for the GNURadio program that was used in this thesis. Marcus Leech continue to contribute to the GNURadio community and continues to provide support for these functions as well[Leech (2006)].

While parts of the GNURadio program was derived from the work from Marcus Leech, additional features were added such as filtering an offending signal, offending signal detection and built in noise generator. Additional work was done on refining some elements of the GUI interface including adding a waterfall display which is covered in more detail in this thesis.

Grand Valley State University. Another example of a software defined radio used as a radiometer is from the University of Illinois and Grand Valley State University built to listen to emissions from Jupiter[Behnke et al. (2013)]. This software defined radio was built using an Analog Devices AD9460 and a Xilinx Spartan-3E-500 FPGA to build the SDR itself. A RF front end was also built to filter and amplify the signal coming into the SDR. Finally, this group also used GNURadio to interface to and talk to the SDR and used both Python and wxGUI to build a working interface. The students reported that the SDR radiometer worked well and was able to do so at a low price point.

The Grand Valley State University radiometer used GNURadio to talk and interface to

the SDR which is the same as the radiometer used in this thesis. However in this case the radiometer was built by the students using the Spartan-3E-500 FPGA and an Analog Devices analog to digital converter. The goal of this thesis however, is to use off the shelf components to make it more accessible to users.

2.3 Radio Frequency Interference (RFI) Mitigation

Radio Frequency Interference (RFI) is a RF signal that is unwanted or undesirable and causes either reduced performance or masks the information that we want. Radio Frequency Interference (RFI) is a common problem with nearly all radiometers because they are highly sensitive receivers so even small signals can cause large problems for the radiometer. It is for this reason why certain frequency bands have been designated for radiometer use by the international community. However, not all entities abide by this and continue to transmit on these protected frequencies. This problem is often exasperated with satellite radiometers since those radiometers are exposed to more potential transmitters from the ground. Current missions such as the Soil Moisture Ocean Salinity (SMOS) satellite has had numerous issues with RFI [Kerr (2012)] that skews the data that they collect and in some cases made the data unusable for soil moisture measurements [Richaume (2012)].

Therefore there is ongoing work in detecting and mitigating this problem to help reduce the problem caused by RFI. Most radiometers do not retain any frequency information of the signal and so different methods have been developed to detect RFI [Forte et al. (2013)] by use of the kurtosis statistic method[De Roo et al. (2007)] or polarization signature method. To mitigate the offending signal, mechanical filters are used to selectively filter out the offending signals [Misra et al. (2012)].

Using mechanical filters is effective in removing the offending signal but does add both weight and complexity to the radiometer. It also adds cost as it requires multiple filters in order to select the narrow band that contains the offending signal.

CHAPTER 3. BACKGROUND

This chapter will examine background information on the basic operation of a radiometer, a software defined radio and background on the development platform used to build a software defined radiometer. We begin with an overview of a traditional radiometer and how this type of radiometer works. This is followed by high level examination on how a software defined radio operates. Finally we will discuss both the hardware and software selected and used for our development platform to create this software defined radiometer.

3.1 Radiometer Basics

The primary task of a radiometer is to measure power. While that statement sounds easy, there are many factors that go in to how well a radiometer can measure the power it observes. A better statement is that a radiometer's primary goal is to accurately measure the correct power within a certain degree of accuracy. In order to accurately and within a high degree of precision measure power a radiometer must take into account factors such as the system noise, the bandwidth of the signal and the stability of the system as a whole[Evans and McLeish (1977)].

The amount of noise that is generated by the object of interest is due to the thermal agitation of the charge carriers, usually the electrons, which is directly correlated to the physical temperature of the source [Nyquist (1928)]. It is because of this correlation that we often refer to the amount of noise received as the noise temperature and it is measured in Kelvins.

In all radiometers there are six stages that is common in all radiometers. They are:

1. Source (antenna or T_A)
2. Amplification (Gain or G),

3. Bandwidth restriction (filtering or β),
4. Power detection (X^2),
5. Integration (τ),
6. Output.

Figure 3.1 shows each stage as they relate to each other. We begin with our source which is often an antenna. Amplification or gain in the system then amplifies this signal so that it is easier to measure changes in our noise. Filtering can occur before or after the amplification stage and in some cases filtering is done before and after the filtering stage. Power detection extracts the power information from the signal. Because the signal is often noisy, we use integration to smooth our signal and improve our sensitivity. Finally our output is the total power information. This information may be as simple as a voltage output or may be a calibrated noise temperature.

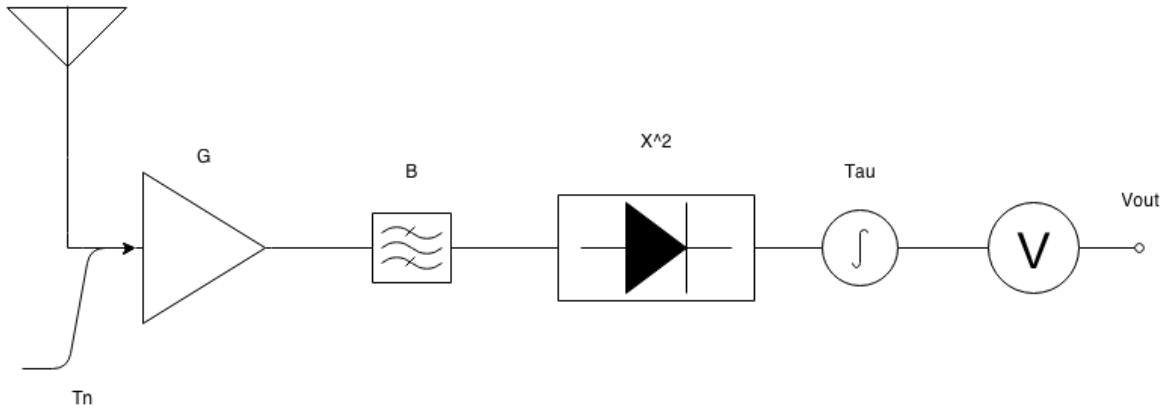


Figure 3.1 A total power radiometer block diagram

In figure 3.1 we have one more component that is common in all radiometers but is not a physical component or stage in the radiometer. That item is additional noise that gets added to the system from within the system itself is called system noise or T_N .

In addition to the system noise, another very common problem with radiometers is stability. Most instabilities in a radiometer are due to fluctuations that occur in the amplification stage

or the Low Noise Amplifiers (LNAs) used. These gain fluctuations can be controlled to a point by closely monitoring and controlling both the voltage and the temperature of the LNAs. However, this is not an easy task and in some cases is not practical. Therefore, other methods have been developed to compensate for these fluctuations. There are three common types of radiometers designed to adjust for gain fluctuations. They are:

1. Dicke radiometer,
2. Noise injection radiometer,
3. Polarimetric or correlating radiometer.

One of the first methods used is the Dicke radiometer which switches between the measurement of the antenna and a known source[Dicke (1946)]. By referencing this known source very quickly the Dicke radiometer can account for and greatly reduce gain fluctuations.

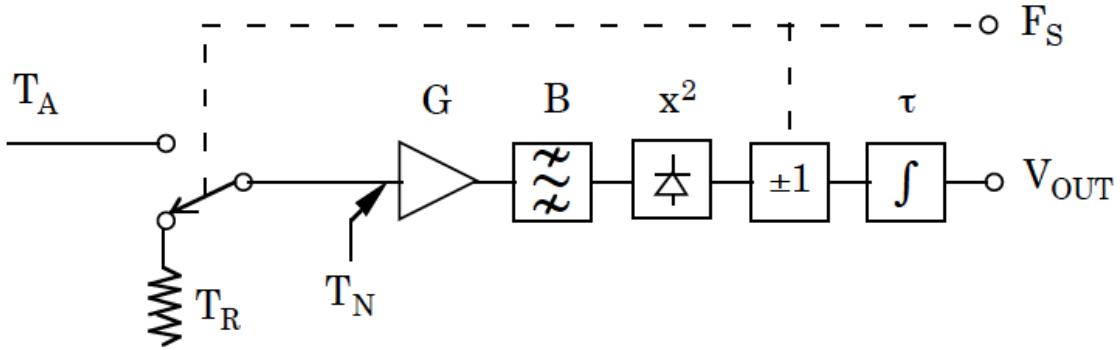


Figure 3.2 A block diagram of a Dicke radiometer

While a Dicke radiometer greatly reduces the fluctuation in the gain of the radiometer and improves stability, it does so at the cost of not seeing the object of interest while it is referencing the known source. This decreases the sensitivity of the radiometer.

A noise injection radiometer is a variation of the Dicke radiometer where a variable noise source is used and injected into the RF chain as seen in Figure 3.3. The output of this noise source is then adjusted so that the noise input plus the signal from our source is equal to the

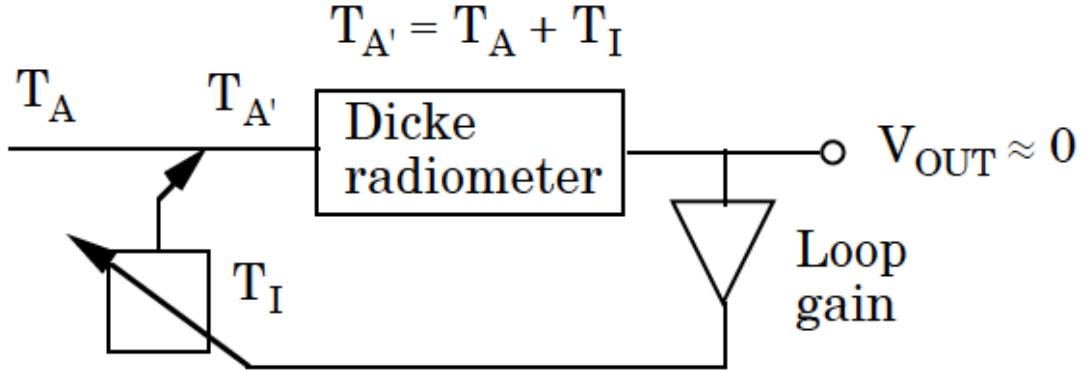


Figure 3.3 A block diagram of a Noise Injection radiometer

reference noise. This completely eliminates fluctuations however increases the system noise which also reduces our sensitivity of the radiometer.

Our source signal is often times assumed to be not polarized. However, this is not the case and an incoming signal often has some polarization. A correlating radiometer [Fujimoto (1964)] uses a dual polarization antenna to split the vertical and horizontal polarization of the signal. Each of these signals is then fed into a radiometer and are correlated. Since gain fluctuations or uncorrelated, they are reduced in the system. This reduces gain fluctuations and also helps maintain the sensitivity of the radiometer. However it adds a much greater complexity to the radiometer and requires two identical receivers.

3.1.1 Measuring RF power

To measure power in a radiometer, several factors are taken into consideration. To begin with we have the noise signal coming from the antenna. Our antenna is assumed to be looking at our target of interest and it is assumed that we can relate the antenna noise to the noise from the source. It is often easier to refer to this noise as the brightness temperature. Therefore, the brightness temperature of the source can be related to the brightness temperature at the antenna. We will refer to this brightness temperature as T_A .

Figure 3.4 shows us an ideal radiometer. That is a radiometer that has an input from the



Figure 3.4 The ideal radiometer block diagram

antenna, T_A , a known bandwidth denoted as B or β and a known gain denoted as G . At the end of the block is the detector, which measures the power from the radiometer.

The power coming from the antenna is the combination of the following items:

1. Gain or G of the system,
2. Bandwidth or β of the system,
3. Signal source or T_A .

To calculate our power from the radiometer, we multiply each item with the Boltzmann constant referred to as k . Equation 3.1 gives us the total power, in watts, of an ideal radiometer.

$$P = k * \beta * G * (T_A) \text{ (watts)} \quad (3.1)$$

As discussed in the above section, the system noise is another component of the radiometer that needs to be addressed. Figure 3.5 shows the additional noise that is injected into the system.

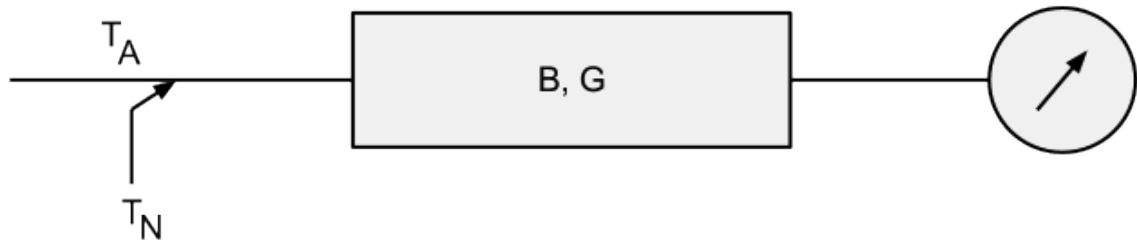


Figure 3.5 A more realistic radiometer model

As it can be seen, this additional noise is added to the noise coming from the antenna source. Therefore, T_N is added to T_A and our final equation for the power measured is shown in equation 3.2.

$$P = k * \beta * G * (T_A + T_N) (\text{watts}) \quad (3.2)$$

Figure 3.1 showed us a more typical radiometer and was discussed in the previous section. However, power detection and the associated voltage output has not been discussed. Power detection is accomplished typically with a square-law detector and this output is often run through an integrator to smooth the output[Leinweber (2001)]. Finally we have the output which is a voltage represented as V_{out} . This results in equation 3.3.

$$V_{OUT} = c * (T_A + T_N) * G \quad (3.3)$$

Here V_{OUT} is shown by the addition of both the noise from the system T_N and the noise from the antenna, T_A and multiplied by the gain in the system, G [Skou and Vine (2006)]. A constant factor c is useful for when we are looking at a radiometer like a Dicke radiometer in which the value of c is $\frac{1}{2}$. In most applications outside of a Dicke radiometer, c is 1 and can be ignored.

The voltage output from this radiometer is then either measured or may also be sampled by an analog to digital converter. This voltage then represents the total power measured by the radiometer, however this measurement is not calibrated.

3.1.2 Integration and filtering

Filtering with a traditional radiometer is usually accomplished by using mechanical filters. Often these are band-pass filters that limit the bandwidth that the radiometer sees. Other filters, such as a low pass filter are also used, but usually to smooth out the output from the square law detector. Another item used to help smooth out the signal from the square law detector is an integrator.

A RC filter is analogous to an integrator where the R and C values determine our time constant and our integration time for the filter[Aitken (1968)]. We know a RC filter is analogous to an integrator by looking at equation 4.3.

$$\frac{1}{RC} \int V_i dt \quad (3.4)$$

To begin with, we look at what an analog RC filter looks like.

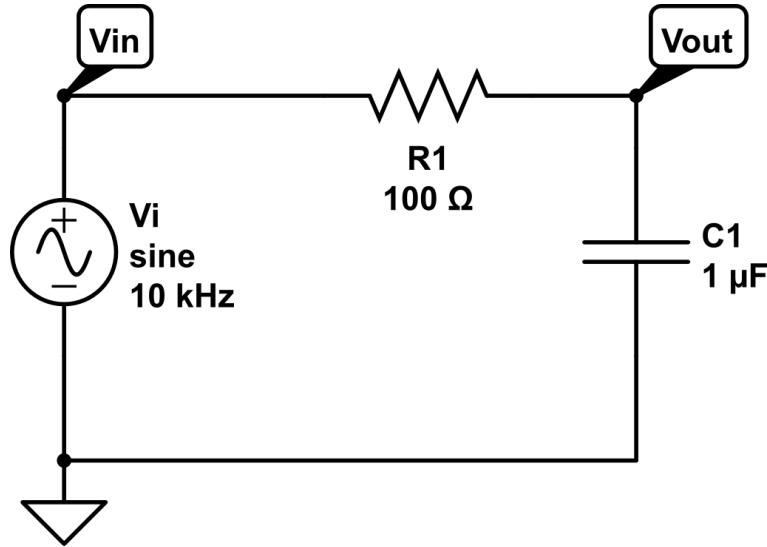


Figure 3.6 A simple RC circuit

This circuit can be represented by equation 4.4.

$$\frac{V_{in} - V_{out}}{R} = C \frac{dV_{out}}{dt} \quad (3.5)$$

This integration smooths out the signal from the square law detector and also improves our sensitivity of the radiometer. This will be further explained the next section.

3.1.3 Radiometer Performance Metrics

There are two criteria that determines how well a radiometer performs. The first criteria is with the sensitivity of the radiometer. This tells us how well the radiometer can differentiate between the information we want and the noise we do not want. The second criteria is stability.

A radiometer needs to be calibrated to give us useful information. Stability refers to how much change or drift a radiometer is experiencing which in turn can affect the accuracy of our data.

Sensitivity. Sensitivity of the radiometer relates to the amount of power that the radio selects from the antenna. This selection is then dependent on the bandwidth that the radiometer is able to listen to. The radiometer however detects not only the signal of interest but also receives a noise signal as well. This noise is added to the signal and can not be separated from the signal. Because this noise is added to the signal, we must be able to determine a change in the signal while the noise signal is also present.

To understand this, let us look at the example of T_A has a value of 200 K and T_N has a value of 800 K. Since T_N is added to our antenna signal, we have a total noise temperature of 1,000 K. This means that if we want to detect a change as small as 1 K, we must be able to measure the difference between 1,000 K and 1,001 K[Skou and Vine (2006)].

The ability of a radiometer to detect these small changes is the radiometer's sensitivity, or the standard deviation of the output signal from the radiometer. This sensitivity is also referred to as the Noise Equivalent Delta (Δ) Temperature or $NE\Delta T$ and is shown in equation 3.6.

$$NE\Delta T = \frac{T_A + T_N}{\sqrt{\beta * \tau}} \quad (3.6)$$

The sensitivity of the radiometer is based on both the bandwidth, β , of the incoming signal and the integration time, τ . As it can be seen in the equation, we want to have as much bandwidth as possible. In a traditional radiometer, this bandwidth is often fixed and is dependent on the band-pass filters used in the radiometer. We can however control τ and a longer integration time will help improve the sensitivity of the radiometer to a certain degree.[Ulaby et al. (1981)]

Stability. Stability and accuracy are additional problems that need to be considered when looking at the radiometer system. To begin we can once again look at the power received equation 3.2.

As we look at this equation, we can see that if k , β , G , and T_N are constant, then stability can be assured. The Boltzman constant k is a known constant and we can also assume that our

bandwidth, β , will also remain constant. Gain and the noise temperature however will vary.

Fluctuations in gain is the largest factor that affects stability in the system and this can be seen in equation 3.7. As discussed earlier, it is this factor that has been a driving force for changes to the design of a radiometer as demonstrated by the Dicke and Correlating radiometer designs.

Gain fluctuations are caused by two factors: the physical temperature of the amplifier and the voltage that feeds the amplifier. High accuracy voltage regulators can help control fluctuations in voltages that can in turn affect the gain. A factor however that is harder to control is temperature. Various methods have been used to control the temperature when the radiometer is used in fluctuating temperature environments such as the outdoors or in space.

$$\Delta T_G = T_{sys} \left(\frac{\Delta G}{G} \right) \quad (3.7)$$

With stability we need to either attempt to stabilize the radiometer as best as we can or compensate for the gain fluctuations. Compensation has been discussed and three types of radiometers have been identified that attempt to compensate for these fluctuations. To control the stability many radiometers will use highly accurate voltage regulators and will control the temperature of the LNA through methods such as thermal electric coolers.

3.2 Software Defined Radios Basics

The basic concept behind a SDR is that it will digitize the RF signal as soon as possible. Once digitized, it is now evaluated by a computer, FPGA, or a dedicated System on Chip (SoC). A canonical software defined radio architecture is one that consists of a power supply, antenna, multi-band RF converter, and a single chip that contains the needed processing and memory to carry out the radio functions in software [Mitola (1995)]. This allows us to extract certain hardware functions, such as filtering, into the digital domain which can then be manipulated by software. Since software is now manipulating the signal, we can rapidly change what functions we execute on the signal by changing the software. This gives SDRs a high amount of flexibility as components that are normally done in hardware can now be done in software and can be

changed by simply uploading new software or firmware to the system. This also gives us a benefit in cost as certain components are no longer needed and changes done in software do not require additional hardware to be added or to be swapped out.

An ideal software defined receiver simply has an antenna connected to an analog to digital converter which sends information to a processing unit such as a Field Programmable Gate Array (FPGA) or computer. For a transmitter, we reverse it and use a digital to analog converter to produce the correct waveform which is then transmitted by the antenna. In reality, SDRs require some additional hardware to make a viable working transceiver. Amplification is still required to amplify the incoming signal and to amplify the signal going to the antenna. Some SDRs also use a mixing stage to move a high frequency signal to a lower frequency signal. This allows for less expensive analog to digital and digital to analog converters to be used.

3.2.1 Software Defined Radio Normal Operation

Software Defined Radios (SDRs) are used for a variety of applications but their primary application has been in the area of communications. They appeal to applications where being able to change a modulation scheme or filter on the fly is desirable. In these areas, SDRs often outperform a traditional hardware only radio with their ability to rapidly change their operations by simply changing their software. Early SDRs were expensive due to the high costs in the analog to digital converters (ADCs) needed and in the high speed Field Programmable Gate Arrays (FPGA) used. In recent years however, the cost of SDRs have decreased due to the cost of these key components decreasing in cost as well. Even though the cost has gone down the performance of SDRs have increased and this has lead to new developments in applications for using SDRs in new and different ways.

Software defined radios have been used in a number of applications. Some applications they have been used in are:

1. Cellular communications,
2. Wireless local area networks,
3. Personal area networks,

4. Digital broadcast.

There are of course other applications than those listed and more applications added as new communication methods continue to evolve [Jondral (2005)]. But this is the strength of a software defined radio, it is capable of performing all of these operations and can be adapted to new ones by simply updating the software that defines it.

3.3 Software Defined Radio Development Platform

The work of this thesis is to use an off the shelf software defined radio (SDR) to perform the same operation or better of a traditional analog radiometer. Using a SDR radio also means that we are able to be more flexible in how the radiometer performs, is capable of frequency agility and adapting to changing conditions such as interference. Using a software defined radio also allows for implantation of different radiometer types such as a correlation radiometer or a polarimetric radiometer that uses Stokes parameters [Wang et al. (2012)]. Normally, these require changes to hardware, but all of these types of radiometers can be implemented in software increasing the flexibility of the system.

3.3.1 Hardware Platform

The key component for a software defined radio radiometer is the hardware that will do most of the work of sampling and processing the signal. The equipment selected for this work is the Ettus Research Group N200 SDR and can be seen in figure 3.7. The N200 has the following features that made it well suited for our specific application:

- Dual 14-bit ADC,
- 50 MS/s Gigabit Ethernet streaming,
- Modular daughter-board system for RF front end

This SDR utilizes daughter boards as the RF front end to the SDR and up to two daughter boards may be installed into a N200. Another important reason the N200 is selected was due

to its ability to handle up to 50 MHz of bandwidth to the computer and up to 25 MHz of RF bandwidth per daughter-board plugged in to the SDR. This means that it is possible to have two receive cards that can stream up to 25 MHz bandwidth each.



Figure 3.7 The USRP N200 from Ettus Research (Image from Ettus Research Website - www.ettus.com)

The N200 utilizes a flexible architecture through the use of daughter-boards for a variety of RF interface systems based on the frequency range desired and if receive and/or transmission is needed. Figure 3.8 shows the overall architecture of the N200 SDR. These daughter boards directly receive the RF signal and then outputs the analog I and Q signals that are then sampled by the N200 A/D converter for reception or receives the I and Q values from the N200 D/A converter for transmission.

The flexible architecture of the N200 and the ability for it to handle large amounts of bandwidth made this hardware ideal for the software defined radio radiometer work done in this thesis.

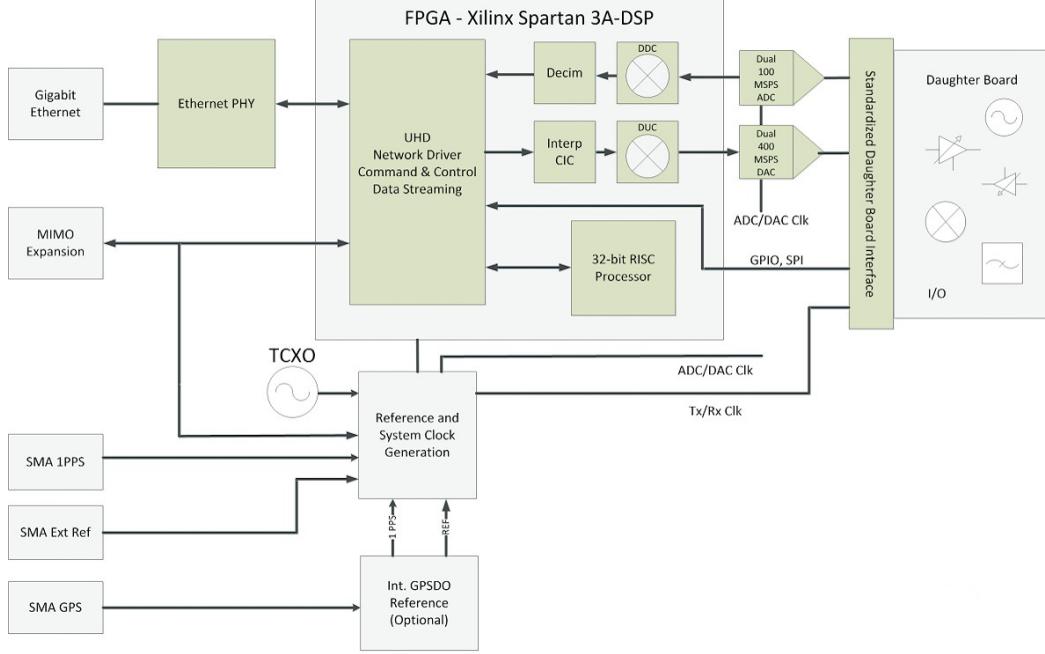


Figure 3.8 A block diagram of the Ettus N200 SDR. (Image from Ettus Research Website - www.ettus.com)

The DBSRX2 Receiver. The daughter board selected is the DBSRX2 card as this card is receive only and operates between 800 MHz and 2.4 GHz. An image of the daughter-board can be seen in figure 3.9. The DBSRX2 also has built in amplification that is adjustable through software. These daughter boards have the required RF hardware for the signal to be processed. In this application it was required that the signal was detected at 1.4 GHz with a bandwidth of 20 MHz. The DBSRX2 receiver met this requirement and was selected to be used with the N200. In this radiometer application transmission is not needed and is illegal in the 1.4 GHz band, which is reserved for passive radiometer applications.

The DBSRX2 has several key components on it that is used to take the analog RF signal and prepare it for digitization by the analog to digital converter. First the signal is amplified through a Programmable Gain Amplifier (PGA). This PGA is accessible from the software and can be configured by the software. Next the signal goes into a direct-conversion integrated circuit that directly converts the RF signal to analog I and Q values. The integrated circuit, a Maxim MAX2112 device, also includes a Low Noise Amplifier (LNA), mixer and Low Pass Filter (LPF). This essentially amplifies the signal, mixes into baseband and then applies a low

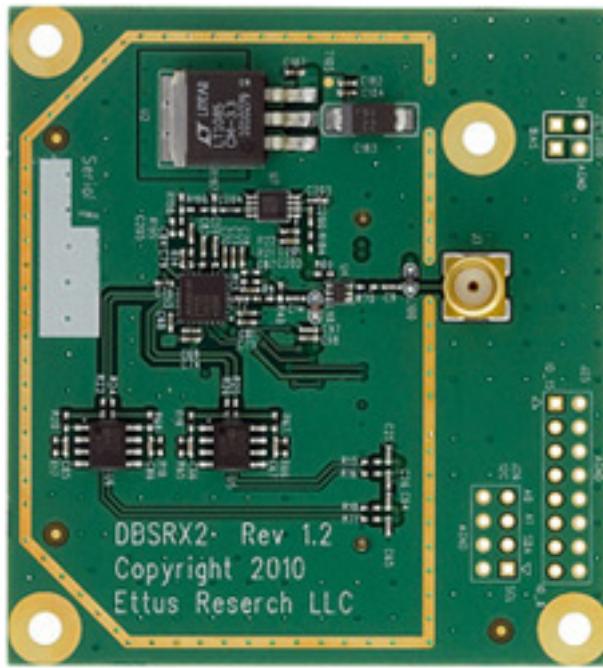


Figure 3.9 The DBSRX2 daughter board from Ettus Research (Image from Ettus Research Website - www.ettus.com)

pass filter.

These analog I and Q values are then passed to the N200 to be sampled by the analog to digital converter. The IQ values are differential signals to minimize noise possible interference.

Since we are using the DBSRX2 after the LNAs that are already in use, the noise temperature added by the DBSRX2 will be small. The DBSRX2 adds approximately 5 dB to the noise factor of the system. Again though, since this is at the end of the RF chain, the total contribution of the DBSRX2 to the overall system noise temperature is small, and has been calculated to be 1.05 dB to the overall noise factor of the system.

3.3.2 Software Platform

Software of course plays a critical role in a software defined radio and also in our software defined radio radiometer. There are two pieces of software that are in play with the software defined radio we are using. The first is the firmware that is used in the FPGA of the N200. This

firmware provides low level processing of the signal so it can be sent to the software located on the computer. It also provides a link for controlling key aspects of the software defined radio such as additional gain, bandwidth and the center frequency. This firmware is already pre-loaded into the FPGA by Ettus Research and can be upgraded through tools provided by Ettus Research.

The second is the software that is running on the host computer. It is this software that provides the calculations on the I/Q data to give us the information we need and also creates a GUI for the user to interface with the radio. For this software, we will be using GNURadio, an open source software program that is used in software defined radios including the N200 SDR that we have.

GNURadio will be used to do all signal processing that is needed. GNURadio is an open source software define radio framework that runs on multiple OSes and offers a rich set of features. In addition, GNURadio is well supported by the Ettus Research group and is the preferred software for interfacing with their hardware.

An easy to use interface was another driving requirement for our implementation of a radiometer in a SDR. GNURadio helps us with this through the use of the GNU Radio Companion or GRC. This was important as it was anticipated that operators of the radiometer have a limited knowledge about programming. GRC uses a simple to use graphical system to design and build radio components in software.

GNURadio Companion works by having the common functions as blocks that can picked and placed on the screen. Once placed, the blocks can be wired up, much like LabView, and the flow of data can be controlled in this fashion. While GNURadio Companion provides most of the essential blocks used in most applications additional blocks can be added if needed. This is because GNURadio is built using Python and the blocks in GNURadio Companion are simply blocks of Python code. To compile a GNURadio Companion flow diagram, you simply run the sheet, which then generates the Python code that is then executed.

GNURadio uses a combination of Python and C++, where Python handles most of the interface and the C++ is most of the drivers and low level interface to the hardware. This allows for an easy to use system but still meets the demanding performance needed for handling

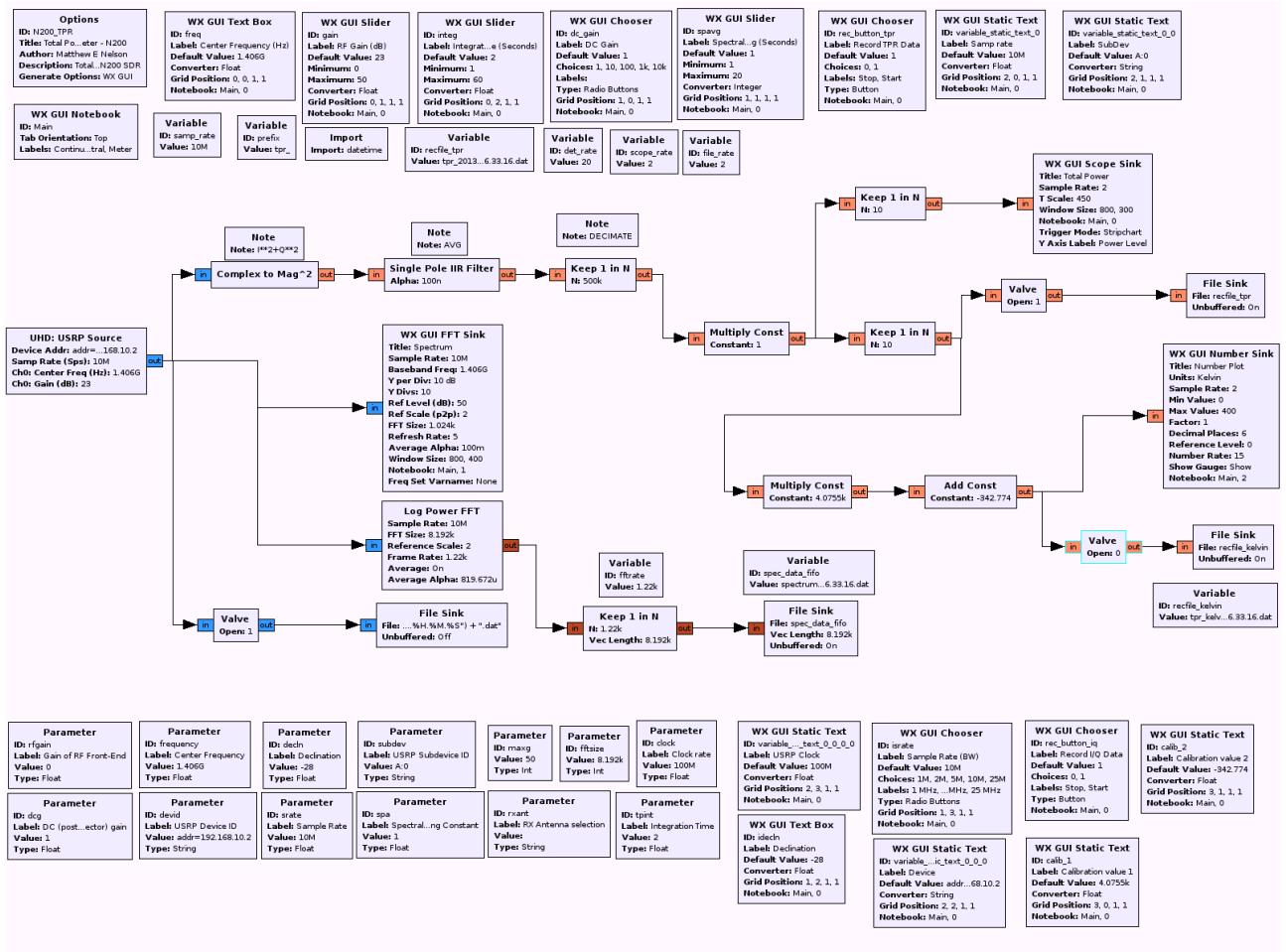


Figure 3.10 Screenshot of the GNURadio Companion editor and the N200 Radiometer block diagram used in the author's experiments

large amounts of data.

GNURadio Companion also includes blocks that allow it to create a GUI type of interface. The typical method it uses for this is using wxGUI although GNURadio Companion does also include blocks that can use QT for generating widgets as well. However, the wxGUI tends to work better and has better support in GNURadio than QT.

Through these blocks we are able to only manipulate the data we need to perform a total power radiometer in software but to also create a user interface that allows us to control the radiometer as well. We are also able to display the information in real time so the user can see changes in power and even monitor spectral information during the operation of the radiometer.

CHAPTER 4. SOFTWARE DEFINED RADIOMETER IMPLEMENTATION

One of the principal goals with this research was to implement a fully functioning total power radiometer within software. The N200 provides us the link between the RF signal captured by the antenna and converts that to a format that can now be used by the to manipulate the signal. Once the signal has been passed to the computer, GNURadio will implement the correct algorithms to detect the power within the signal, filter and output the information.

One of the advantages of course with a SDR is that filtering can also be done within the software. In addition, thanks to the WxGUI that GNURadio uses, we can also build a user interface that can control several key variables that are useful for us. This includes controlling the gain on the programmable gain amplifier on the DBSRX2 card, the sampling or bandwidth of the signal, the center frequency, and also the integration time. All of these can now be controlled in real time as well. GNURadio will also store the power data so that we are able to do further analysis of the data using a software program like Matlab or Python. Because GNURadio is a very flexible system we are able to do more with the signal such as filtering, polarimetric radiometer, and frequency analysis. We are also able to add additional features and improvements through updates to the software.

4.1 Requirements

To help us quantify the required performance of the radiometer, information was provided to us by Dr. Brian Hornbuckle and also derived from existing radiometers. The requirements given are outlined in table 4.1.

These requirements mainly drive the hardware requirements needed for the software defined

Table 4.1 Required Radiometer performance

Parameter	Value	Units
Minimum bandwidth	20	MHz
Operational frequency	1400 - 1420	MHz
$NE\Delta T$	1	Kelvin

radio radiometer. Chapter 3 Section 3.3 provides more in depth information on the hardware used and why it was selected.

4.1.1 Software Requirements

The driving force for the software requirement was to have a system that was easy to use yet powerful enough to handle the amount of data that is required. One reason for the development of this platform is to make radiometers more accessible to other researchers and other programs such as education and even amateur radiometer work. Therefore, ease of use was taken into consideration when selecting the hardware and the associated software used with it.

The data flow model for the hardware selected uses the FPGA to perform low level signal processing on the signal and moves high level processing to the host computer. This allows for less processing requirements on the physical hardware but requires a host computer that is able to process the incoming information. It also requires a software package that is able to process this information efficiently.

Because a requirement is an easy to use system GNURadio was selected as it includes GNU-Radio Companion (GRC). GRC is a supplemental program which uses a graphical interface for creating the radio environment. It also includes options to create a user interface during the operation of the N200 as well. This allowed us to rapidly create both the critical radio components needed for the radiometer and also a control interface.

This software meet the criteria of allowing a simple to use interface to be built and used in the control and data recording of the information required. It also uses a simple interface for making changes to the program. These changes can be both in the GUI and also to how the program processes the information.

4.2 Theory of Operation of a Software Defined Radio Radiometer

A Software defined radio consist of both hardware and software that allow it to perform the operations of a radio or communication channel. A software defined radio used for radiometer applications is identical to a software defined radio used for, as an example, a 802.11b radio with one major difference. Since we need to amplify the signal more than what most communication applications require; we do require more powerful or additional LNAs to boost this signal. In addition, since the first LNA plays a major role in the overall system noise and this system noise does affect performance of the radiometer, the selection of this LNA is important. However, all other components are the same components used in other applications.

A software defined radio radiometer behaves analogous to a more traditional radiometer and thus the application is the same as a traditional radiometer. This includes applications such as radio astronomy that includes applications in Earth Science such as soil moisture and ocean salinity[Ruf and Gross (2010)]. A software defined radio radiometer can also allow for new application development that can expand the remote sensing field. Since we have moved the majority of the hardware to software this allows us to further shrink the size and weight of the radiometer. This allows for other radiometer applications such as Unmanned Aerial Vehicles (UAVs) for scanning soil moisture and ocean salinity remotely[McIntyre and Gasiewski (2007)].

We will now introduce the three major components that make up a software defined radio radiometer.

4.2.1 RF Front End

The RF front end plays a critical role in the radiometer as the LNAs used in the front end has a large impact on the system noise generated by the radiometer itself. A traditional radiometer utilizes both amplification through the LNAs and also includes filtering to the desired bandwidth. A SDR radiometer does not require the filters as we are able to create these in software, however the amplification stages need to remain.

A typical RF front end uses a 3 stage Low Noise Amplifier (LNA) to amplifier the noise while keeping the noise contributed to the system as low as possible. As with any radiometer, the

first LNA is the most critical as it contributes the most to the overall system noise temperature. For this reason a LNA that did not have a large gain but had a low noise figure is chosen. The second and third LNA has higher gain values at the cost of a higher noise figure, although not by much. However, since they are further down the chain, they do not contribute as much to the total system noise.

$$F = F_1 + \frac{F_2 - 1}{G_1} + \frac{F_3 - 1}{G_1 G_2} + \frac{F_4 - 1}{G_1 G_2 G_3} + \cdots + \frac{F_n - 1}{G_1 G_2 G_3 \cdots G_{n-1}} \quad (4.1)$$

Equation 4.1 shows us how the noise factor and gain of the amplifiers affect each other. It can be shown that the first amplifier or first noise figure in the system contributes the most to the overall system noise figure. Additional components contribute, but at a much lower contribution.

4.3 Mapping Traditional Radiometer Functions to a Software Defined Radio Radiometer

In order to recreate a radiometer in software we need to identify the key components of a radiometer and then recreate those components in software. As discussed in chapter three, the three components identified that are key to a radiometer is listed below.

1. Power detection
2. Integration
3. Bandwidth limitation or filtering

The following sections will now examine how these items are mapped from their analog component to the software or digital component.

4.3.1 Power detection

Power detection is a key ability that allows a radiometer to function. At its core a radiometer is a power detector. Therefore, the implementation of power detection is a crucial function of a software defined radio radiometer.

A traditional radiometer uses a square-law detector which takes the input signal and produces a voltage that is proportional to the square of the voltage. This allows us to take an analog RF signal and convert the noise voltage that for all intense and purposes has a mean value of zero, and produce a noise power.

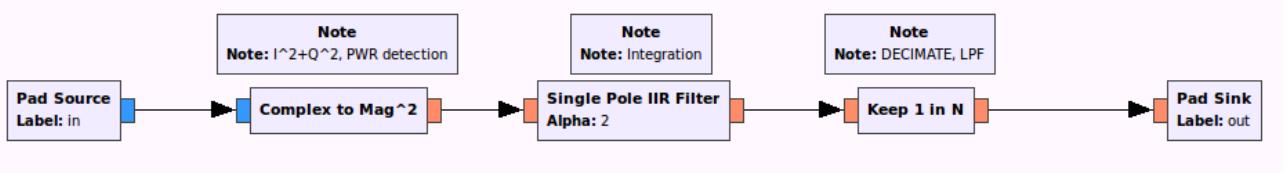


Figure 4.1 A block diagram showing how the radiometer performs the equivalent square law detector in software.

To implement this in software we need to build a square law detector mathematically. A square law detector mathematically is the sum of the squares. Once the signal has been digitized, it is expressed in data bits of I and Q , which represent in-phase and quadrature-phase of the signal. By squaring each term, we get the desired result of the power of the signal [Sarijari et al. (2009)][Rashid et al. (2011)] can be shown in equation 4.2.

$$I^2 + Q^2 = P_{out} \quad (4.2)$$

Therefore, like the analog square-law detector we are taking peak voltage values, which has an equivalent noise voltage and a mean value of zero, and square them to produce a noise power that is proportional to the square of this amplitude. Figure 4.1 shows the GNURadio blocks used to perform this function.

Like the analog square-law detector, this signal will fluctuate rapidly and to improve the sensitivity of the radiometer we wish to integrate this signal. We now want to look at how we can replicate a RC filter or integrator in the software defined radio.

4.4 Integration

Another step that we typically do in a traditional radiometer is to integrate the signal over time. This gives us an average of the signal and smooths out the output. This also helps to

improve the sensitivity of the radiometer by equation 3.6.

In a traditional radiometer, we can integrate by using a simple integrator circuit, which consists of an op-amp, resistor and capacitor. This circuit configuration is also equivalent to a low pass filter circuit as well, and the two are interchangeable. We can then look at how we filter in the digital domain, and this is done with an infinite impulse response filter or IIR. We can use this digital filter to then integrate the signal for our total power radiometer.

To begin, we will examine how a RC filter is analogous to an integrator where the R and C values determine our time constant and our integration time for the filter[Aitken (1968)]. We know a RC filter is analogous to an integrator by looking at equation 4.3. A SDR however operates in the digital domain at discrete intervals. One type of filter that can be used is the Infinite Impulse Response (IIR) filter.

$$\frac{1}{RC} \int V_i dt \quad (4.3)$$

To begin with, we look at what an analog RC filter looks like.

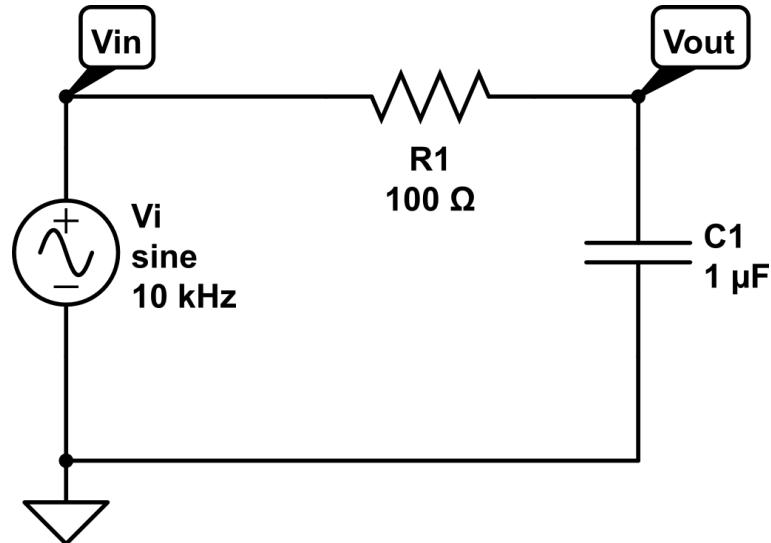


Figure 4.2 A simple RC circuit

This circuit can be represented by equation 4.4.

$$\frac{V_{in} - V_{out}}{R} = C \frac{dV_{out}}{dt} \quad (4.4)$$

A Finite Impulse Response (FIR) filter is a digital filter that can take an impulse signal and decays to zero after a finite number of iterations. This type of digital filter can be represented by equation 4.5 which mathematically expresses the FIR Filter.

$$y_n = \sum_{i=0}^{P-1} c_i x_{n-i} \quad (4.5)$$

This simply says that the nth output is a weighted average of the most recent P inputs.

An Infinite Impulse Response (IIR) filter is the same as the FIR filter, except that we add a summation term which feeds back the previous output.

$$y_n = \sum_{i=0}^{P-1} c_i x_{n-i} + \sum_{j=1}^Q d_j y_{n-j} \quad (4.6)$$

Equation 4.6 shows that a FIR filter is a IIR filter, except that $Q = 0$ [Cross (1998)].

To get a better understanding on how the digital IIR filter relates to the RC filter analog, we can look at the Fourier Transform and the relationship of the input to the output in the frequency domain.

$$H(f) = \frac{\sum_{j=0}^{P-1} c_j e^{-2\pi i j f T}}{1 - \sum_{k=1}^Q d_k e^{-2\pi i k f T}} \quad (4.7)$$

In equation 4.7, f is our frequency in Hz and T is the time between samples in seconds and is related to our sampling frequency.

We now want to show the link between our analog RC circuit and the IIR filter. Looking at equation 4.4, which represents the differential equation relating the input voltage V_{in} to the output voltage V_{out} , we can substitute for input and output of our IIR filter. Since we are now in the time domain, we need to define what T is and we can do that using equation 4.8.

$$T = \text{time between samples} = \frac{1}{\text{sampling rate}} \quad (4.8)$$

We can now relate our input voltage to the input to our IIR filter and the output voltage to the output of our IIR filter.

$$x_n = v_{in}(nT) \quad (4.9)$$

$$y_n = v_{out}(nT) \quad (4.10)$$

We can now rewrite our difference equation with x_n and y_n .

$$\frac{x_n - y_n}{R} = C \frac{y_n - y_{n-1}}{T} \quad (4.11)$$

Finally, we can solve for y_n which results in our final equation for showing how a IIR filter is related to an RC filter.

$$y_n = \frac{T}{T + RC} x_n + \frac{RC}{T + RC} y_{n-1} \quad (4.12)$$

It can be seen that an IIR filter can have the same frequency response as we expect from an analog RC filter. As our sampling rate approaches infinity, the approximation gets closer to the original response from the analog RC circuit.

For the cutoff frequency of a RC circuit, we know that it has the relationship shown in equation 4.13.

$$f_c = \frac{\sqrt{3}}{2\pi RC} \rightarrow RC = \frac{\sqrt{3}}{2\pi f_c} \quad (4.13)$$

The RC term gives us our time constant of the circuit and can be used to calculate out our coefficients. We are not concerned about the actual values of R and C with our IIR filter, instead we just need the product of R and C .

In GNURadio most of the work is done for us. We can simply enter in our desired cutoff frequency and GNURadio will calculate our IIR filter coefficients. However, this shows that an IIR filter works much like an analog RC low pass filter.

4.5 System overview

Like a traditional radiometer, the SDR will use an antenna to look at the target of interest. SDRs still use a RF stage that takes the power from the source and amplifies it. The difference

though begins after that. A SDR will then sample and generate I and Q values that represents the amplitude and phase of the signal. From there, this data is sent to a computer to be processed. We can then use this information to calculate the power that is being seen. In addition, we can manipulate the signal in other ways such as applying a filter to filter out an unwanted source.

As we have shown the two of the major components of a traditional radiometer, the power detection and integration of the signal can be replicated in software and therefore can be implemented in a software defined radio. The information can now be stored, displayed or both for further analysis.

There is one component of the software defined radio that we are not able to implement in software and that is with the signal amplification. This however does play a major role in the performance of the radiometer and is a key element that should not be overlooked. While this is not implemented in software, it still plays a critical role in our software defined radio radiometer.

Next we will look at the software used in defining the radiometer. It will also be shown what the impact is of this software on the performance of the radiometer.

4.5.1 Control of the SDR Hardware through GNURadio

The N200 sends all data across the 1 Gbps connection to be read in by a host computer running GNURadio. This data is the raw I/Q values that is read by the on board A/D and processed by the on board FPGA. An example of a very simple GNURadio software implementation would simply take this data and store the data to a hard drive in a file. This can be very handy if we want to simply record the data and then process it later. However, depending on the sample rate, it can consume a large amount of storage. A short recording can easily consume 1-2 GB with a sample rate of 10 Msps. It also does not give us any immediate feedback on the radiometer and it does not give us controls of the radiometer such as frequency, integration time or other key variables. Fortunately GNURadio has tools that allows us to build up a very rich application that is able to give us the data we need and control the software defined radio as well.

The GNURadio Companion allows us to create python code that is used to not only receive the data from the SDR but also perform signal processing on the incoming information. Additional controls are added that allow for tuning of the signal processing parameters and control of the radio functions. This allows us to build up an application that can be run on any computer that is capable of running GNURadio.

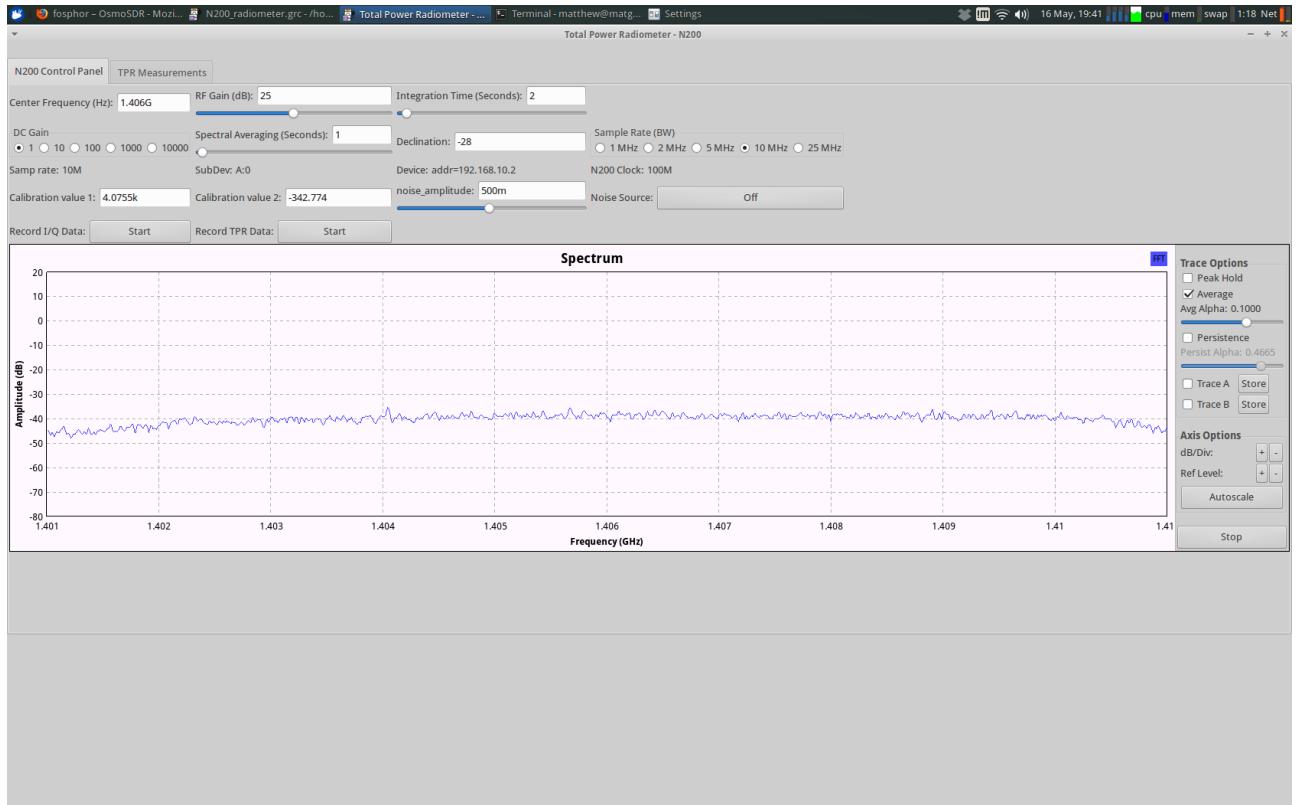


Figure 4.3 A screenshot of the interface made for communication with and controlling the software defined radio

Through this interface we are able to control several key aspects of not only the radio hardware within the SDR but also with the behavior of the radiometer as well. Hardware control of the SDR hardware allows us to change frequency and also adjust gain values within the N200. Bandwidth is another parameter that can be altered from here as well. Bandwidth affects the bandwidth for the RF signal, but also has an impact on the radiometer sensitivity as well. Additional controls allows for alterations to the radiometer and includes being able to adjust the integration time of the radiometer.

4.5.2 Impact of the Controls Related to Radiometry

The controls that have been added for controlling the radiometer can have a large impact on the performance of the radiometer. There is a reason why these controls were added to the GUI for the radiometer and that is they play key roles in how the radiometer performs.

For any radiometer noise temperature is a large consideration and is critical to the design of the radiometer. One method to determine how well a radiometer is to look at the sensitivity of the radiometer. We can do this by looking at the smallest change in temperature the radiometer can see. We will call this the Noise Equivalent ΔT or $NE\Delta T$ of the radiometer. The equation for this is shown below.

$$NE\Delta T = \frac{T_A + T_{sys}}{\sqrt{\beta * \tau}} \quad (4.14)$$

β can be changed by changing the sample rate of the SDR. The sample rate effectively controls the bandwidth in which the SDR is operating at. This also gives us a band-pass filter as well, since the SDR will not respond to frequencies outside of this bandwidth.

τ is the integration time for the radiometer. This parameter is set by the user through the GUI and allows us to change the integration time in seconds.

4.5.3 GNURadio Data Handling

Once we have the data that has been processed by the software defined radio we will want to display this information and be able to store the data so we can analyze it later if needed. Data display can be handled GNURadio where we can plot the total power over time. This allows the user to be able to visualize the total power and be able to determine if the total power has increased or decreased over the time window shown.

Although not usually needed for a total power radiometer we also have the ability to look at the signal in terms of a frequency versus amplitude. This allows looking for any unusual signals that may be interfering with the system or causing erroneous data with our radiometer.

Finally, we will want to store the data so we can do additional analyses on it at a later time. The GNURadio program allows us to store the data in two formats. The first format

is storing the raw I/Q data from the radiometer. This format allows us to playback the data through GNURadio at a later time. This can be useful for if we wish to change parameters in GNURadio such as bandwidth or integration time. It is also a good diagnostic tool as we can check that the signal coming in is clean or if we need to apply additional filters to remove an unwanted signal. This file may be quite large, consuming several gigabytes of data for a 20 MHz wide signal in a matter of minutes of record time.

The second format is the total power that has been calculated by the radiometer. This file is much smaller since much of the signal information has now been reduced to simple power versus time information. This allows for easy manipulation through any math program such as Matlab for analysis.

4.5.4 GNURadio Data Display

The information from the software defined radio can be displayed through GNURadio to show a number of things. Since we have both frequency and magnitude information we can display this information. We are able to also display the information that shows the total power that is being seen by the radiometer as well.

We are not limited to just total power from the radiometer. If the radiometer has been calibrated, those calibration points can be entered and GNURadio can calculate the calibrated noise temperature. Additional information may also be added as needed. For example, we are able to view the full spectrum that the radiometer sees. This can be a useful tool for looking at potential RFI issues.

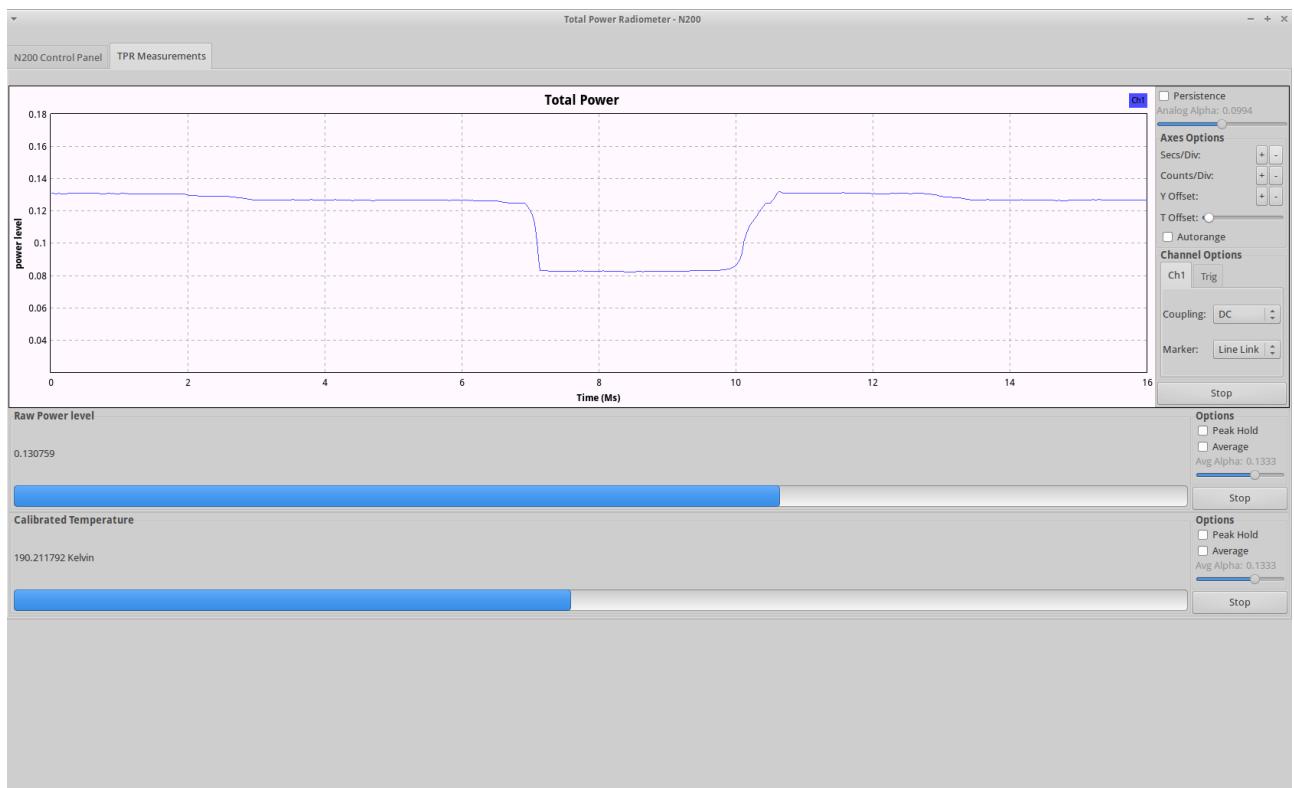


Figure 4.4 A screenshot showing the ticker tape display for the total power readings. In addition, raw and calibrated noise temperature is shown below.

CHAPTER 5. EVALUATION SETUP AND EXPERIMENTAL DESIGN

The experiments outlined in this chapter is to demonstrate and verify that a software defined radiometer is able to perform as well or better than a traditional radiometer. In addition, it will be demonstrated that a software defined radiometer can add additional functionality not typically found in most radiometers.

Experiment one verifies the operation of a software defined radiometer by comparing typical operations to a square-law detector, a typical device used in radiometry and calibration of the radiometer. Experiment two verifies the expected sensitivity and stability of a software defined radiometer. Experiment three will demonstrate a software defined radiometer mitigating an interfering signal and compare the results to a traditional radiometer with no interfering signal mitigation. Finally, Experiment four will examine the affect filtering and bandwidth of the signal affects performance.

5.1 Experiment I - Software Defined Radiometer Verification and Calibration

In this experiment we will verify the operation of a software defined radiometer. This experiment will verify that a software defined radiometer is able to perform as expected and will compare our results to a square-law detector which is commonly used in traditional radiometers. To verify the results of the information that the software defined radio is obtaining a square-law detector is used to measure the power of the incoming signal in parallel to the software defined radio. This signal is split using a power divider so that the information will be the same to both devices with the except for the 3 dB plus insertion loss the power divider adds. This allows us to verify the software defined radio with a proven system.

5.1.1 Data Collection

Two sets of data is produced with this experiment. First, data is generated from the software defined radio using GNURadio. Second, data is generated from a data acquisitions device that is attached to the square-law detector. Data from each one is stored to a local computer running the appropriate software.

Software Defined Radio Data. The data from the software defined radio is stored in files generated from GNURadio. GNURadio uses a sink block to output the data to either a screen, socket connection such as TCP/IP or a file. In our case we will use a file sink to output the data to a file. The flow of data to this sink is controlled by a valve block. This allows for the user to turn on and off recording of the data. A block diagram of the file blocks is in figure 5.1.



Figure 5.1 The File Sink block used in GNURadio. Source: GNURadio

There are two types of files that the SDR generates. The first type of file is the complete I and Q data points recorded from the SDR. This file is stored as a little-indian format and as complex values. Due to the sampling rate, it is not uncommon for this file to grow to be quite large, usually several gigabytes of data for a 10-15 minute run. However, this file can then be

feed back through GNURadio later to be played back if needed and it contains the information needed to completely recreate the signal.

The second file type is the total power values generated from the total power block in GNURadio. A block diagram of this block can be found in Appendix A in figure A.1 and the source code for this can also be found in Appendix A. This file is also a little-endian format however this file only has float values. This file is also much smaller than the file that contains the I and Q data points. A typical file size is 50 - 100 kB for a 10 to 15 minute run.

Square-law detector data. The square-law detector outputs information as an analog voltage that is linearly proportional to the RF power measured. This required a system that can capture the analog signal from the ADL5902 and then send the data to be stored. The National Instruments USB-6009 data acquisition unit was selected as it met the requirements for an easy to use yet high enough resolution to obtain accurate information. The USB-6009 unit has 8 analog inputs that can sample up to 48 ksps with a resolution of 14-bits which is more than adequate for the experiments. To use the USB-6009 a fairly simple Lab View program was created to obtain, display and store the data from the ADL5902. This program retrieved the information from the USB-6009 and stored the data in both Labview's binary format and in a more human friendly ASCII format. The USB-6009 then connects to a host computer through the USB interface. This made obtaining the data and using device straightforward to use.

The Labview program created generates a user interface to display the data and allow for turning on and off the recording. This allowed for easy monitoring of the square-law data while the experiment was being run. A screen shot of the GUI created in Labview is shown in figure 5.2.

This program uses the National Instruments DAQ assistant which allows for quick configuration and setup for the computer to talk to a number of NI devices. Labview also includes blocks that allows us to easily record the data to a file and also use a low pass filter. These blocks made up most of the program and resulted in a program that was quickly made. Figure 5.3 shows the blocks used and the wiring of the blocks.

GNURadio Testing. Before testing the program with the N200, testing was done with the built in noise generator in GNURadio. This was used to test its ability to measure small changes

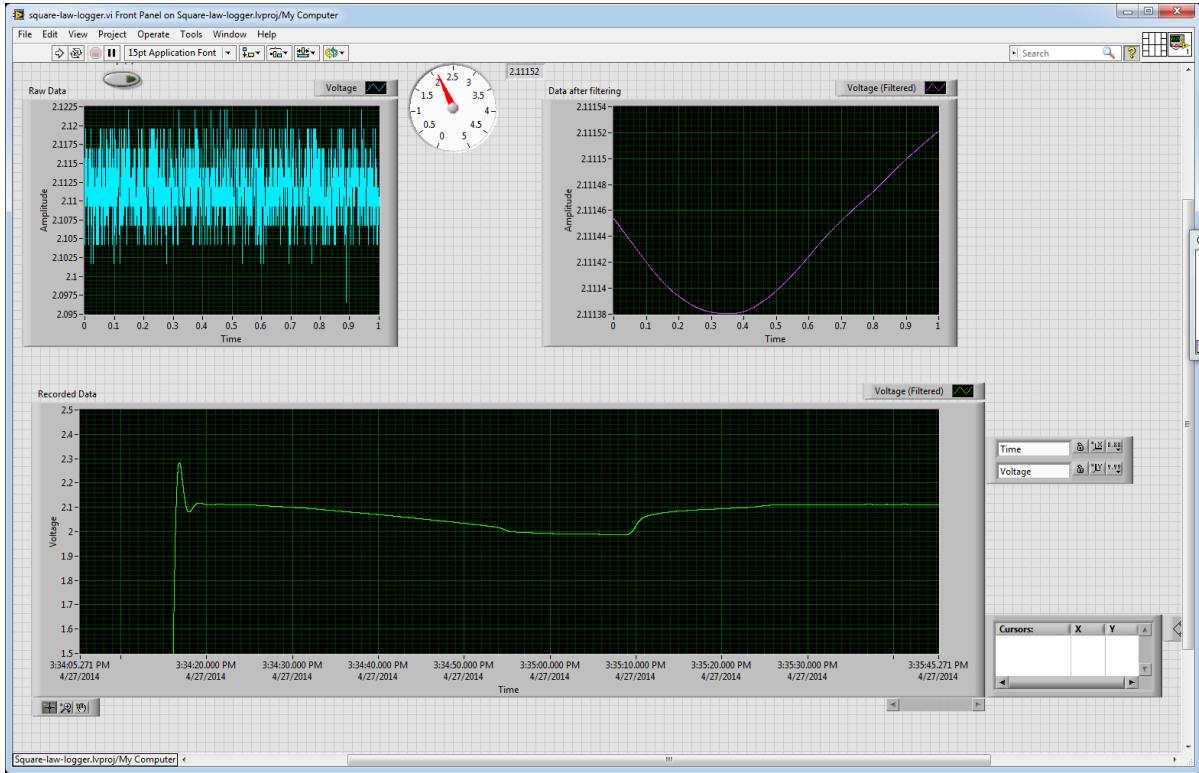


Figure 5.2 A screenshot of the Labview GUI interface. Source: Labview

in noise. This simulated noise verified that the program written was able to detect changes in noise power using a simulated Gaussian source. It was also desired to also use a hardware based noise generator, however a suitable noise generator could not be located on campus.

After this was done additional tweaks were made to both the GUI interface and to the program that performs the calculations for the total power radiometer. The essential components were moved to a custom block which helped to clean up the block diagram.

5.1.2 Experimental setup

For this experiment, the following hardware was used:

1. N200 Software Defined Radio with DBSRX2 Daughter-board
2. ADL5902 Square-law detector
3. National Instruments USB-6009 Data Acquisition Unit

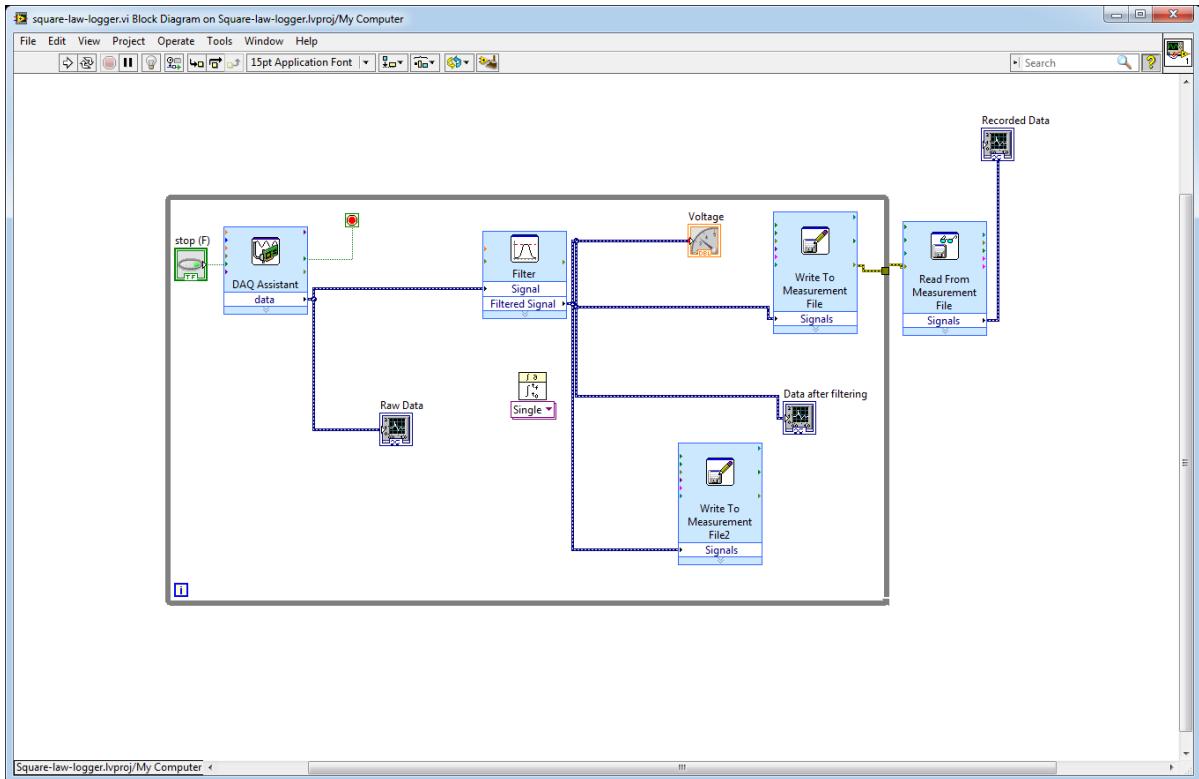


Figure 5.3 A screenshot of the Labview block diagram. Source: Labview

4. ZN2PD-20-S+ Power Divider
5. 50-ohm matched load
6. Rigol DP832 Power Supply
7. ISU Radiometer RF Front End

The ISU Radiometer, as shown in figure 5.5 was used as it has the necessary Low Noise Amplifiers (LNAs) to provide the needed amplification. The only item powered for these experiments was the LNAs, all other components are passive or were deactivated. The ISU Radiometer uses the following hardware:

1. 4 x Integrated Microwave Bandpass filters (1400 - 1425 MHz)
2. 2 x Miteq AMF-3F-01400147-30-10P LNA

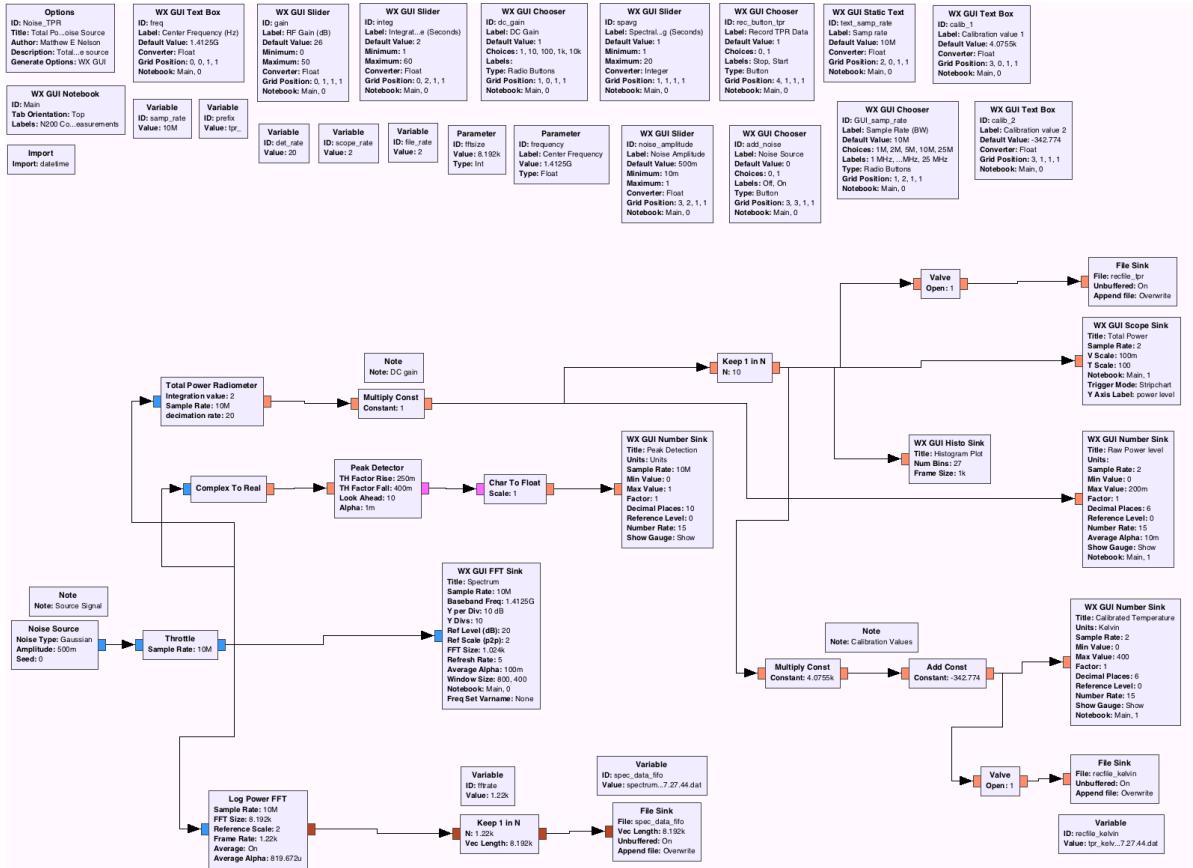


Figure 5.4 A screenshot of the GNURadio program using a noise generator block. Source: GNURadio

3. 1 x Miteq AMF-2F-01400147-04-10P LNA

Figure 5.6 shows a block diagram of the experimental setup. A matched load is used to simulate our antenna or source. This matched load is then submerged in temperature baths. The physical temperature of this matched load is then the noise temperature the radiometer sees and can be used to calibrate the radiometer. The ISU radiometer provides the amplification needed for our experiments and finally we divide that signal between the ADL5902 square-law detector and the N200 Software defined radio. The N200 is then connected to a personal computer running XUbuntu Linux and GNURadio.

The 50 ohm load was then submersed into a hot or cold bath. These baths were temperature controlled first by using Liquid Nitrogen (LN2) which is known to boil at 77 Kelvin and an ice

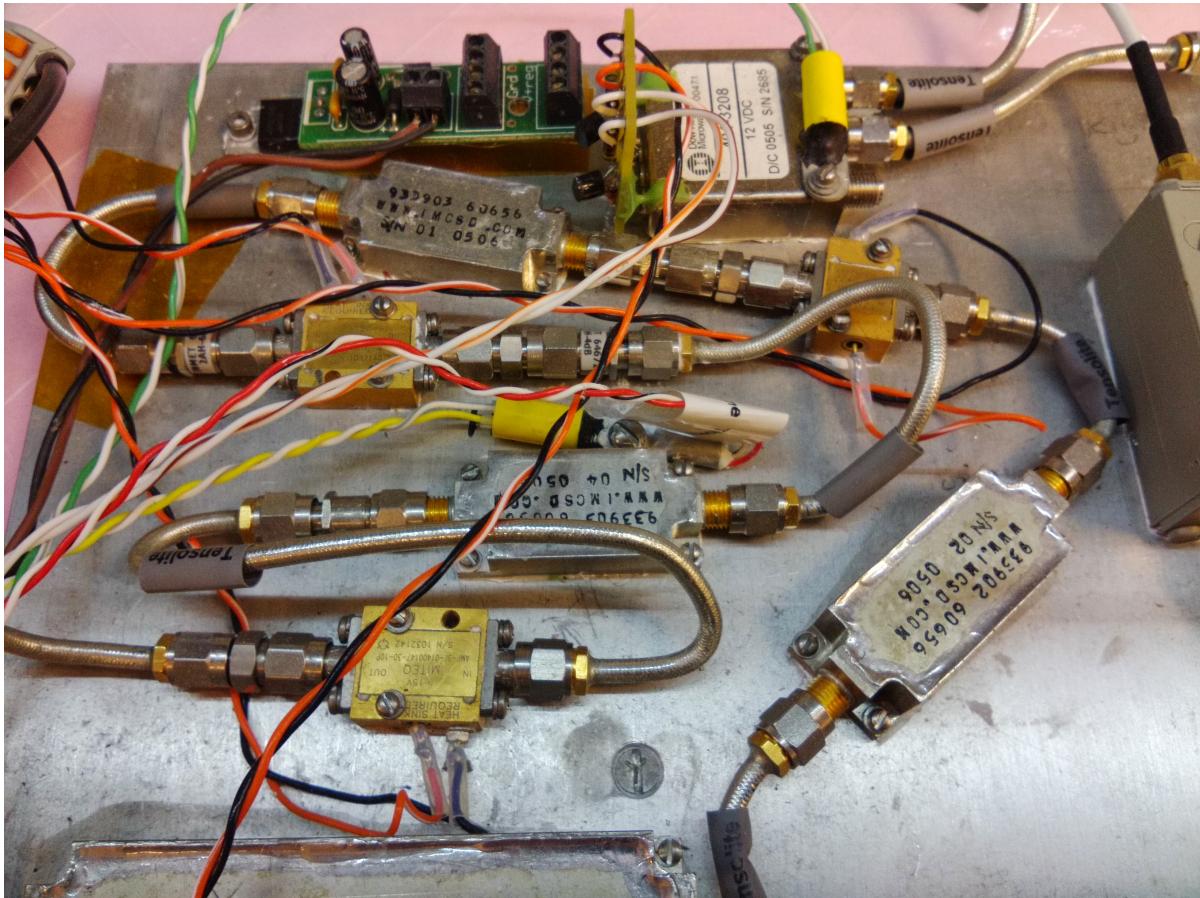


Figure 5.5 The ISU radiometer's LNAs and bandpass filters used in most experiments. Source: Matthew E. Nelson

water bath which is known to be at 273.15 Kelvin. The temperature of these could easily be monitored and maintained. The load was submersed in each bath for a minimum of 2 minutes to allow for it to reach the same temperature as the bath. The noise temperature measured was then recorded using the setup described.

5.2 Experiment II - Verification of sensitivity and stability

In this experiment we will verify the sensitivity and stability of a software defined radiometer. This experiment will verify that a software defined radiometer is able to meet the expected sensitivity and stability within an acceptable range.

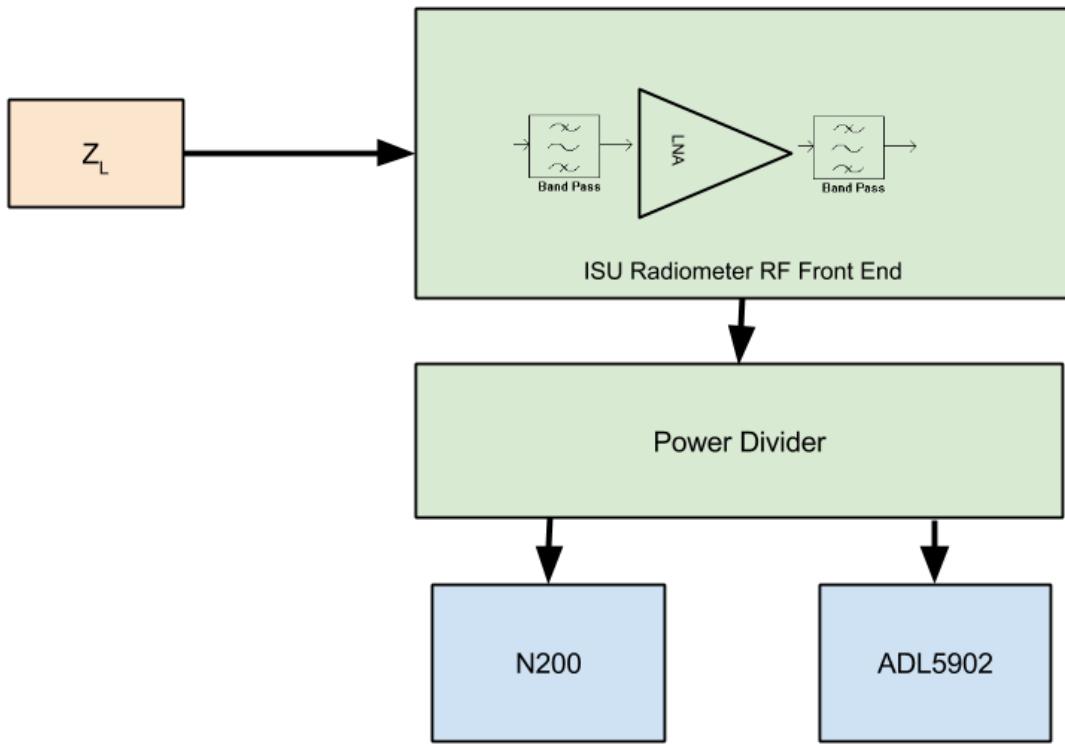


Figure 5.6 Block diagram of Experiment 1 setup. Source: Matthew E. Nelson

5.2.1 Data Collection

The data collected for this experiment was the total power measurements made from the software defined radiometer. These measurements uses the same method as outlined in section 5.1.1.

5.2.2 Experimental setup

5.3 Experiment III - Interfering Signal Mitigation

In this experiment we will generate an interfering signal and then mitigate the signal using a software defined filter. A square-law detector will be hooked up in parallel to measure the same signal but will have no mitigation methods for the signal. We will then compare the two signals to verify that the software defined radiometer can mitigate an interfering signal and still make total power measurements.

The addition of an unwanted signal can be a determinant to the radiometer and has an adverse effect on how the radiometer operates. In today's world though, it is getting more difficult to control intentional radiators as the RF spectrum becomes crowded with more devices [Ellingson et al. (2003)]. This problem becomes even a greater issue with radiometers used in orbiting space crafts as they are able to see large areas[Misra et al. (2012)]. Even though the band we are working in of 1.4 GHz is an internationally protected frequency, there can still be both intentional and unintentional radiators that cause interference.

RFI detection and mitigation is not a new topic in radiometry and there have been other methods in both the detection and mitigation of these signals. [Forte et al. (2013)][McIntyre and Gasiewski (2012)][De Roo et al. (2007)].

This test was designed to test whether or not a software defined radiometer would be able to cope with an interfering signal. This test injected a known signal at 1.406 GHz to interfere with the normal operation of the radiometer. The amplitude of this signal is then incremented and decremented at various times in the test. This was done to reflect a possible real world scenario and to make it easy to identify the interfering signal with the square-law detector which only measures power.

In order to mitigate this offending signal, a filter is designed to remove the offending signal. An initial design of the filter was designed using a program that is part of the GNURadio software package, GNU Radio Filter Design tool. Figure 5.7 shows a screen shot of the tool when designing the band-reject filter for this application. This tool generates filter values also called taps that GNURadio will use for defining the filter. The GUI program shown in figure 5.7 allows us to interactively create a filter. Because this tool is part of the GNURadio package, it also includes a command line interface to the program. This allows us to call the program from within GNURadio to integrate this functionality into our software designed radiometer.

5.3.1 Data Collection

The data collected for this experiment includes both the total power measurements from the software defined radiometer and the square-law data. Section 5.1.1 explains in detail the setup and configuration of the equipment used to collect this data.

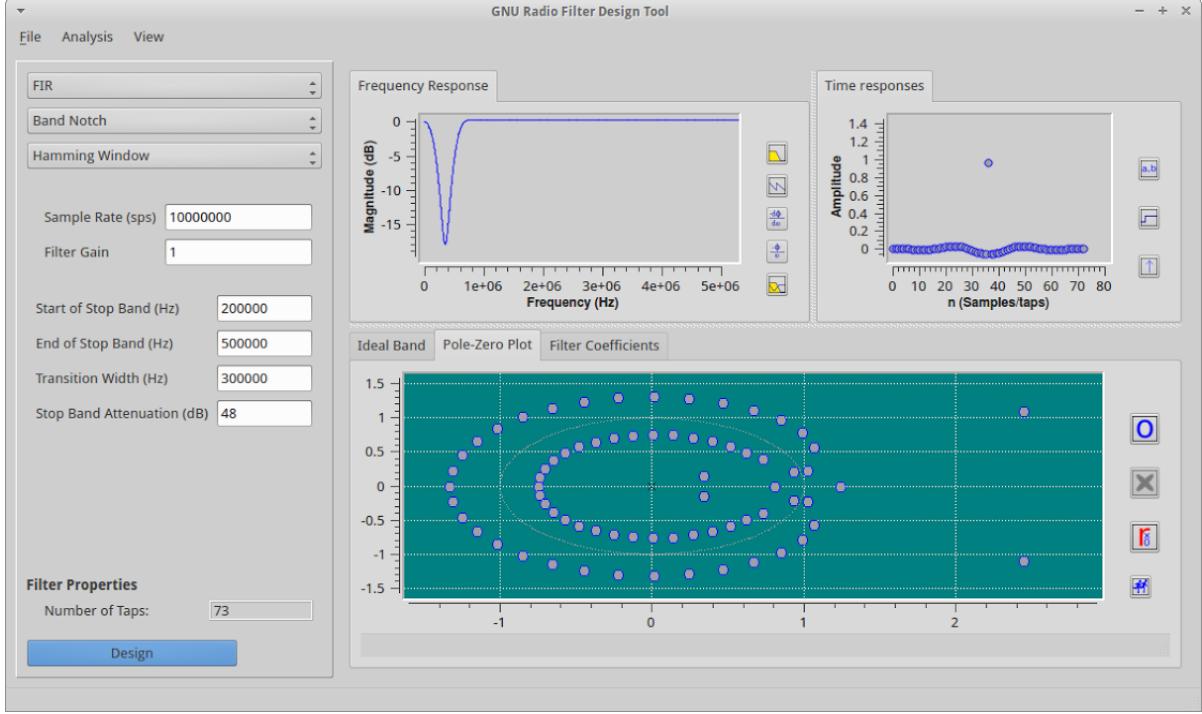


Figure 5.7 Image of the GNU Radio Filter Design tool

5.3.2 Experimental setup

The experimental setup for this experiment is similar to the setup outline in section 5.1. In addition to the setup outline in that section, another software defined radio was added and injected into the RF signal chain. This software defined radio is then configured to operate like a signal generator and generates the offending signal.

Our signal generator is the HackRF One shown in figure 5.8. This SDR is a cheaper SDR and thus has lower specifications than the Ettus Research N200 used as the software defined radiometer. However, for the task of a signal source, it will suit our needs.

For this experiment, the HackRF will generate a sinusoidal wave at a fixed frequency and with the amplitude changed at different times during the experiment. This is controlled from a program called *osmocomsiggen*. *Osmocom* was originally developed to communicate with *OsmocomSDR* hardware.

In this experiment we will examine the impact of filtering an interfering signal on the performance of the software defined radiometer. We will specifically look at the NE Δ T of

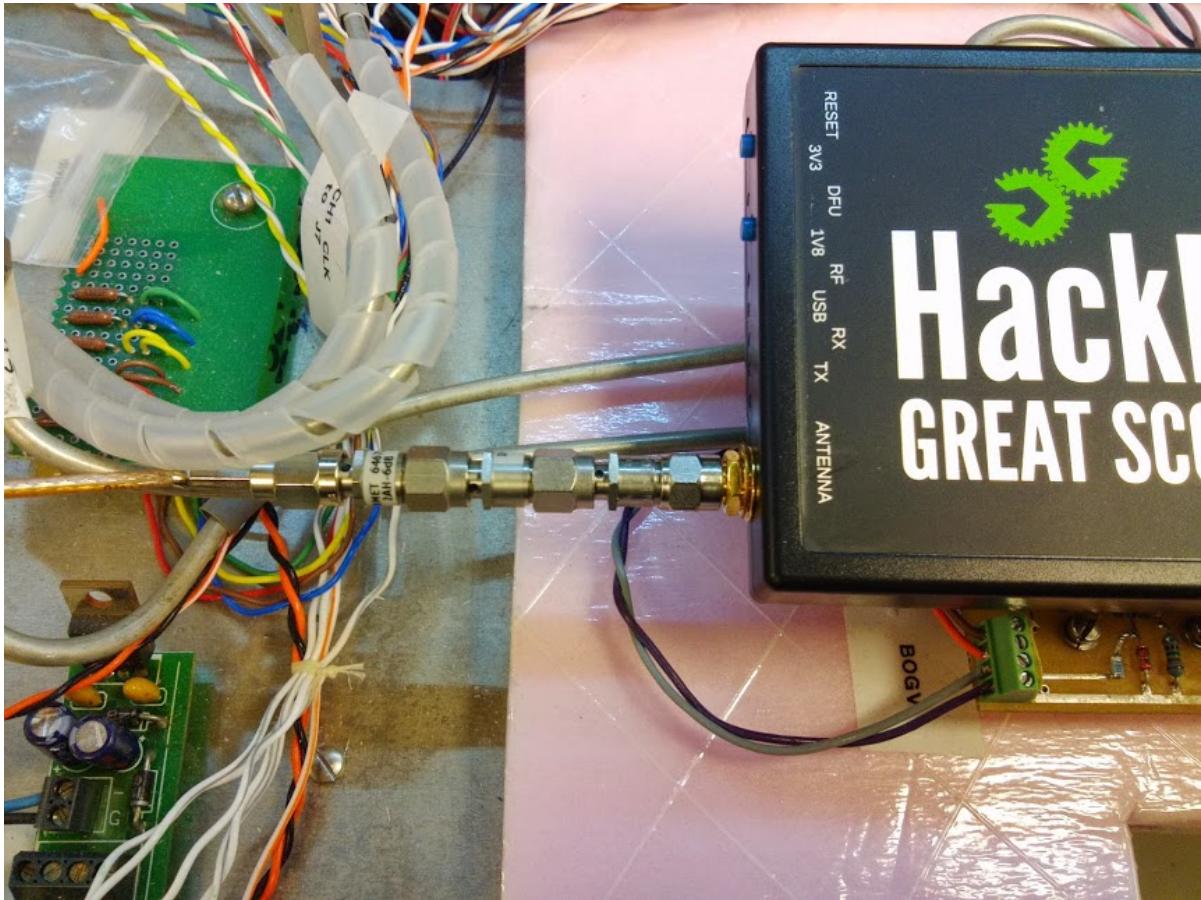


Figure 5.8 Image of the HackRF used to generate the offending signal

the signal and also how reducing our overall bandwidth affects our total power received to the radiometer.

5.4.1 Data Collection

The data collected for this experiment is the total power reading measurements obtained from the software defined radio. This data is identical to the method used in section 5.1.1.

5.4.2 Experimental setup

There is a trade-off when using filters to remove the offending signal. The larger the filter is that is used, the more information that is lost in the signal. Equation 5.1 shows how the filter impacts the $NE\Delta T$ for the system. Figure 5.1 shows that as the bandwidth of the filter

goes up, the $NE\Delta T$ also increases which results in the radiometer to be not as sensitive and may not be able to detect a fine enough change in the noise temperature. Figure 5.1 also shows that if we put a limit on what we expect our sensitivity to be, then that limits how much we can filter out. In the case shown in 5.1, we placed a limit of .2 Kelvin as the maximum $NE\Delta T$ we wished to have. This meant that we can only filter up to 5 MHz before we exceed this value.

$$NE\Delta T = \frac{T_A + T_N}{\sqrt{(\beta - \beta_{filter}) * \tau}} \quad (5.1)$$

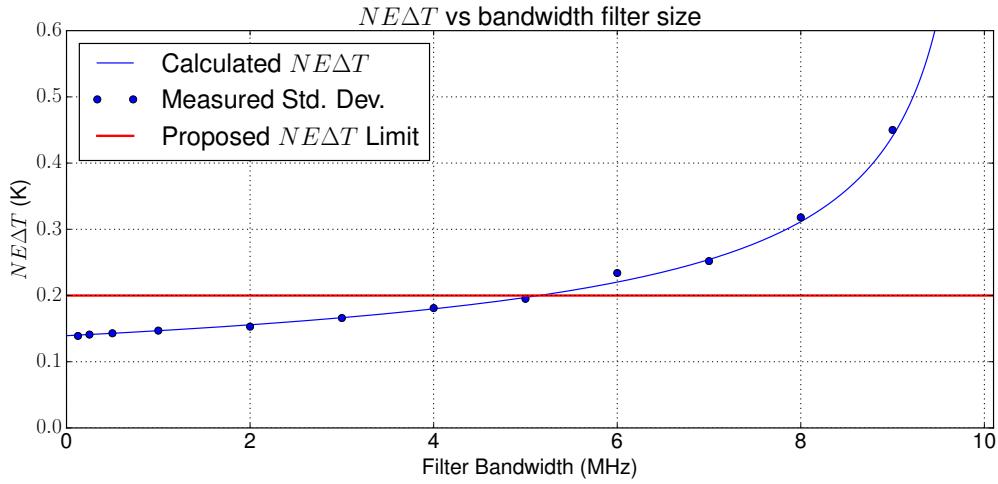


Figure 5.9 Graph showing the impact of the filter to the $NE\Delta T$ performance of the radiometer

Another impact that the filter has is how much power we expect to see at the radiometer. We will refer to equation ?? that shows the power received as a function of the bandwidth, gain, system noise temperature and measured noise temperature. This equation is now re-written in 5.2 to show that the bandwidth decreases as our filter bandwidth increases.

$$P = k * (\beta - \beta_{filter}) * G * (T_A + T_N) \quad (5.2)$$

A simple test was conducted that looked at what the theoretical power measured based on equation 5.2 and what was measured with our SDR radiometer configuration. Figure 5.10 shows the results of this test.

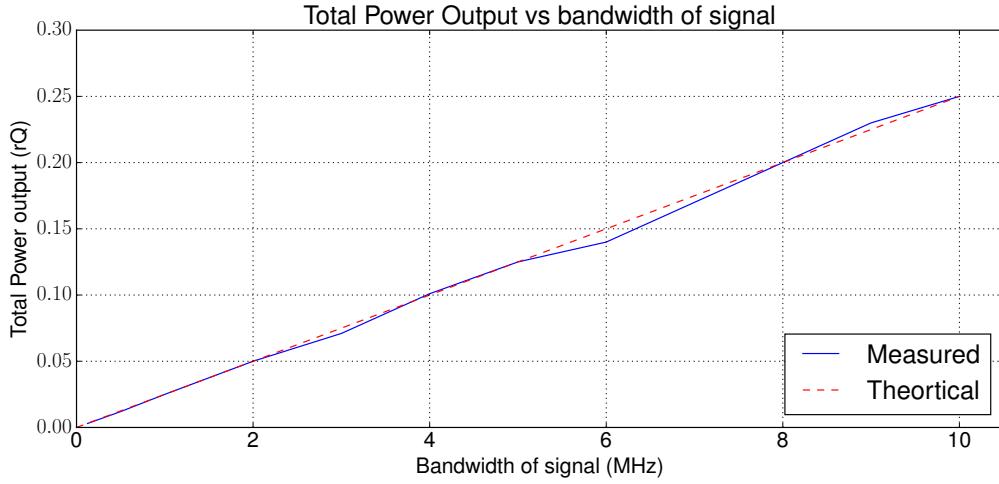


Figure 5.10 Graph showing both theoretical and experimental data on the impact to the total power as a function of the bandwidth of the filter

As expected as the system increases the bandwidth of the filter, the total bandwidth seen is decreased and this results in a decrease in the total power that we might expect at the radiometer. This means that at a certain point we may not be able to obtain the power information we need.

Filtering has other impacts to the signal that needs to be addressed. Filtering a signal does have a cost associated with it in terms of the computational cost. This depends on how tightly we need to have the filter and the sampling rate at which we are running the software defined radio. In an ideal world we would have filters that have very steep frequency response rates. However, this is computationally very expensive and so we often compromise by having less steep frequency response due to computational limits.

Another factor with filters that does not directly impact the system but does need to be discussed is delays. In the experiments in this thesis the graphs were compensated for delays due to the filters used. A FIR filter that is used to filter out the offending signal has a very predictable delay and is computed in equations. Another type of filter that is used is the IIR filter used in the power detection stage of the total power radiometer software. An IIR filter however does not have a linear delay time and is instead based on the filter and the frequency it is filtering. There are tools however that can be used to calculate this delay and this allows for us to properly compensate any delay with filtered signal. This is important for doing comparison

graphs with the square-law detector which does not use as many filters and does so at different frequencies.

5.4.3 Test Results

We begin by looking at what happens to our total power readings with no filter applied. As we stated in the Lab Setup portion, the frequency of the offending signal will not change, but the amplitude will. This will mean that we should see clear indications of the total power changing as the amplitude of the offending signal changes.

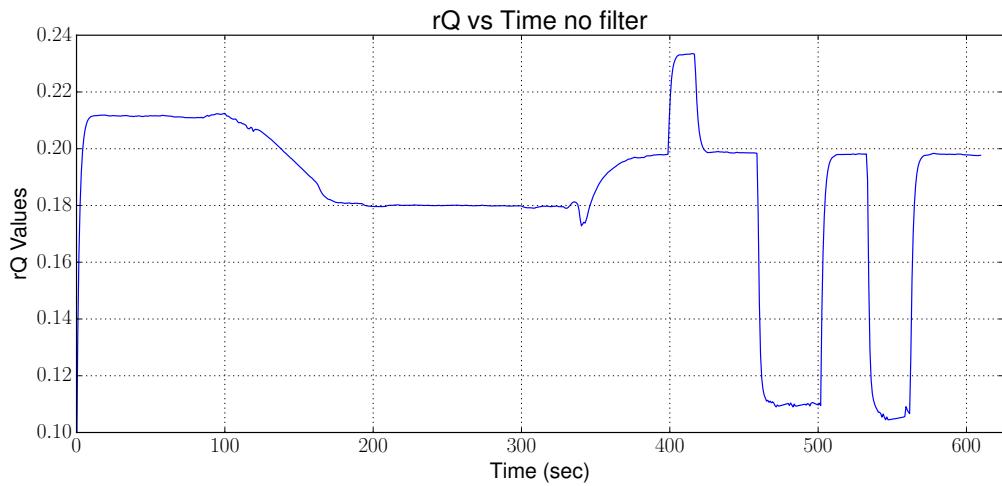


Figure 5.11 Graph showing the unfiltered total power measurements on the software defined radio

We can see that there are pulses that occur in the graph in figure 5.11 and that changes in the amplitude of the offending signal affect our total power readings. These same pulses can also be seen in the square-law detector data as well which can be seen in figure 5.12.

We can clearly see in both the software defined radio and the square-law detector that there is an interfering signal. If we now look at the spectrum view on the software defined radio, we can see the signal in question which is at 1.405 GHz. Figure 5.13 shows us what the software defined radio sees which is a spike at 1.405 GHz. The square-law detector of course has no frequency information, so our only method to detect an interfering signal is by looking at the total power readings. In our case there is elevated readings and we can see the spikes in the

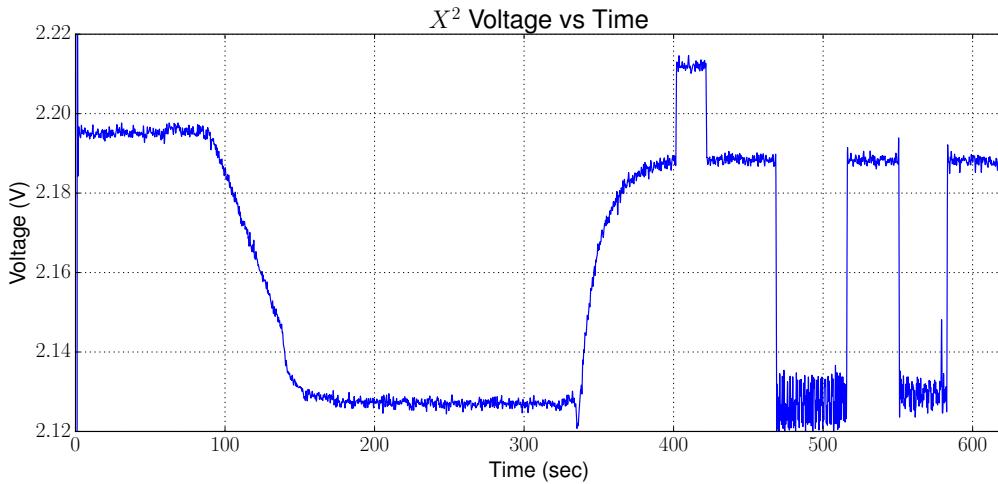


Figure 5.12 Graph showing the raw total power read from the square-law detector with an interfering signal.

square-law data. However, we do not know where in the spectrum the signal is at.

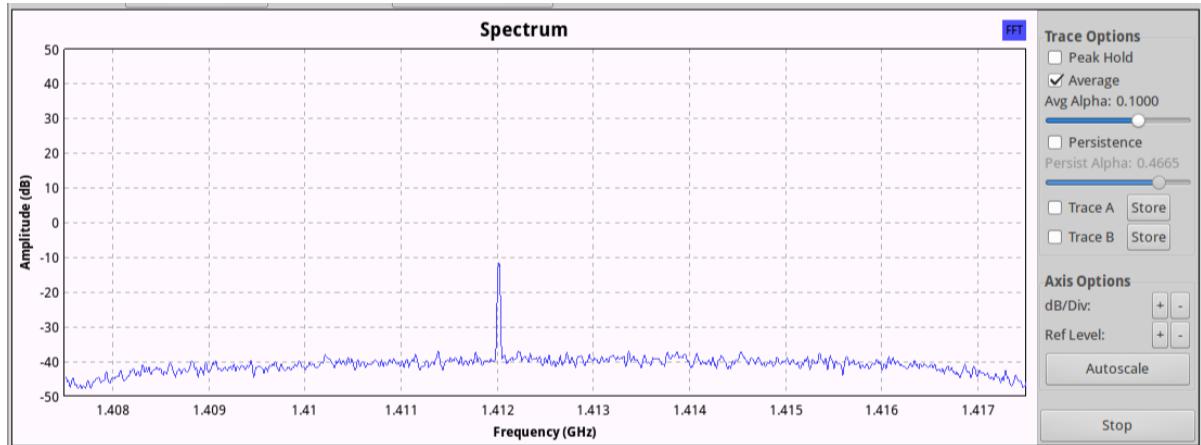


Figure 5.13 Image showing the spectrum view from the software defined radio

Since we know where the offending signal is, we can now design our filter to filter out this signal. In our GNURadio program we can specify both the frequency and the bandwidth our band-reject filter that we desire. Ideally we want to keep the bandwidth of the filter as tight as possible to the offending signal but we also want to make sure our filter is effective in removing the signal. Figure 5.14 shows the spectrum display from the software defined radio with the filter now turned on and filtering the offending signal.

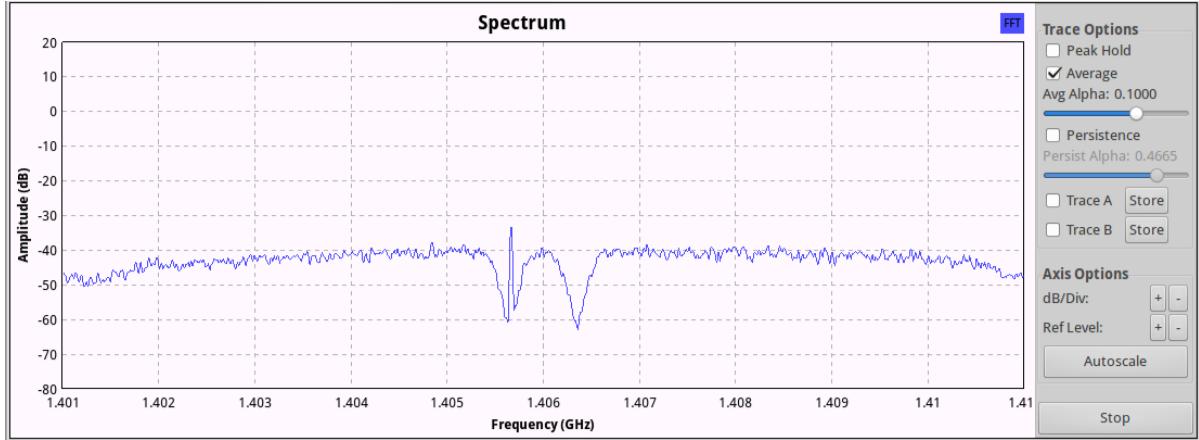


Figure 5.14 Image showing the software defined radio with the filter on and filtering the offending signal

Since we have now removed the offending signal, we will want to re-run our experiment and once again compare the difference between the software defined radio and the square law detector. We can begin by looking at the software defined radio total power readings. In figure 5.15 we can see a calibrated graph of the noise temperature seen by the software defined radio. This graph is very similar to the graphs we expect from our total power radiometer. However, we want to compare this to our square-law detector as well.

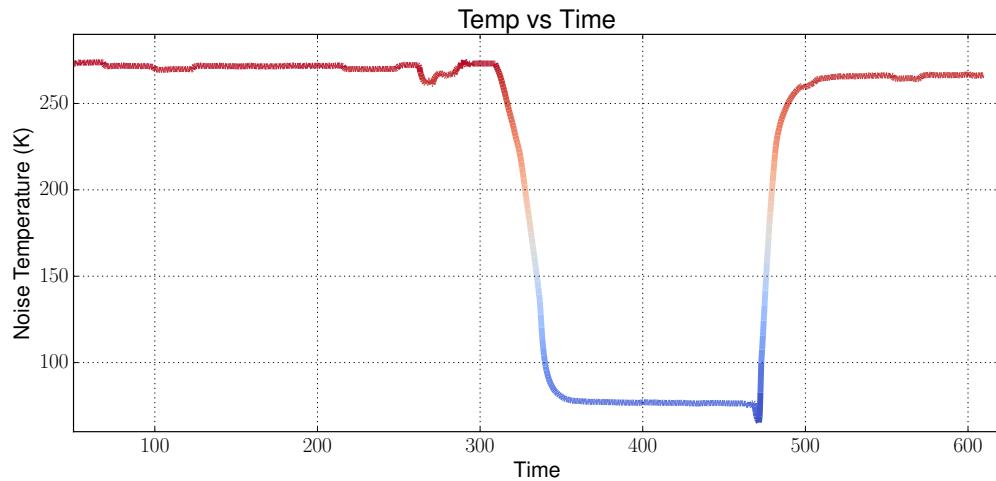


Figure 5.15 Graph showing the calibrated total power readings with the filter removing the offending signal

Figure 5.16 now shows both the software defined radio and the square-law detector calibrated

total power readings. In this graph you can see that the software defined radio is able to make normal readings where the square-law detector still shows the changes in amplitude which would make both calibration and obtaining useful data difficult.

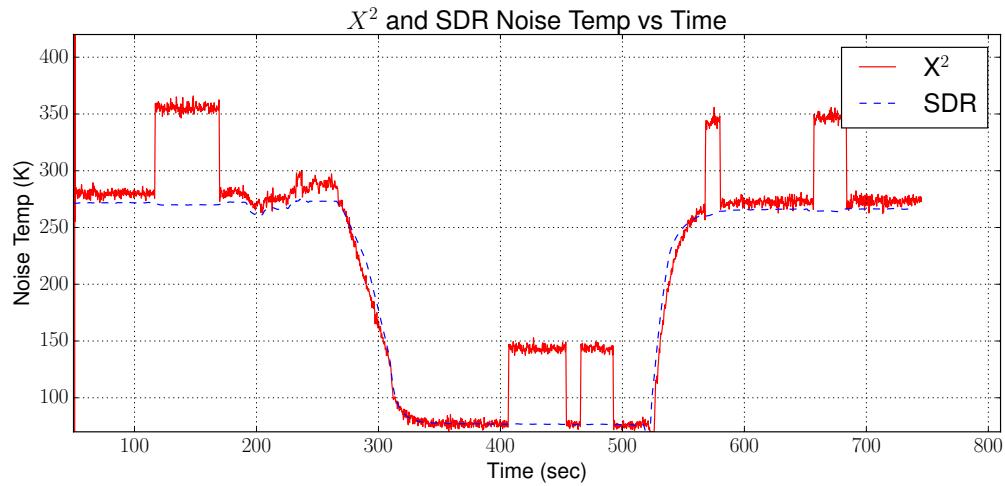


Figure 5.16 Image of the offending signal being filtered out by the SDR. It can be seen that the signal is no longer visible.

CHAPTER 6. RESULTS AND ANALYSIS

This chapter will display the results obtained from the experiments outlined in Chapter 5. We will then give an analysis of the data included data interpretation and the information learned from each experiment. Next we will do an analysis of a traditional radiometer vs our software defined radiometer in terms of price, weight, and quality. Finally we will summarize our results and analysis from this chapter.

6.1 Experiment I - Software Defined Radiometer Verification and Calibration

As outlined in Chapter 5, Section 5.1, this experiment is to verify the operation of a software defined radiometer. This is done by performing experiments that are similar to verification and calibration methods used on a traditional radiometer. By comparing these results to a square-law detector, receiving the same signal, we can verify the operation of the software defined radiometer.

6.1.1 Data Collected

The first data we will look at with be with the software defined radio. The SDR records the total power measurements to a binary file that either Matlab or Python can read. In our results, we used iPython Notebook to read and generate the graphs used in this thesis. We will begin by looking at the raw total power readings which we will call raw Q values or rQ values. These are the uncalibrated total power readings from the software defined radio.

Figure 6.1 shows the total power reading versus time and is also marked when the matched load was submerged in Ice Water, LN2, or a hot water bath. Since we know what the tempera-

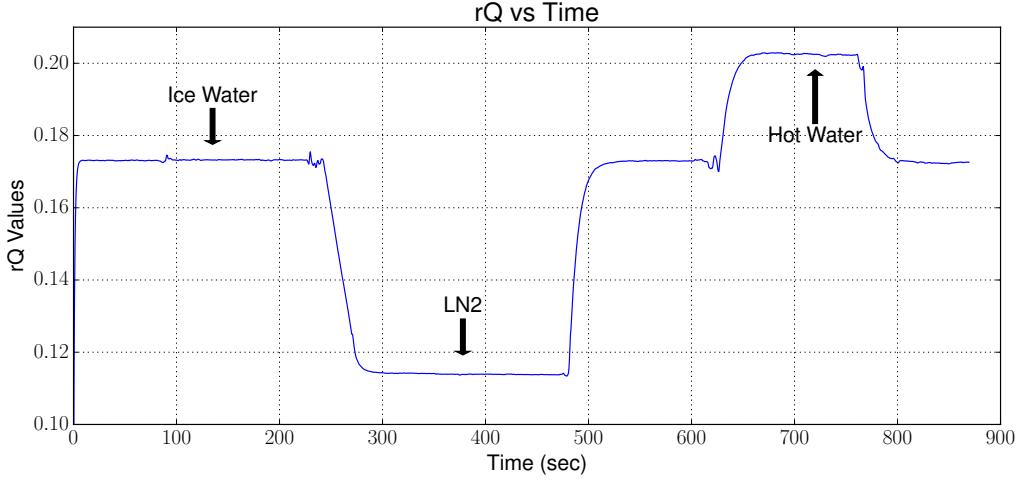


Figure 6.1 Graph of the un-calibrated rQ values of Experiment 1

tures are for the ice water and LN2, we can now calibrate these readings to a noise temperature reading. This is done by reading in a calibration file we have stored in csv format and performing linear algebra to solve the slope of the line. This was done in our iPython Notebook using the following code.

```
a = numpy.array ([[rQ_values[0], 1.0], [rQ_values[1], 1.0]], numpy.float32)
b = numpy.array ([temp_values[0], temp_values[1]])
z = numpy.linalg.solve(a, b)
```

Now that we have our calibration points, we can now re-graph this data but now calibrated as a reference noise temperature. Figure 6.2 shows a calibrated graph of the data in relation to noise temperature. In addition, we have colorized this graph to show warmer to cooler temps. This is helpful as we often refer to these as noise temperatures.

Now that we have looked at the software defined radio data, we want to look at the square-law detector with the end result of comparing the two. The square-law detector gives us power information as a voltage, so once again we will need to calibrate this to the known temperature references. However, before that we need to do one other step with the data. Unlike the software defined radio, there is no filter or integrator to help smooth out the data, so the square-law data is very noisy. Our first step will be to filter the data. Figure 6.3 shows our raw data that we

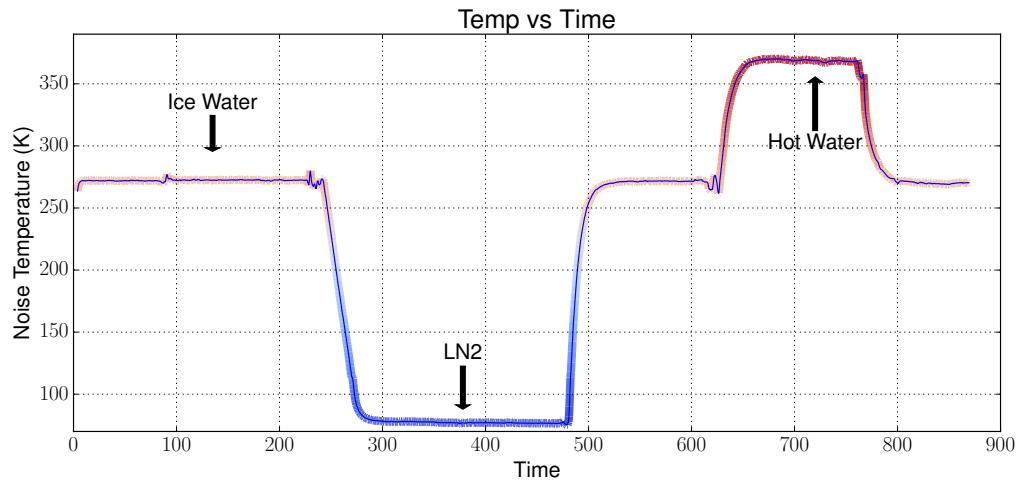


Figure 6.2 Graph of the calibrated noise temperature of Experiment 1

get from the square-law detector.

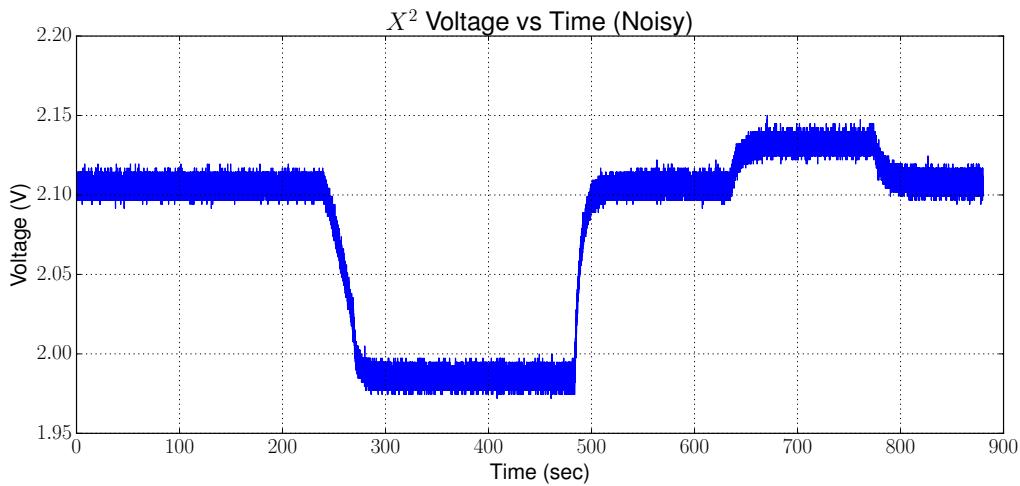


Figure 6.3 Raw and noisy graph from the square-law detector used in Experiment 1

Once again we can use Python to process this information and specifically we can use SciPy which includes several useful signal processing modules. For our use, we will use a low pass filter to clean up the signal. The following code allows us to do just that.

```
from scipy import signal
N=100
Fc=2000
```

```

Fs=1600
h=scipy.signal.firwin(numtaps=N, cutoff=40, nyq=Fs/2)
x2_filt=scipy.signal.lfilter(h, 1.0, x2_voltage)

```

Figure 6.4 now shows our data that has been filtered by the low pass filter. This is similar to the low pass filter that is also used by the software defined radio.

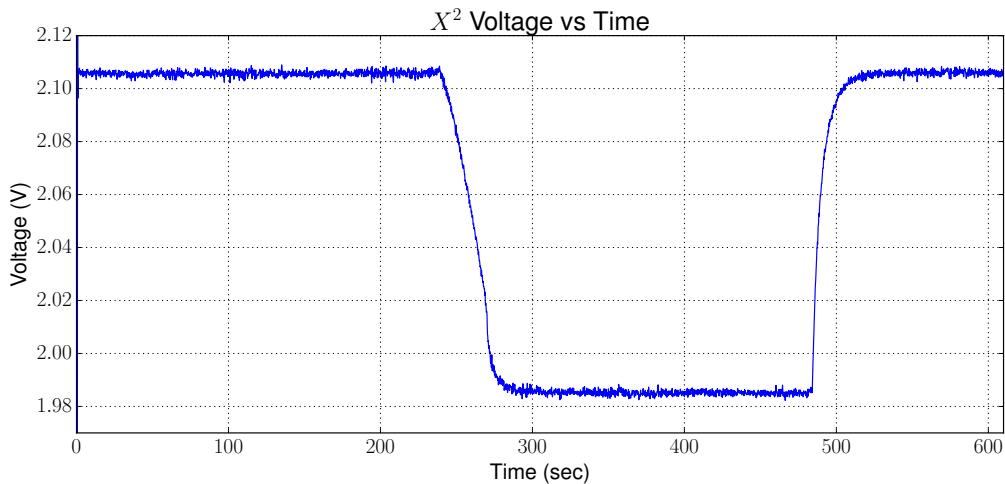


Figure 6.4 Filtered data from the square-law detector used in Experiment 1

Using the same technique as earlier, we can now calibrate the raw voltages from the square-law detector to the noise temperature. Like the software defined radio data, we can calibrate the voltages from the square-law detector to the physical temperature that our matched load is placed in. *Figure 6.5* shows the data from the square-law detector calibrated to our known temperature points. This now allows us to directly compare the square-law detector to the software defined radio data since we have a common reference point in which to compare the data.

We now want to compare both the Software Defined Radio and the square-law to make sure that they match up. Since both the SDR and the square-law are now calibrated to a noise temperature, we can easily graph both of the data and compare to see how well they match up.

We can see in *Figure 6.6* that both the software defined radio and the square-law detector match up very nicely. This shows that both the square-law detector and the software defined radio agree when properly calibrated. This verifies that the software defined radio can indeed

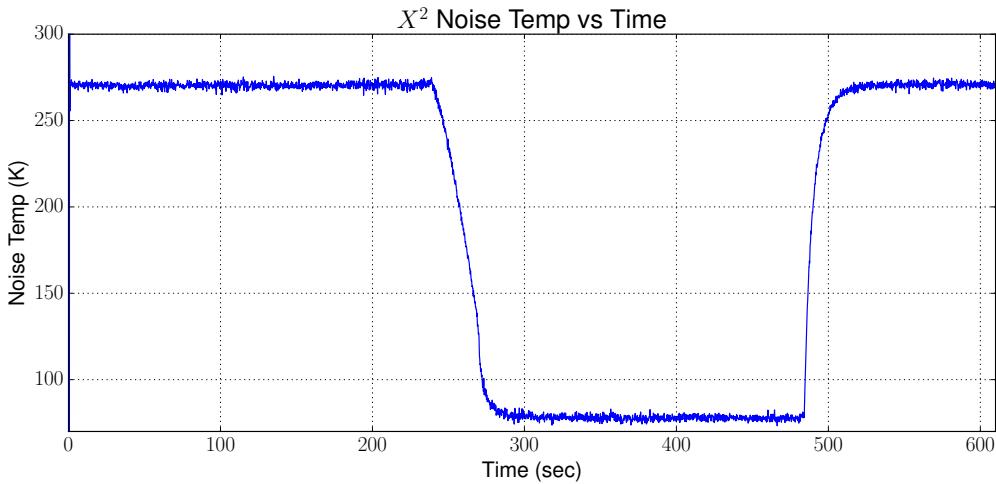


Figure 6.5 Calibrated data from the square-law detector used in Experiment 1

operate as a total power radiometer and the data we obtain from this setup agrees with an analog and more traditional radiometer.

6.2 Data Analysis

Matlab is one tool we can use to process the information that is stored by GNURadio. Appendix A contains the Matlab source that will read the total power file generated by GNURadio. It then calculates information such as the $NE\Delta T$ and the calibration points based on the user input. We can also use Matlab to graph this information as well.

While Matlab is one tool, other tools can be used. Python for example is also capable of reading in these files and when paired with NumPy and SciPy can be used to perform analysis on the data as well [Uengtrakul and Bunnjaweht (2014)]. In addition, the open source mathematical program Octave should also be able to read and work with these files. For this thesis both Matlab and Python was used to provide analysis on the data.

Most of the graphs generated in this thesis was generated using iPython notebooks. IPython notebooks uses Python but allows it to be executed in a web browser either locally or on a server. Using iPython notebooks however also allows us to add additional information using Markdown and basic HTML. This allowed the author to paint a complete picture of the experimental results illustrating pictures of the setup and the code and steps used to analyze the data. You can find

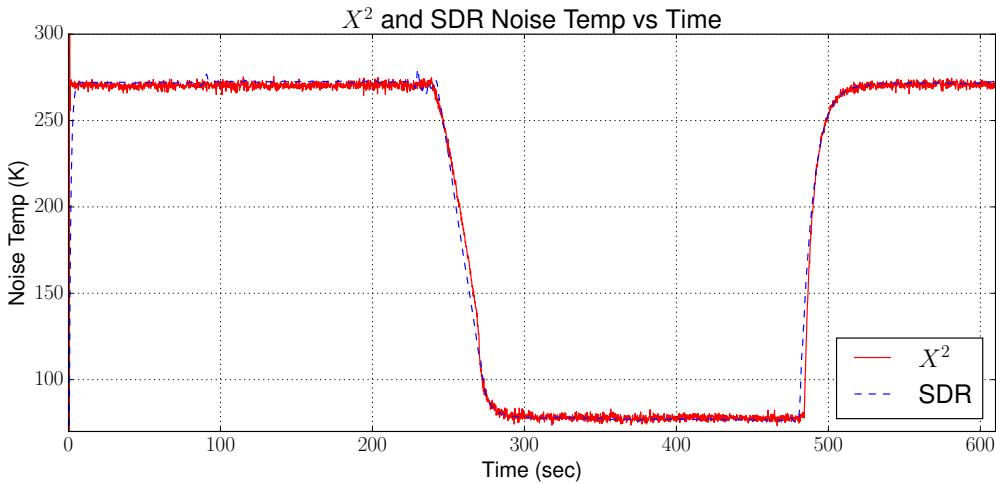


Figure 6.6 Figure showing both the SDR and square-law noise temperature data in Experiment 1

these notebooks on the author's Github site and can use NBViewer to view them. An example link is .

6.2.1 Application with Soil Moisture Readings

A common application of radiometers is in the measurement of soil moisture. All items naturally emit RF energy due to the random excitation by the electrons in the object. The amount of noise that gets generated varies by temperature, but the amount that reaches the antenna varies by the amount of moisture in the soil. If we can calibrate the radiometer to two known soil conditions, then we can measure the various levels of soil moisture in the soil. At this time, we will simply look at the percentage of moisture in the soil, which will vary from zero percent or dry soil to one hundred percent or very wet soil. The drier the soil, the more thermal noise we receive and the "warmer" the noise temperature. Wet soil on the other hand attenuates the thermal noise and shows up as a "cooler" noise temperature.

Using the Software Defined Radio, we can set up two methods to visually look at the noise temperature and thus the soil moisture percentage. Since we do not have an antenna hooked up, we will simulate this by using two reference temperatures. In this case the ice water bath and the LN2 that we just used and was shown earlier.

```
In [3]: tpr = 'tpr_2014.06.12.Lab0.dat'
Uses SciPy to open the binary file from GNURadio

In [4]: f = scipy.fromfile(open(tpr),dtype=scipy.float32)
Because of the valve function in GNURadio, there are zeros that get added to the file. We want to trim out those zeros.

In [5]: f = numpy.trim_zeros(f)
Create an index array for plotting

In [6]: y = numpy.linspace(0,1,numpy.size(f))
```

Plot the data

```
In [7]: plot(y,f)
xlabel('time')
ylabel('rQ Values')
title('rQ vs Time')
grid(True)
```

Figure 6.7 A screenshot showing the iPython notebook code and related graphs generated for parsing GNURadio data

A unique visual aid we can use with GNURadio is a waterfall display. This display gives us information that includes frequency, amplitude and time. It is referred to as a waterfall display due to the fact the display continually moves from top to bottom and looks like a waterfall. Amplitude information is given by mapping the range of amplitudes to a color bar. Frequency is given on the x-axis and time is on the y-axis. Figure 6.8 shows a screen shot of the waterfall display in the GUI created in GNURadio Companion for this experiment. This is the same program we have used but with the waterfall display now added. The data for the waterfall is pulled directly from our source block or in our case the N200 software defined radio.

Let's take a look at two screen shots of the waterfall. We will put them side by side so we can better compare the display when the thermal load is in the ice water bath and when the load is in LN2.

In figure 6.9 we can see that the ice water screen shot appears warmer than the right side

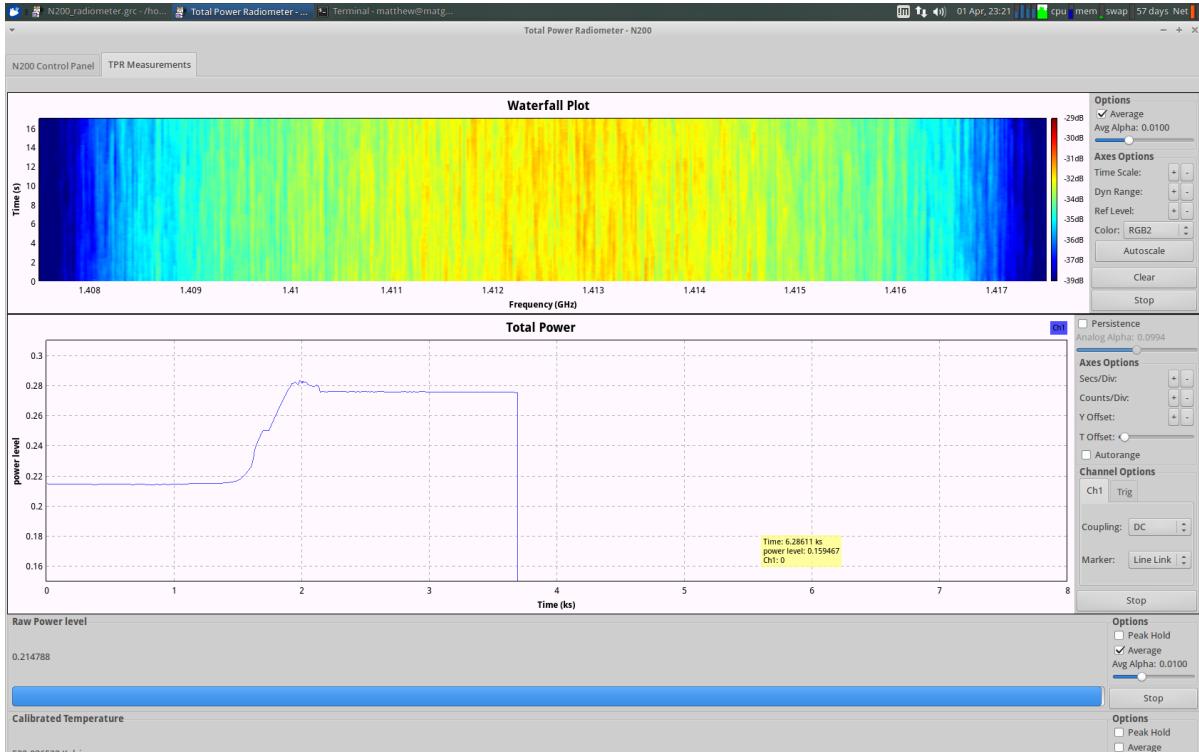


Figure 6.8 Screenshot of the waterfall display used in Experiment 1

screen shot which appears cooler. There is a limitation in GNURadio and the waterfall display that limits what the range for the power readings are. The overall power that we see only changes by about 3 dB and our current range in the waterfall display is set to 10 dB. If we could reduce the range, the color change would be even greater and more pronounced. However, this does show that a change can be seen visually with color to indicate the noise temperature.

If we assume that our LN2 is dry soil and that our ice water bath is wet soil, we can now

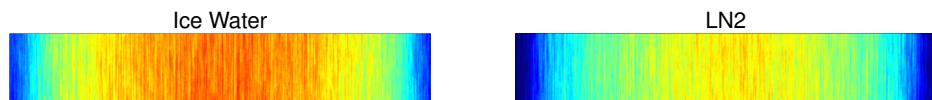


Figure 6.9 Side by side comparison of the waterfall display for Experiment 1

interpolate the data to this scale and show the information we obtained in Experiment one as both a noise temperature and soil moisture. Figure 6.10 shows the same data we looked at earlier but we have now added a scale to show soil moisture as a percentage.

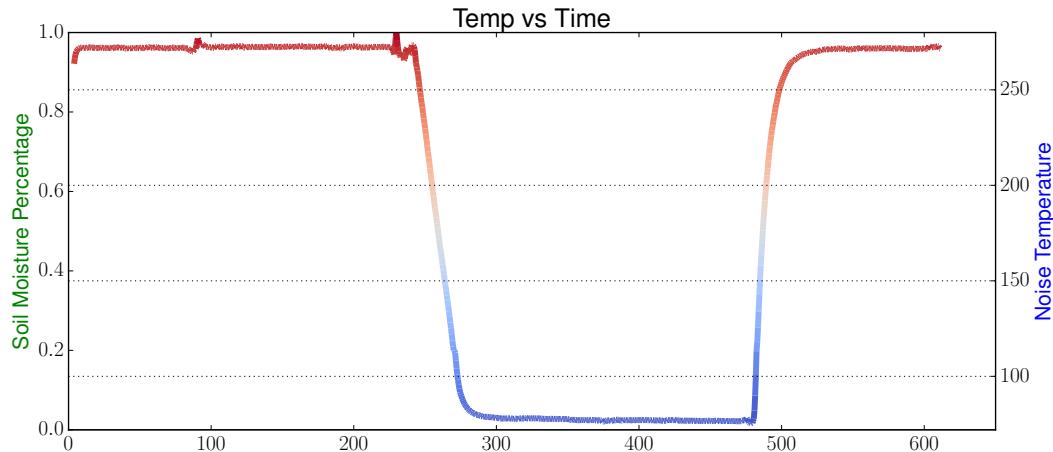


Figure 6.10 Plot of the noise temperature of Experiment 1 with soil moisture percentage

While this demonstrates that we can collate our total power readings with a soil moisture percentage, we would use actual field tests to calibrate the radiometer. In addition, we could also calibrate to soil moisture content instead of a percentage if desired. Both methods have been done with traditional radiometers[Jonard et al. (2011)][Shi et al. (2003)].

6.3 Experiment II - Verification of sensitivity and stability

6.3.1 Data Collected

6.3.2 Data Analysis

To verify stability of the radiometer, we look to see how much change the radiometer records over a relatively long period of time. To test this a matched load was submerged in a liquid nitrogen bath for an extended period of time, in this case for fifteen minutes. The readings were then looked at to study the trend of the data. The data is graphed in figure 6.11.

As it can be seen in figure 6.11, the amount of change over a period of one hour is quite

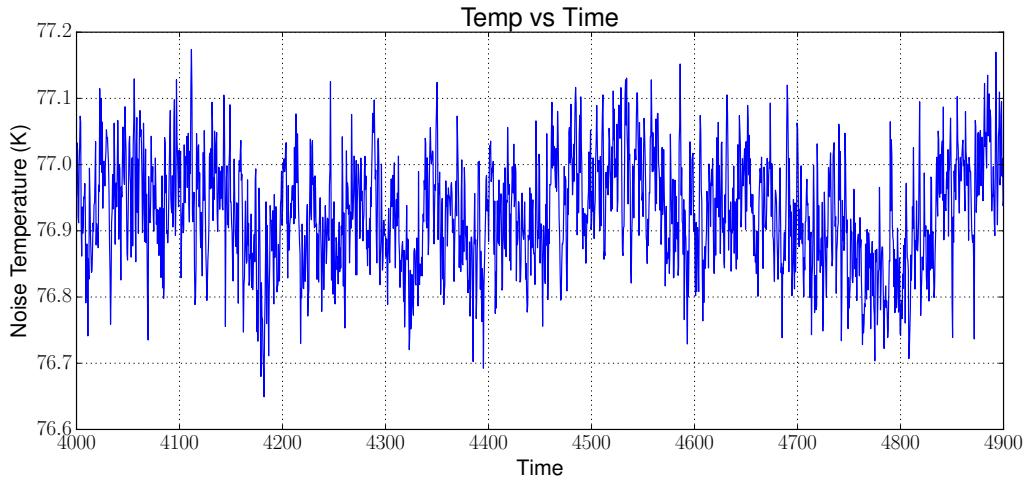


Figure 6.11 Graph of the calibrated total power over a period of fifteen minutes.

small. The standard deviation for this sample is 0.09 kelvin. The $NE\Delta T$ calculated using 10 MHz for the bandwidth, an integration time of 2 seconds and with our sample at 77 Kelvin is calculated to be 0.10 Kelvin with a system temperature of 350 Kelvin. Therefore, our system is behaving as we expect it to for this stability test.

Figure 6.12 shows a graph of the total readings but with the standard deviation now plotted. This shows that the system is stable and is operating as expected within our expected $NE\Delta T$.

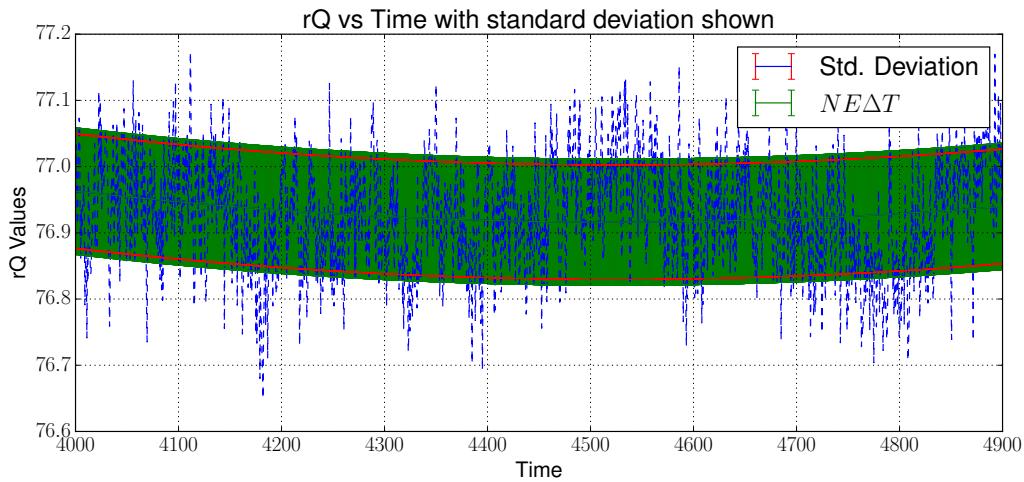


Figure 6.12 Graph of the calibrated total power with the standard deviation plotted.

6.4 Experiment III - Interfering Signal Mitigation

6.4.1 Data Collected

6.4.2 Data Analysis

6.5 Experiment IV - Performance impact of interfering signal mitigation

6.5.1 Data Collected

6.5.2 Data Analysis

6.6 Benefits to Software Defined Radio Radiometer

A study was conducted on what benefits a software defined radio radiometer would have over a more traditional radiometer. This was focused on looking at three main areas; cost, weight and size, and the value a SDR radiometer can add over traditional radiometers.

6.6.1 Cost Benefits

Software defined radios have become more commonplace in recent years and this has generated a number of Commercial Off The Shelf (COTS) solutions. A COTS solution is often a

lower cost solution due to the mass manufacturing that takes place. This has driven the cost of many SDRs to under one thousand dollars while still having excellent performance characteristics. The N200 SDR purchased for this research cost fifteen hundred dollars and the daughter-board cost one hundred and fifty dollars approximately. Other software defined radios however have come out on the market since then. Ettus for example has some that are below one thousand dollars and the author has also obtained the HackRF One SDR that now sells for three hundred dollars. The main difference with the different software defined radios on the market is with both the resolution, or how many bits the ADC is, and the bandwidth they are able to handle.

Table 6.1 Cost Analysis

<i>Device</i>	<i>Quantity</i>	<i>Cost</i>
<i>SDR Solution</i>		
<i>N200 SDR</i>	<i>1</i>	<i>\$1515</i>
<i>LNA at \$60 ea.</i>	<i>3</i>	<i>\$180</i>
<i>DBSRX2 Daughter-board</i>	<i>1</i>	<i>\$152</i>
<i>GNURadio</i>	<i>1</i>	<i>\$0</i>
<i>Total</i>		<i>\$1847</i>
<i>ISU Radiometer</i>		
<i>LNA, FPGA, ADC, Microcontroller and power supplies</i>	<i>1</i>	<i>\$10,000¹</i>
<i>Commercial Off the Shelf Unit</i>		
<i>Spectracyber 1420 MHz Hydrogen Line Spectrometer</i>	<i>1</i>	<i>\$2,650</i>

As seen in table 6.1, even the higher cost Ettus research equipment is a lower cost option than the custom built ISU radiometer purchased from University of Michigan and even a comparable off the shelf radiometer. It should be noted that the radiometer from the University of Michigan is also a dual polarization radiometer so there are two RF front ends and two ADCs that feed into a FPGA board. It would be quite easy to add dual polarization to the Ettus N200 SDR as it does support two daughter-boards. This would increase the cost to \$2,179 for the additional LNAs and daughter-board.

The largest cost benefit is that key components that you find in a radiometer, the filters and square-law detector can now be all done in software instead of needed additional equipment.

¹ Purchase price in 2005

The system is also much more frequency agile, which means it can work on a broader range of frequencies than most traditional radiometers with very little change in hardware and in some cases may require no change in hardware. Some of this does depend on the SDR hardware however. The Ettus N200 for example uses daughter-boards to provide the RF interface. While these boards provide a high quality in the RF signal, it does come at a cost and are usually designed for certain bands of frequencies. Other low cost SDRs however are also very wide range in the frequencies they will work in. The HackRF for example works from 10 MHz to 6 GHz, but does so at the cost of lower resolution, less gain in its front end and a lower bandwidth that it can handle.

6.6.2 Weight and component size benefits

A typical radiometer has many components that are involved in the design of the radiometer. This includes filters, LNAs and the power detection or square-law detector used. These components add both weight, size and costs to the radiometer. A software defined radio however digitizes the signal and we are able to replace the filters and square-law detector with their software equivalent. While a software defined radio does add both the ADC and usually a FPGA to do the processing on the signal, advances in semiconductor technology has continued to shrink these components. These components are also lighter than the filters often used in radiometers.

Table 6.2 Weight Analysis

Device	Mass
SDR Solution	
N200 SDR	1.2 kg
LNA at .03 kg ea	.09 kg
DBSRX2 Daughter-board	.1 kg
Total	1.39 kg
ISU Radiometer	
LNA, FPGA, ADC, Microcontroller and power supplies	22.7 kg
Commercial Off the Shelf Unit	
Spectracyber 1420 MHz Hydrogen Line Spectrometer	6 kg ²

Size is another benefit as since semiconductor technology has continued to shrink compo-

²Estimated, no data available

nents. Again, since items like the filters and square-law detector are removed and done in software this helps to reduce the overall size.

6.6.3 Value added benefits

A software defined radio radiometer adds additional value for two reasons. One, it is able to work with both frequency and magnitude where most radiometers do not. This allows for additional analysis on the signal and can help identify issues such as an interfering signal that was demonstrated in this thesis.

Second, we are able to have an agile system that is able to adapt to changing conditions with very little or no change to hardware. Different types of radiometers can be implemented such as a Dicke radiometer, dual polarization radiometer or a radiometer that can perform Stokes parameters. In addition, since we have both frequency and power information we can create a system that is able to adapt to changing conditions such as dealing with an interfering signal.

6.7 Disadvantages of a SDR Radiometer

Although we have outlined a number of advantages of using a COTS SDR Radiometer and how a SDR can add additional value to the radiometer system, there are some disadvantages to a SDR Radiometer.

6.7.1 Power Consumption

One of the largest drawbacks to a SDR radiometer can be in the power consumption of the SDR. With the move to perform functions such as power detection and filtering we now require additional computational power to perform these tasks. With those computational cycles additional power is now required. The use of FPGAs and SoC however can help to minimize these power concerns as they are more efficient than using a full scale x86 based processor and on board computer system.

Power and CPU requirements also increase as we add additional functionality such as filtering an offending signal. While these additions may not require additional hardware, it can

require additional processor or computational requirements. This will cause additional strain on the processor and also in the memory requirements for the SDR as well.

6.7.2 Bandwidth constraints

While SDR technology has advanced, bandwidth is still a constraint that affects SDRs and in turn a SDR Radiometer. Bandwidth plays a critical role in the radiometer's sensitivity as explained in this thesis, therefore the fact that many SDRs are limited in bandwidth does create a disadvantage. In many cases this bottleneck takes place in both the transport and processing of large bandwidth systems. This also relates to the power consumption disadvantage since larger bandwidth also means requiring additional computational cycles as well.

In contrast, a square-law detector usually has a very large bandwidth, as much as one gigahertz, and is why we usually need to filter to the frequency band of interest.

6.8 Results Summary

CHAPTER 7. CONCLUSION AND FUTURE WORK

7.1 Conclusion

In this thesis we have shown that an off the shelf SDR can be used to perform as a radiometer. Using a SDR has several advantages such as a more flexible system and can result in a less expensive system by using commercial off the shelf components. Since a SDR offers high flexibility, changes to the system can be done very quickly and helps in future proofing the system.

7.2 Future work

The work in this thesis demonstrated a basic radiometer system. There are other more complex radiometer systems that can be implemented in software. These include a correlating radiometer or a Dicke radiometer that can improve stability of the radiometer without the need of additional equipment such as thermal electric coolers.

7.2.1 Improvements

Several improvements can be done to further enhance this software defined radiometer. One improvement is to remove the personal computer (PC) and move the software defined radio to a Field Programmable Gate Array (FPGA) or an Application Specific Integrated Circuit (ASIC) solution. A second improvement is to reduce the dependency on stable Low Noise Amplifiers (LNAs) by being able to compensate for instabilities in the system.

Removing the PC. For this thesis we focused on the software that would run on a PC or comparable computer system running a full operating system like Linux. This aids speeding up development by using software tools such as GNURadio to rapidly develop the software used

to create a radiometer in software. While this works just fine for testing the theory on an off the shelf SDR acting as a radiometer, it does require hardware that is capable of running a full operating system and the associated software. For some applications of a radiometer, this is not a huge concern. In the case for the ISU radiometer, the concern is not that large since the radiometer is not designed to be portable and requires additional support equipment such as a generator anyway. However, other remote sensing applications, such as space based applications, would require a more efficient method. It is very possible to move the software generated in GNURadio into the firmware of the N200. This will help to offload the work needed by the computer and would allow for the computer or similar system to act as more of a control method and for data storage.

Improving stability. One of the largest challenges with a radiometer is improving the stability of the radiometer. Drifts in temperature can greatly affect the gains from the LNAs and also change how much noise all components in the radiometer contribute. A software defined radio can help as we are digitizing the signal as soon as possible. This helps in eliminating the analog components for power detection and even for filtering, but does not eliminate all physical hardware, mainly the LNAs. In this thesis, we did not focus on this issue since the tests were done in a controlled lab environment.

However, a more compact, lower cost and easier setup would be to just have the LNAs attach directly to the SDR without any temperature compensation. While this can be done, we have now lost stability in the LNAs and we need to compensate for that. Several methods have been discussed to handle instability in a traditional radiometer. Some of these methods would be suitable for implementation in a software defined radiometer.

A very traditional method is a Dicke radiometer. A Dicke radiometer switches between the antenna and a known noise source. A future work for a software defined radiometer would be to use a digitally generated noise source, such as a Gaussian noise source, and then switch between the antenna and this known noise source. This noise source can also be adjusted and is performed in software, therefore stability in the noise source is not an issue.

Correlation. Another method to improve stability and sensitivity is to correlate the information with another input which can be another antenna looking at the same source or can be two

polarization from the same antenna[Clapp and Maxwell (1967)][Aitken (1968)]. This results in two receiver systems looking at the same source and you have two signals, S_1 and S_2 . Since we are looking at the same source, both signals will be correlated in time, and when multiplied they will provide an output proportional to the strength of the source signal. The noise introduced by each receiver will then have a lower correlation due to the random nature of the noise. This results in a radiometer with a greater sensitivity due to the reduction of the noise even though two receivers are used [Fujimoto (1964)].

The N200 software defined radio was chosen as it is capable of having two different daughter-cards plugged in. Therefore, it is possible to have both sources enter the software defined radio and once digitized we can sum the magnitudes of the two incoming sources. This is quite easy to do and is shown in figure 7.1.

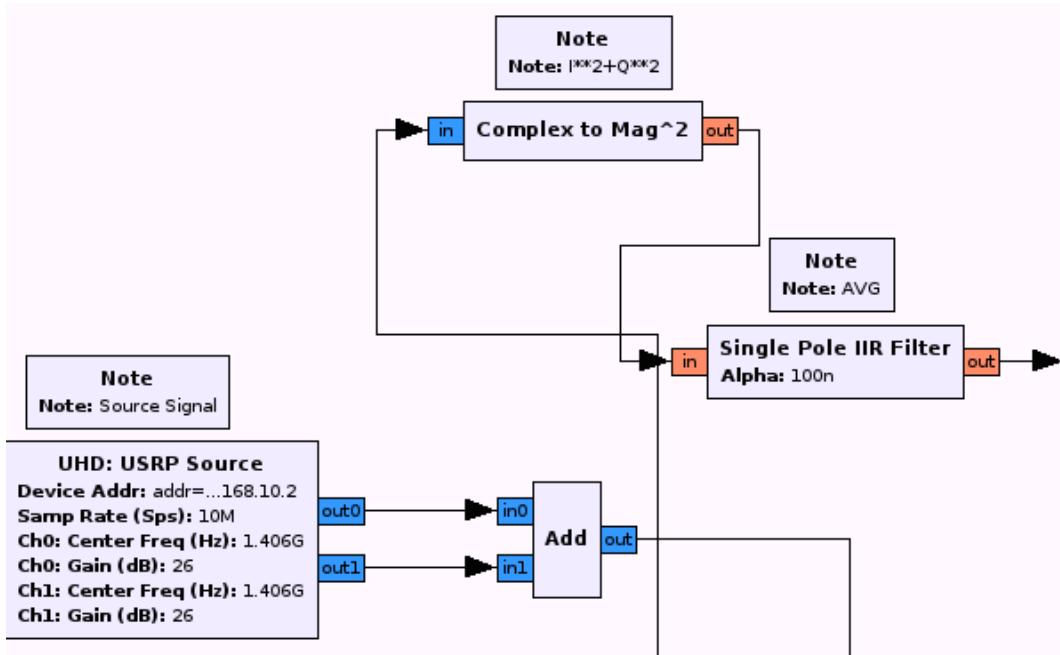


Figure 7.1 The key blocks used for creating a correlating radiometer in software. The key blocks is the USRP source, which allows us to address both daughter-boards and the ADD block which sums the signals.

Although figure 7.1 shows a correlating software defined radio radiometer, it has not been tested. In theory, this should correlate the signal and improve the sensitivity of the radiometer, however further testing is needed.

7.2.2 Further testing

The improvements and additional features outlined in Section 7.2 will require additional testing and verification. While it has been shown that a software defined radiometer operates identically to a traditional radiometer, further testing is needed to verify different operating modes of a software defined radiometer.

7.3 Closing statement

This thesis demonstrated that it is possible to use off the shelf components and a software defined radio to implement a working radiometer that can be used in various radiometer applications such as soil moisture, ocean salinity and radio astronomy. Use of a software defined radiometer can potentially allow radiometers to be used by a wider audience of users by creating an easy to user GUI and reducing the cost and hardware complexity that most radiometers require. The result is more locations that are able to use radiometers as remote sensing tools to learn more about our planet and even the cosmos.

APPENDIX A. Source code

The following is the source code to several programs that make the radiometer possible. The first is the Python code used with the Ettus N200 software defined radio. This code is a heir block, which means it can be imported into GNURadio and used as a block. Once imported you may run your source into this block and use any sink needed to either display or log the data. This code was generated using GNURadio Companion.

The third code supplied is Python code that can be used to read the data generated and plot it. In many ways it mimics the functions of the Matlab script buy uses Python, NumPy and SciPy to perform the mathematical functions. This may be a better option for those that wish to look at the data but do not have access to Matlab since Python is free to download.

Finally, this code has been included in this thesis as a point of reference. It may be out of date and some other pieces of code was also used for the experimentation used in this thesis. Copies of this thesis source L^AT_EXcode, some experimental data, and any other code used can be found on the author's GitHub repository, <https://github.com/matgyver/Radiometer-SDR-Thesis>.

Python code for total power radiometer

The main code that acts as the total power radiometer is the TPR.py code. This module adds a custom block to GRC that can then be called and used like any other block in GRC. A screenshot of the blocks used in this is shown below.

Total Power Radiometer Block

```
#!/usr/bin/env python
```



Figure A.1 Blocks used for creating a total power radiometer in software. Source: GNURadio Companion

```

#####
# Gnuradio Python Flow Graph
#
# Title: Total Power Radiometer
#
# Author: Matthew Nelson
#
# Description: Blocks for power detection, integration and LPF
#               for a total power radiometer
#
# Generated: Sun Apr 12 23:03:59 2015
#####

from gnuradio import blocks
from gnuradio import filter
from gnuradio import gr
from gnuradio.filter import firdes

class TPR(gr.hier_block2):

    def __init__(self, integ=1, samp_rate=1, det_rate=1):
        gr.hier_block2.__init__(
            self, "Total Power Radiometer",
            gr.io_signature(1, 1, gr.sizeof_gr_complex*1),
            gr.io_signature(1, 1, gr.sizeof_float*1),
        )

```

```

#####
# Parameters
#####
self.integ = integ
self.samp_rate = samp_rate
self.det_rate = det_rate

#####
# Blocks
#####
self.single_pole_iir_filter_xx_0 = filter.
    single_pole_iir_filter_ff(1.0/((samp_rate*integ)/2.0),
        1)
(self.single_pole_iir_filter_xx_0).set_processor_affinity
([1])
self.blocks_keep_one_in_n_4 = blocks.keep_one_in_n(gr.
    sizeof_float*1, samp_rate/det_rate)
self.blocks_complex_to_mag_squared_1 = blocks.
    complex_to_mag_squared(1)

#####
# Connections
#####
self.connect((self.blocks_complex_to_mag_squared_1, 0), (
    self.single_pole_iir_filter_xx_0, 0))
self.connect((self.blocks_keep_one_in_n_4, 0), (self, 0))
self.connect((self, 0), (self.
    blocks_complex_to_mag_squared_1, 0))

```

```

self.connect((self.single_pole_iir_filter_xx_0, 0), (self
    .blocks_keep_one_in_n_4, 0))

def get_integ(self):
    return self.integ

def set_integ(self, integ):
    self.integ = integ
    self.single_pole_iir_filter_xx_0.set_taps(1.0/((self.
        samp_rate*self.integ)/2.0))

def get_samp_rate(self):
    return self.samp_rate

def set_samp_rate(self, samp_rate):
    self.samp_rate = samp_rate
    self.single_pole_iir_filter_xx_0.set_taps(1.0/((self.
        samp_rate*self.integ)/2.0))
    self.blocks_keep_one_in_n_4.set_n(self.samp_rate/self.
        det_rate)

def get_det_rate(self):
    return self.det_rate

def set_det_rate(self, det_rate):
    self.det_rate = det_rate
    self.blocks_keep_one_in_n_4.set_n(self.samp_rate/self.
        det_rate)

```

det_rate)

Python code for analyzing data

*I*Python notebooks was used to perform an analysis on the data. The code presented here is the export of this notebook. The Markdown language as well as some HTML is used to create easy to read pages that include the python code, generated graphs and descriptive text.

Total Power Radiometer

```
#-*- coding: utf-8

#Radiometer Parsing Function

#This code shows an example of reading in and plotting data that
is outputted from a GNURadio GRC file.

#In this example a Total Power Radiometer is developed in
GNURadio GRC and uses the File Sink function to store the data

.

#The plot then shows the total power output from the radiometer
as a matched load is submerged in Liquid Nitrogen,
#then Ice Water and then left to dry.

# - - -
#
#### Read the data

# Import Needed functions

# Import needed libraries
from pylab import *
import pylab
import scipy
```

```
import numpy
import scipy.io as sio
import csv

# Use this to set the filename for the data file and CSV
# Calibration file.

tpr = 'tpr_2014.06.12.Lab0.dat'
calib = 'tpr_calib_2014.06.12.Lab0.csv'
x2_data = 'tpr_x2_2014.06.12.Lab0.csv'

# Uses SciPy to open the binary file from GNURadio
f = scipy.fromfile(open(tpr), dtype=scipy.float32)

# Because of the value function in GNURadio, there are zeros that
# get added to the file. We want to trim out those zeros.

# In [5]:
f = numpy.trim_zeros(f)
```

```

# Create an index array for plotting. Also, since we know the
interval the data is taken, we can convert this to an actual
time.

# In [6] :

y = numpy.linspace(0,(len(f)*.5),numpy.size(f))

#### Plot the data

# In [7] :

plot(y,f)
xlabel('Time (sec)')
ylabel('rQ Values')
title('rQ vs Time')
grid(True)

pylab.show()

# ## Calibration

# The rQ values are the raw values from the total power
radiometer and are uncalibrated. While the graph shows the
change in the total power recorded and shows that the
radiometer can detect changes in noise temperature, it has no
other meaning than that. What we want is to show what the

```

total power is in relation to a noise temperature. Since we have recorded the values of the rQ at fixed and known teperatures, we can create a calibration line and calibrate the radiometer. For this experiment, we found that the following values matched our two known temperatures.

```

#
# /rQ Value/X^2 Voltage/Temperature
# /-----/-----
# .0977    / 1.9617      /77 K
# .1507    / 2.085       /273.15 K
#
# We can now solve for y = mx + b since we have two equations and
# two unknowns.
#
# To work with this, a calibration file is created. This is a
# very simple CSV file that contains 3 values: The raw rQ value,
# the raw voltage from the square-law detector (discussed later
# ) and the observed temperature. The table above would then
# look like the following in the file.
#   ''
# .0977,1.9617,77
# .1507,2.085,273.15
#   ''
#   - - -
#
# We need to read in the values from our CSV file that contains
# the values
```

```
# In [67]:
# In [67]:



read_csv = open(calib, 'rb')
csvread = csv.reader(read_csv)
rQ_values = []
temp_values = []
voltage = []

for row in csvread:
    rQ, volt, temp = row
    rQ_values.append(float(rQ))
    voltage.append(float(volt))
    temp_values.append(float(temp))
read_csv.close()

a = numpy.array([[rQ_values[0], 1.0], [rQ_values[1], 1.0]], numpy.
float32)
b = numpy.array([temp_values[0], temp_values[1]])

z = numpy.linalg.solve(a, b)
print z

# Now we apply these values to the array that holds our raw rQ
values

g = f*z[0]+z[1]
```

```

# Now we can re-plot the graph but this time with the calibrated
noise temperatures

plt.figure()
plot(y,g)
xlabel('Time')
ylabel('Noise Temperature (K)')
title('Temp vs Time')
grid(True)

pylab.show()

# This is looking better, but the time at the bottom doesn't have
much meaning. Since we know the sample rate of the Software
Defined Radio, we can calculate the time interval between each
sample.

# - - -
# # Square-law data
#
# We now want to look at the data from the Square-Law detector to
verify the operation of the SDR. In the experiment that was
conducted above, a power splitter was used to split the RF
signal so that one went to the SDR and the other to a square-
law detector (with a 3.1 dB loss though). Therefore both data
should be the same. Let's read and then plot this data.

```

```

read_csv = open(x2_data, 'rb')
csvread = csv.reader(read_csv)
dummy = []
x2_voltage = []

for row in csvread:
    dummy, x2voltage = row
    x2_voltage.append(float(x2voltage))
read_csv.close()

# Like the SDR data, we want to have a time reference at the
bottom.

w = numpy.linspace(0,(len(x2_voltage)*.01),numpy.size(x2_voltage)
)

plt.figure()
plot(w,x2_voltage)
xlabel('Time (sec)')
ylabel('Voltage (V)')
title('X^2 Voltage vs Time (Noisy)')
grid(True)

pylab.show()

# The Square-law detector doesn't have a filter on it unlike the
data we get from the SDR. The GNURadio program takes the data

```

*and applies a Low Pass Filter to "clean up" the information.
We need to do the same with our Square-law data.*

```

from scipy import signal

N=100

Fc=2000

Fs=1600

h=scipy.signal.firwin(numtaps=N, cutoff=40, nyq=Fs/2)
x2_filt=scipy.signal.lfilter(h, 1.0, x2_voltage)

plt.figure()
plot(w,x2_filt)
xlabel('Time (sec)')
ylabel('Voltage (V)')
title('X^2 Voltage vs Time')
axis([0, 610, 1.94, 2.12])
grid(True)

pylab.show()

# Now we wish to calibrate this data as well. We will use the
# same file and use the calibration points in that file.

a = numpy.array([[voltage[0],1.0],[voltage[1],1.0]],numpy.float32
)
b = numpy.array([temp_values[0],temp_values[1]])

```

```

z = numpy.linalg.solve(a,b)
print z

x2_calib = x2_filt*z[0]+z[1]

plt.figure()
plot(w,x2_calib)
xlabel('Time (sec)')
ylabel('Voltage (V)')
title('X^2 Noise Temp vs Time')
axis([0, 610, 70, 300])
grid(True)

pylab.show()

# This looks to be the same as our SDR graph, but let's overlay
# them to make sure

plt.figure()
plot(w,x2_calib,'r',label='X^2')
plot(y,g,'b',label='SDR')
xlabel('Time (sec)')
ylabel('Voltage (V)')
title('Noise Temperature vs Time')
axis([0, 610, 70, 300])
grid(True)
legend(loc='lower right')

```

```
# We have some timeshift due to two reasons. One, the timing isn't always perfect when starting the collection of the two data sets. And two, we get a timeshift from filtering the square-law data
```

```
pylab.show()
```

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