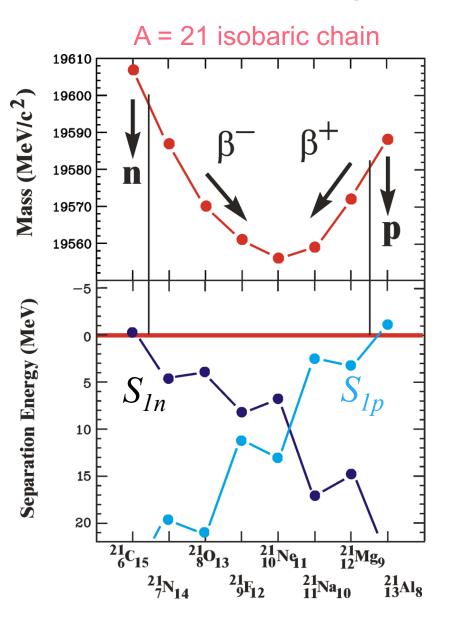


Separation energies



one-nucleon separation energies

$$S_{1n} = B(N, Z) - B(N - 1, Z)$$

$$S_{1p} = B(N, Z) - B(N, Z - 1)$$

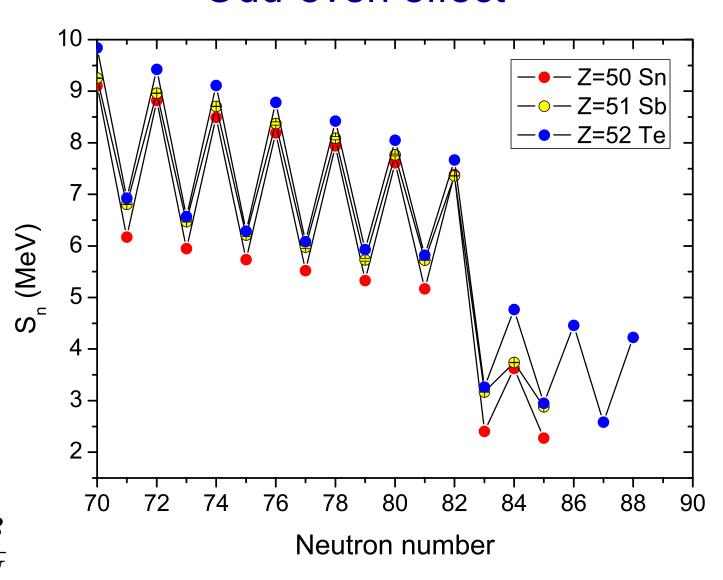
two-nucleon separation energies

$$S_{2n} = B(N, Z) - B(N - 2, Z)$$

$$S_{2p} = B(N, Z) - B(N, Z - 2)$$

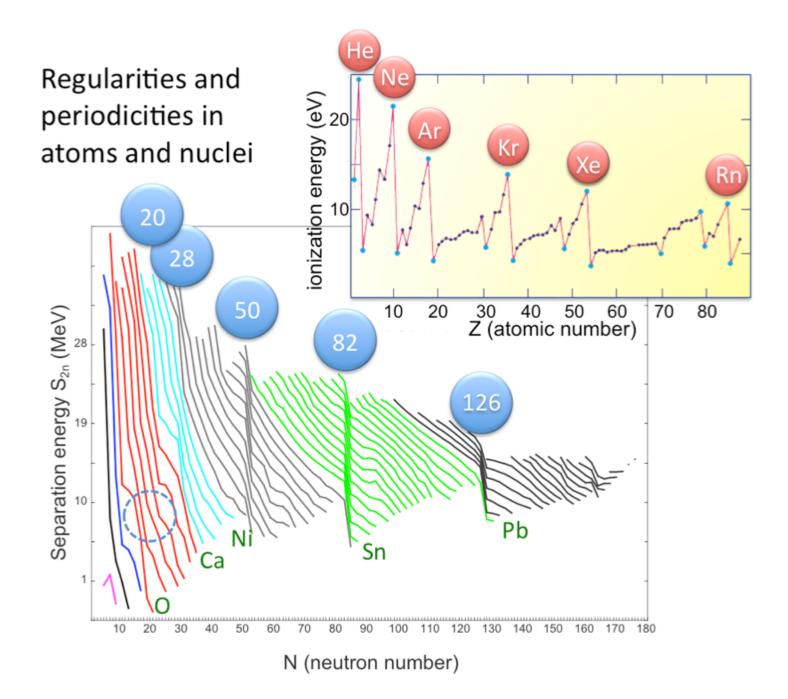
http://www.nndc.bnl.gov/chart/

Odd-even effect



$$\lambda = -\frac{dB}{dN}$$

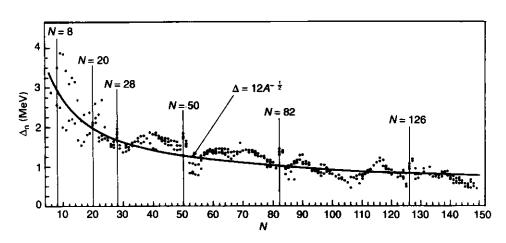
J. Phys. G 39 093101 (2012)



Pairing energy

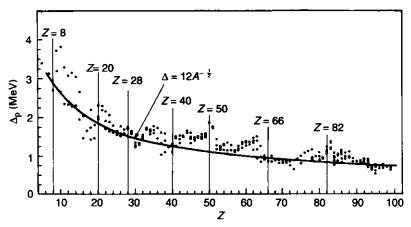
$$\Delta_n = \frac{B(N+1,Z) + B(N-1,Z)}{2} - \frac{N\text{-odd}}{B(N,Z)}$$

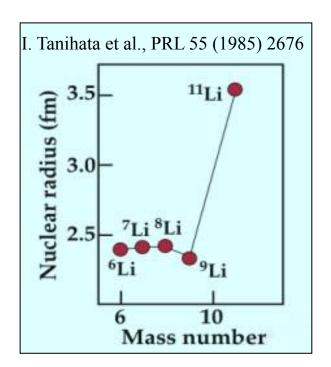
$$\Delta_p = \frac{B(N,Z+1) + B(N,Z-1)}{2} - B(N,Z)$$
Z-odd



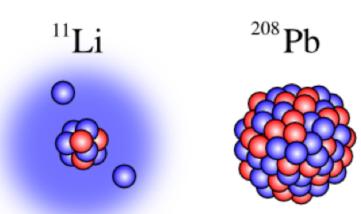
A common phenomenon in mesoscopic systems!

Example: Cooper electron pairs in superconductors...





Halos



Consider a spherical square-well potential such as in the figure. The energy of the last occupied neutron state is $\varepsilon < 0$ and its quantum numbers are n and $\ell = 0$. We assume: $\varepsilon \approx 0$.

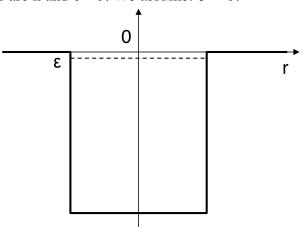
Prove that the r.m.s radius of the neutron orbit, defined as

$$R_{rms} = \sqrt{\langle \hat{r}^2 \rangle}$$

obeys the following relation:

$$R_{rms} \sim (-\varepsilon)^{-1/2}$$

Discuss the result.



<u>Hint</u>: The asymptotic behavior of the Bessel function at large distance is:

$$H_{\ell}(z) \xrightarrow[r \to +\infty]{} 1 \times 3 \times \dots (2\ell-1)z^{-\ell-1}$$

For super achievers: