

1. a. $\phi_\nu(\nu) = \frac{1}{\sqrt{\pi}} \frac{1}{\nu_0} \frac{c}{b} e^{-\nu^2/b^2}$

from Draine, $\lambda_0 = c/\nu_0$
and $\frac{v}{c} = \frac{\lambda - \lambda_0}{\lambda_0}$ from the Doppler formula

so $v = c \frac{\lambda - \lambda_0}{\lambda_0} = cx$.

$\therefore \phi(x) = \frac{\lambda_0}{\sqrt{\pi} b} e^{-(cx/b)^2}$

$\tau_\nu = \frac{\pi e^2}{m_e c} f_{lu} N_l \phi_\nu \left[1 - \frac{N_u/g_u}{N_l/g_l} \right]$ neglect stim. emission

$\therefore \tau(x) = \frac{\pi e^2}{m_e c} f_{lu} N_l \left(\frac{\lambda_0}{\sqrt{\pi} b} e^{-c^2 x^2/b^2} \right)$

$$\tau(x) = \frac{\sqrt{\pi} e^2 \lambda_0}{m_e c b} f_{lu} N_l e^{-c^2 x^2/b^2}$$

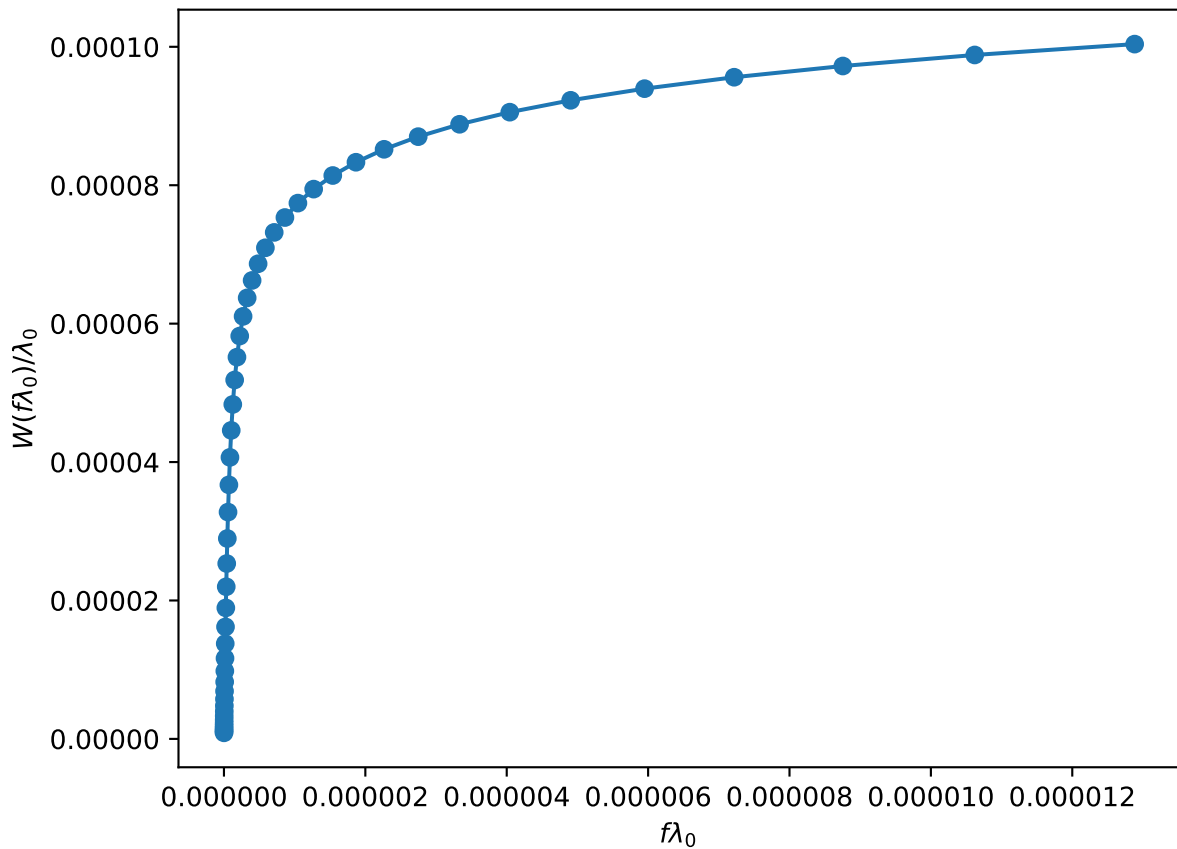
b. $W = \int_0^\infty (1 - e^{N \lambda_0 \phi_\nu}) \frac{d\nu}{\nu_0}$

$x = \nu_0 \left(\frac{1}{\nu} - \frac{1}{\nu_0} \right) = \frac{\nu_0}{\nu} - 1 \Rightarrow dx = -\frac{\nu_0}{\nu^2} d\nu$

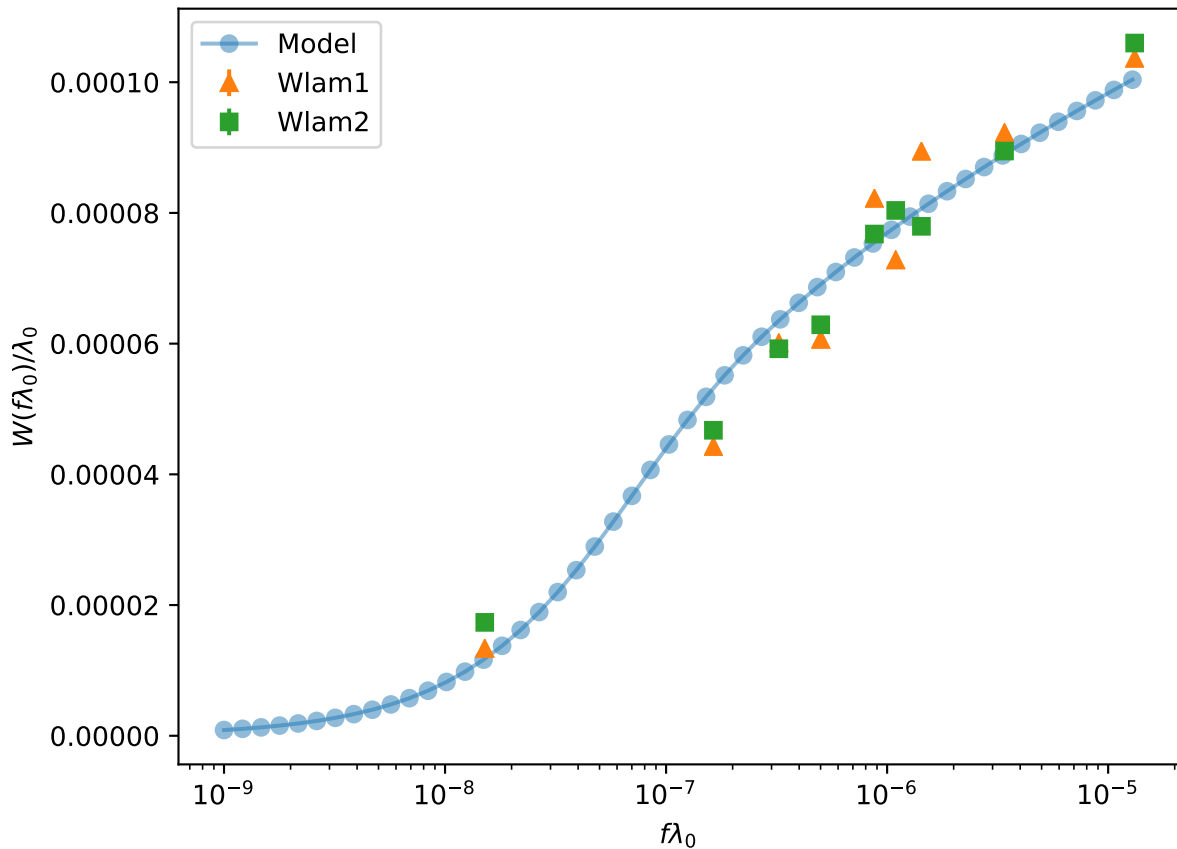
$\nu = \frac{\nu_0}{(x+1)} \quad \therefore d\nu = -\frac{\nu^2 dx}{\nu_0} = -\frac{\nu_0 dx}{(x+1)^2}$

$$\frac{W_\lambda}{\lambda_0} = \int_{-1}^{\infty} \frac{1 - e^{-\tau(x)}}{(x+1)^2} dx$$

$N=1.0\text{e}+15 \text{ 1 / cm}^2$, $b=6.0\text{e}+05 \text{ cm / s}$



$N=1.0\text{e}+15 \text{ 1 / cm}^2$, $b=6.0\text{e}+05 \text{ cm / s}$



```

import astropy.constants as c
import astropy.units as u
import numpy as np
import matplotlib.pyplot as plt
import pdb

N = 10**15 * u.cm**-2
b = (6 * u.km/u.s).cgs
x = np.linspace(-5, 5) * (b/c.c.cgs).value
flam = np.logspace(-9, -4.89)

def tau(x, flam):
    prefac = np.sqrt(np.pi)*c.e.esu**2/c.m_e.cgs/c.c.cgs
    (prefac * flam * N / b * np.exp(-(c.c.cgs*x/b)**2)).value

def W(x, flam):
    dx = np.diff(x)[0]
    sol_array = np.zeros(np.shape(flam))
    for i, f in enumerate(flam):
        sol_array[i] = np.sum((1 - np.exp(-tau(x, f)))/(x+1)**2*dx)
    sol_array

plt.plot(flam, W(x, flam), marker='o')
plt.xlabel('$f\lambda_{0}$')
plt.ylabel('$W(f\lambda_{0}) / \lambda_{0}$')
plt.title('N={:.1e}, b={:.1e}'.format(N, b))
plt.tight_layout()
plt.show()

data = np.loadtxt('cog.data.txt', skiprows=1)
lam, f, wlam1, wlam1_err, wlam2, wlam2_err = data.T * u.angstrom.to(u.cm)
f = f/u.angstrom.to(u.cm)

plt.plot(flam, W(x, flam), marker='o', label='Model', alpha=0.5)
plt.errorbar(f*lam, wlam1/lam, yerr=wlam1_err, marker='^', linestyle='None',
label='Wlam1')
plt.errorbar(f*lam, wlam2/lam, yerr=wlam2_err, marker='s', linestyle='None',
label='Wlam2')
plt.xlabel('$f\lambda_{0}$')
plt.ylabel('$W(f\lambda_{0}) / \lambda_{0}$')
plt.title('N={:.1e}, b={:.1e}'.format(N, b))
plt.xscale('log')
plt.legend()
plt.tight_layout()
plt.show()

```