

# ARCS scenario 1 data inversions

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This document describes the mathematical formulation of the ARCS scenario 1 data inversion problem. The measurements being made for scenario 1 will result in production of maps of magnetic field fluctuations  $\delta\mathbf{B}$ , and flows  $\mathbf{v}$ . These will be converted, as part of standard data processing, into parallel current density and electric field:

$$\mathbf{J} = \nabla \times \left( \frac{\delta\mathbf{B}}{\mu_0} \right) \quad (1)$$

$$\mathbf{E} = -\mathbf{v} \times \mathbf{B} \quad (2)$$

In the case of scenario 1 data, the key unknown physical parameters are the ionospheric Pedersen and Hall conductances. Poynting flux may also be calculated from the scenario 1 data:

$$\mathbf{S} = \mathbf{E} \times \frac{\delta\mathbf{B}}{\mu_0} \quad (3)$$

## 1 Physical constraints

The current continuity equation, here used in field-integrated form, relates electric fields and conductances to current:

$$\boxed{J_{\parallel} = \Sigma_P \nabla \cdot \mathbf{E}_{\perp} + \nabla \Sigma_P \cdot \mathbf{E}_{\perp} - \nabla \Sigma_H \cdot (\mathbf{E}_{\perp} \times \hat{\mathbf{b}})} \quad (4)$$

Note that this equation effectively has two unknown fields  $\Sigma_P, \Sigma_H$ , but represents only one physical constraint; hence additional information is needed. This is found in the form of the Poynting theorem:

$$\frac{\partial w}{\partial t} + \nabla \cdot \mathbf{S} = -\mathbf{J} \cdot \mathbf{E} \quad (5)$$

Similar to the assumptions made to produce Equation 4 we neglect time-dependent terms and proceed to integrate the equation along a geomagnetic field line:

$$S_{\parallel, top} - S_{\parallel, bottom} + \nabla_{\perp} \cdot \mathcal{S}_{\perp} = -\Sigma_P E^2 \quad (6)$$

where  $\mathcal{S}_{\perp}$  is the column integrated perpendicular Poynting flux. If we further assume that there is no Poynting flux through the bottom of the ionosphere or the lateral sides of our volume of interest (i.e. net incoming D.C. Poynting flux is dissipated) a simple relation between parallel Poynting flux and Pedersen conductance.

$$\boxed{S_{\parallel} = -\Sigma_P E^2} \quad (7)$$

## 2 Estimating conductances

Several different procedures can be developed for converting the maps of electric field and Poynting flux into conductances. Two approach are discussed here.

Equation 7 fully specifies the Pedersen conductance given quantities that are measurable by scenario 1 experiments, so the most obvious path would be to then provide the Pedersen conductance to Equation 4. Superficially, the equation allows solution for the gradient of the Hall conductance and in principle one would need to compute a line integral of this quantity to solve for Hall conductance:

$$\Sigma_H(\mathbf{r}_2) - \Sigma_H(\mathbf{r}_1) = \int_{\mathbf{r}_1}^{\mathbf{r}_2} \nabla \Sigma_H \cdot d\mathbf{r} \quad (8)$$

Moreover, one would also need the value of the Hall conductance at some reference point  $\mathbf{r}_1$  to complete the solution for Hall conductance. While it may be possible to choose a point with low density and assume zero Hall conductance at that reference point there is a more serious issue with this approach and with the set of physical constraints being used, more generally. Equation 4 only provides constraints on the derivative of the Hall conductance *in the direction of the  $\mathbf{E} \times \mathbf{B}$  drift*. Thus, there is information about the Hall conductance (namely the variation in the direction of the electric field) that is completely unconstrained by current continuity. As a result, the Hall conductance lies partly in the null space of the problem defined by Equations 4 and 7 and some additional assumptions/information/regularization will be required to solve the inverse problem.

Another approach to the inverse problem would be to view the conservation laws as constraints to be combined together with other prior information in the form of, e.g., smoothness constraints. Here we rewrite the physical constraints in a matrix form to facilitate application of results from linear inverse theory. Field quantities can be “flattened” into vectors using column major ordering and then operators can be represented through matrix operations. The latter step can be understood as a decomposition of the derivative operations into finite difference matrices:

$$\underline{j} = \underline{I} \underline{p} (\nabla \cdot \mathbf{E}_\perp) + \underline{L}_x \underline{p} E_x + \underline{L}_y \underline{p} E_y - \underline{L}_{E \times B} \underline{h} E_\perp \quad (9)$$

$$\underline{s} = -E^2 \underline{I} \underline{p} \quad (10)$$

Concatenating the unknown conductances into a single vector we get:

$$\underline{x} \equiv \begin{bmatrix} \underline{p} \\ \underline{h} \end{bmatrix} \quad (11)$$

The left-hand sides of each conservation law (i.e. measurements) are similarly stacked:

$$\underline{b} \equiv \begin{bmatrix} \underline{j} \\ \underline{s} \end{bmatrix} \quad (12)$$

Finally the right-hand side operations may be expressed in block diagonal form:

$$\underline{A} \equiv \begin{bmatrix} \underline{I} (\nabla \cdot \mathbf{E}_\perp) + \underline{L}_x E_x + \underline{L}_y E_y & - \underline{L}_{E \times B} E_\perp \\ -E^2 \underline{I} & \underline{0} \end{bmatrix} \quad (13)$$

Yielding our full set of constraints as:

$$\underline{A} \underline{x} = \underline{b} \quad (14)$$

As discussed previously this system will not be full-rank, but serves as a starting point for a suitable generalized inverse for this problem. As a final note the full system has size  $2 \cdot N \cdot M \times 2 \cdot N \cdot M$ ; where  $N, M$  are the  $x, y$  size of the data maps provided by instrument teams.

### 3 Maximum likelihood estimator

The maximum likelihood estimator, assuming Gaussian-distributed noise is (note we drop the underline notation here for brevity):

$$\hat{x}_{ML} = (A^T A)^{-1} A^T b \quad (15)$$

The matrix to be inverted here is singular for reasons noted previously; we adopt a Tikhonov regularization scheme to mitigate this:

$$\hat{x} = (A^T A + \lambda I)^{-1} A^T b \quad (16)$$

where  $\lambda$  is a regularization parameter.

FIXME: Need to add error covariance calculations.

### 4 Connections to precipitating electrons

Conductances are ultimately driven by electron precipitation, here encapsulated in terms of total energy flux  $Q$  and average energy  $E_{av}$  - these precipitation parameters are what is needed to ultimately drive the GEMINI simulations.

One of the simplest parameterizations of conductance is the Robinson formulas:

$$\Sigma_P = \frac{40E_{av}}{16 + E_{av}^2} \sqrt{Q} \quad (17)$$

$$\frac{\Sigma_H}{\Sigma_P} = 0.45 E_{av}^{0.85} \quad (18)$$

Using these to constrain conductances creates a physical correlations between them that does not exist using just the constraints from conservation laws.