

ASTR507 HW2

problem 2

```
In [2]: from numpy import pi,exp,sqrt,conj,array
        from astropy import units as u
```

Initialize Constants and test out units package.

```
In [3]: m = 9.1 * 10**(-28)*u.gram #mass of electron
        k = 1.4 * 10**(-16)*u.erg / u.Kelvin #boltzmann constant
        T = 10**(6)*u.Kelvin #temperature
        #V = (40 * 10**(3)*u.parsec**3).to(u.centimeter**3)
        n = 0.01 * u.centimeter**(-3) #density
        c = 2.99 * 10**10 *u.cm / u.s #speed of light
```

Define the normalization constants.

```
In [4]: A = (4*pi*m**(3./2.)*n)/(2*pi*k*T)**(3./2.)
        b = 0.5*m/(k*T)
```

Using the given solution to the integral we are solving equation: $\frac{Axe^{-bx^2}}{2b} = 1$ since we are solving for the single fastest electron in the distribution.

```
In [5]: print "A = ",(A*(u.cm/u.s)**2)
        print "b = ",(b)

A = 1.32223940717e-28 g(3/2) / (cm erg(3/2) s2)
b = 3.25e-18 g / erg
```

Stuff in front:

```
In [6]: A/(2*b)
```

```
Out[6]: 2.0342145 × 10-11  $\frac{\text{g}^{1/2}}{\text{erg}^{1/2} \text{ cm}^3}$ 
```

Via Wolfram, the root to this equation is:

```
In [7]: r = 1.95*10**(8) #real
        j = 1.1*10**(9) #imaginary
```

therefore the speed is the modulus of (r-j)

```
In [8]: v = sqrt(r**2 + j**2)
```

```
In [9]: print "speed is",v
        print "v = {}c".format(v/c.value)
```

```
speed is 1117150392.74
v = 0.0373628893893c
```

so no, the cosmic rays cannot be of thermal origin, since this speed is so much less than the speed of cosmic rays.

problem 3d

for H_2 :

```
In [229]: #constants
          T = 1000 #K
          g = 980 #cm/s^2
          r = 637.1 * 10**6 # cm
          m = 2 * 1.6737236 * 10**(-24) #for Hydrgen Molecule = 2 * m_H
          G = 6.7*10**(-8) #cgs
          M = 5.972 * 10**27 #g

          #composite constants
          sigma = pi * (10**(-8))**2
          n = m*g / (sigma * k.value * T)
          v_s = sqrt(2*k.value*T/m)
          v_esc = sqrt(2*G*M / r)
          lam = (v_esc/v_s)**2
```

```
In [230]: C = (n * v_s) / (2. * sqrt(pi))
```

```
In [231]: Flux = C *exp(-lam)*(lam + 1)
```

```
In [232]: print Flux # num particles / s cm^2
```

```
29322950.6884
```

```
In [233]: t = (1 * 10**9) * (365*24*60*60) # s/Gyr
```

Rate of loss of hydrogen over the entire area of exosphere:

```
In [234]: N_exo = Flux * 4*pi*r**2 #amount lost thermally over entire area of exosphere
```

Amount of hydrogen lost thermally over 1Gyr

```
In [235]: N_esc = N_exo * t
```

```
In [236]: print "number of H2 molecules lost thermally over 1Gyr over entire area of exosphere is", N_esc
print "number of H atoms lost (2*N_H2)/N_H2_atmosphere", (2 * N_esc/10**43)
```

```
number of H2 molecules lost thermally over 1Gyr over entire area of exosphere is 4.71671191201e+42
```

```
number of H atoms lost (2*N_H2)/N_H2_atmosphere 0.943342382402
```

number of H2 lost thermally over 1Gyr over entire area of exosphere
4.71671191201e+42 molecules

number of H atoms lost is $2N_{H2}$

$$\frac{N_{H_{atomsescaped}}}{N_{H_{atomsin atmosphere}}} \approx 0.94$$

problem 3e

for O_2 :

```
In [245]: mo = 2. * 2.6567626 * 10**(-23) # O2 in g
          v_s = sqrt(2.*k.value*T/mo)
          v_esc = sqrt(2.*G*M / r)
          lam = (v_esc/v_s)**2
```

```
In [246]: n = mo*g / (sigma * k.value * T)
```

```
In [247]: C = (n * v_s) / (2 * sqrt(pi))
          Flux_o2 = C * exp(-lam)*(lam + 1.) #num particles / s cm^2
```

```
In [248]: print Flux_o2 #num particles / s cm^2

1.75087645658e-88
```

Rate of loss of molecular Oxygen over the entire area of exosphere:

```
In [250]: N_exo_o = Flux_o2 * 4.*pi*r**2. #amount escaping over entire exosphere / s
          print N_exo_o

8.93059875204e-70
```

Number of O2 lost over 1Gyr

```
In [224]: N_o = N_exo_o * t
```

```
In [251]: print "Number of O2 molecules escaping over 1Gyr over entire exosphere is",N_o

Number of O2 molecules escaping over 1Gyr over entire exosphere is
2.81635362244e-53
```

$$N_{O_{2\text{escaped}}} \approx 0$$

```
In [252]: 4.71671191201e+42 / 2.81635362244e-53
```

```
Out[252]: 1.6747584090394123e+95
```

$$\text{Flux of } H_2 \text{ escaping} \approx 3 \times 10^7 \frac{\text{particles}}{\text{cm}^2 \text{s}}$$

$$\text{Flux of } O_2 \text{ escaping} \approx 2 \times 10^{-53} \frac{\text{particles}}{\text{cm}^2 \text{s}}$$

Assuming no sources or sinks of H_2 or O_2 the amount of O_2 should stay the same while the amount of H_2 will decrease rapidly unless it is locked up into H_2O very quickly.

In []: