

Canonical Quantum Field Theory

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1 Canonical Quantization of a Scalar Field

A scalar field $\phi(x)$ that obeys the Klein-Gordon equation

$$(\partial^2 + m^2)\phi(x) = 0 \quad (1)$$

also obeys the following *equal-time commutation relations* (ETCRs):

$$[\phi(t, \underline{x}), \dot{\phi}(t, \underline{x}')] = i\delta(\underline{x} - \underline{x}') \quad (2)$$

$$[\phi(t, \underline{x}), \phi(t, \underline{x}')] = [\dot{\phi}(t, \underline{x}), \dot{\phi}(t, \underline{x}')] = 0 \quad (3)$$

A scalar field may be expanded in terms of mode operators:

$$\phi(t, \underline{x}) = \phi^+(x) + \phi^-(x) \quad (4)$$

$$= \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega(\underline{p})} (a(\underline{p})e^{-ip \cdot x} + a(\underline{p})^\dagger e^{ip \cdot x}) \quad (5)$$

where the creation and annihilation operators in momentum space obey the commutation relations

$$[a(\underline{p}), a^\dagger(\underline{p}')] = 2\omega(\underline{p})(2\pi)^3 \delta(\underline{p} - \underline{p}') \quad (6)$$

$$[a(\underline{p}), a(\underline{p}')] = [a^\dagger(\underline{p}), a^\dagger(\underline{p}')] = 0. \quad (7)$$

To prevent infinite contributions to the energy, we require that all terms be *normal ordered* - that is, all annihilation operators stand to the right of all creation operators. This results in the convenient property

$$\langle 0 | : \phi(x) \phi(y) : | 0 \rangle = 0.$$

We may simplify covariant commutation relations for the Klein-Gordon field by introducing three invariant functions

$$\Delta^+(x) \equiv -i \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega(\underline{p})} e^{-ip \cdot x} \quad (8)$$

$$\Delta^-(x) \equiv \Delta^+(x)^* = i \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega(\underline{p})} e^{ip \cdot x} \quad (9)$$

$$\Delta(x) \equiv \Delta^+(x) + \Delta^-(x). \quad (10)$$

All three satisfy the Klein-Gordon equation, and we may furthermore write Δ in a manifestly covariant form:

$$\Delta(x) = -i \int \hat{d}^4p \hat{\delta}(p^2 - m^2) \theta(p_0) e^{-ip \cdot x}. \quad (11)$$

Δ also satisfies the microcausality condition

$$[\phi(x), \phi(y)] = 0, \quad (x - y)^2 < 0.$$

2 Interactions and Scattering in ϕ^3 Theory

The Lagrangian for this theory is given by

$$\mathcal{L} \equiv \mathcal{L}_0 + \mathcal{L}_I \quad (12)$$

$$= \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{1}{3!} g \phi^3. \quad (13)$$

We may write the Euler-Lagrange equation as

$$(\partial^2 + m^2) \phi = -\frac{1}{2} g \phi^2. \quad (14)$$

Since this is a nonlinear partial differential equation, we would like to proceed by solving it in terms of powers of coupling g .

We define the S-matrix as the time evolution operator between two states in the far past and the far future at which the particles involved may be considered non-interacting:

$$|\Psi, \infty\rangle = S |\Psi, -\infty\rangle. \quad (15)$$

Since in general, we must consider many possible initial and final states, we consider the S -matrix elements defined as

$$S_{fi} = \langle f | S | i \rangle \quad (16)$$

Note that the S -matrix is unitary ($S^\dagger S = S S^\dagger = 1$) even when particles in the state are destroyed and/or created.

Furthermore, in the future we will work in the interaction picture of quantum mechanics - a mix of the Schrödinger and Heisenberg pictures. We split the Hamiltonian into free and non-interacting parts $H = H_0 + H_I$ and let the time evolution of the state be governed by H_I , and the time evolution of the operators by H_0 :

$$\frac{\partial}{\partial t} O(t) = i[H_0, O(t)] \quad (17)$$

$$i \frac{\partial}{\partial t} |\psi, t\rangle = H_I |\psi, t\rangle. \quad (18)$$

Integrating Eq. 18, we obtain a recursion relation for $|\Psi, t\rangle$:

$$|\Psi, t\rangle = |\Psi, -\infty\rangle + \int_{-\infty}^t dt_1 \frac{\partial}{\partial t} |\Psi, t_1\rangle \quad (19)$$

$$= |i\rangle - i \int_{-\infty}^t dt_1 H_I(t_1) |\Psi, t_1\rangle \quad (20)$$