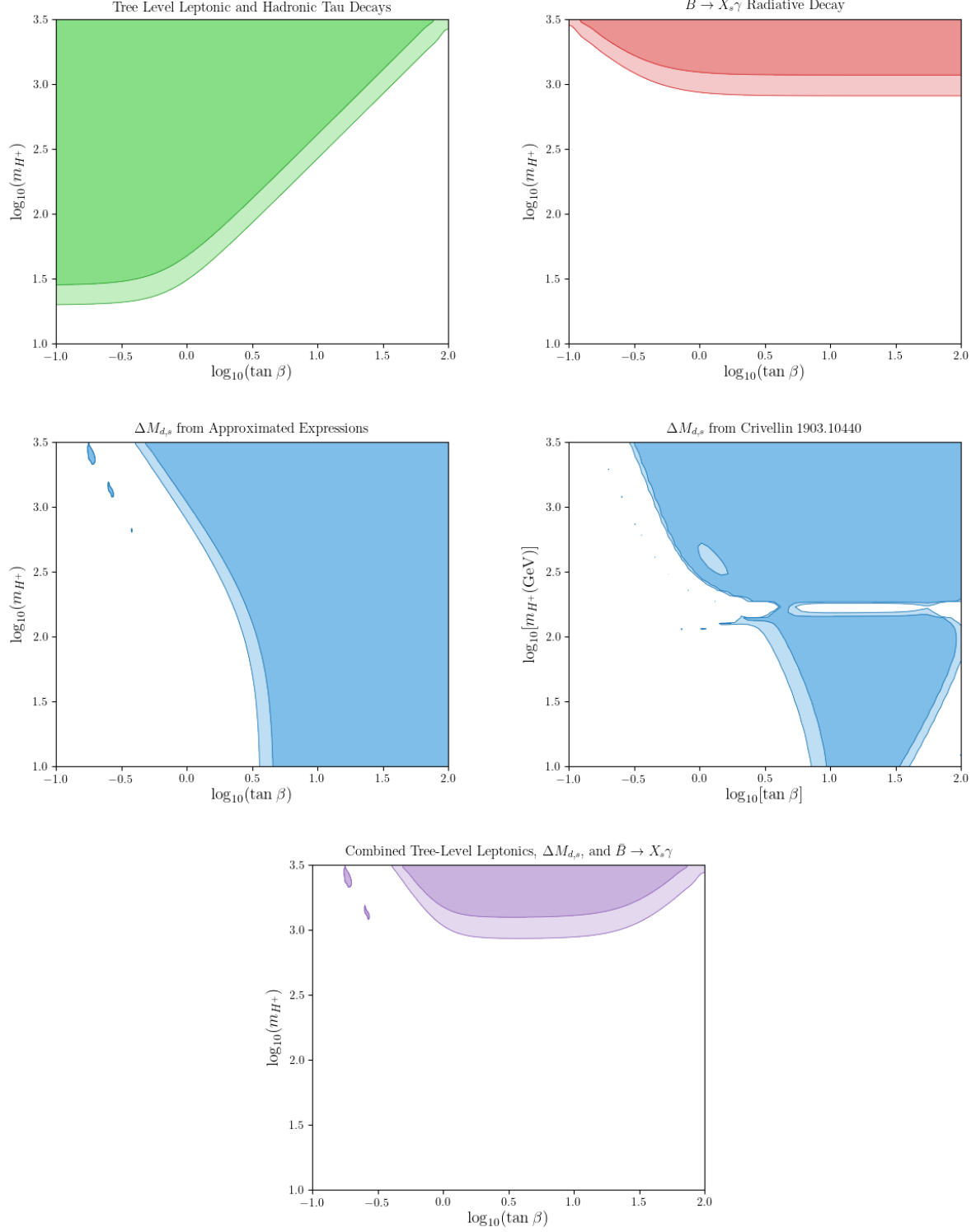


# Project Notes

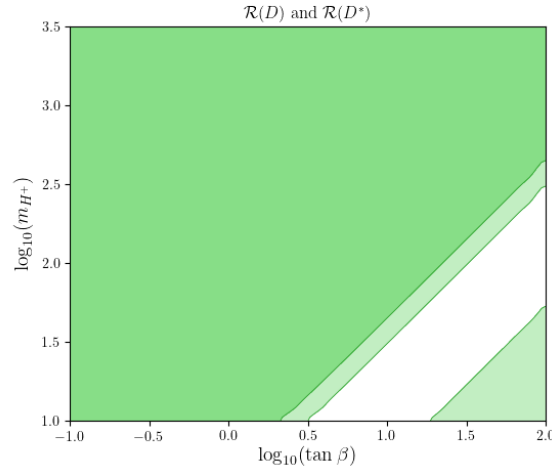
## 1 Replicating results

- Here are my plots for replicating what we did in first term:

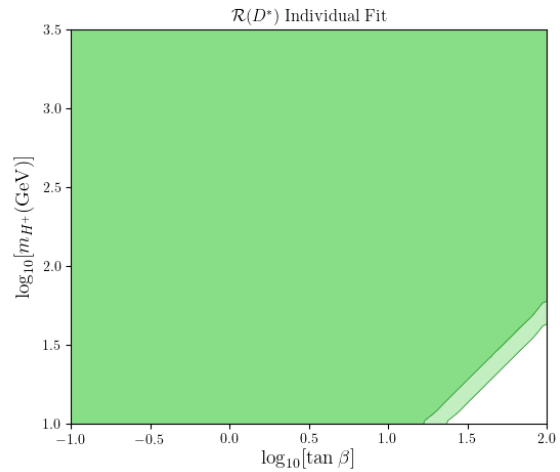
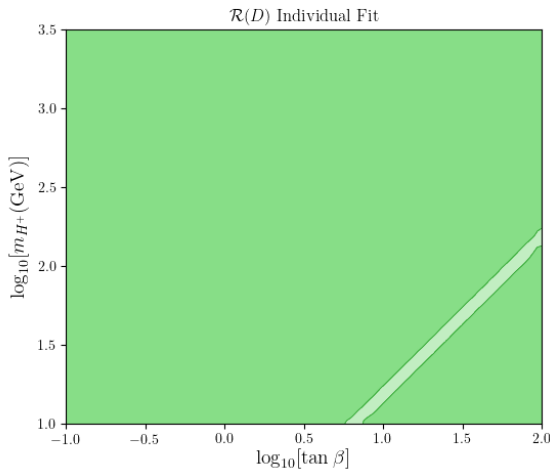


- Like Oliver, the main difference seems to be with  $b \rightarrow s \gamma$  where it fits higher than before – **the combined fit yields  $m_{H^+} \gtrsim 970 \text{ GeV}$  at  $2\sigma$  compared to  $m_{H^+} \gtrsim \text{GeV}$  in my previous fits**
- *Note: all the plots here are showing contours for  $1, 2\sigma$*

- I have modified the SM value in flavio for  $b \rightarrow s\gamma$  to fit the current result of  $(3.40 \pm 0.17)e^{-4}$  but still got this higher result
- It seems like it could be down to how flavio fits, but we would need to confirm this if we want to leave it as is
- I have also included the fit for B mixing from higher-order expressions in [Crivellin 1903.10440](#) to show the difference
- Adding in  $B_{s,d} \rightarrow \mu\mu$  and  $\mathcal{R}(D^{(*)})$  yields similar results to what I had before



- $\mathcal{R}(D)$  and  $\mathcal{R}(D^*)$  are fitting simultaneously it seems right now, which wasn't what we had before, so looking at individual plots:

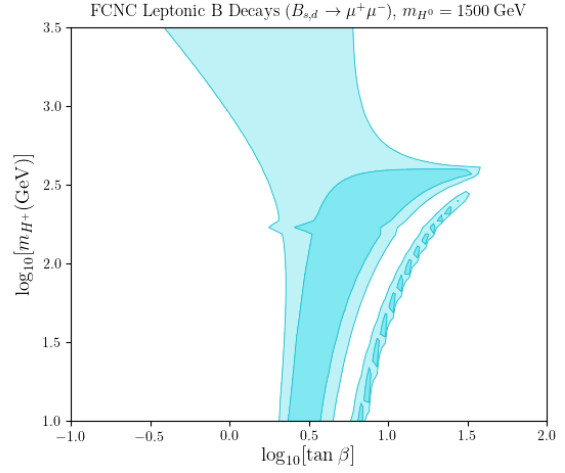
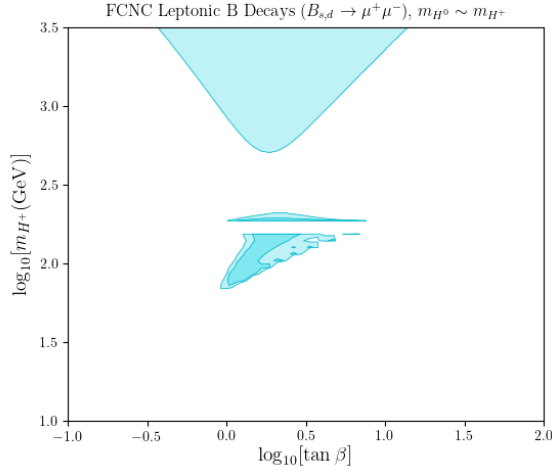


- It appears that both of them do fit individually which is not what we had before – I'm not quite sure why this is the case currently
- For  $B_{s,d} \rightarrow \mu\mu$ , the left diagram is approximating  $m_{H^+} \sim m_{H^0}$  which is the rough limit James has from the obliques, and the right diagram is fixing  $m_{H^0} = 1500$  GeV
- Using the convention from 1903.10440 for the trilinear couplings means that instead of using  $M = m_{12}/(\sin \beta \cos \beta)$  as I did previously, you use  $\lambda_3$  from the 2HDM potential
- The two trilinear couplings we have to consider are

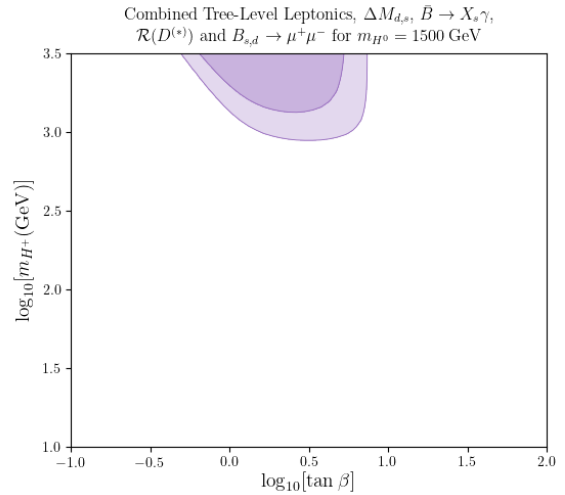
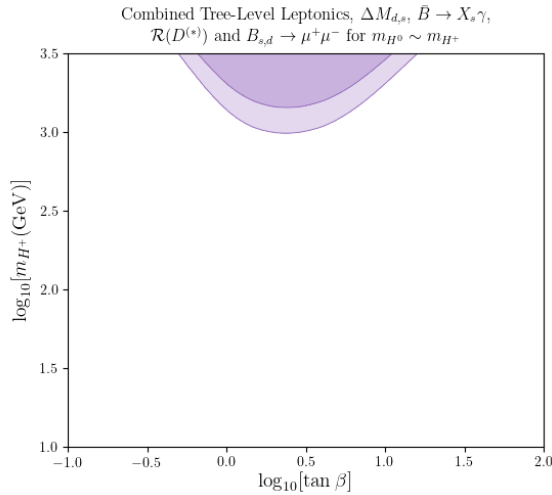
$$\lambda_{h^0 H^+ H^-} = v \sin(\beta - \alpha) \lambda_3,$$

$$\lambda_{H^0 H^+ H^-} = v \cos(\beta - \alpha) \lambda_3$$

- In the alignment limit which I have been using so far for these,  $\sin(\beta - \alpha) = 1$ , so we only have to consider  $\lambda_{h^0 H^+ H^-}$



- The contributions from the trilinear coupling seems to be minimal anyway, varying  $\lambda_3$  from  $0.01 \rightarrow 2$  doesn't change the results to any noticeable level, so for now I have set  $\lambda_3 = 0.1$ , but it's probably better to fit it properly at some point
- The coupling only contributes to  $C_S$  and  $C_P$  operators so far which only impact  $B \rightarrow \mu\mu$
- Checking the combined fit for all these observables to compare to my overall project work (bar the inclusion of James' obliques):



- Left is approximating  $m_{H^+} \sim m_{H^0}$  as above for  $B \rightarrow \mu\mu$  (can't compare to project directly as didn't use this limit); right is fixing  $m_{H^0} = 1500$  GeV
- For the right, we have  $m_{H^+} \gtrsim 970$  GeV and  $\tan \beta \lesssim 7$ ; in my project for this, I had  $m_{H^+} \gtrsim 500$  GeV and  $\tan \beta \lesssim 21$

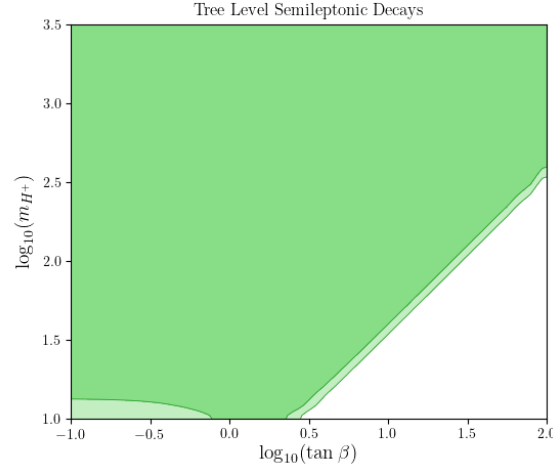
## 2 New Observables and To Do

- Also started looking at the tree-level semileptonic decays
- For semileptonics and leptonics, I think the WC contributions work out the same, e.g. for  $b \rightarrow u\mu\nu_\mu$  (using the subscript convention from [flavio's WET basis](#))

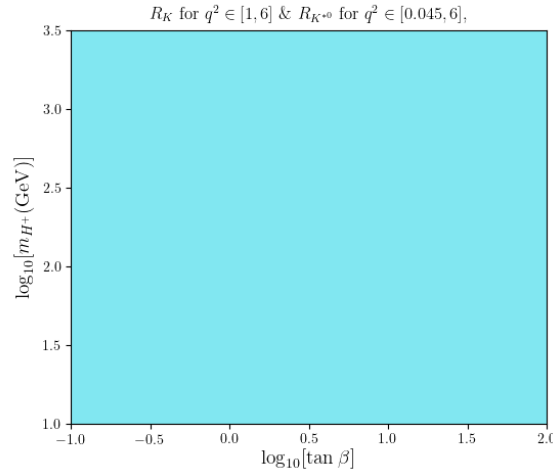
$$\mathcal{O}_{SR} = -\frac{4G_F}{\sqrt{2}} V_{ub}(\bar{u}_L b_R)(\bar{\mu}_R \nu_{\mu L}) \rightarrow C_{SR} = \frac{m_u}{m_{H^+}^2}$$

$$\mathcal{O}_{SL} = -\frac{4G_F}{\sqrt{2}} V_{ub}(\bar{u}_R b_L)(\bar{\mu}_L \nu_{\mu R}) \rightarrow C_{SL} = \frac{m_b \tan^2 \beta}{m_{H^+}^2}$$

- For the tree-level leptonics, this transforms to  $r_H$  from 0907.5135, and looks to give the right results for semileptonics (including being used for the  $\mathcal{R}(D)$ s above)
- So providing the SM calculations for the semileptonics are fine, it's quite simple to fit these too:



- The next thing is to look at  $R_K$  and  $R_{K^*0}$
- From 1704.05340, the operators needed for the  $R_K$ s are  $C_7, C_7', C_9, C_9', C_{10}, C_{10}'$  for both  $bs\mu\mu$  and  $bsee$
- All these can be got from 1903.10440 – already have the formulae for  $C_{10}^{(\prime)}$ s from  $B \rightarrow \mu\mu$  calculations
- For both  $m_{H^+} \sim m_{H^0}$  and  $m_{H^0} = 1500 \text{ GeV}$  as I looked at  $B \rightarrow \mu\mu$ , we get no constraint in our parameter space from  $R_K$  and  $R_{K^*0}$  (although I still need to check through all the formulae again for mistakes):



- I want to check in literature as well for the  $R_K$ s in 2HDM Type II to see if there's any fits that include these to see what results have been found historically to give some at least a rough idea of what should be expected
- I will also look at fitting  $b \rightarrow s\gamma$  using the expressions for  $C_7, C_8$  in 1903.10440 since  $C_7$  is in  $R_K$ s anyway
- I need to do the fits including  $B \rightarrow \mu\mu$  in the wrong sign limit too and compare to my previous results