Parity Asymmetry in the CMB UF REU under Dr. Zachary Slepian

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Outline

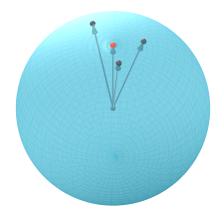
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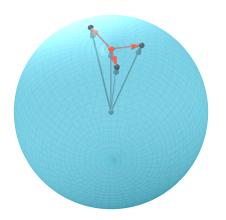
Overview of the problem

The goal is to determine whether or not parity asymmetry can be detected using the Cosmic Microwave Background. To do so we are modelling the 4-point correlation function as a function of 4 points on a sphere, defined by vectors $\vec{x}_0, \vec{x}_1, \vec{x}_2, \vec{x}_3$



Overview of the problem

We then define a set of vectors $\vec{r}_1, \vec{r}_2, \vec{r}_3$ where $\vec{r}_i = \vec{x}_i - \vec{x}_0$



Overview of the probelm

If we then take two correlation functions:

$$\zeta_E(\vec{x}_0,\vec{x}_1,\vec{x}_2,\vec{x}_3)$$

which is the correlation function of the vectors from the Earth the the CMB and

$$\zeta_{x_0}(\vec{r_1},\vec{r_2},\vec{r_3})$$

which is the correlation function centered on point x_0 , then our goal is to write one as a function of the other in order to help constrain the possible parity of both

Start by expanding the two correlation functions in terms of isotropic basis functions:

$$\zeta_{X_0}(\vec{r}_1, \vec{r}_2, \vec{r}_3) = \sum_{\Lambda} \tilde{\zeta}_{\Lambda}(r_1, r_2, r_3) \mathcal{P}_{\Lambda}(\hat{r}_1, \hat{r}_2, \hat{r}_3)
\zeta_{E}(\vec{x}_0, \vec{x}_1, \vec{x}_2, \vec{x}_3) = \sum_{\Lambda'} \tilde{\zeta}_{\Lambda'}(x_0, x_1, x_2, x_3) \mathcal{P}_{\Lambda'}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3)$$

Where Λ is the set $\{l_1, l_2, l_3\}$ and Λ' is the set $\{l_0', l_1', l_{01}', l_2', l_3'\}$

We can then set the two functions equal since they are correlation functions describing the same points

$$\zeta_{x_{0}}(\vec{r}_{1}, \vec{r}_{2}, \vec{r}_{3}) = \zeta_{E}(\vec{x}_{0}, \vec{x}_{1}, \vec{x}_{2}, \vec{x}_{3})
\Downarrow
\sum_{\Lambda} \zeta_{\Lambda}(r_{1}, r_{2}, r_{3}) \mathcal{P}_{\Lambda}(\hat{r}_{1}, \hat{r}_{2}, \hat{r}_{3}) = \sum_{\Lambda'} \zeta_{\Lambda'}(x_{0}, x_{1}, x_{2}, x_{3}) \mathcal{P}_{\Lambda'}(\hat{x}_{0}, \hat{x}_{1}, \hat{x}_{2}, \hat{x}_{3})$$

We now multiply both sides by $\mathcal{P}^*_{\Lambda''}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3)$ and then integrate over all orientations $d\hat{x}_0 d\hat{x}_1 d\hat{x}_2 d\hat{x}_3 = d\hat{x}$

$$\int d\hat{x} \sum_{\Lambda} \zeta_{\Lambda}(r_1, r_2, r_3) \mathcal{P}_{\Lambda}(\hat{r}_1, \hat{r}_2, \hat{r}_3) P^*_{\Lambda''}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3) =$$

$$\int d\hat{x} \sum_{\Lambda'} \zeta_{\Lambda'}(x_0, x_1, x_2, x_3) \mathcal{P}_{\Lambda'}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3) P^*_{\Lambda''}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3)$$

Using the orthogonality of isotropic basis functions which states

$$\int d\hat{R} \; \mathcal{P}_{\Lambda}(\hat{R}) \mathcal{P}_{\Lambda'}^*(\hat{R}) = \delta_{\Lambda_1 \Lambda_1'} \delta_{\Lambda_2 \Lambda_2'} ...$$

we can simplify the right hand side of the previous equation:

$$\int d\hat{x} \sum_{\Lambda'} \zeta_{\Lambda'}(x_0, x_1, x_2, x_3) \mathcal{P}_{\Lambda'}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3) \mathcal{P}_{\Lambda''}(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3)$$

$$= \zeta_{\Lambda''}(x_0, x_1, x_2, x_3)$$

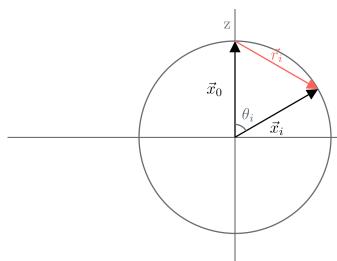
Plugging this back into the previous equation we get the key relation:

$$\frac{\zeta_{\Lambda''}(x_0, x_1, x_2, x_3) = \int d\hat{x} \sum_{\Lambda} \zeta_{\Lambda}(r_1, r_2, r_3) \mathcal{P}_{\Lambda}(\hat{r}_1, \hat{r}_2, \hat{r}_3) P_{\Lambda''}^*(\hat{x}_0, \hat{x}_1, \hat{x}_2, \hat{x}_3)}{(1)}$$

Solving the integral on the right will give an expression for the coefficients of the correlation function centered on earth in terms of the correlation function centered at x_0

Solving for \vec{r} in terms of \vec{x}

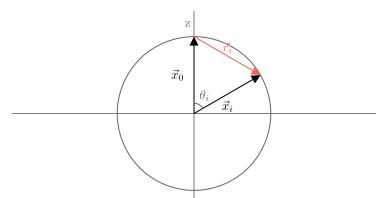
Consider a great circle going through the poles at an angle ϕ_i to the x-axis where ϕ_i is the ϕ angle of vector $\vec{x_i}$ for i=1,2,3



Solving for \vec{r} in terms of \vec{x}

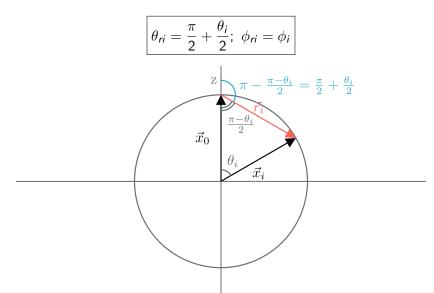
Using the law of Cosines and defining the distance to the cmb to be d^* we get:

$$|\vec{r_i}| = \sqrt{d^{*2} + d^{*2} - 2d^*d^*\cos(\theta_i)}$$
$$r_i = \sqrt{2}d^*\sqrt{1 - \cos(\theta_i)}$$



Solving for \vec{r} in terms of \vec{x}

Similarly, we can solve for the angular dependence of $\vec{r_i}$



Solving for the coefficients

Plugging those back into equation 1 we get:

$$\zeta_{\Lambda''}(x_{0}, x_{1}, x_{2}, x_{3}) =
\int d\hat{x} \sum_{\Lambda} \zeta_{\Lambda}(\sqrt{2}d^{*}\sqrt{1 - \cos(\theta_{1})}, \sqrt{2}d^{*}\sqrt{1 - \cos(\theta_{2})}, \sqrt{2}d^{*}\sqrt{1 - \cos(\theta_{3})})
\times \mathcal{P}_{\Lambda}(\frac{\pi}{2} + \frac{\theta_{1}}{2}, \phi_{1}; \frac{\pi}{2} + \frac{\theta_{2}}{2}, \phi_{2}; \frac{\pi}{2} + \frac{\theta_{3}}{2}, \phi_{3})
\times \mathcal{P}_{\Lambda''}^{*}(\theta_{0}, \phi_{0}; \theta_{1}, \phi_{1}; \theta_{2}, \phi_{2}; \theta_{3}, \phi_{3})$$

Different approaches

- Directly solving the integral
- Rotating the spherical harmonics
- Using an addition formula
 - Associated Legendre Polynomials
 - Jacobi Polynomials

Directly solving the integral

We can expand the integral using the definitions of the isotropic basis functions:

$$\mathcal{P}_{l_1 l_2 l_3}(\hat{r}_1, \hat{r}_2, \hat{r}_3) = (-1)^{l_1 + l_2 + l_3} \sum_{m_1 m_1 m_3} \begin{pmatrix} l_1 & l_2 & l_3 \\ m_1 & m_2 & m_3 \end{pmatrix} \times Y_{l_1 m_1}(\hat{r}_1) Y_{l_2 m_2}(\hat{r}_2) Y_{l_3 m_3}(\hat{r}_3)$$

$$\mathcal{P}_{l_1 l_2(l_{12}) l_3 l_4}(\hat{r}_1, \hat{r}_2, \hat{r}_3, \hat{r}_4) = (-1)^{l_1 + l_2 + l_3 + l_4}$$

$$\times \sum_{m_{12}} (-1)^{l_{12} - m_{12}} \sum_{m_1 m_2 m_3 m_4} \begin{pmatrix} l_1 & l_2 & l_{12} \\ m_1 & m_2 & -m_{12} \end{pmatrix} \begin{pmatrix} l_{12} & l_3 & l_4 \\ m_{12} & m_3 & -m_4 \end{pmatrix}$$

$$\times Y_{l_1 m_1}(\hat{r}_1) Y_{l_2 m_2}(\hat{r}_2) Y_{l_3 m_3}(\hat{r}_3) Y_{l_4 m_4}(\hat{r}_4)$$

Directly solving the integral

Looking at just the angular dependent parts of the integral we have:

$$\int d\theta_0 d\phi_0 d\theta_1 d\phi_1 d\theta_2 d\phi_2 d\theta_3 d\phi_3 sin(\theta_0) sin(\theta_1) sin(\theta_2) sin(\theta_3)$$

$$\times Y_{l_0'm_0'}(\theta_0, \phi_0) Y_{l_1'm_1'}(\theta_1, \phi_1) Y_{l_2'm_2'}(\theta_2, \phi_2) Y_{l_3'm_3'}(\theta_3, \phi_3)$$

$$\times Y_{l_1m_1}(\frac{\theta_1}{2} + \frac{\pi}{2}, \phi_1) Y_{l_2m_2}(\frac{\theta_2}{2} + \frac{\pi}{2}, \phi_2) Y_{l_3m_3}(\frac{\theta_3}{2} + \frac{\pi}{2}, \phi_3)$$

Directly solving the integral

We can further separate the integral into ϕ and θ dependent parts by using the following definition of sperical harmonics:

$$Y_{l}^{m}(\theta,\phi) = \sqrt{\frac{2l+1}{4\pi} \frac{(l-m)!}{(l+m)!}} P_{l}^{m}(\cos\theta) e^{im\phi}$$

Solving the ϕ integral

Looking at just the ϕ dependent terms:

$$\int_{0}^{2\pi} d\phi_{0} d\phi_{1} d\phi_{2} d\phi_{3} e^{im'_{0}\phi_{0}} e^{im'_{1}\phi_{1} + im_{1}\phi_{1}} e^{im'_{2}\phi_{2} + im_{2}\phi_{2}} e^{im'_{3}\phi_{3} + im_{3}\phi_{3}}$$

$$= \left[\frac{1}{im'_{0}} e^{im'_{0}\phi_{0}} \Big|_{0}^{2\pi} \right] \left[\frac{1}{i(m'_{j} + m_{j})} e^{i(m'_{j} + m_{j})\phi_{j}} \Big|_{0}^{2\pi} \right] \text{ for } j = 1, 2, 3$$

Each of the above integrals will always be zero since all m's are integers unless $m_0'=0$ and $m_j'+m_j=0$ in which case each term becomes 2π so the total ϕ integral becomes:

$$16\pi^4\delta^K_{m_0',0}\delta^K_{m_1,-m_1'}\delta^K_{m_2,-m_2'}\delta^K_{m_3,-m_3'}$$

Examining the 3j symbols

Our original expression had three 3j symbols:

$$\begin{pmatrix} l'_0 & l'_1 & l'_{01} \\ m'_0 & m'_1 & -m'_{01} \end{pmatrix} \begin{pmatrix} l'_{01} & l'_2 & l'_3 \\ m'_{01} & m'_2 & m'_3 \end{pmatrix} \begin{pmatrix} l_1 & l_2 & l_3 \\ m_1 & m_2 & m_3 \end{pmatrix}$$

If we now plug in the constraints given by the delta functions from the ϕ integral we get:

$$\begin{pmatrix} l'_0 & l'_1 & l'_{01} \\ 0 & -m_1 & -m'_{01} \end{pmatrix} \begin{pmatrix} l'_{01} & l'_2 & l'_3 \\ m'_{01} & -m_2 & -m_3 \end{pmatrix} \begin{pmatrix} l_1 & l_2 & l_3 \\ m_1 & m_2 & m_3 \end{pmatrix}$$

Since the bottom elements of the 3j symbol must sum to zero, this adds the additional constraint that $m_{01}^\prime=-m_1$

Solving the θ integral

Looking at the θ integral, we have:

$$\int_0^{2\pi} d\theta_0 d\theta_1 d\theta_2 d\theta_3 sin(\theta_0) sin(\theta_1) sin(\theta_2) sin(\theta_3)$$