
Non-perturbative aspects of gauge theories

Exercise sheet 12 – Quantum Gravity in the Einstein-Hilbert truncation

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Exercise 20: Quantum Gravity in the Einstein-Hilbert truncation

In this exercise we investigate quantum gravity in the Einstein-Hilbert truncation,

$$\Gamma_k = 2\kappa^2 Z_k \int d^4x \sqrt{g} [-R + 2\Lambda_k] . \quad (1)$$

As a further simplification we only take contributions from the transverse-traceless spin-two mode of the graviton into account, i.e., we neglect the other graviton as well as the ghost modes. Due to this approximation, we never have to specify a gauge-fixing action. Start from transverse-traceless graviton two-point function

$$\Gamma_{h^{\text{TT}}h^{\text{TT}}}^{(2)} = \frac{Z_k}{32\pi} \left(\bar{\Delta} - 2\Lambda_k + \frac{2}{3}\bar{R} \right) , \quad (2)$$

define the regulator as

$$R_k = \Gamma_{h^{\text{TT}}h^{\text{TT}}}^{(2)} \Big|_{\Lambda_k=\bar{R}=0} \cdot r_k \left(\frac{\bar{\Delta}}{k^2} \right) , \quad (3)$$

with the Litim-type cutoff

$$r_k(x) = \left(\frac{1}{x} - 1 \right) \Theta(1 - x) . \quad (4)$$

Evaluate now the trace over the Laplace operator on the right-hand side of the Wetterich equation

$$\text{Tr} \frac{1}{\Gamma_{h^{\text{TT}}h^{\text{TT}}}^{(2)} + R_k} \partial_t R_k , \quad (5)$$

with heat-kernel techniques, see next page as well as Appendix G.1 of the lecture notes for details.

Turn now to the left-hand side of the Wetterich equation and take a scale derivative of (1). Compare the terms proportional to \sqrt{g} and $\sqrt{g}\bar{R}$ from the left-hand side with the result from the right-hand side, (5). Deduce from this the flow equations of the Newton coupling and the cosmological constant. The resulting flow equations are

$$\begin{aligned}\partial_t g_k &= (2 + \eta_g) g_k, \\ \eta_g &= -\frac{5}{6\pi} g_k \left(2 \frac{1 - \frac{1}{6}\eta_g}{(1 - 2\lambda_k)^2} + \frac{1 - \frac{1}{4}\eta_g}{1 - 2\lambda_k} \right), \\ \partial_t \lambda_k &= -4\lambda_k + \frac{\lambda_k}{g_k} \partial_t g_k + \frac{5}{4\pi} g_k \frac{1 - \frac{1}{6}\eta_g}{1 - 2\lambda_k}.\end{aligned}\tag{6}$$

Bonus question 1:

Why is the spin-two approximation rather good in most cases?

Bonus question 2:

Use mathematica and find numerically the non-Gaussian fixed point of (6) and determine the eigenvalues of the stability matrix.

Heat-kernel techniques

Heat-kernel techniques are used to evaluate the trace of a function that depends on the Laplace operator on a curved background. You can use the formula

$$\text{Tr } f(\Delta) = \frac{1}{(4\pi)^2} \left[B_0(\Delta) Q_2[f(\Delta)] + B_2(\Delta) Q_1[f(\Delta)] \right] + \mathcal{O}(R^2),\tag{7}$$

with the definition

$$Q_n[f(x)] = \frac{1}{\Gamma(n)} \int dx x^{n-1} f(x).\tag{8}$$

The B_n are called heat-kernel coefficients and often written as

$$B_n(\Delta) = \int d^4x \sqrt{g} \text{Tr } b_n(\Delta).\tag{9}$$

The values of the heat-kernel coefficients depend on the field. For the transverse-traceless tensor (TT), transverse vectors (TV) and scalars (S) on a sphere, they are given by

	TT	TV	S
Tr b_0	5	3	1
Tr b_2	$-\frac{5}{6}R$	$\frac{1}{4}R$	$\frac{1}{6}R$

Solution

We need to evaluate

$$\begin{aligned}
\text{Tr } G_{h\text{TT}} \partial_t R_k &= \text{Tr } \frac{\Delta(\partial_t r_k(\Delta) - \eta_g r_k(\Delta))}{\Delta(1 + r_k(\Delta)) - 2\Lambda_k + \frac{2}{3}\bar{R}} \\
&= \text{Tr } \frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{\Delta(1 + r_k(\Delta)) - 2\Lambda_k} - \frac{2}{3}\bar{R} \text{Tr } \frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{(\Delta(1 + r_k(\Delta)) - 2\Lambda_k)^2}
\end{aligned} \tag{10}$$

With the heat-kernel formulas we obtain for example for the first term

$$\begin{aligned}
\text{Tr } \frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{\Delta(1 + r_k(\Delta)) - 2\Lambda_k} &= \frac{1}{(4\pi)^2} \left(B_0(\Delta) Q_2 \left[\frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{\Delta(1 + r_k(\Delta)) - 2\Lambda_k} \right] \right. \\
&\quad \left. + B_2(\Delta) Q_1 \left[\frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{\Delta(1 + r_k(\Delta)) - 2\Lambda_k} \right] \right) \\
&= \frac{1}{(4\pi)^2} \int d^4x \sqrt{g} \left[5 \Phi_2^1(-2\Lambda_k) - \frac{5}{6}\bar{R} \Phi_1^1(-2\Lambda_k) \right]
\end{aligned} \tag{11}$$

where the Φ_m^n are the threshold functions

$$\begin{aligned}
\Phi_n^p(\omega) &= \frac{1}{\Gamma(n)} \int_0^\infty dz z^{n-1} \frac{z(-2z r_k(z) - \eta_g r_k(z))}{(z(1 + r_k(z)) + \omega)^p} \\
&= \frac{1}{\Gamma(n)} \frac{1}{(1 + \omega)^p} \left(\frac{2}{n} - \frac{\eta_g}{n(n+1)} \right).
\end{aligned} \tag{12}$$

In the last line we have evaluated the threshold functions for the Litim-type cutoff (4). So the term in (11) becomes

$$5 \Phi_2^1(-2\Lambda_k) - \frac{5}{6}\bar{R} \Phi_1^1(-2\Lambda_k) = 5 \frac{1 - \frac{\eta_g}{6}}{1 - 2\lambda_k} - \frac{5}{6}\bar{R} \frac{2 - \frac{\eta_g}{2}}{1 - 2\lambda_k}. \tag{13}$$

For the second term in (10)

$$\begin{aligned}
-\frac{2}{3}\bar{R} \text{Tr } \frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{(\Delta(1 + r_k(\Delta)) - 2\Lambda_k)^2} &= -\frac{2}{3} \frac{\bar{R}}{(4\pi)^2} B_0(\Delta) Q_2 \left[\frac{\Delta(-2\frac{\Delta}{k^2} r'_k(\Delta) - \eta_g r_k(\Delta))}{(\Delta(1 + r_k(\Delta)) - 2\Lambda_k)^2} \right] \\
&= -\frac{2}{3} \frac{\bar{R}}{(4\pi)^2} \int d^4x 5 \Phi_2^2(-2\Lambda_k) \\
&= -\frac{10}{3} \frac{\bar{R}}{(4\pi)^2} \int d^4x \frac{1 - \frac{\eta_g}{6}}{(1 - 2\lambda_k)^2}.
\end{aligned} \tag{14}$$

We have found all the terms proportional to \sqrt{g} and $\sqrt{g}R$ and dropped all term of $\mathcal{O}(R^2)$.

From (1) we get

$$\begin{aligned}\partial_t \Gamma_k^{\text{grav}} &= 2\kappa^2 \int d^4x \sqrt{g} [- (\partial_t Z_k) R + 2 (\partial_t Z_k \Lambda_k)] \\ &= 2Z_k \kappa^2 \int d^4x \sqrt{g} [\eta_g R + 2 (k^2 \partial_t \lambda_k + 2\Lambda_k - \eta_g \Lambda_k)] ,\end{aligned}\tag{15}$$

Further we know from $g_k = G_k k^2 = G k^2 / Z_k$ that

$$\beta_g = \partial_t g_k = (2 + \eta_g) g_k .\tag{16}$$

So the comparison of \sqrt{g} terms leads to

$$\begin{aligned}\partial_t \lambda_k &= (-2 + \eta_g) \lambda_k + \frac{1}{4\kappa^2} (\text{right-hand side}) \\ &= -4\lambda_k + \frac{\lambda_k}{g_k} \partial_t g_k + 8\pi g_k \left(\frac{5}{(4\pi)^2} \frac{1 - \frac{\eta_g}{6}}{1 - 2\lambda_k} \right) \\ &= -4\lambda_k + \frac{\lambda_k}{g_k} \partial_t g_k + \frac{10}{4\pi} g_k \frac{1 - \frac{\eta_g}{6}}{1 - 2\lambda_k} .\end{aligned}\tag{17}$$

And the comparison of $\sqrt{g}R$ terms leads to

$$\begin{aligned}\eta_g &= \frac{1}{2\kappa^2} (\text{right-hand side}) \\ &= 16\pi \left(-\frac{5}{6} \frac{1}{(4\pi)^2} \frac{2 - \frac{\eta_g}{2}}{1 - 2\lambda_k} - \frac{10}{3} \frac{1}{(4\pi)^2} \frac{1 - \frac{\eta_g}{6}}{(1 - 2\lambda_k)^2} \right) \\ &= -\frac{5}{3\pi} \left(\frac{1 - \frac{\eta_g}{4}}{1 - 2\lambda_k} + 2 \frac{1 - \frac{\eta_g}{6}}{(1 - 2\lambda_k)^2} \right) .\end{aligned}\tag{18}$$

Bonus question 1:

Why is the spin-two approximation rather good in most cases?

The TT-mode contains 5 degrees of freedom. The graviton vector mode 3 and the two scalar graviton modes each 1. The ghost modes are -8 degrees of freedom. In Landau gauge $\alpha \rightarrow 0$, the graviton vector mode and one scalar are exactly cancelled by the corresponding ghost mode. In summary, one has $5 + 1 - 4$ degrees of freedom. Furthermore the graviton degrees of freedom are usually enhanced by the cosmological constant ($1 - 2\lambda_k$ in the denominator of the graviton propagator). So with the TT-approximation we are taking the most of the degrees of freedom into account.

Bonus question 2:

Use mathematica and find numerically the non-Gaussian fixed point of (6) and determine the eigenvalues of the stability matrix.

See 'Background-TT-equations.nb' in the git. The fixed-point values are

$$(g_k^*, \lambda_k^*) = (0.86, 0.18) .\tag{19}$$

And the critical exponents are given by

$$\theta_{1,2} = 2.9 \pm 2.6 i . \tag{20}$$