

**Department of Physics and Astronomy**  
**Heidelberg University**

Bachelor thesis in Physics  
submitted by

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# **Functional Renormalization and Quantum Gravity**

This bachelor thesis has been carried out by

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at the

**Institute for Theoretical Physics**

at

**Heidelberg University**

under the supervision of

**Prof. Dr. Jan M. Pawłowski**



# **Functional Renormalization and Quantum Gravity**

Mathieu Kaltschmidt

## **Abstract**

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## **Zusammenfassung**

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## Chapter 1.

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# Introduction

Throughout this thesis we use units such that  $\hbar = c = G \equiv 1$ .

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## Functional methods in Quantum Field Theory

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### 2.1. Generating Functionals and Correlation Functions

We consider a theory setting of  $N$  scalar fields  $\varphi_a(x)$ ,  $a \in \{1, \dots, N\}$  in  $d$ -dimensional Euclidean space. The corresponding partition sum in presence of sources  $J_a(x)$  reads

$$Z[J] = \int \mathcal{D}\varphi e^{-\mathcal{S}[\varphi] + J \cdot \varphi}. \quad (2.1)$$

The information content of the partition sum results mainly from the classical action functional  $\mathcal{S}[\varphi]$ , which determines the classical field equations

$$\frac{\delta \mathcal{S}}{\delta \varphi(x)} = 0. \quad (2.2)$$

**Notation:** The scalar product sums over field components and integrates over all space ...

$$J \cdot \varphi = \int_x J_a(x) \varphi_a(x) = \int_p \tilde{J}_a(p) \tilde{\varphi}_a(p) \quad (2.3)$$

with

$$\int_x = \int_{\mathbb{R}^d} d^d x \quad \text{and} \quad \int_p = \int_{\mathbb{R}^d} \frac{d^d p}{(2\pi)^d} \quad (2.4)$$

Mean field description:

$$\phi := \langle \varphi \rangle = \frac{1}{Z} \frac{\delta Z}{\delta J} \Big|_{J=0} = \int \mathcal{D}\varphi \varphi e^{-\mathcal{S}[\varphi] + J \cdot \varphi} \quad (2.5)$$

Higher correlations:

$$\langle \varphi_1 \cdots \varphi_n \rangle := \langle \varphi^n \rangle = \frac{1}{Z} \frac{\delta^n Z}{\delta^n J} = \int \mathcal{D}\varphi \overbrace{\varphi_1 \cdots \varphi_n}^{:= \varphi^n} e^{-S[\varphi] + J \cdot \varphi} \quad (2.6)$$

We obtain the Schwinger functional by taking the logarithm:

$$W[J] = \ln Z[J] \quad (2.7)$$

For the special case of  $n = 2$  the correlation function yields the connected 2-point function which is also known as the propagator  $G_{ab}(x, y)$  correlating the field  $\varphi_a$  at spacetime point  $x$  with the field  $\varphi_b$  at  $y$ .

$$\begin{aligned} G_{ab}(x, y) &= \frac{\delta^2 W[J]}{\delta J_a(x) \delta J_b(y)} = \frac{\delta}{\delta J_a(x)} \left( \frac{1}{Z} \frac{\delta Z}{\delta J_b(y)} \right) \\ &= \frac{1}{Z} \left( \frac{\delta^2 Z}{\delta J_a(x) \delta J_b(y)} \right) - \frac{1}{Z^2} \left( \frac{\delta Z}{\delta J_a(x)} \right) \left( \frac{\delta Z}{\delta J_b(y)} \right) \\ &= \langle \varphi_a(x) \varphi_b(y) \rangle - \phi_a(x) \phi_b(y) = \langle \varphi_a(x) \varphi_b(y) \rangle_c \end{aligned} \quad (2.8)$$

The Effective Action:

The effective action can be obtained by performing a Legendre transform of the Schwinger functional, i. e.:

$$\Gamma[\phi] = \sup_J \left\{ \int_x J(x) \phi(x) - \mathcal{W}[J] \right\} = \int_x J_{\text{sub}}(x) \phi(x) - \mathcal{W}[J_{\text{sub}}] \quad (2.9)$$

Quantum equation of motion:

$$\frac{\delta \Gamma[\phi]}{\delta \phi(x)} = J(x) \quad (2.10)$$

Dyson-Schwinger equation:

$$\frac{\delta \Gamma[\phi]}{\delta \phi(x)} = \frac{\delta \mathcal{S}}{\delta \varphi(x)} \left[ \varphi = G \cdot \frac{\delta}{\delta \phi} + \phi \right] \quad (2.11)$$

## 2.2. The Functional Renormalization Group

- Kadanoff Block-Spin model

- maybe visualization of Ising model + phase transitions

### 2.2.1. Renormalization group consistency

Cutoff independence of the full quantum effective action:

$$\Lambda \frac{d\Gamma}{d\Lambda} = 0 \quad (2.12)$$

Full effective action in a generic representation:

$$\Gamma[\phi] = \mathcal{D}_\Lambda[\phi] + \Gamma_\Lambda[\phi] \quad (2.13)$$

Formal discussion:

$$\Gamma_k[\phi] = \Gamma_\Lambda[\phi] + \int_\Lambda^k \frac{dk'}{k'} \mathcal{F}_{k'}[\phi] \quad (2.14)$$

## 2.3. The Renormalisation Group Flow for the Effective Action

We introduce the RG time scale  $t$ :

$$\partial_t = \frac{\partial}{\partial \ln(k/\Lambda)} = \frac{k}{\Lambda} \frac{\partial}{\partial (k/\Lambda)} = k \partial_k \quad (2.15)$$

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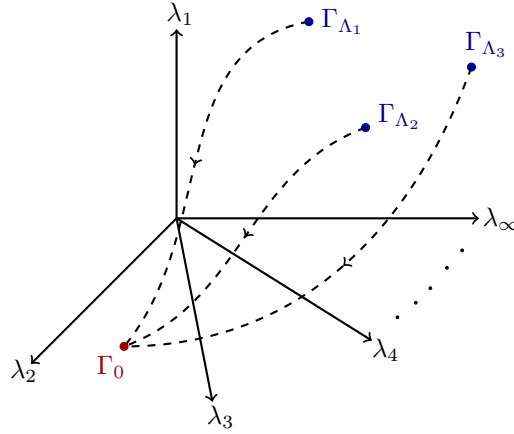
$$\begin{aligned} \partial_t \Gamma_k[\phi] &= \frac{1}{2} \text{Tr} \left[ \frac{1}{\Gamma_k^{(2)}[\phi] + R_k} \partial_t R_k \right] \\ &= \frac{1}{2} \int_p \frac{1}{\Gamma_k^{(2)}[\phi] + R_k} (p, -p) \partial_t R_k(p^2) \end{aligned} \quad (2.16)$$

This translates directly into the following diagrammatic representation:

$$\partial_t \text{ (shaded circle) } = \frac{1}{2} \text{ (circle with } \otimes \text{) }$$

where  $\otimes = \partial_t R_k$  represents the insertion of the respective regulator.

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**Figure 2.1.:** Flow of  $\Gamma_k$  through infinite-dimensional theory space for different regulators, inspired by [5]

## Fundamentals of General Relativity

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### 3.1. The Einstein Equations

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The Einstein-Hilbert action:

$$\mathcal{S}_{\text{EH}}[g_{\mu\nu}] = \frac{1}{16\pi G} \int_x \sqrt{-\det g_{\mu\nu}} (\mathcal{R} - 2\Lambda) \quad (3.1)$$

Varying this action as usual yields the Einstein equations in absence of matter:

$$G_{\mu\nu} + \Lambda g_{\mu\nu} = 0 \quad (3.2)$$

where we used  $G_{\mu\nu} = \mathcal{R}_{\mu\nu} - \frac{1}{2}g_{\mu\nu}\mathcal{R}$ .

Diffeomorphism invariance, Lie derivatives:

$$\mathcal{L}_\omega \phi = \omega^\mu \partial^\mu \phi = \omega^\mu \nabla^\mu \phi \quad (3.3)$$

Now we include matter.

Energy-Momentum Tensor:

$$T_{\mu\nu} = \frac{-2}{\sqrt{-\det g_{\mu\nu}}} \frac{\delta \mathcal{S}_{\text{matter}}}{\delta g^{\mu\nu}} \quad (3.4)$$



Matter part of the action for a minimally coupled scalar field  $\phi$ :

$$\mathcal{S}_{\text{matter}}[g_{\mu\nu}, \phi] = -\frac{1}{2} \int_x \sqrt{-\det g_{\mu\nu}} (g^{\mu\nu} \nabla_\mu \phi \nabla_\nu \phi - g_{\mu\nu} V(\phi)) \quad (3.5)$$

From this, we get the Einstein equations including matter by demanding the variation  $\sqrt{-\det g_{\mu\nu}} \frac{\delta \mathcal{S}}{\delta g^{\mu\nu}}$  to vanish. This yields:

$$\frac{1}{8\pi G} \left[ \mathcal{R}_{\mu\nu} - \frac{1}{2} (\mathcal{R} - 2\Lambda) g_{\mu\nu} \right] = T_{\mu\nu} \quad (3.6)$$

### 3.2. Perturbative Non-Renormalizability of Gravity

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# Quantum Gravity in the Einstein-Hilbert Truncation

## 4.1. RG approach to Quantum Gravity

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## 4.2. Truncations of the theory space

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## 4.3. The effective action for Quantum Gravity

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## **4.4. Non-Gaussian Fixed Points**

# Conclusions and Outlook

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## Mathematical Appendix

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### A.1. Heat Kernel techniques

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### A.2. York decomposition

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## Numerical Implementation

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### B.1. Determination of the Fixed Points

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# List of Figures

- 2.1. Flow of  $\Gamma_k$  through infinite-dimensional theory space for different regulators 6



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## Declaration of Authorship

I hereby certify that this thesis has been composed by me and is based on my own work, unless stated otherwise.

Heidelberg, \_\_\_\_\_