

**Assignment #6**  
 Introduction to GR, 2020 Fall  
 International Centre for Theoretical Sciences  
 Instructor: Prof. Bala Iyer, Tutor: Md Arif Shaikh<sup>1</sup>  
 Due on Dec 9, 2020 11:59 PM.

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1. (a) Show that under an infinitesimal coordinate transformation  $x^{\text{new}\mu} = x^{\text{old}\mu} + \xi^\mu$ ,  $h_{\mu\nu}$  transforms as

$$h_{\mu\nu}^{\text{new}} = h_{\mu\nu}^{\text{old}} - \partial_\mu \xi_\nu - \partial_\nu \xi_\mu \quad (1)$$

and the trace reversed metric perturbation transforms as

$$\bar{h}_{\mu\nu}^{\text{new}} = \bar{h}_{\mu\nu}^{\text{old}} - \partial_\mu \xi_\nu - \partial_\nu \xi_\mu + \eta_{\mu\nu} \partial_\alpha \xi^\alpha \quad (2)$$

- (b) Assuming solution of the form  $\bar{h}_{\mu\nu} = A_{\mu\nu} \exp(ik_\alpha x^\alpha)$  and  $\xi_\mu = B_\mu \exp(ik_\alpha x^\alpha)$ , express eq. 2 in terms of  $A_{\mu\nu}$  and  $B_\mu$ . Find the Lorenz gauge condition on  $A_{\mu\nu}$ . Show that  $A_{\mu\nu}^{\text{new}}$  satisfies the gauge condition if  $A_{\mu\nu}^{\text{old}}$  does.
- (c) Find the constrain on  $B_\mu$  in order to make  $A_{\mu\nu}$  traceless.
- (d) Show that the condition  $A_{\mu\nu} u^\nu = 0$ , imposes only three constraints on  $B_\mu$ . ( $u^\mu$  is a constant time-like four-velocity). Do this by showing that the combination  $k^\mu (A_{\mu\nu} u^\nu)$  vanishes for any  $B_\mu$ .
- (e) Use the results of the previous two questions to solve for  $B^\mu$  as a function of  $k^\mu$ ,  $A_{\mu\nu}^{\text{old}}$  and  $u^\mu$ .
- (f) Show that it is possible to choose  $\xi^\alpha$  in eq. 1 to make any superposition of plane waves satisfy the equations  $A_{\mu\nu} u^\mu = 0$  and  $A_\mu^\mu = 0$ , so that these are generally applicable to gravitational waves of any sort.
- (g) Show that we cannot achieve  $A_{\mu\nu} u^\mu = 0$  and  $A_\mu^\mu = 0$  for a static solution, i.e. one for which  $\omega = 0$ .
2. The equation  $(\nabla^2 + \Omega^2)B_{\mu\nu} = -16\pi S_{\mu\nu}$  in the vacuum region outside the source – i.e. where  $S_{\mu\nu} = 0$  – can be solved by separation of variables. Assume a solution for  $B_{\mu\nu}$  of the form  $\sum_{lm} A_{\mu\nu}^{lm} f_l(r) Y_{lm}(\theta, \phi)/\sqrt{r}$ , where  $Y_{lm}$  is the spherical harmonic.
- (a) Show that  $f_l(r)$  satisfies the equation

$$f_l'' + \frac{1}{r} f_l' + (\Omega^2 - \frac{(l + \frac{1}{2})^2}{r^2}) f_l = 0 \quad (3)$$

- (b) Show that the most general spherically symmetric solution is given by

$$B_{\mu\nu} = \frac{A_{\mu\nu}}{r} e^{i\Omega r} + \frac{Z_{\mu\nu}}{r} e^{-i\Omega r}. \quad (4)$$

3. Show that

$$\oint \vec{n} \cdot \nabla B_{\mu\nu} dS \approx -4\pi A_{\mu\nu}. \quad (5)$$

(Assume that the source is nonzero only inside a sphere of radius  $\epsilon \ll 2\pi/\Omega$ .)

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