

Frequentist Inference

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Winter 2022

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Statistics and probability seem like the same topic, but there are very fundamental differences. In probability, we are given some distribution and must compute certain probabilities. In statistics, we are given the results (the data) and must infer what distribution it came from. Some notes to put later.

The three pillars of frequentist inference is hypothesis testing, point estimation, and confidence intervals. After we establish some fundamental with hypothesis testing, we talk about point estimation, which is the most popular problem in statistics and machine learning. In estimation, the bias variance tradeoff is a central idea. We introduce three methods in creating good point estimators, starting with the most naive method of moments. This is improved with the more modern maximum likelihood estimation, which is most popular for parametric models. For nonparametric models, minimax estimation is popular. Then we finish off with confidence intervals for uncertainty quantification.

Next, we talk about model selection, focusing on the cross validation and information criteria.

1 Hypothesis Testing

This was done for the first time as early as 1710, when a researcher tried to measure sex ratios in a population [Arb10].

A significance test is a method used to decide whether the data at hand sufficiently supports a particular hypothesis. The hypothesis to be tested is called the **alternative hypothesis**, denoted H_1 or H_a , and the status quo is called the **null hypothesis**, denoted H_0 . Assuming that H_0 is true, we compute the likelihood of the data happening. If the sample is not too unlikely (past some significance level), we fail to reject H_0 , and if there is strong evidence, we reject H_0 . H_0 and H_a can be devised in countless ways.

Example 1.1 ()

There are countless test statistics we can build, but here are some common examples,

1. Proportion: Company A produces circuit boards, but 10% of them are defective. Company B claims that they produce fewer defective circuit boards.

$$H_0 : p = 0.10 \text{ versus } H_a : p < 0.10 \quad (1)$$

2. Means: It is known that the average height of boys in KIS is 176cm. Ben claims that the average height is lower than this.

$$H_0 : \mu = 176 \text{ versus } H_a : \mu < 176 \quad (2)$$

3. Difference of Means: If μ_1 and μ_2 denote the true average breaking strengths of the same type of twine produced by two different companies. Jenny claims that the $\mu_1 - \mu_2 > 5$.

$$H_0 : \mu_1 - \mu_2 = 0 \text{ versus } H_a : \mu_1 - \mu_2 > 5 \quad (3)$$

1.1 Sampling Distributions

Definition 1.1 (Population, Parameters)

When conducting a statistical study, there is a set of items or events which is of interest for some experiment. This can be modeled with some probability space $(\Omega, \mathcal{F}, \mathbb{P})$. Usually, we are interested in some numerical property of this population, and so we implicitly define a random variable $X : \Omega \rightarrow \mathbb{R}$ that induces some distribution $X \sim P$, which we call the **population**. A statistical population can be a group of existing objects or a hypothetical and potentially infinite group of objects conceived as a generalization from experience.

With this, we can interpret the population $X \sim P$ as a random variable, and we often call this the **parent distribution**. We are often interested in its population **parameters**, which can be any measured quantity of a population that summarizes or describes an aspect of it. In generality, the parameter of population X is denoted θ , and it is a fixed value.

Example 1.2 (Populations)

Here are some examples of populations:

1. We let Ω be the discrete sample space of all hands in poker, and our random variable will assign a numerical ranking to each hand 0 for no hand, 1 for pairs, 2 for two pairs, etc.
2. Ω is the sample space of all individuals in the U.S. and we can construct a random variable X that assigns to each individual their height. Even though Ω is finite, we can interpret it as continuous, which leads to a continuous distribution $X \sim P$.

Example 1.3 (Parameters)

Some population parameters can be:

1. The true mean of the population $\mu_X = \mathbb{E}[X]$
2. The true variance of the population $\sigma_X^2 = \text{Var}[X]$
3. In the height example, we can set $X = (X_1, X_2, X_3)$ and try to construct a linear regression model that predicts X_3 from X_1, X_2 . Theoretically, this is $\mathbb{E}[X_3 \mid X_1, X_2]$, and we must find the best function of form $x_3 = a + b_1x_1 + b_2x_2$ that is closest to the conditional expectation. There does exist a unique one, and so a, b_1, b_2 are all population parameters.

In general, the population is the total set of all relevant things that we are interested in. The specific quantity of the actual population is called the **population parameter**, e.g. the true mean μ or the true variance σ^2 of X , usually denoted with θ . But usually, these parameters are not known since the population is too big to experimentally measure, so we must try and estimate it with samples. This is the entire point of statistics; otherwise, we would already know everything we want to know.

Definition 1.2 (Samples)

From the population $X \sim P$ (which still has unknown distribution), we can take n **samples** by considering iid $x_1, x_2, \dots, x_n \sim P$.

1. We should note that the samples x_i are random variables themselves. Not fixing them yet and still considering them in generality as random objects allows us to do more theoretical calculations.
2. Once these samples have been realized (i.e. $\omega \in \Omega$ is realized, and all x_i 's are also realized), we can treat them as fixed values.

Sometimes, we may not assume independence, but for most cases we do. A common rule is that if the sample size n is less than 10% of the population size, then we can assume independence.

Definition 1.3 (Empirical Distribution)

Now given that we have these iid samples, we can construct the **empirical distribution** $\hat{X} \sim \hat{P}$, defined as the discrete distribution that assigns probability $1/n$ to each value x_i for $i \in [n]$. In other words, we have

$$\mathbb{P}(\hat{X} = x) = \frac{1}{n} \text{ for } x \in \{x_1, \dots, x_n\} \quad (4)$$

We can write the CDF of the empirical distribution, called the **empirical distribution function**, as the sum of indicators

$$F_X(x) = \frac{1}{n} \sum_{i=1}^n \mathbb{I}_{[x_i, +\infty)}(x) \quad (5)$$

As expected, we would expect the empirical distribution to converge to the actual distribution.

Theorem 1.1 (Glivenko–Cantelli theorem)

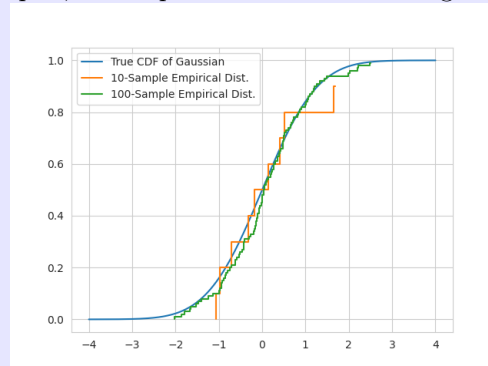
The empirical distribution of iid samples $x_1, \dots, x_n \sim P_n$ converges almost surely to $X \sim P$ as $n \rightarrow \infty$. More specifically, given that the CDF of X is F and the CDF of P_n is the step function F_n , we have

$$\|F_n - F\|_\infty = \sup_{x \in \mathbb{R}} |F_n(x) - F(x)| \rightarrow 0 \quad (6)$$

almost surely as $n \rightarrow \infty$.

Example 1.4 (Empirical Distribution of Standard Gaussian)

We expect the empirical distribution of the standard Gaussian to converge. Indeed, numerical results show that for 10 and 100 samples, the empirical CDF does converge to the true CDF.

**1.2 Concentration of Measure**

Now let's move on to concentration inequalities, which say that the probability that a random variable is greater than something is bounded by something. These probability bounds are extremely useful in of themselves. It allows us to talk about convergence theory, which tells us what happens to a statistic, such as \bar{X} , as I get more and more data. The first inequality exploits the fact that the tails of a Gaussian RV decay very quickly, and a lot of concentration inequalities attempt to mimic this exponential bound but for non-Gaussian distributions.

Theorem 1.2 (Gaussian Tail Inequality)

Given $X \sim \mathcal{N}(0, 1)$, the inequality says that the probability of X taking values past a certain t decays exponentially.

$$\mathbb{P}(|X| > t) \leq \frac{2e^{-t^2/2}}{t}$$

If we have $x_1, \dots, x_n \sim \mathcal{N}(0, 1)$, then

$$\mathbb{P}(|\bar{X}| > t) \leq \frac{2}{\sqrt{nt}} e^{-nt^2/2}$$

We can assume that the coefficient is less than 1 if n is large. The above tells us that this bound exponentially decays with t but also with the number of samples n .

Theorem 1.3 (Markov's Inequality)

Given a nonnegative random variable $X > 0$, we have

$$\mathbb{P}(X > t) \leq \frac{\mathbb{E}[X]}{t}$$

Proof.

We have

$$\begin{aligned}
 \mathbb{E}[X] &= \int_0^\infty xp_X(x) dx \\
 &\geq \int_t^\infty xp_X(x) dx \\
 &= t \int_t^\infty p_X(x) dx \\
 &= t\mathbb{P}(X > t)
 \end{aligned}$$

Theorem 1.4 (Chebyshev's Inequality)

Given a random variable X with mean $\mu = \mathbb{E}[X]$, we have

$$\mathbb{P}(|x - \mu| > t) \leq \frac{\text{Var}(X)}{t^2}$$

Theorem 1.5 (Hoeffding's Inequality)

Let x_1, x_2, \dots, x_n be independent (not necessarily identical) random variables s.t. $a_i \leq X_i \leq b_i$ almost surely. Consider the random variable $\bar{X} = \frac{1}{n}(x_1 + \dots + x_n)$. Then, for all $t > 0$,

$$\mathbb{P}(|\bar{X} - \mathbb{E}[\bar{X}]| \geq t) \leq 2 \exp\left(-\frac{2n^2t^2}{\sum_{i=1}^n (b_i - a_i)^2}\right)$$

Now in addition to bounding probabilities, we would like to bound expectations.

Theorem 1.6 (Cauchy-Schwartz)

Given random variables X, Y , it is often hard to compute the expectation of XY since it is hard to compute the distribution of it (sums are easy). But we can bound it as

$$|\mathbb{E}[XY]| \leq \mathbb{E}[|XY|] \leq \sqrt{\mathbb{E}[X]\mathbb{E}[Y]}$$

Theorem 1.7 (Jensen's Inequality)

Given g a convex function and X a random variable, we have

$$\mathbb{E}[g(X)] \geq g(\mathbb{E}[X])$$

1.3 Kullback Leibler Divergence

Now a popular metric between PMFs/PDFs is the KL divergence.

Definition 1.4 (Kullback-Leibler Divergence)

Given random variable X and Y ,

1. If they are discrete with PMFs P and Q , the KL-divergence is defined

$$D_{KL}(P \parallel Q) := \sum_x P(x) \log \left(\frac{P(x)}{Q(x)} \right) = \mathbb{E} \left[\log \frac{P(x)}{Q(x)} \right]$$

where we can interpret the expectation as $X \sim p$.

2. If they are continuous with PDFs p and q , the KL-divergence is defined

$$D_{KL}(p \parallel q) := \int p(x) \log \left(\frac{p(x)}{q(x)} \right) dx = \mathbb{E} \left[\log \frac{P(x)}{Q(x)} \right]$$

where we can interpret the expectation as $X \sim p$.

1. The fact that $D_{KL}(p \mid p) = 0$ is obvious.
2. To prove that $D_{KL}(p \mid q) \geq 0$, we use Jensen's inequality

$$-D_{KL}(p \mid q) = \mathbb{E} \log \frac{q(X)}{p(X)} \leq \log \mathbb{E} \frac{q(X)}{p(X)} = \log \int \frac{q(x)}{p(x)} p(x) dx = \log(1) = 0$$

This is not a metric though.

1.4 Bounding Maximum of Random Variables

Given that we have n samples $x_1, \dots, x_n \sim P$, it is conventional to index them with open brackets to denote order

$$X_{(1)} \leq X_{(2)} \leq \dots \leq X_{(n)}$$

Our goal is now to find the distribution of $X_{(n)} = \max_i X_i$. If we know the distribution of P , $\mathbb{P}(X_{(n)} \leq x)$ is just the probability that all the X_i 's are less than x , and by independence we can product out the CDFs, differentiate to get the PDF, and compute. So it's not too hard to do this theoretically, but in practice this is hard to do since we don't exactly know P .

So if you didn't know the X_i 's, the best you can assume is that

$$\mathbb{E} \max\{x_1, \dots, x_n\}$$

is going to grow like n . But if we can bound the MGF with $\mathbb{E} e^{tX} \leq e^{t^2 \sigma^2} 2$, then we can show that $\mathbb{E} \max X_i$ doesn't grow like n , but rather like $\log n$.

$$\mathbb{E} \max X_i \leq \sigma \sqrt{2 \log n}$$

1.5 Big-O, Little-O Notation

Going back to calculus, if we have a function $f : X \rightarrow \mathbb{R}$, we can say that

1. $f(x) = O(g(x))$ if f is of the same order as g . That is, they grow at the same rate

$$\frac{f(x)}{g(x)} \rightarrow c \text{ as } x \rightarrow \infty$$

for some constant c .

2. $f(x) = o(g(x))$ if f is negligible w.r.t. g . That is, f is infinitesimal w.r.t. g .

$$\frac{f(x)}{g(x)} \rightarrow 0 \text{ as } x \rightarrow \infty$$

for some constant c .

Now there is a probabilistic notation as well. The concept of boundedness translates to being able to capture most of the mass of the random variable within some interval, and infinitesimality translates to the probability mass concentrating around 0.

Definition 1.5 (O_p, o_p Notation)

Let x_1, x_2, \dots be a sequence of random variables.

1. $x_n = o_p(1)$ if

$$\mathbb{P}(|Y_n| > \epsilon) \rightarrow 0 \text{ as } n \rightarrow \infty$$

for all $\epsilon > 0$. This means that x_n gets more and more concentrated around 0.

2. $x_n = O_p(1)$ if for all $\epsilon > 0$, then there exists a C s.t.

$$\mathbb{P}(|x_n| > C) \leq \epsilon$$

for all large n . That is, we can always trap the majority of the probability mass of x_n within the interval $[-C, C]$. This must hold for all x_n with $n > N$, so the mass can't "escape" to infinity. We can think of it as the distribution is "settling down" and not shooting off to somewhere.

3. $x_n = o_p(a_n)$ means that

$$\frac{x_n}{a_n} = o_p(1)$$

4. $x_n = O_p(a_n)$ means that

$$\frac{x_n}{a_n} = O_p(1)$$

Theorem 1.8 ()

Given $Y_1, \dots, Y_n \sim \text{Bernoulli}(p)$, let $\hat{p} = \frac{1}{n} \sum Y_i$. Then

$$\hat{p}_n - p = o_p(1)$$

which is also written $\hat{p}_n = p + o_p(1)$, which means that the random variable \hat{p}_n is some constant p plus a random variable that is going to 0.

Proof.

By Hoeffding's inequality,

$$\mathbb{P}(|\hat{p}_n - p| > \epsilon) \leq 2e^{-2n\epsilon^2}$$

which goes to 0 as $n \rightarrow \infty$.

Example 1.5 ()

Given $Y_1, \dots, Y_n \sim \text{Bernoulli}(p)$, let $\hat{p} = \frac{1}{n} \sum Y_i$. Then,

$$\hat{p} - p = O_p\left(\frac{1}{\sqrt{n}}\right)$$

1.6 One Sample Z and T Tests

Let us have some population $X \sim P$ and a null hypothesis that claims $H_0 : \mu = \theta_0$. Since we are interested in the mean, we would like to use CLT or some other theorem to determine what the distribution of the mean of n samples \bar{x}_n looks like (either Normal or Student T centered around θ_0 and scaled down by factor of \sqrt{n}). When we actually sample, the value $\bar{x}_n = \hat{\theta}$ is realized, and we would like to see if sampling $\hat{\theta}$ from

the distribution centered around θ_0 is likely, usually after normalizing. If it isn't, then we reject H_0 .

How do we decide whether to use the z-test or the t-test? It is known that $\text{StudentT}(n-1)$ converges to $\mathcal{N}(0, 1)$ in distribution as $n \rightarrow +\infty$. Therefore, depending on the context of the problem, at a certain point N (usually $N = 30$ or perhaps higher for skewed distributions), the difference between these two are negligible.

1. Z-test: if we know the population variance σ^2 , but it is rarely the case that we actually know σ^2 .
2. T-test: if we do not know the population variance σ^2 , which we then substitute for the sample variance S^2 .
3. Z-test: if we do not know the population variance (which we substitute for S^2), but our sample size is greater than N , then we can approximate the t -distribution with our normal, allowing us to use the Z-test again.

In general, the alternative to the null hypothesis $H_0 : \theta = \theta_0$ will look like one of the following three assertions:

1. Two-Sided Test: $H_a : \theta \neq \theta_0$
2. One-Sided Test: $H_a : \theta > \theta_0$ (in which case the null hypothesis is $\theta \leq \theta_0$)
3. One-Sided Test: $H_a : \theta < \theta_0$ (in which case the null hypothesis is $\theta \geq \theta_0$)

Now we must still quantify *how* unlikely our sample mean θ must be compared to θ_0 in order to reject the null hypothesis. This is where we specify our **significance level**, denoted by α (common values 0.10, 0.05, 0.01). This specifies the tail-regions in which θ will land in with probability α . Usually, working with general normal/t distributions is tedious, so we can rescale them and use their z/t-scores.

Definition 1.6 (Z-score)

Given a value x sampled from distribution $X \sim \mathcal{N}(\mu, \sigma^2)$, its **z-score** is defined to be the number of standard deviations away from the mean.

$$z := \frac{x - \mu}{\sigma} \quad (7)$$

Now given a significance level $\alpha \in [0, 1]$, let z_α be the value such that the measure of a standard normal distribution past z_α is $1 - \alpha$ (i.e. the 100α percentile). z_α is called the **critical z-value**.

Definition 1.7 (T-score)

Given a value x sampled from distribution $X \sim \text{StudentT}(n)$, its **t-score** is defined to be the number of standard deviations away from the mean.

Example 1.6 ()

A factory has a machine that dispenses 80mL of fluid in a bottle. An employee believes the average amount of fluid is not 80mL. Using 40 samples, he measures the average amount dispensed by the machine to be 78mL with a sample standard deviation of 2.5.

1. Let the true mean be μ and true standard deviation be σ . The null hypothesis is $H_0 : \mu = 80$ and the alternative is $H_1 : \mu \neq 80$, making this a two-sided test.
2. We don't know the true standard deviation σ , so we must use the sample standard deviation S . This requires us to use the t -test, but since $n > 30$, we can invoke CLT and state that \bar{x}_{40} is (approximately) Gaussian with mean μ and standard deviation S/\sqrt{n} . So, we use the z -test.
3. At a 95% confidence level, we have $\alpha = 0.05$, and our rejection region is $(-\infty, z_{0.025}] \cup [z_{0.975}, +\infty)$. Since we are looking at a standard Gaussian, we have by symmetry $z_{0.025} = -1.96$

and $z_{0.975} = 1.96$, and our critical z-value is $z^* = 1.96$.

4. So the z-score for 78 is

$$z = \frac{\bar{x} - \mu_0}{s/\sqrt{n}} = \frac{78 - 80}{2.5/\sqrt{40}} = -5.06 \quad (8)$$

which is definitely in the reject region. So this tells us that we can reject the null hypothesis with a 95% level of confidence.

Example 1.7 ()

A company manufactures car batteries with an average life span of 2 or more years. An engineer believes this value to be less. Using 10 samples, he measures the average life span to be 1.8 years with a standard deviation of 0.15.

1. Let the true mean be μ and true standard deviation be σ . The null hypothesis is $H_0 : \mu \geq 2$ and the alternative is $H_1 : \mu < 2$, making this a one-sided test.
2. We don't know the true standard deviation σ , so we must use the sample standard deviation S . This requires us to use the t -test, especially since $n = 10$ is not large enough for us to invoke CLT.
3. At a 99% confidence level, we have $\alpha = 0.01$, and our rejection region is $(-\infty, t_{0.01}] = (-\infty, -2.82]$.
4. The t-score for the observed mean value is

$$t = \frac{\bar{x} - \mu_0}{s/\sqrt{n}} = \frac{1.8 - 2}{0.15/\sqrt{10}} = -4.22 \quad (9)$$

which is definitely in the reject region. So this tells us that we can reject the null hypothesis with a 99% level of confidence.

We may have to account for errors. There is always a chance that our evidence leads us to an incorrect conclusion, and we have names for this.

Definition 1.8 (Errors)

Given a hypothesis test where we look for evidence supporting our alternative claim,

1. A **type 1 error** is when the null hypothesis is rejected, but it is true (false positive).
2. A **type 2 error** is when we fail to reject the null hypothesis, when it is false (false negative).

1.7 Power of a test

1.8 Common tests (t-test, z-test, chi-square test, F-test)

1.9 Multiple testing problem

2 Point Estimation

Definition 2.1 (Sample Statistic, Estimators and Estimates)

Now given a population X , we would like to use the n iid samples x_1, \dots, x_n to estimate a parameter θ of interest with our own random variable/value $\hat{\theta}_n$, called a **sample statistic**. We must note the dual nature of the sample statistic as a random variable and a value is similar to that of samples.

1. The statistic $\hat{\theta}_n$ is a random variable itself, referred to as the **estimator**. More specifically, it is a function $\hat{\theta}_n : \mathbb{R}^n \rightarrow \mathbb{R}$ of the n samples, i.e. a transformation of random variables

$$\hat{\theta}_n = \hat{\theta}_n(x_1, x_2, \dots, x_n) \quad (10)$$

This makes $\hat{\theta}_n$ also a random variable, which attempts to estimate the true θ , which is some unknown fixed value. Since $\hat{\theta}_n$ is a random variable, it has its own distribution, called the **sampling distribution** of $\hat{\theta}_n$.

2. Once these samples x_i have been realized, the estimator realizes and the value realized is now called the **estimate**.

This sampling distribution is a distribution of the statistic $\hat{\theta}_n$, and this forms a separate distribution with its own mean and variance.

1. the mean of the sampling distribution is denoted $\mu_{\hat{\theta}_n}$
2. the standard deviation of the sampling distribution is denoted $\sigma_{\hat{\theta}_n}^2$, also called the **standard error**.

We would want these estimators to have three properties:

1. unbiasedness
2. consistency
3. efficiency

We would like the sampling distribution of our statistic to give us good estimate in two ways. $\hat{\theta}_n$ should not be too far off from the actual parameter θ (bias is small), and $\hat{\theta}_n$ should not fluctuate too widely (variance of $\hat{\theta}_n$ should be small).

Definition 2.2 (Bias, Variance of Estimator)

Given an estimator $\hat{\theta}$ of a sample x_1, \dots, x_n estimating population parameter θ , the **sampling bias** refers to

$$\text{Bias}(\hat{\theta}) = |\mathbb{E}[\hat{\theta}] - \theta| \quad (11)$$

and the **sampling variance** refers to

$$\text{Var}(\hat{\theta}) = \mathbb{E}[(\hat{\theta} - \mathbb{E}[\hat{\theta}])^2] \quad (12)$$

A good rule of thumb to remember is that statistics is about replacing expectations with averages.

$$\mathbb{E} \mapsto \frac{1}{n} \sum_i \quad (13)$$

This is really the fundamental quality of statistics. Then after that we can do some fancy things, like minimizing something or manipulating another, but every single time we see an expectation just replace it with an average.

Definition 2.3 (Sample Mean)

Given a population X with $\mu = \mathbb{E}[X]$ and $\sigma^2 = \text{Var}(X)$, our estimator for μ is simply the average of the n samples x_1, \dots, x_n , called the **sample mean** or the **sampling distribution of the sample mean**.

$$\bar{x}_n = \hat{\mu}_n = \frac{1}{n}(x_1 + \dots + x_n) \quad (14)$$

This gives us the sampling distribution of the sample means. The mean and standard deviation (i.e. standard error) of \bar{x}_n is denoted $\mu_{\bar{x}_n}$ and $\sigma_{\bar{x}_n}$.

1. The mean of \bar{x}_n is μ .

$$\mu_{\bar{x}_n} = \mu \quad (15)$$

because

$$\mathbb{E}[\bar{x}_n] = \mathbb{E}\left[\frac{1}{n} \sum_{i=1}^n x_i\right] = \frac{1}{n} \sum_{i=1}^n \mathbb{E}[x_i] = \mathbb{E}[x] = \mu \quad (16)$$

2. The variance of \bar{x}_n is σ^2/n , i.e. the standard error of \bar{x}_n is $\sigma_{\bar{x}_n} = \sigma/\sqrt{n}$.

$$\sigma_{\bar{x}_n} = \frac{\sigma}{\sqrt{n}} \quad (17)$$

because

$$\sigma_{\bar{x}_n}^2 = \frac{1}{n^2} \sum_{i=1}^n \text{Var}(x_i) = \frac{1}{n} \text{Var}(x) = \frac{\sigma^2}{n} \quad (18)$$

Practically, this tells us that when trying to estimate the value of a population mean, due to the factor of $1/\sqrt{n}$, reducing the error on the estimate by a factor of 2 requires acquiring 4 times as many observations in the sample. But realistically, the true standard deviation σ is unknown, and so the standard error of the mean is usually estimated by replacing σ with the sample standard deviation S instead.

$$\sigma_{\bar{x}_n} \approx \frac{S}{\sqrt{n}} \quad (19)$$

3. By CLT, \bar{x}_n converges to $\mathcal{N}(\mu, \sigma^2/n)$ in distribution as $n \rightarrow +\infty$ (but in practicality, we assume this for $n \geq 30$). The fact that its mean and variance is μ and σ^2/n isn't that impressive. What is really impressive is that no matter what the distribution of x is, the sampling distribution of the mean will be Gaussian.

Example 2.1 (Sample Means)

Here are some figures of sample means. Note that with a uniform parent distribution, the sampling distribution of its mean looks like a Gaussian even without a large n . However, this is not necessarily true for different parent distributions, such as the exponential.

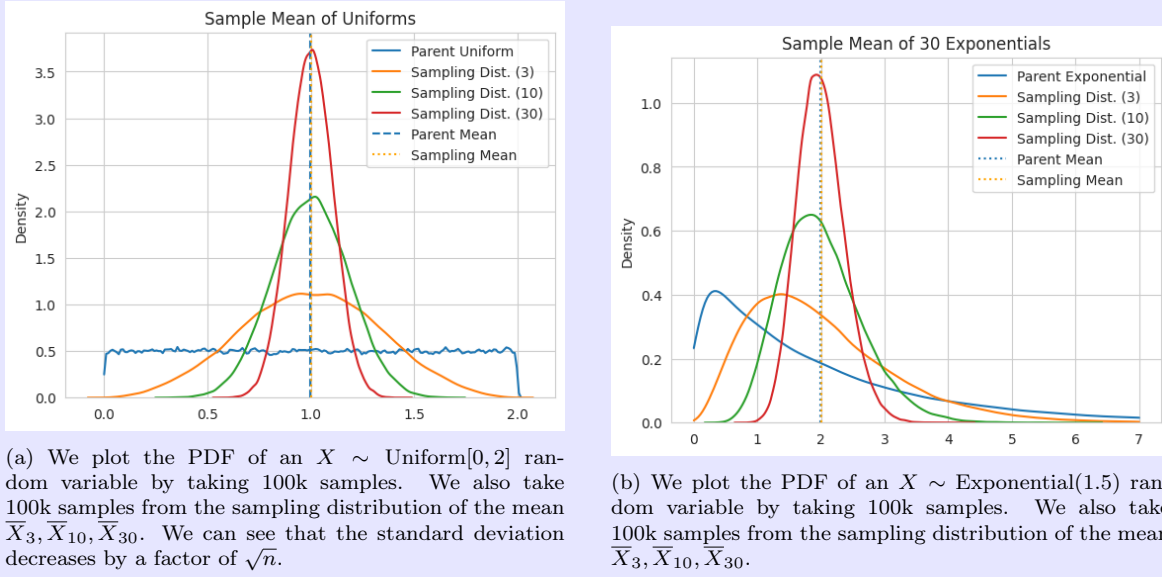


Figure 1

If the parent distribution is normal, then we don't even need CLT to claim that the sampling distribution of the sample mean is normal, since sums of normals are normal.

Now the variance of the population is defined to be $\sigma^2 = \mathbb{E}[(X - \mathbb{E}[X])^2]$, and by our rule of thumb, we can replace the expectations with sample means, by first setting $\mathbb{E}[X] = \hat{\mu}$ and averaging out the values $(X - \hat{\mu})^2$.

Definition 2.4 (Sample Variance)

Given a population X , our estimator for $\sigma^2 = \mathbb{E}[(X - \mathbb{E}[X])^2]$ is simply the average of the squared distances of the n samples $\{(x_i - \hat{\mu})^2\}_{i=1}^n$.

$$S_n^2 = \hat{\sigma}_n^2 = \frac{1}{n} \sum_{i=1}^n (x_i - \bar{x}_n)^2 \quad (20)$$

The mean and standard deviation of S_n^2 is denoted $\mu_{S_n^2}$ and $\sigma_{S_n^2}$. Note that there is a small difference that the sum for variance is divided by $n - 1$ rather than n , since we want it to be unbiased, but we will correct this later.

While the CLT states that the sampling distribution of the sample mean will look approximately Gaussian, we do not have this luxury when looking at the sampling distribution of sample variance.

Example 2.2 (Sample Variance)

Take a look at the following sampling distributions of the sample variance. There does not seem to be strong signs of convergence to a Gaussian. Their means do not align either.

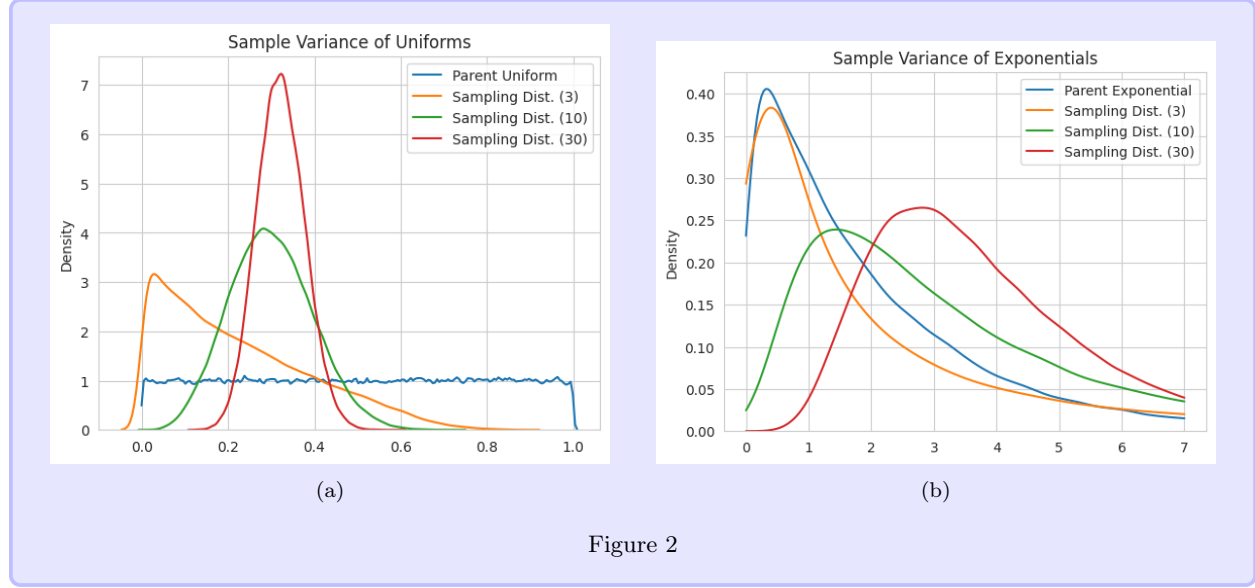


Figure 2

2.1 Sampling from Gaussians

Now if we assume that the parent distribution is Gaussian, then we can conclude some extra things and more kinds of distributions arise. Let $x_1, \dots, x_n \sim \mathcal{N}(\mu, \sigma^2)$, with \bar{x}_n the sample mean and S_n^2 the sample variance. Say that we want to find the distribution of \bar{x}_n .

1. In the unrealistic case where we know the true σ^2 , we don't even need to consider the sample variance. From the basic property of Gaussians, we know that $\bar{x}_n \sim \mathcal{N}(\mu, \sigma^2/n)$, or after standardizing,

$$\frac{\bar{x}_n - \mu}{\sigma/\sqrt{n}} \sim \mathcal{N}(0, 1) \quad (21)$$

2. In the realistic case where we don't know the true σ^2 , we should replace it with our sample variance S^2 , and it turns out that because of this extra uncertainty in the variance, our sampling distribution follows the student-t distribution, which can be interpreted as a mixture of Gaussians with differing variances.

$$\frac{\bar{x}_n - \mu}{S/\sqrt{n}} \sim \text{StudentT}(n-1) \quad (22)$$

Now if we are interested in finding the distribution of S_n^2 :

1. In the unrealistic case where we know the true μ , we don't need to consider the sampling distribution of \bar{x}_n . We have

$$S_n^2 = \frac{1}{n} \sum_{i=1}^n (x_i - \mu)^2 \sim \text{Gamma}\left(\frac{n}{2}, \frac{n}{2\sigma^2}\right) \quad (23)$$

2. In the realistic case where we don't know μ , we have

$$\frac{n-1}{\sigma^2} S_n^2 = \frac{1}{\sigma^2} \sum_{i=1}^n (x_i - \bar{x}_n)^2 \sim \chi^2(n-1) \quad (24)$$

2.2 Bias Variance Noise Decomposition

Let's do some further analysis on this. When you take a supremum over a function class, it decomposes into 3 terms.

1. One of which quantifies how big the function class is (more variance).

2. One of which quantifies the distance between the truth and the function class (bias).
3. One is the noise term, which is the irreducible error.

Example 2.3 (Bias and Variance Tradeoff in Polynomial Regression)

Let's motivate this by trying to fit a polynomial on some data.

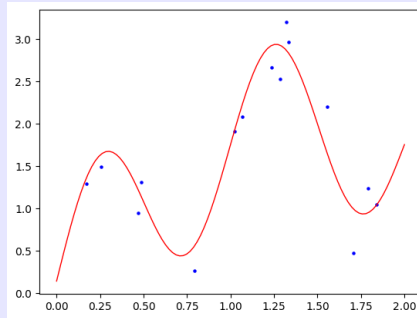


Figure 3: A sample of $|\mathcal{D}| = 15$ data points are generated from the function $f(x) = \sin(2\pi x) + 2 \cos(x - 1.5)$ with Gaussian noise $N(0, 0.3)$ on the interval $[0, 1]$.

If we try to fit a polynomial function, how do we know which degree is best? Well the most simple thing is to just try all of them. To demonstrate this even further, I generated 10 different datasets \mathcal{D} of size 15 taken from the same true distribution. The best fitted polynomials for each dataset is shown below.

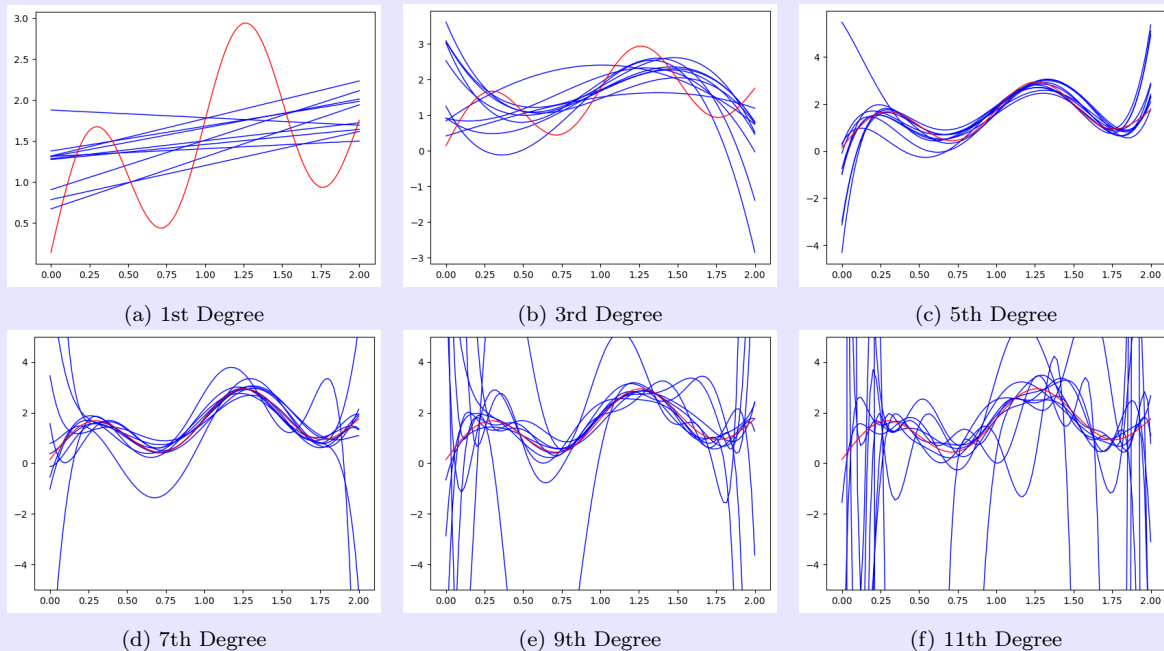


Figure 4: Different model complexities (i.e. different polynomial degrees) lead to different fits of the data generated from the true distribution. The lower degree best fit polynomials don't have much variability in their best fits but have high bias, while the higher degree best fit polynomials have very high variability in their best fits but have low bias. The code used to generate this data is [here](#).

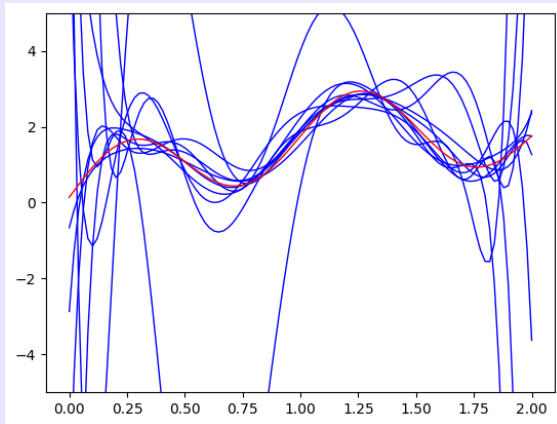
We already know that the 5th degree approximation is most optimal, and the lower degree ones are

underfitting the data, while the higher degree ones are **overfitting**. As mentioned before, we can describe the underfitting and overfitting phenomena through the bias variance decomposition.

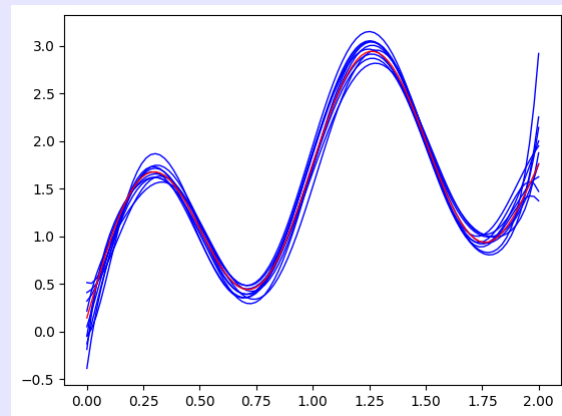
1. If we underfit the data, this means that our model is not robust and does not capture the patterns inherent in the data. It has a high bias since the set of function it encapsulates is not large enough to model $\mathbb{E}[Y | X]$. However, it has a low variance since if we were to take different samples of the dataset \mathcal{D} , the optimal parameters would not fluctuate.
2. What overfitting essentially means is that our model is too complex to the point where it starts to fit to the *noise* of the data. This means that the variance is high, since different samples of the dataset \mathcal{D} would cause huge fluctuations in the optimal trained parameters θ . However, the function set would be large, and thus it would be close to $\mathbb{E}[Y | X]$, leading to a low bias.

Example 2.4 (Polynomial Regression Continued)

Another way to reduce the overfitting problem is if we have more training data to work with. That is, if we were to fit a 9th degree polynomial on a training set of not $N = 15$, but $N = 100$ data points, then we can see that this gives a much better fit. This makes sense because now the random variable \mathcal{D} , as a function of more random variables, has lower variance. Therefore, the lower variance in the dataset translates to lower variance in the optimal parameter.



(a) $M = 9, N = 15$



(b) $M = 9, N = 100$

Figure 5: Increasing the number of data points helps the overfitting problem. Now, we can afford to fit a 9th degree polynomial with reasonable accuracy.

3 Method of Moments

4 Maximum Likelihood Estimation

5 Minimax Estimation

6 Confidence Intervals

Recall that the central limit theorem says that given a sequence of iid random variables x_1, \dots, x_n coming from a random variable with true mean μ and variance σ^2 , the sample mean is similar to a $\mathcal{N}(\mu, \sigma^2/n)$ random variable. That is, the sample mean converges in distribution

$$\bar{X}_n \xrightarrow{\text{dist}} \mathcal{N}\left(\mu, \frac{\sigma^2}{n}\right) \quad (25)$$

as $n \rightarrow \infty$. Another way to state it is that the normalized sample mean is similar to a standard Gaussian.

$$\frac{\bar{x}_n - \mu}{\sigma_{\bar{x}_n}} = \frac{\bar{x}_n - \mu}{\sigma/\sqrt{n}} \xrightarrow{\text{dist}} \mathcal{N}(0, 1) \quad (26)$$

So, given that we have enough samples, I will perfectly understand its fluctuations. Now let's introduce some definitions that will allow us to unify some ideas into simpler notation: the realized value x , the number of standard deviations it is away from the mean, and the probability that it takes that value (or more extreme).

Definition 6.1 (z-score)

Given a $\mathcal{N}(\mu, \sigma^2)$ distribution, the **z-score** of a number $x \in \mathbb{R}$ is defined to be the number of standard deviations away from the mean.

$$z = \frac{x - \mu}{\sigma} \quad (27)$$

Definition 6.2 (Percentile)

Given $X \sim \mathcal{N}(0, 1)$ and significance level $\alpha \in [0, 1]$, let us define $q_\alpha \in \mathbb{R}$ as the point where

$$\mathbb{P}(X \geq q_\alpha) = \alpha \quad (28)$$

i.e. the 100α th percentile of the standard normal. Note that given $X \sim \mathcal{N}(0, 1)$, we have

$$\mathbb{P}(|X| > q_{\alpha/2}) = \alpha \quad (29)$$

Now given x_1, \dots, x_n from a population X with mean μ and standard deviation σ , let \bar{x}_n be the sampling distribution of the mean. By virtue of the central limit theorem, we can write

$$\mathbb{P}\left(\left|\frac{\bar{X}_n - \mu}{\sigma/\sqrt{n}}\right| \geq q_{\alpha/2}\right) \approx \alpha \iff \mathbb{P}\left(\left|\frac{\bar{X}_n - \mu}{\sigma/\sqrt{n}}\right| \leq q_{\alpha/2}\right) \approx 1 - \alpha \quad (30)$$

which implies that with probability $1 - \alpha$, we have

$$\bar{X}_n \in \left[\mu - q_{\alpha/2} \frac{\sigma}{\sqrt{n}}, \mu + q_{\alpha/2} \frac{\sigma}{\sqrt{n}}\right] \iff \mu \in \left[\bar{X}_n - q_{\alpha/2} \frac{\sigma}{\sqrt{n}}, \bar{X}_n + q_{\alpha/2} \frac{\sigma}{\sqrt{n}}\right] \quad (31)$$

This is how we construct a confidence interval. In other words, as n becomes large (ideally at least 30), the probability that an interval around our sample mean contains the actual mean μ can be approximated by a Gaussian. But note that CI requires to know the actual standard deviation σ . There are three ways to deal with this:

1. This may actually be known from the start, especially if we are working with calibrated devices with standard devices that have been experimentally verified.
2. We can simply bound σ , depending on what kind of random variable we are working with. For example, given $X \sim \text{Bernoulli}(p)$, its standard deviation is bounded by $\sigma = \sqrt{p(1-p)} \leq \frac{1}{2}$, so we can create a confidence interval that is larger than any other confidence interval we can make if we had known the true σ .

$$p \in \left[\bar{X}_n - q_{\alpha/2} \frac{1}{2\sqrt{n}}, \bar{X}_n + q_{\alpha/2} \frac{1}{2\sqrt{n}}\right] \quad (32)$$

3. We can approximate σ with the sample standard deviation S , which turns out to be an unbiased estimator.

Example 6.1 (Proportion of Right-Side Kissers)

We have observed 80 out of 124 right-side kisses, resulting in a sample estimate of $\hat{p} = 0.645$. Given that we want a confidence interval of 95%, we want an $\alpha = 0.05$, implying a the value $q_{\alpha/2} = q_{0.025} = 1.96$. So, with probability 0.95, we have

$$p \in \left[0.645 - \frac{1.96}{2\sqrt{124}}, 0.645 + \frac{1.96}{2\sqrt{124}} \right] = [0.56, 0.73] \quad (33)$$

If we had, say 3 observations, rather than 124, we would have a 95% confidence interval of $p \in [0.10, 1.23]$, which is terrible, but in this case even CLT is not valid.

Example 6.2 (Proportion of Voters)

Given that we sample $n = 100$ people from a city's population to ask whether they support candidate A or B, we have 54 people who support candidate A, so $\hat{p} = 0.54$. Say that we want a 95% confidence interval, which leads to $q_{\alpha/2} = q_{0.025} = 1.96$. So, with probability 0.95, we have

$$p \in \left[0.54 - 1.96 \frac{\sigma}{\sqrt{100}}, 0.54 + 1.96 \frac{\sigma}{\sqrt{100}} \right] \quad (34)$$

and by substituting σ for $S = \sqrt{0.54(1 - 0.54)} \approx 0.5$, we get

$$p \in \left[0.54 - 1.96 \frac{0.284}{\sqrt{100}}, 0.54 + 1.96 \frac{0.284}{\sqrt{100}} \right] = [0.44, 0.64] \quad (35)$$

An interpretation of confidence intervals is that if you keep on sampling \bar{x} or \hat{p} and construct 95% CIs, then 95% of the time these intervals will contain the true mean μ or proportion p (or more if we had bounded the CI with a bigger interval).

Example 6.3 ()

We survey 6250 teachers to ask whether they think computers are essential for teaching. 250 were randomly selected and 142 felt that they were essential. Let's construct a 99% confidence interval for the proportion of teachers who felt that computers were essential. We would like to construct a CI for the true $\mu = p$, and we have $\bar{x} = 142/250 = 0.568$.

1. 99% confidence corresponds to $\alpha = 0.01$, which corresponds to a z-score of $q_{\alpha/2} = 2.576$.
2. The parent distribution is Bernoulli(p), with $\mu = p$ and $\sigma = \sqrt{p(1-p)}$. The sampling distribution of \bar{x} has $\mu_{\bar{x}} = p$ also and $\sigma_{\bar{x}} = \sigma/\sqrt{n}$.
3. We need to know the details of the sampling distribution, but we don't know σ , which is needed to calculate $\sigma_{\bar{x}}$. However, we can estimate it using the sample standard deviation $S = \sqrt{0.568(1 - 0.568)} = 0.5$.
4. Our sampling distribution has standard deviation $\sigma_{\bar{x}} \approx S/\sqrt{n} = 0.5/\sqrt{250} = 0.031$, and our z-score was 2.576, so our 99% confidence interval is 2.576 standard deviations from our mean. That is, with probability 0.99,

$$p \in [0.568 - 2.576 \cdot 0.031, 0.568 + 2.576 \cdot 0.031] = [0.488144, 0.647856] \quad (36)$$

6.1 CIs for means, proportions, and variances

6.2 Bootstrap confidence intervals

7 Cross Validation

We have understood the theoretical foundations of overfitting and underfitting with the bias variance decomposition. But in practice, we don't have an ensemble of datasets; we just have one. Therefore, we don't actually know what the bias, the variance, or the noise is at all. Therefore, how do we actually *know* in practice when we are underfitting or overfitting? Easy. We just split our dataset into 2 different parts: the training set and testing sets.

$$\mathcal{D} = \mathcal{D}_{train} \sqcup \mathcal{D}_{test} \quad (37)$$

What we usually have is a **training set** that allows us to train the model, and then to check its performance we have a **test set**. We would train the model on the training set, where we will always minimize the loss, and then we would look at the loss on the test set. Though we haven't made a testing set, since we know the true model let us just generate more data and use that as our testing set. For each model, we can calculate the optimal θ , which we will denote θ^* , according to the **root mean squared loss**

$$h_{\theta^*} = \underset{h_{\theta}}{\operatorname{argmin}} \sqrt{\frac{1}{N} \sum_{i=1}^N (y^{(i)} - h_{\theta}(\mathbf{x}^{(i)}))^2} \quad (38)$$

where division of N allows us to compare different sizes of datasets on equal footing, and the square root ensures that this is scaled correctly. Let us see how well these different order models perform on a separate set of data generated by the same function with Gaussian noise.

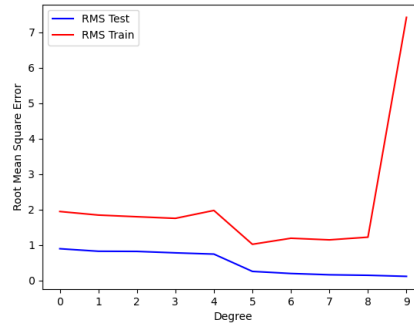


Figure 6: We can see that the RMS decreases monotonically on the training error as more complex functions become more fine-tuned to the data. However, when we have a 9th degree polynomial the RMS for the testing set dramatically increases, meaning that this model does not predict the testing set well, and performance drops.

Now we know that a more complex model (i.e. that captures a greater set of functions) is not necessarily the best due to overfitting. Therefore, researchers perform **cross-validation** by taking the training set $(\mathcal{X}, \mathcal{Y})$. We divide it into S equal pieces

$$\bigcup_{s=1}^S D_s = (\mathcal{X}, \mathcal{Y}) \quad (39)$$

Then, we train the model \mathcal{M} on $S - 1$ pieces of the data and then test it across the final piece, and do this S times for every test piece, averaging its performance across all S test runs. Therefore, for every model \mathcal{M}_k , we must train it S times, for all K models, requiring KS training runs. If data is particularly scarce, we set $S = N$, called the **leave-one-out** technique. Then we just choose the model with the best average test performance.

The following result shows that cross-validation (data splitting) leads to an estimator with risk nearly as good as the best model in the class.

Theorem 7.1 (Gyorfi, Kohler, Krzyak, Walk (2002))

Let $\mathcal{M} = \{m_h\}$ be a finite class of regression estimators indexed by a parameter h , with m being the true risk minimizer, $m_{\hat{h}}$ being the empirical risk minimizer over the whole dataset \mathcal{D} , and m_H being the empirical risk minimizer over the test set $\mathcal{D}_{\text{test}}$ for ordinary least squares loss.

$$m_H = \operatorname{argmin}_{m_h} \frac{1}{N} \sum_{i \in \mathcal{D}_{\text{test}}} (y_i - m_h(x_i))^2 \quad (40)$$

$$m_{\hat{h}} = \operatorname{argmin}_{m_h} \frac{1}{N} \sum_{i \in \mathcal{D}} (y_i - m_h(x_i))^2 \quad (41)$$

If the data Y_i and estimators are bounded by L , then for any $\delta > 0$, we have

$$\mathbb{E} \int |m_H(x) - m(x)|^2 d\mathbb{P}(x) \leq (1 + \delta) \mathbb{E} \int |m_{\hat{h}}(x) - m(x)|^2 d\mathbb{P}(x) + \frac{C(1 + \log |M|)}{n} \quad (42)$$

where $c = L^2(16/\delta + 35 + 19\delta)$.

Code 7.1 (Minimal Example of Train Test Split in scikit-learn)

To implement this in scikit-learn, we want to use the `train_test_split` class. We can also set a random state parameter to reproduce results.

```
1 from sklearn.model_selection import train_test_split
2
3 # Split into training (80\%) and test (20\%) data
4 X_train, X_test, Y_train, Y_test = train_test_split(X, Y, test_size=0.2,
    random_state=66)
```

However, this process requires a lot of training runs and therefore may be computationally infeasible. Therefore, various **information criterion** has been proposed to efficiently select a model.

7.1 Leave 1 Out Cross Validation**7.1.1 Generalized (Approximate) Cross Validation****7.1.2 Cp Statistic****7.2 K Fold Cross Validation**

8 Information Criteria

References

[Arb10] J. Arbuthnot. An argument for divine providence. *Philosophical Transactions*, 27:186–190, 1710.