

# Searches for Ultra Long-Lived Particles

with **MATISSE**

TRIUMF COLLOQUIUM

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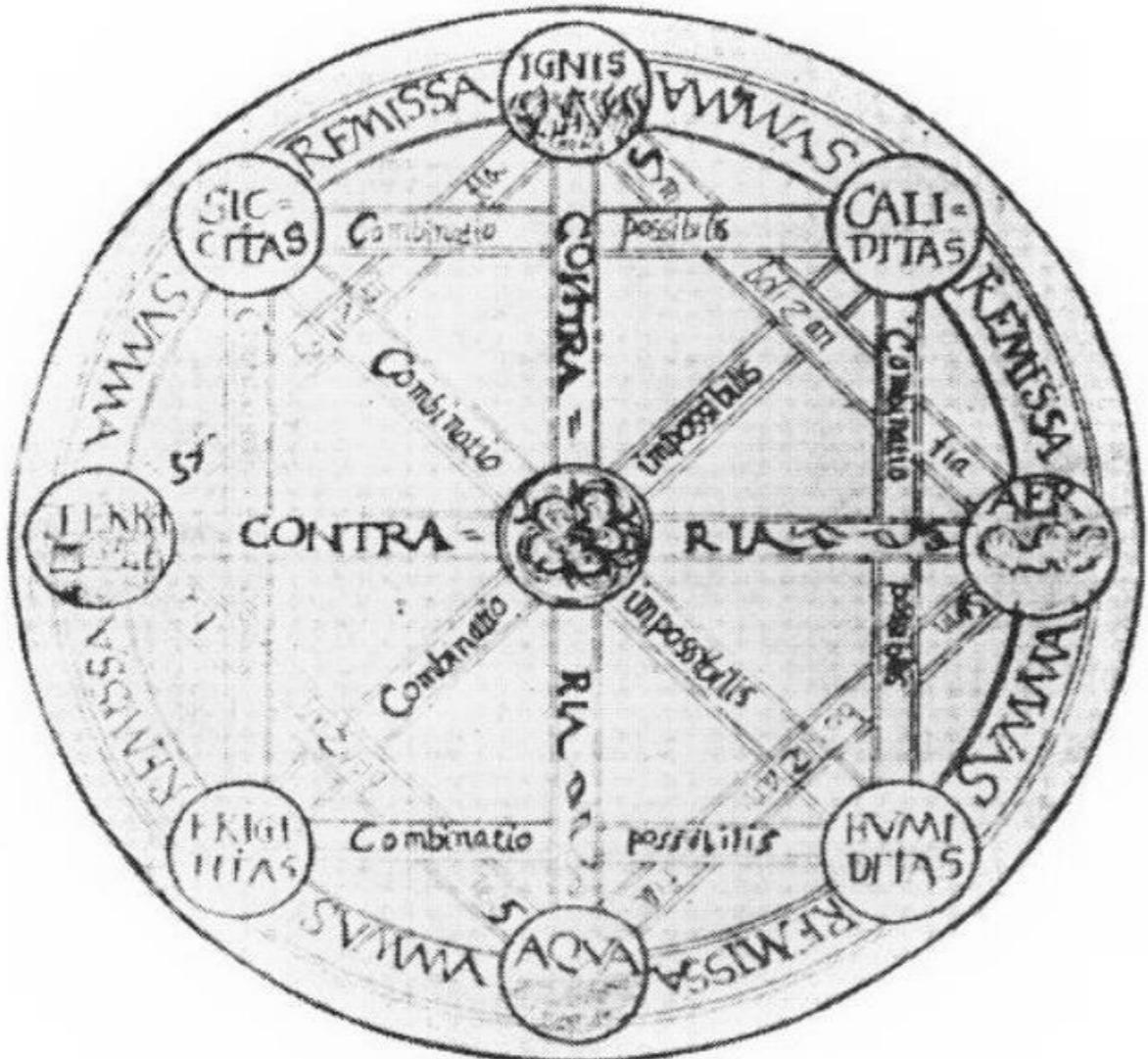
# Outline

- Fundamental Motivations for Long Lived Particles
- LLP Searches at the Large Hadron Collider
- MATHUSLA
  - Basic Concept
  - LLP Sensitivity
  - Cosmic Ray Telescope
- Detector Design
  - Test Stand
- Next Steps & Opportunities

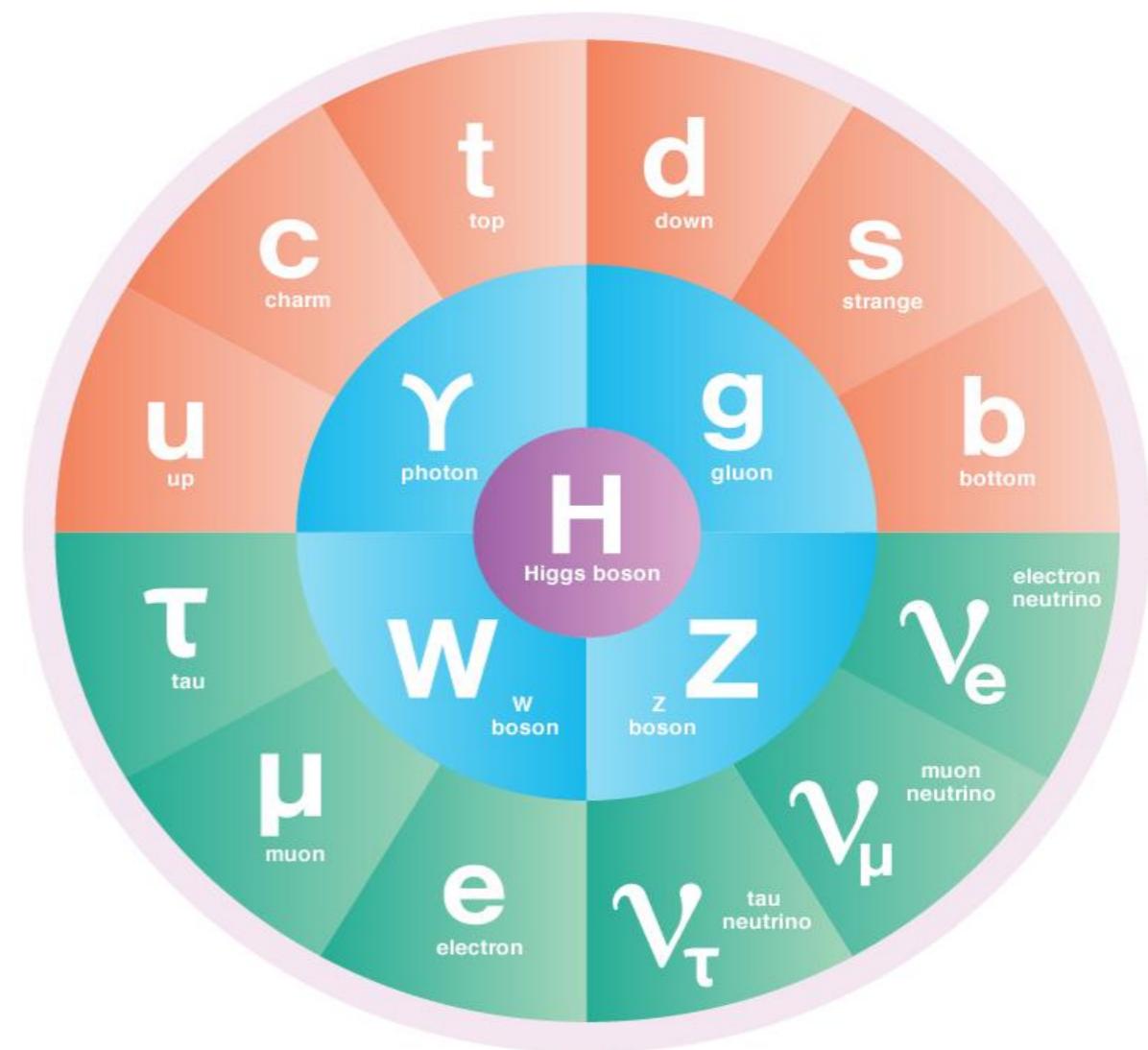
# Fundamental Motivations for LLPs

# The Standard Model ...

2500 years ago



Today

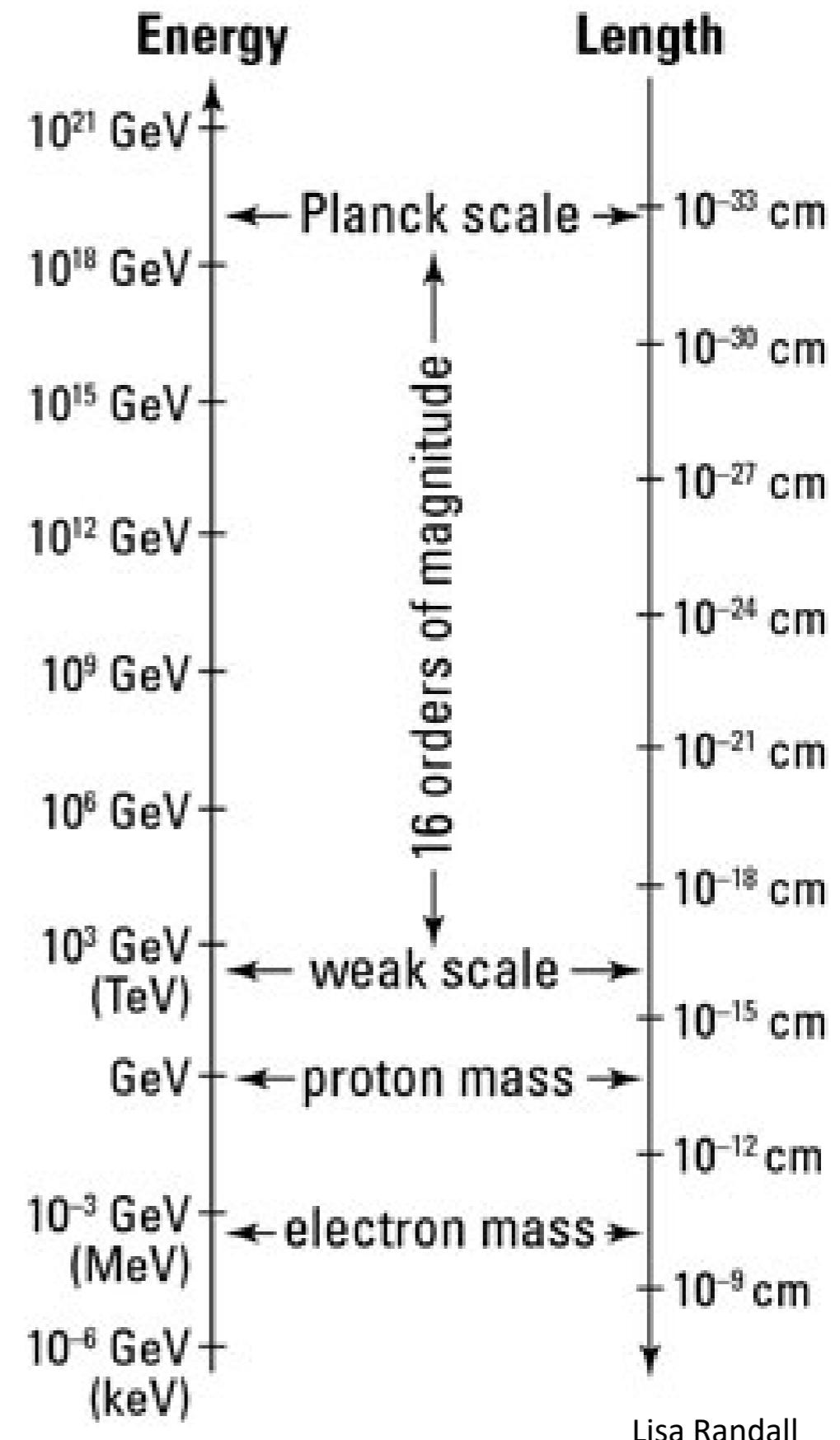
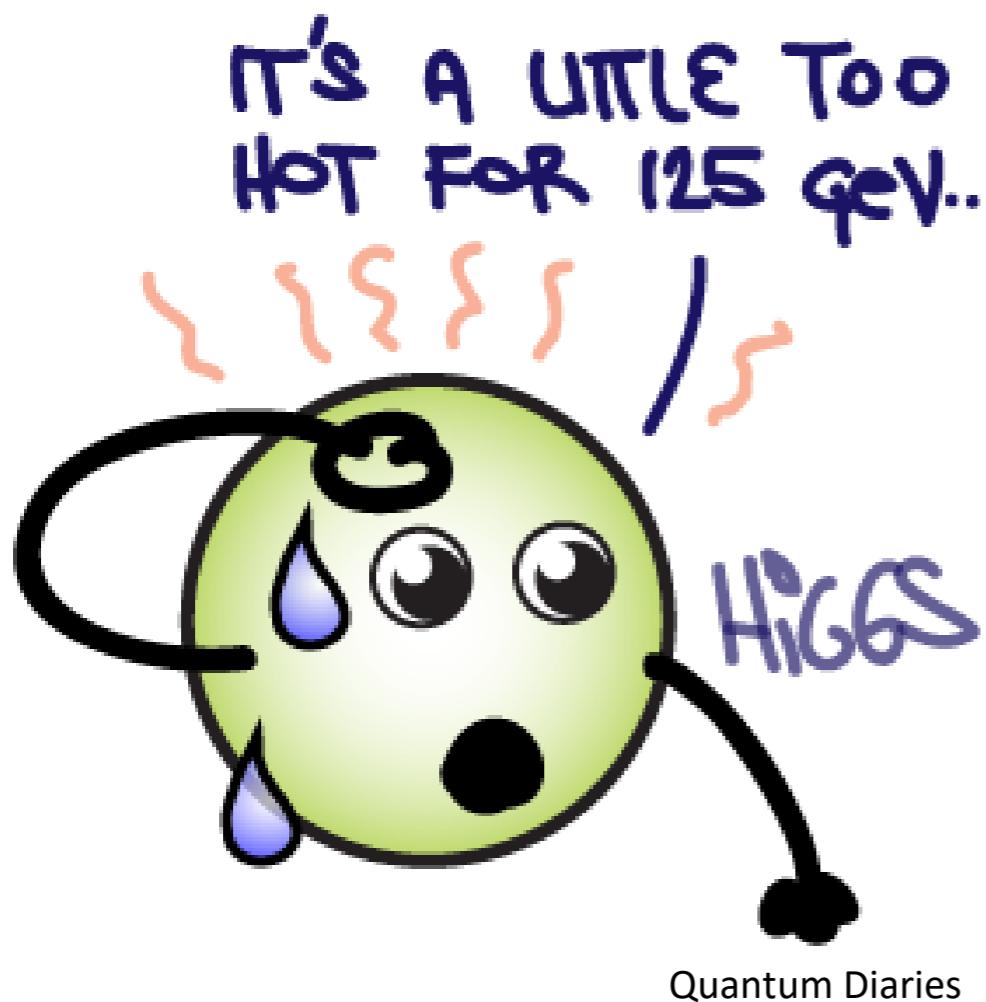


● QUARKS ● LEPTONS ● BOSONS ● HIGGS BOSON

Symmetry Magazine: Sandbox Studio, Chicago

# ... is incomplete!

**Hierarchy Problem & Naturalness:**  
why is the Higgs Boson mass so  
low, and what stabilizes its mass?



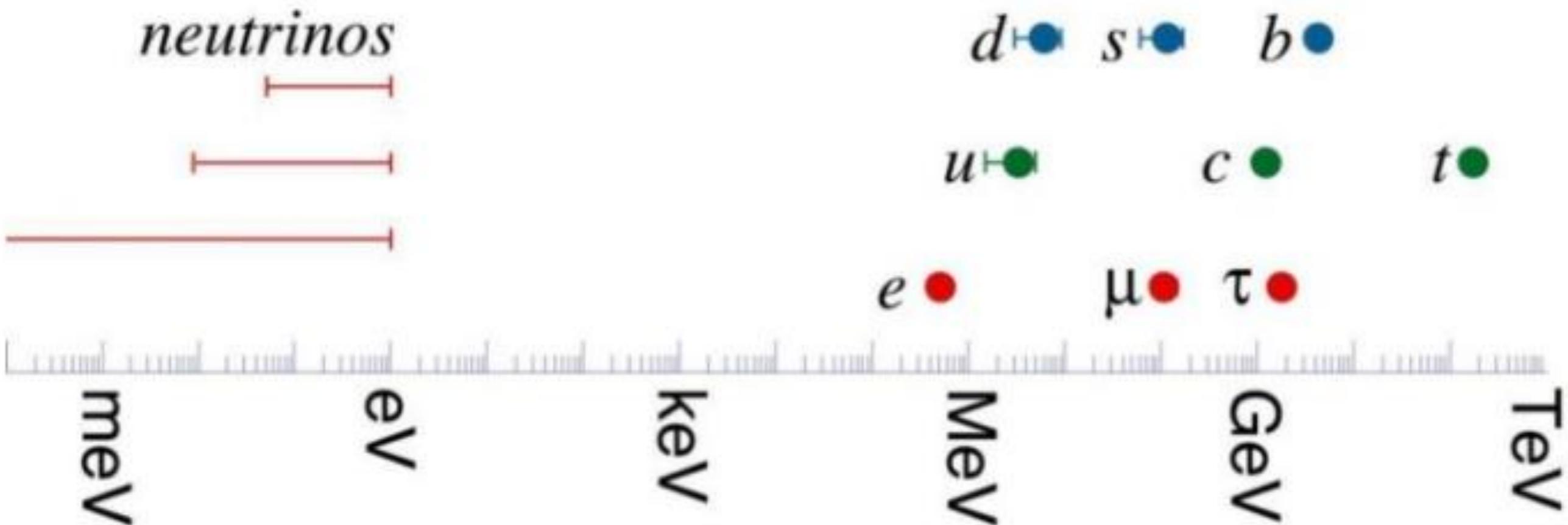
Lisa Randall

# ... is incomplete!

**Neutrinos:** where do their tiny masses come from?

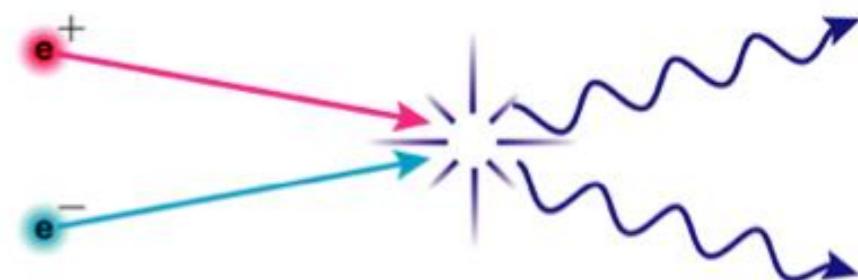


Symmetry Magazine: Sandbox Studio, Chicago with Corinne Mucha



Hitoshi Murayama

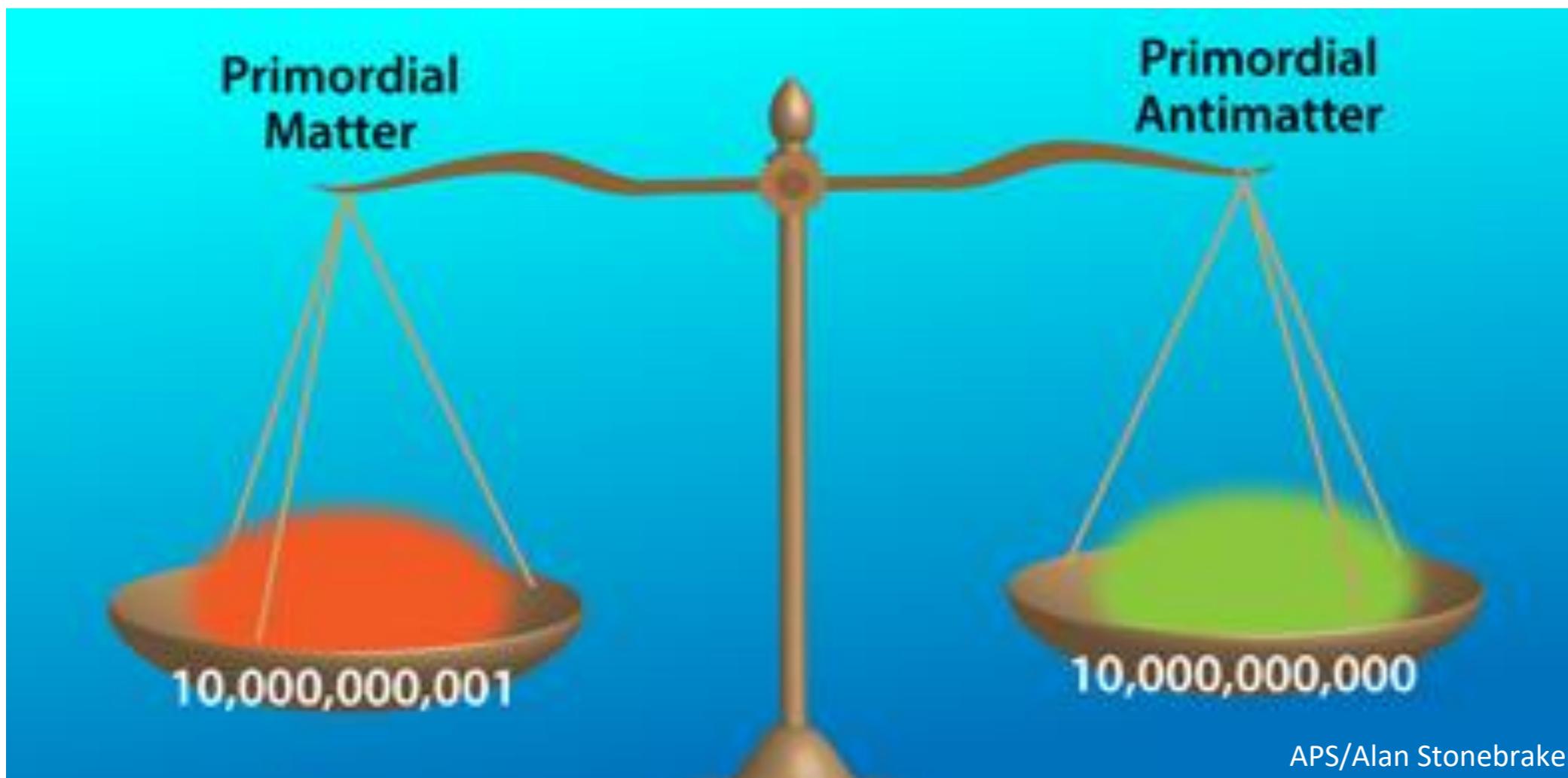
# ... is incomplete!



**Matter-antimatter asymmetry:** How did the universe become matter-dominated?



Addison Wesley



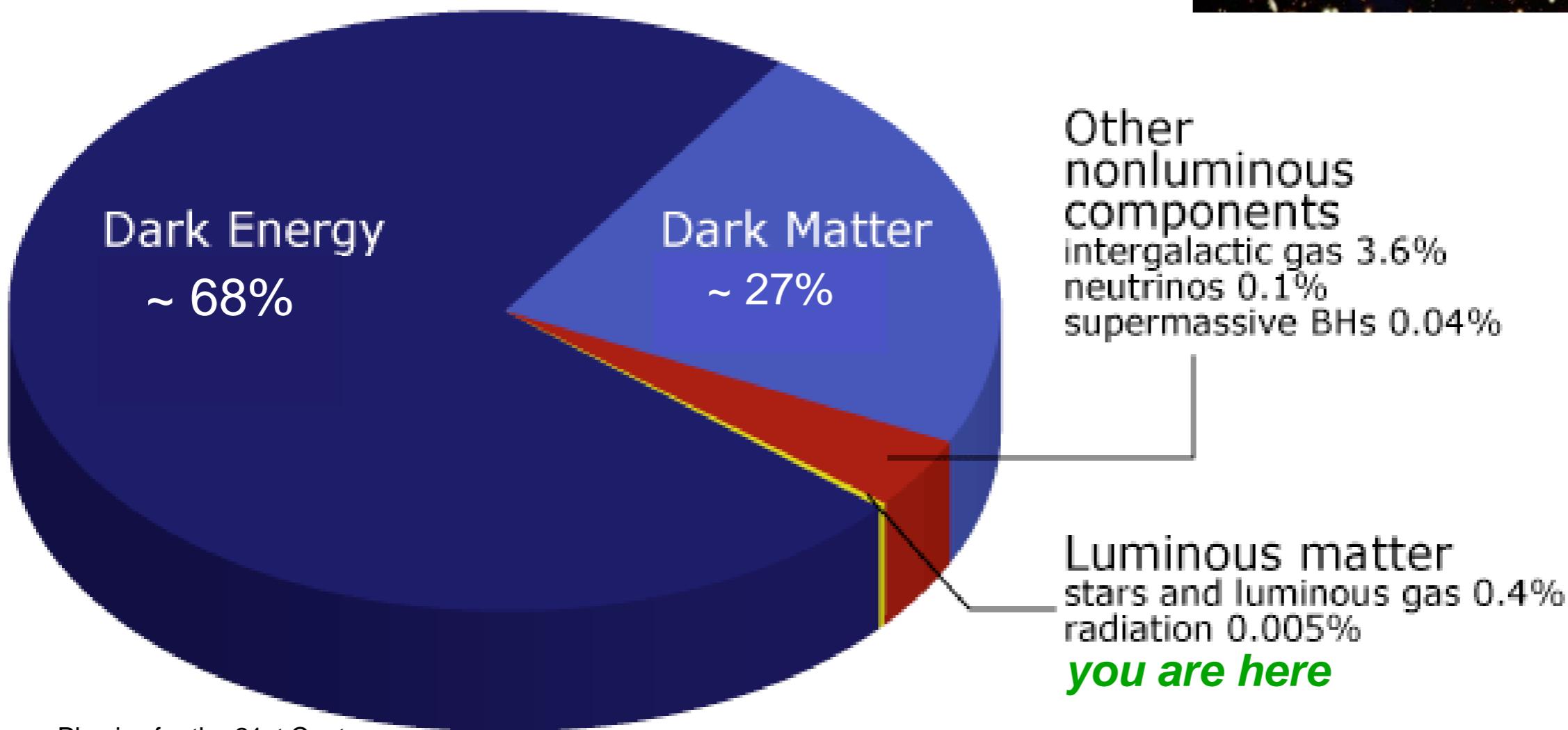
APS/Alan Stonebraker

# ... is incomplete!

**Dark Matter:** What is it, and what sets the relative amount of it in the universe?



Chandra



# Theoretical motivation for LLPs

**Seeking to go Beyond the Standard Model (BSM) motivates the possibility of so-far-undiscovered LLPs**

**"Top-down"**: Various BSM theories (e.g. supersymmetry) constructed to explain the Fundamental Mysteries naturally include new LLPs

**"Bottom-up"**: LLPs occur in the SM (e.g. muons), and can occur via similar mechanisms when adding new particles to the model

# Top-down LLP motivation

Motivation	Top-down Theory	IR LLP Scenario
Naturalness	RPV SUSY GMSB mini-split SUSY Stealth SUSY Axinos Sgoldstinos  Neutral Naturalness <small>UV theory</small> Composite Higgs Relaxion	<b>BSM=/<math>\rightarrow</math>LLP</b> <small>(direct production of BSM state at LHC that is or decays to LLP)</small>  <b>Hidden Valley</b> <small>overlapping system</small> ALP <small>EFT</small>  SM+S
Dark Matter	Asymmetric DM Freeze-In DM SIMP/ELDER Co-Decay Co-Accretion Dynamical DM	SM+V (+S)  <b>exotic Z decays</b>
Baryogenesis	WIMP Baryogenesis Exotic Baryon Oscillations Leptogenesis	 <b>exotic Higgs decays</b>
Neutrino Masses	Minimal RH Neutrino with $U(1)_{\text{B-L}}$ $Z'$ with $SU(2)_R$ $W_R$ long-lived scalars with Higgs portal from ERS	HNL  <b>exotic Hadron decays</b>
	Discrete Symmetries	

# Bottom-up LLP motivation

**Example: “Hidden Sector” of BSM particles & forces not directly accessible to us**

The Standard Model is only ~5% of the universe.

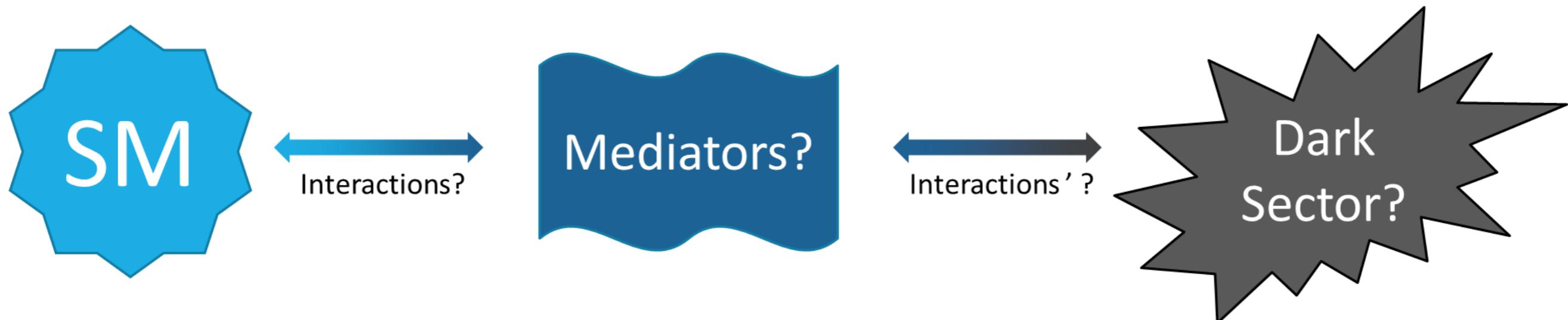
It includes 3 forces.

Why should the ~25% that is Dark Matter be any simpler?

Dark Forces?

How would DM interact with the SM?

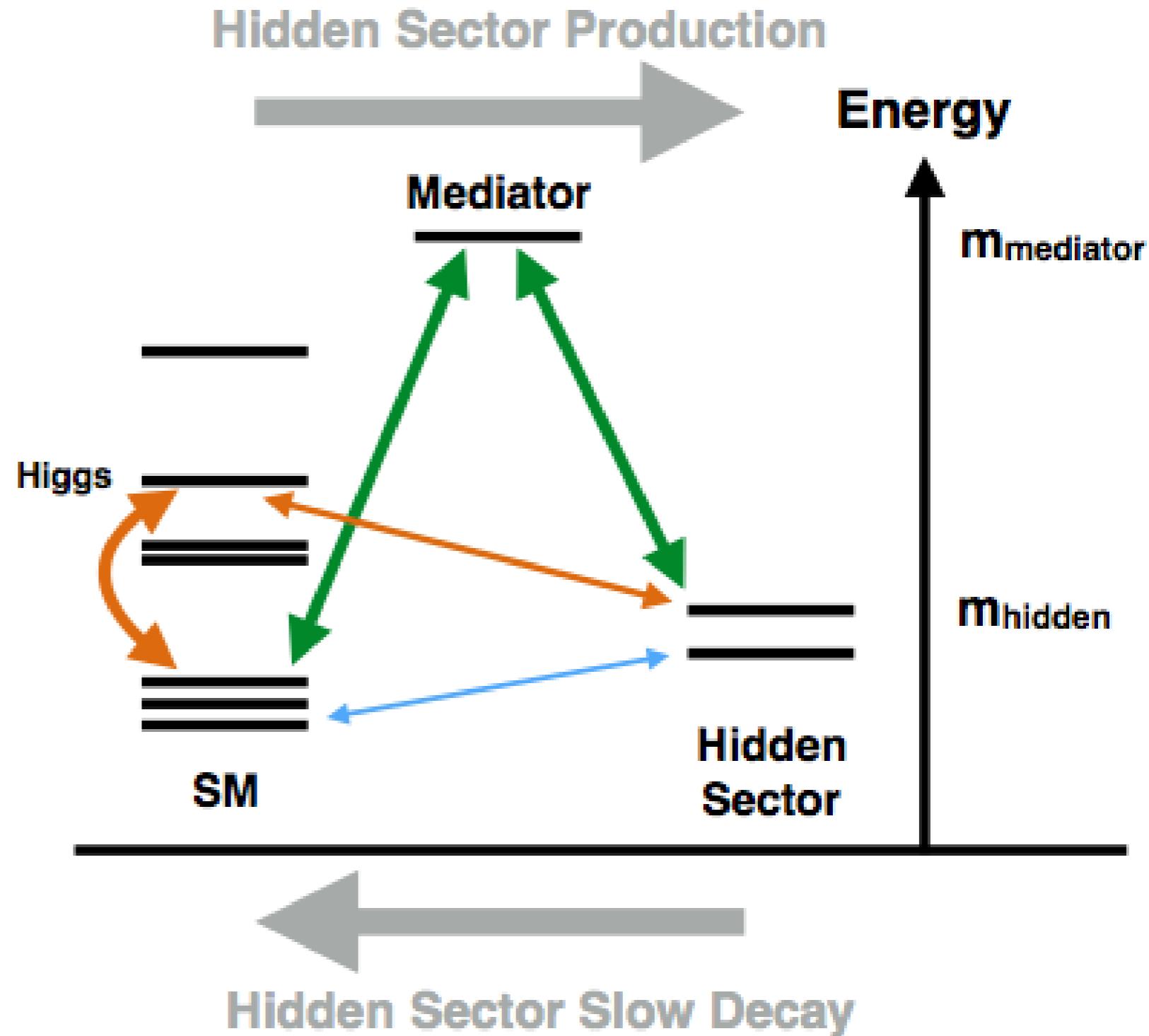
Mediator particles?



# Bottom-up LLP motivation

Dark Matter lives in the Hidden Sector

Metastable Hidden Sector states can only decay to SM via mediators (e.g. extra Higgs Bosons) with small couplings, so they are long-lived



# LLP Searches at the HL-LHC

# Large Hadron Collider at CERN

Probing the fundamental mysteries with 4 main detectors since 2010

Proton-proton collisions with 7, 8, then 13 TeV centre-of-mass energy

Announced the Higgs Boson discovery in July 2012

Currently in a shutdown period following “Run 2”



# High-Luminosity Large Hadron Collider

Upgrading the world's most powerful supercollider:  
10x the original rate of particle collisions

Planned for mid-2020s, following “Run 3”



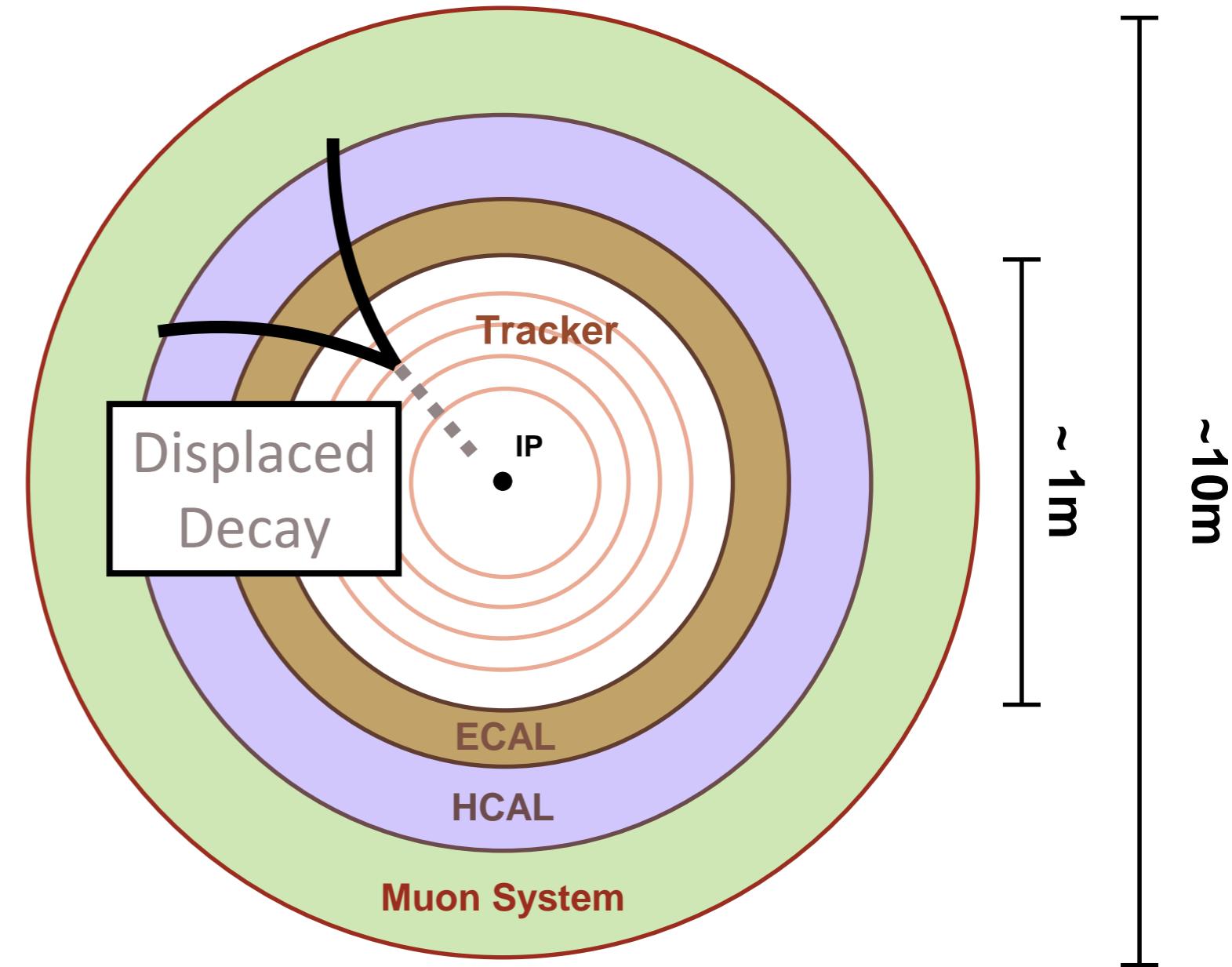
# LLPs at the LHC

**The problem of long lifetimes: LHC could be making LLPs that are invisible to its main detectors!**

- If the LLP has  $c \cdot \text{lifetime} \gg \text{detector size}$ , most escape the detector entirely
- Even LLPs that decay in the detector, but a significant distance away from the Interaction Point, are difficult to spot
- If the LLPs decay in the detector with only a tiny rate, they get swamped by backgrounds

# LLPs at the LHC

Neutral LLPs that decay in ATLAS or CMS are *spectacular* signatures that are missed by most standard searches, since **trigger & event reconstruction** are designed for decays near the Interaction Point



Comprehensive LLP search program at LHC has been ramping up over the past few years

**Significant progress has been made, but LLP searches are very labour-intensive**

# Why search for LLPs at HL-LHC at all?

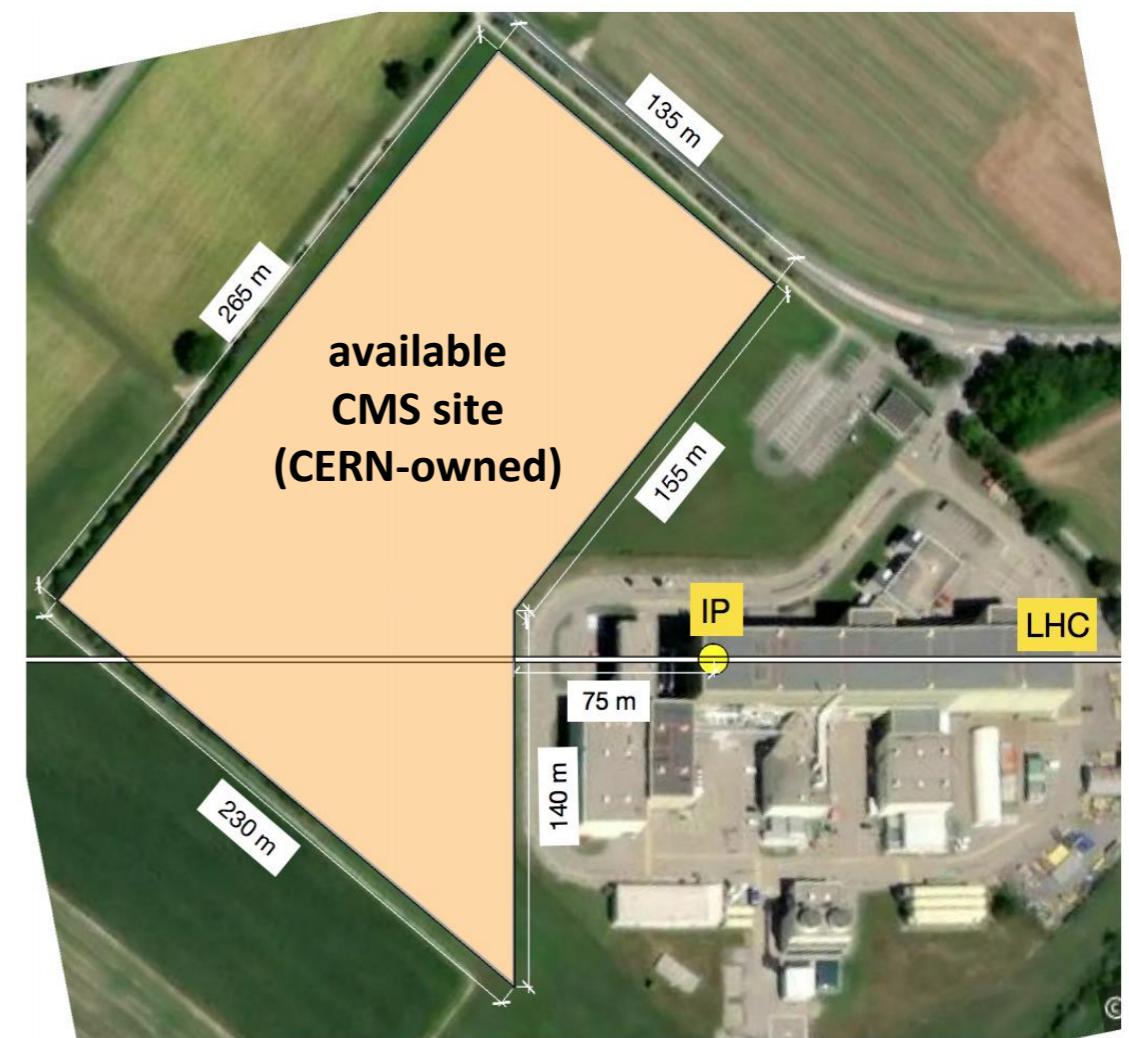
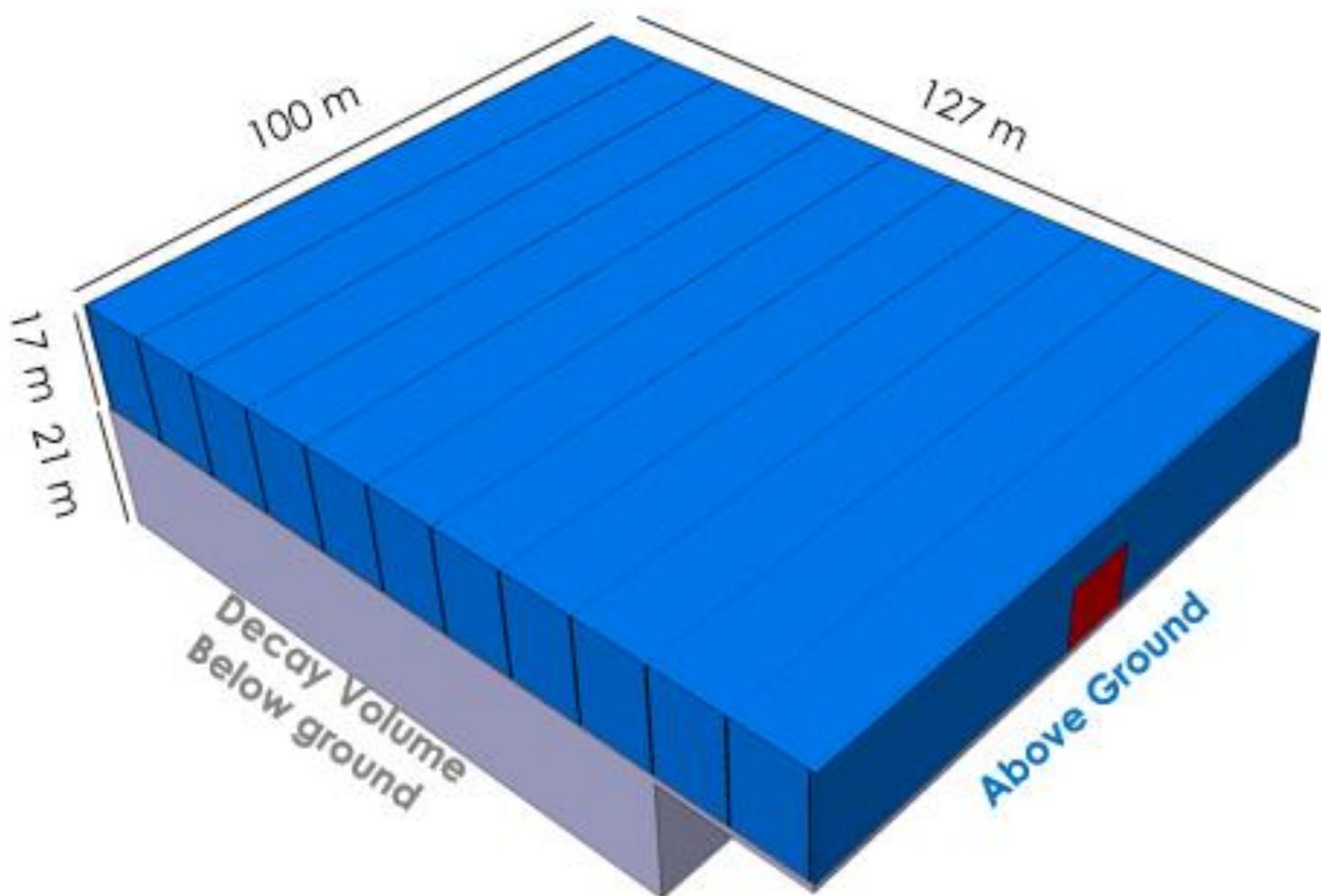
Searches at other facilities have also been proposed, but HL-LHC has many advantages:

- High center of mass energy gives access to heavy states that may be coupled to LLPs (e.g. Higgs) not easily produced at lower energies
- Large data sample
- Enormous existing investment in machine and experiments

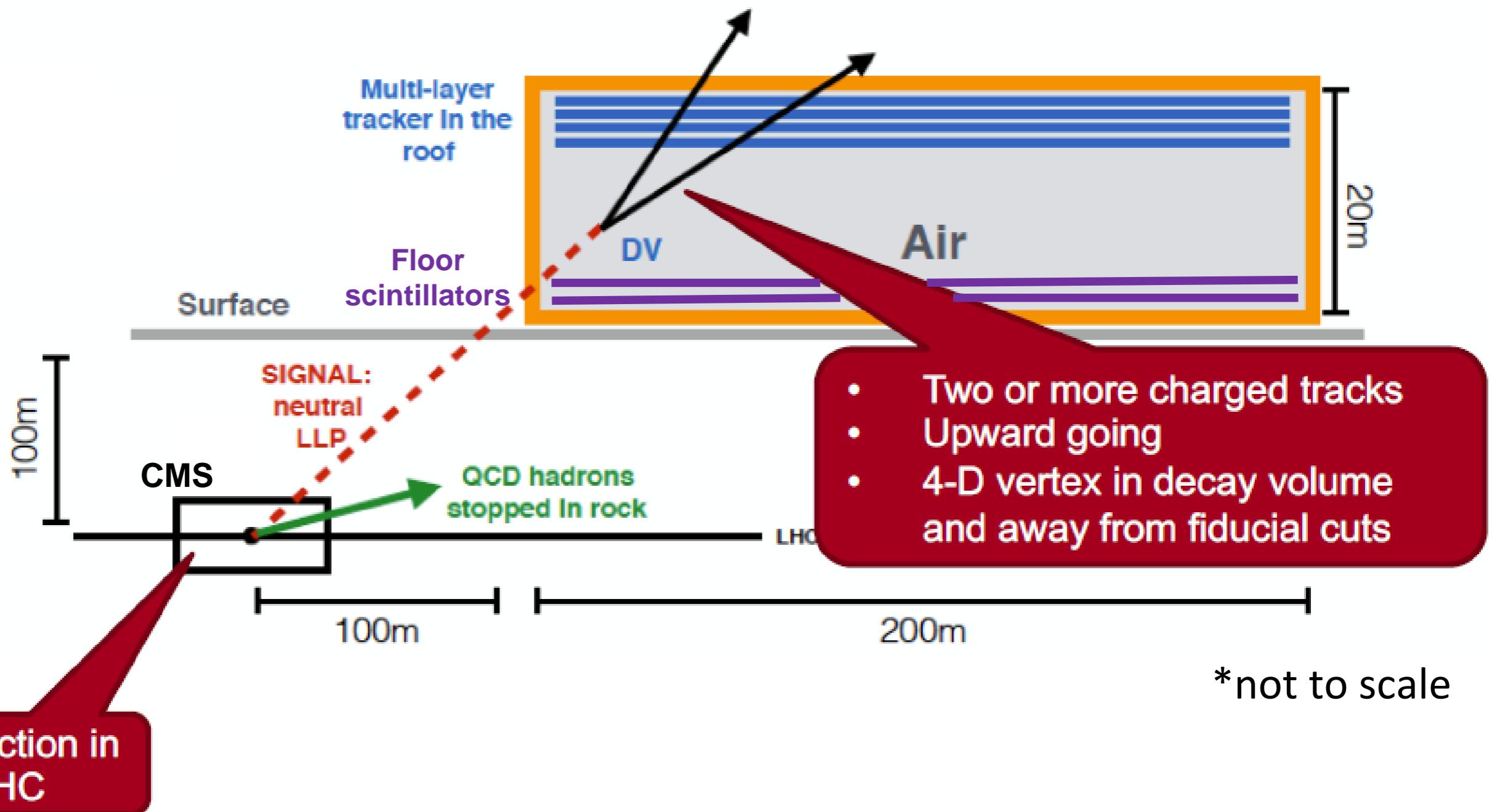
LLP searches have been identified by the [HL-]LHC community as a growing priority

The Concept of  
**MATRUSÁ**

# An external LLP detector for HL-LHC

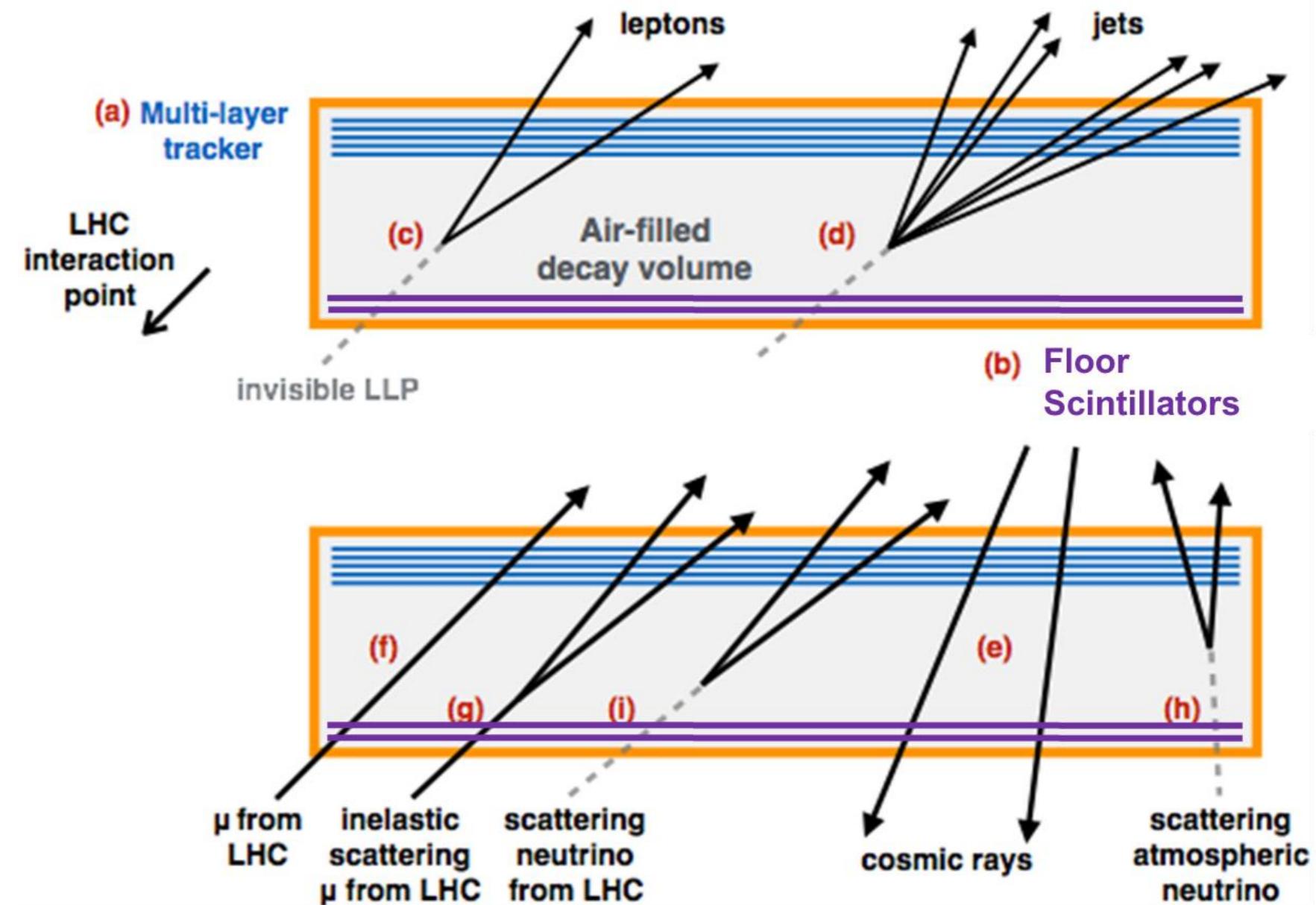


# An external LLP detector for HL-LHC



# LLP identification with MATHUSLA

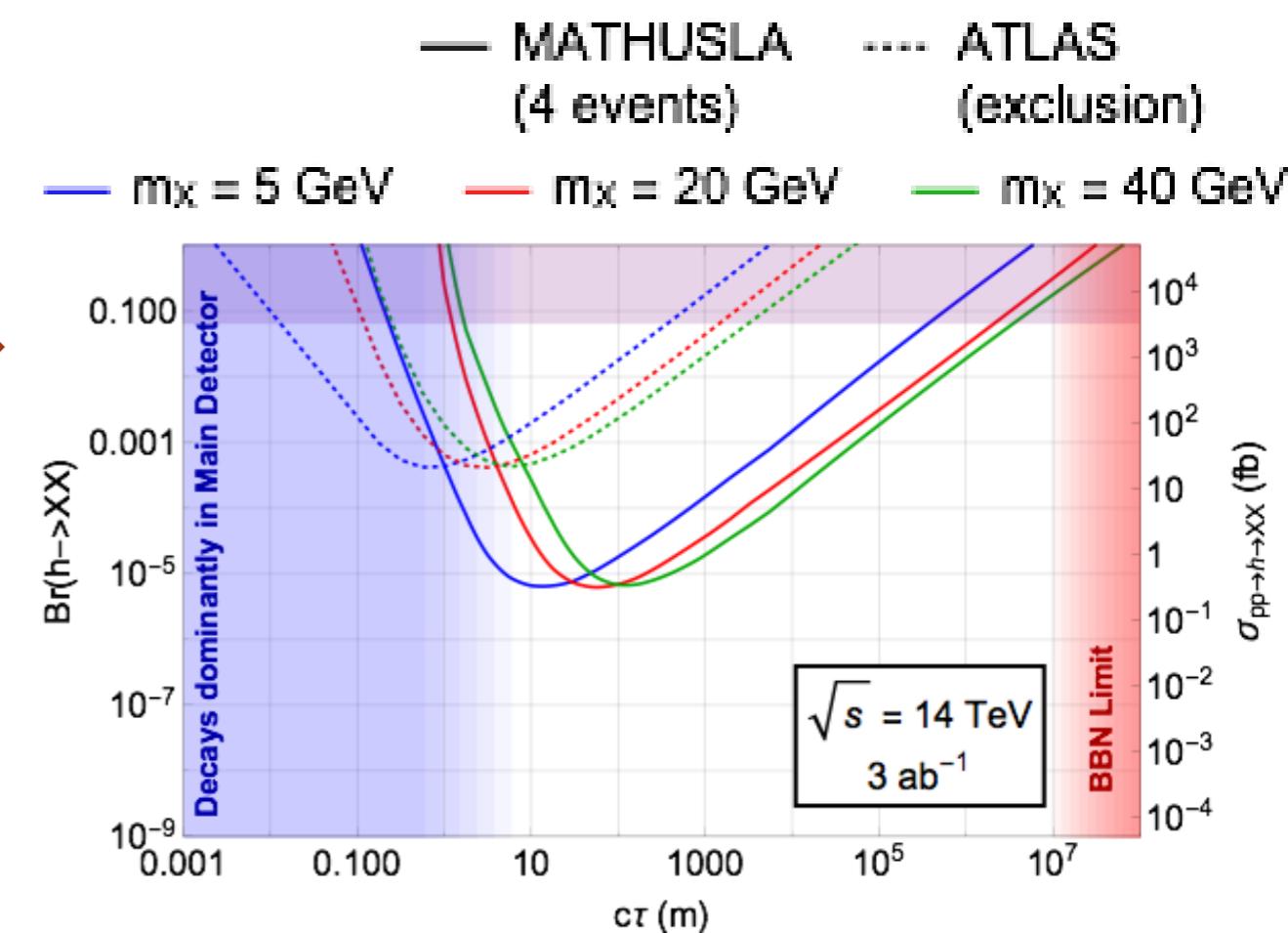
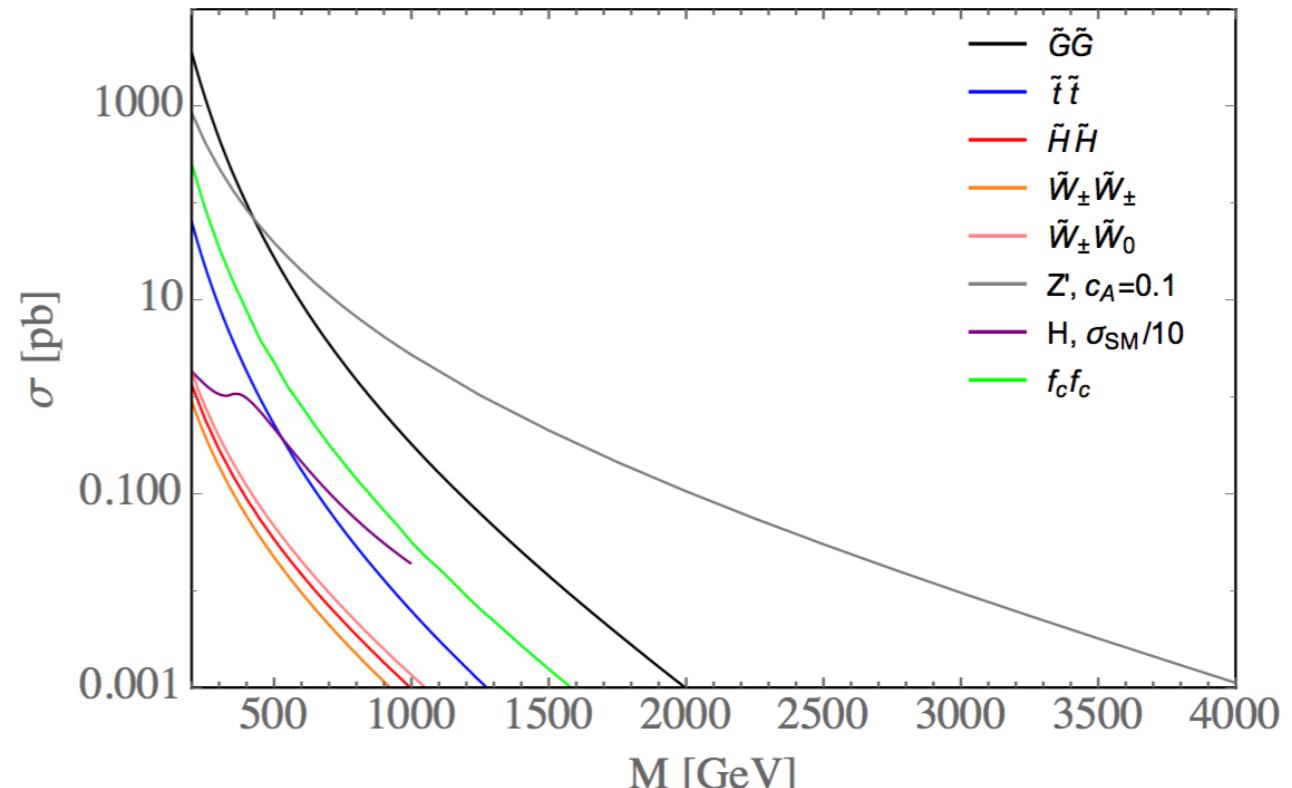
- Strict geometry & timing cuts will achieve near-zero backgrounds for neutral LLP decays!



- MATHUSLA standalone: Determination of LLP decay mode and boost
- MATHUSLA + CMS: Also LLP production mode, mass range and spin

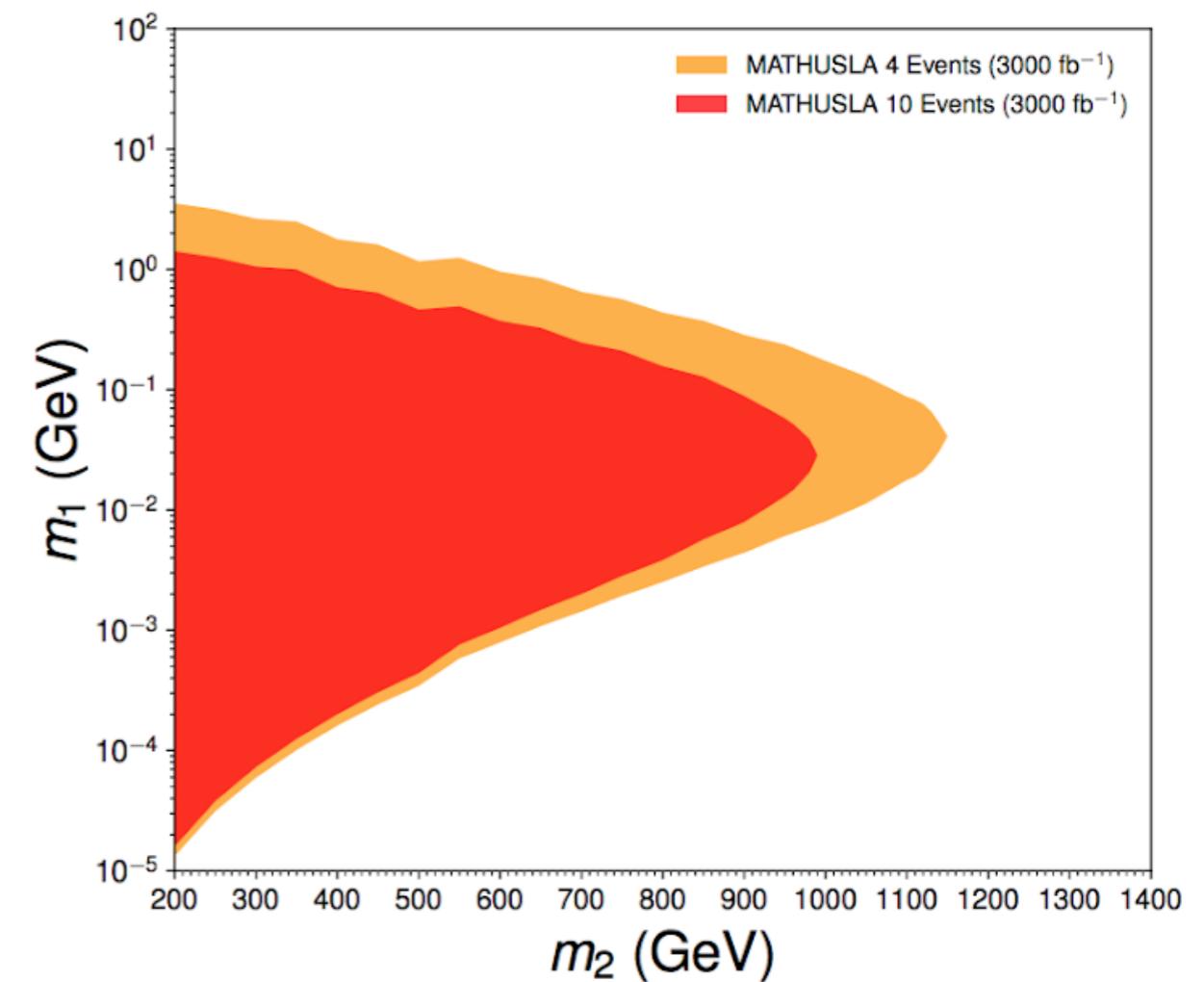
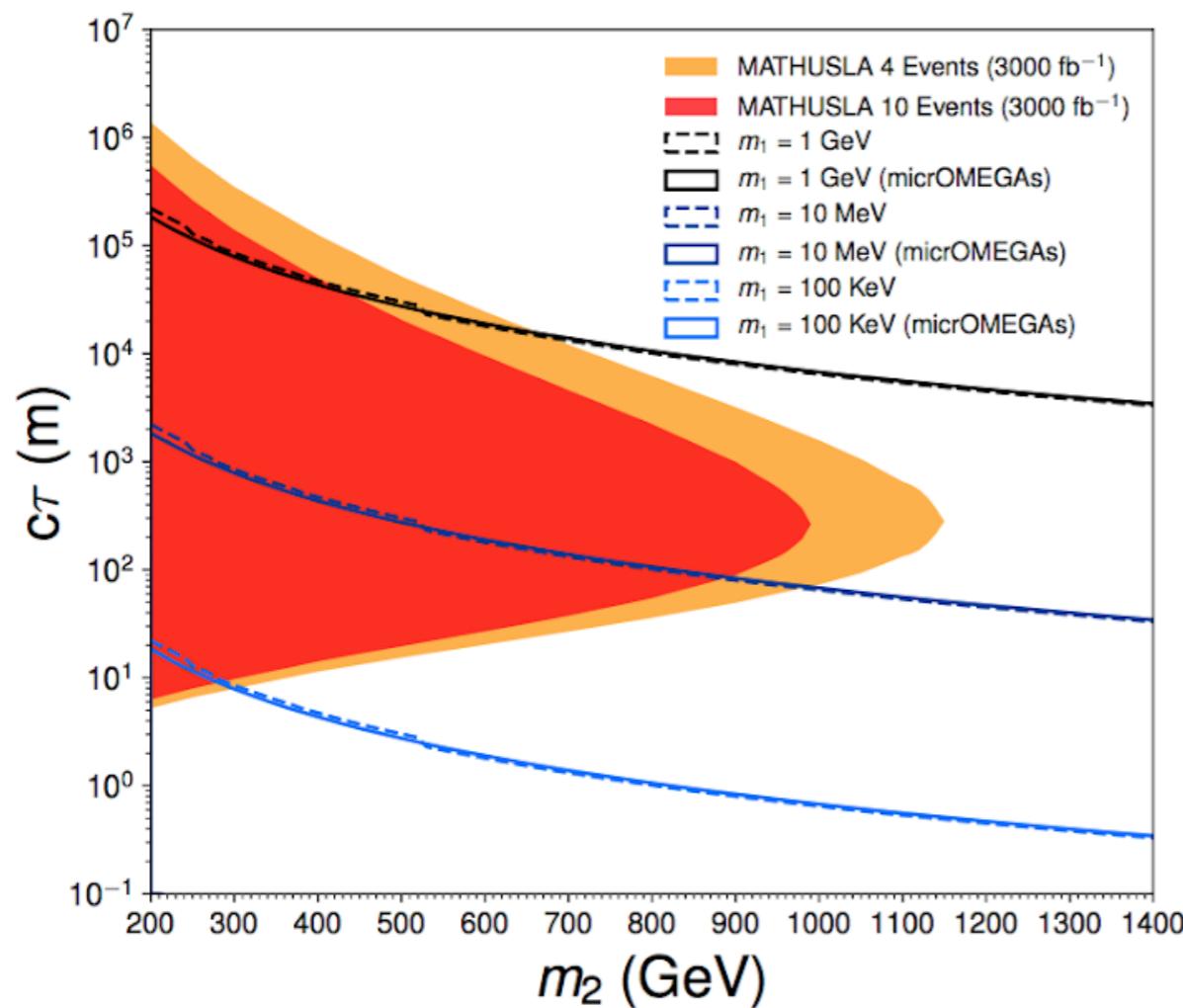
# LLP sensitivity

- Sensitivity to any LLP production process with  $\sigma > \text{fb}$  
- Can probe scales from  $>\text{TeV}$  all the way down to  $\text{GeV}$
- Scenarios include:
  - Exotic Higgs decay (up to 1000x better sensitivity than LHC main detectors) 
  - Heavy sterile neutrinos
  - Dark photons
  - Axions



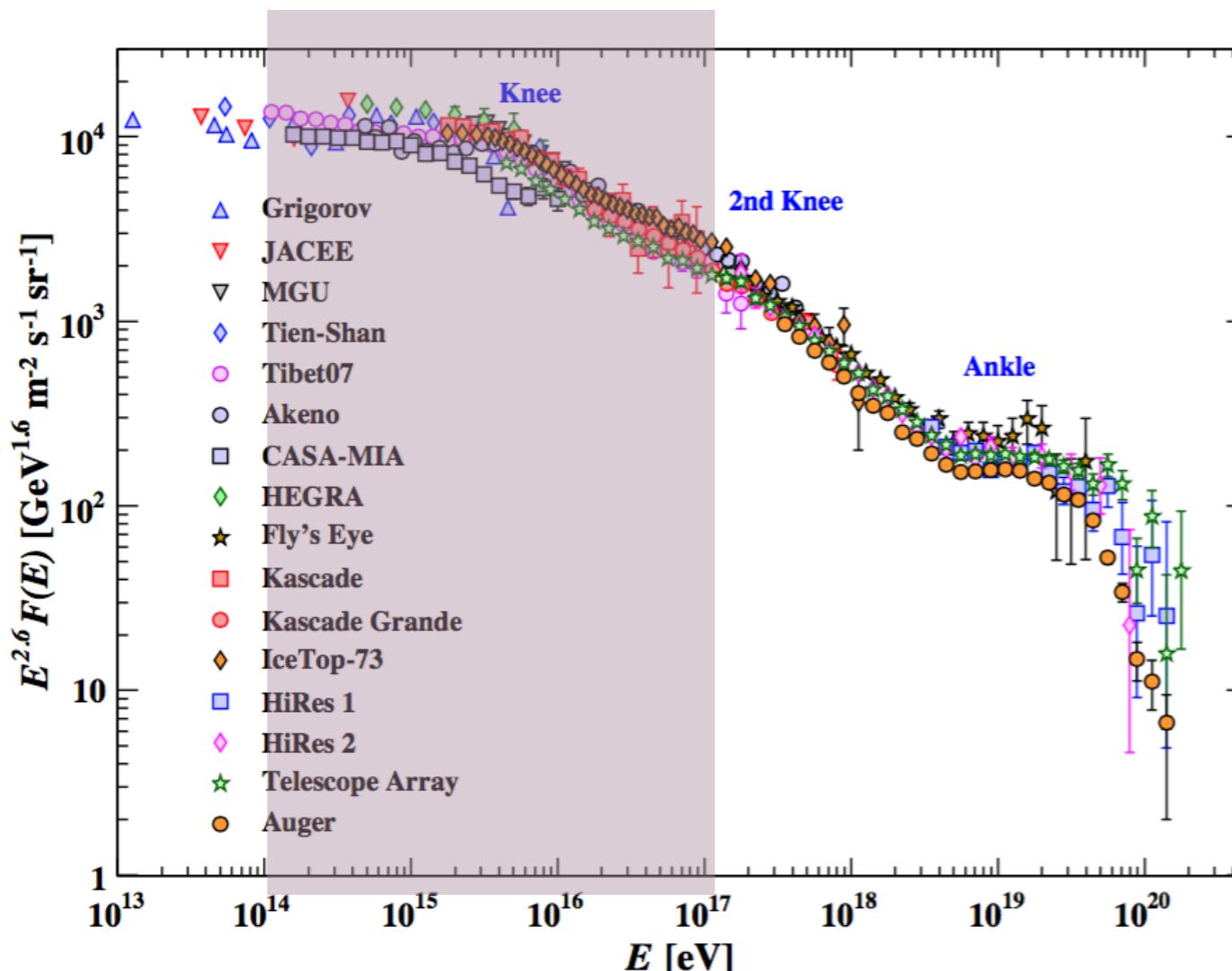
# LLP sensitivity

- Scenarios where  $\text{LLP} \rightarrow \text{DM} + \text{SM}$  decay is the only way to see the DM
- e.g. model with BSM  $(\chi_1, \chi_2)$  where DM candidate  $\chi_1$  is very difficult to see due to low mass and possibly very small coupling. We look for LLP  $\chi_2$  which eventually decays to  $\chi_1$



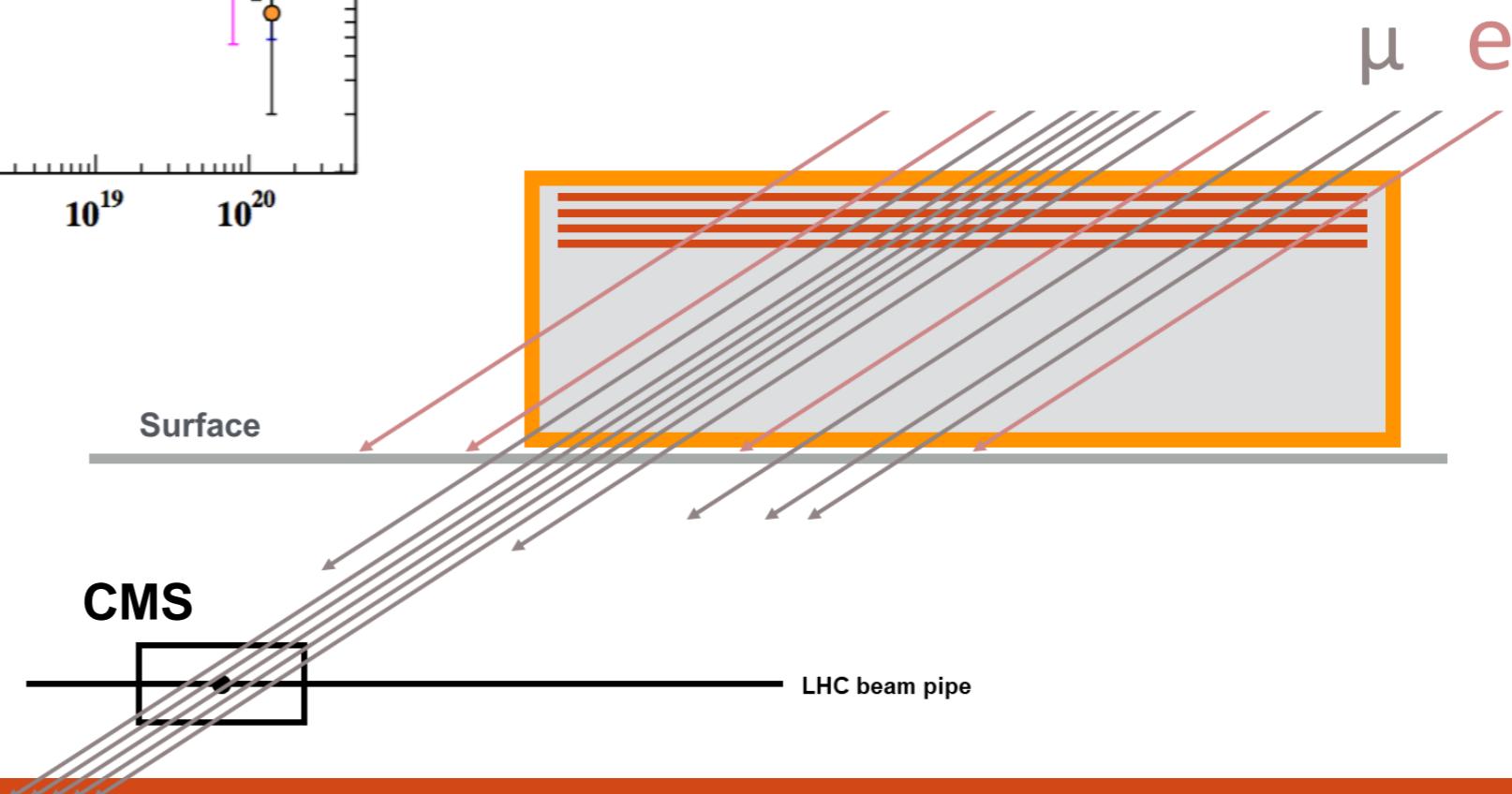
(lines yield observed DM relic density)

# MATHUSLA as a Cosmic Ray Telescope



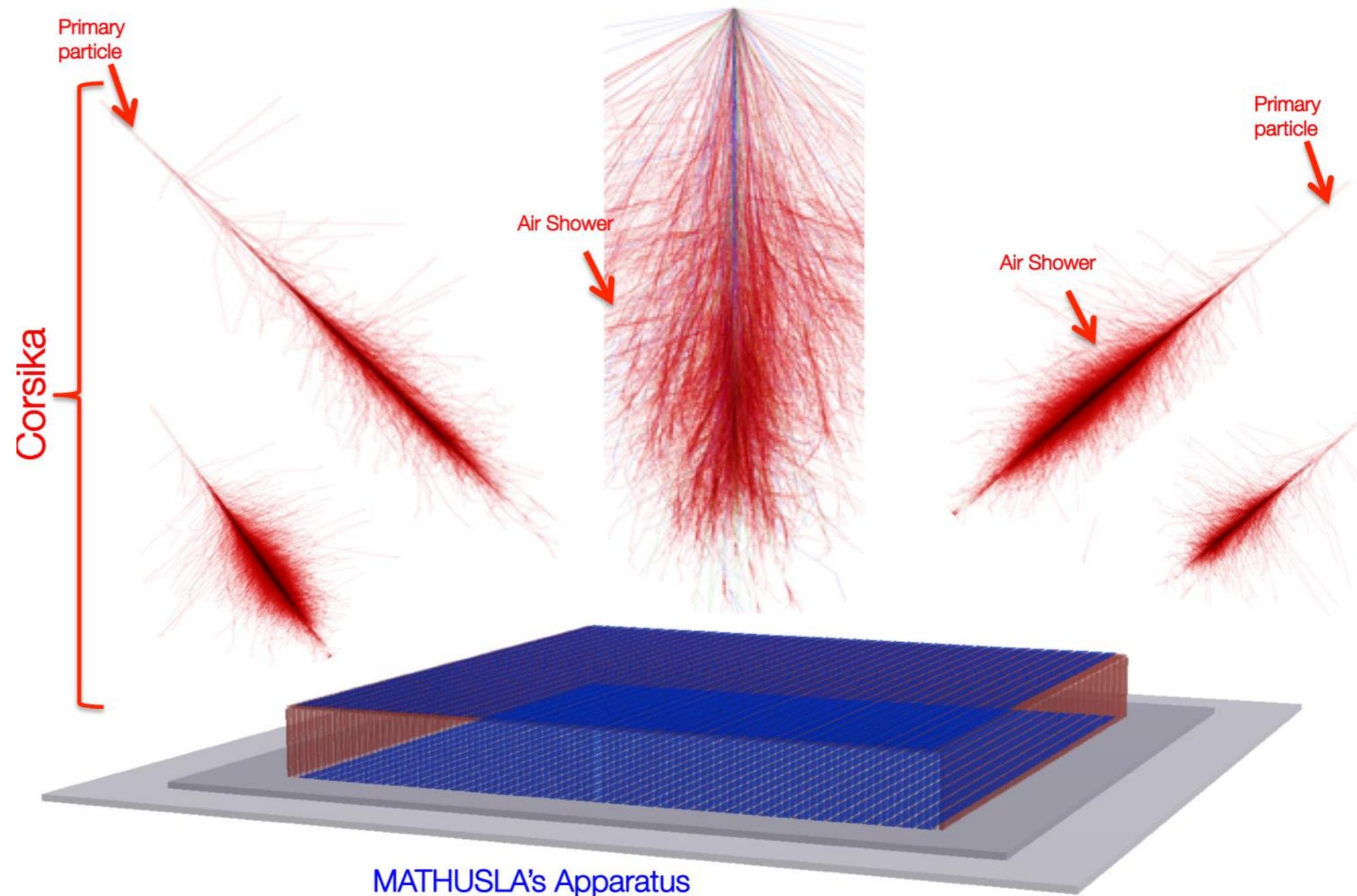
Probes highly interesting  
region of CR spectrum

“Guaranteed Physics Return”  
Unique abilities in CR  
experimental ecosystem  
(precise resolution,  
directionality, large-area  
coverage)



# MATHUSLA as a Cosmic Ray Telescope

Standalone:  
reconstruction of  
shower core,  
direction, total #  
charged particles,  
slope of radial  
particle density  
distribution



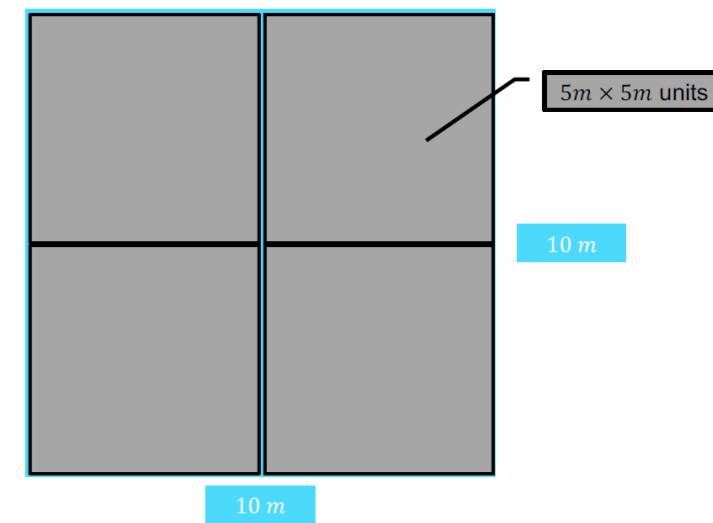
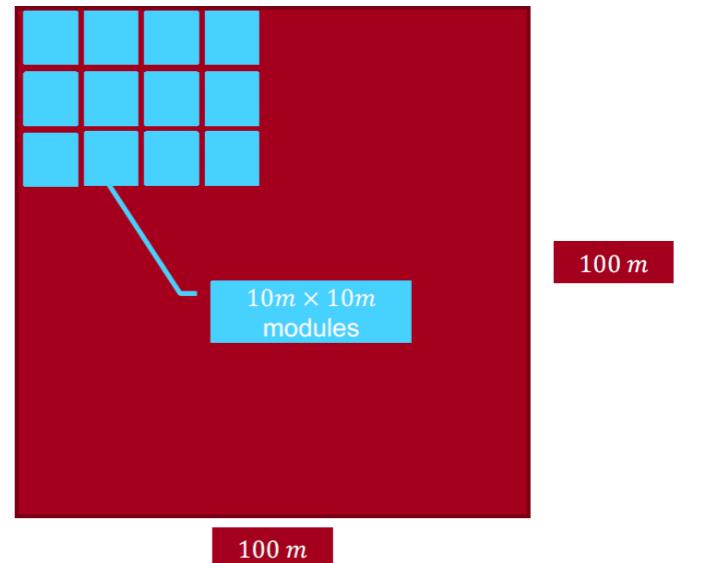
+ CMS: analysis of muon bundles traversing both detectors, probing heavy primary CR spectra and astrophysical acceleration mechanisms

# MATRÍCULA

## Detector Design

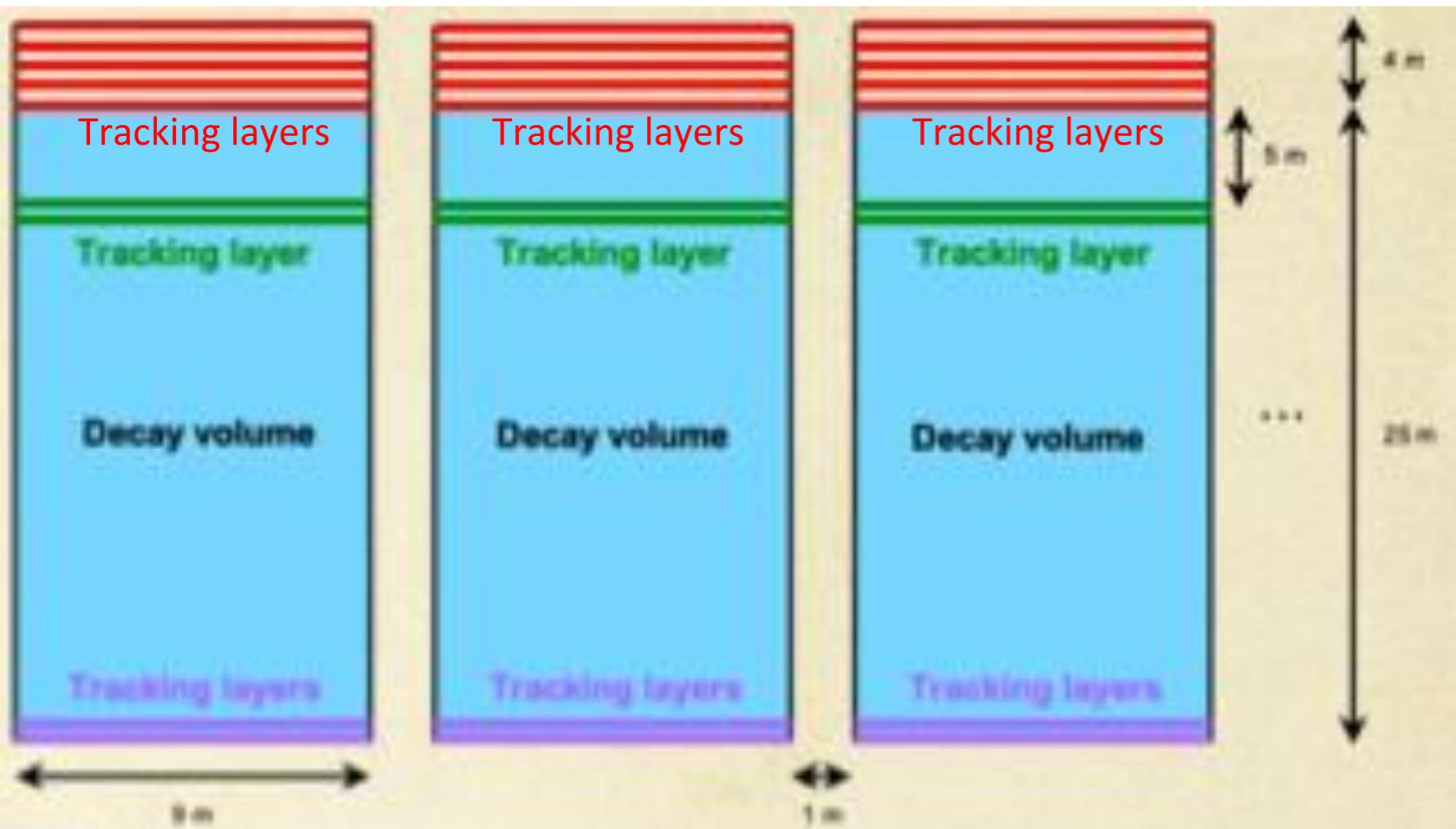
# Detector design

Modular design facilitates staged construction  
and commissioning



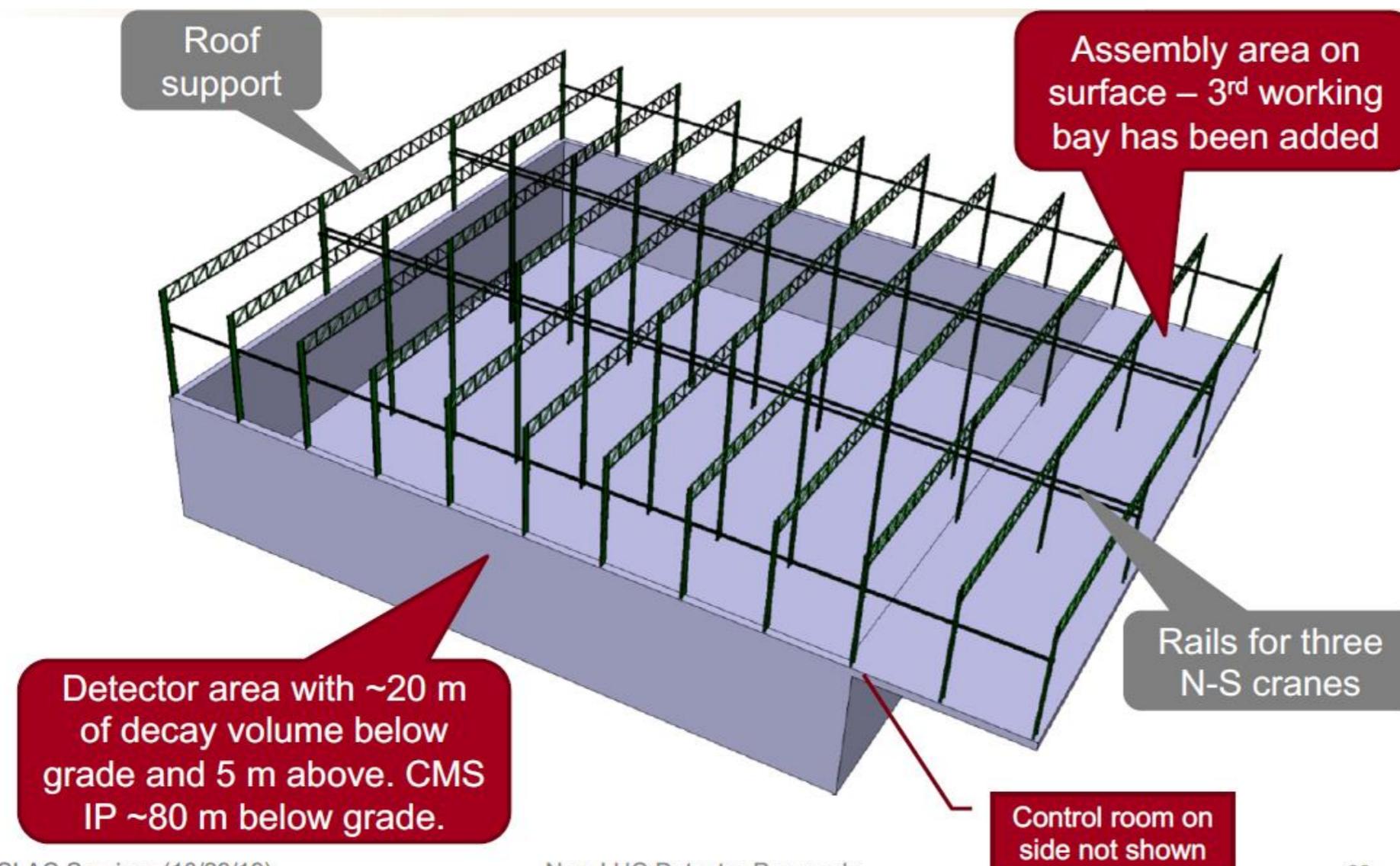
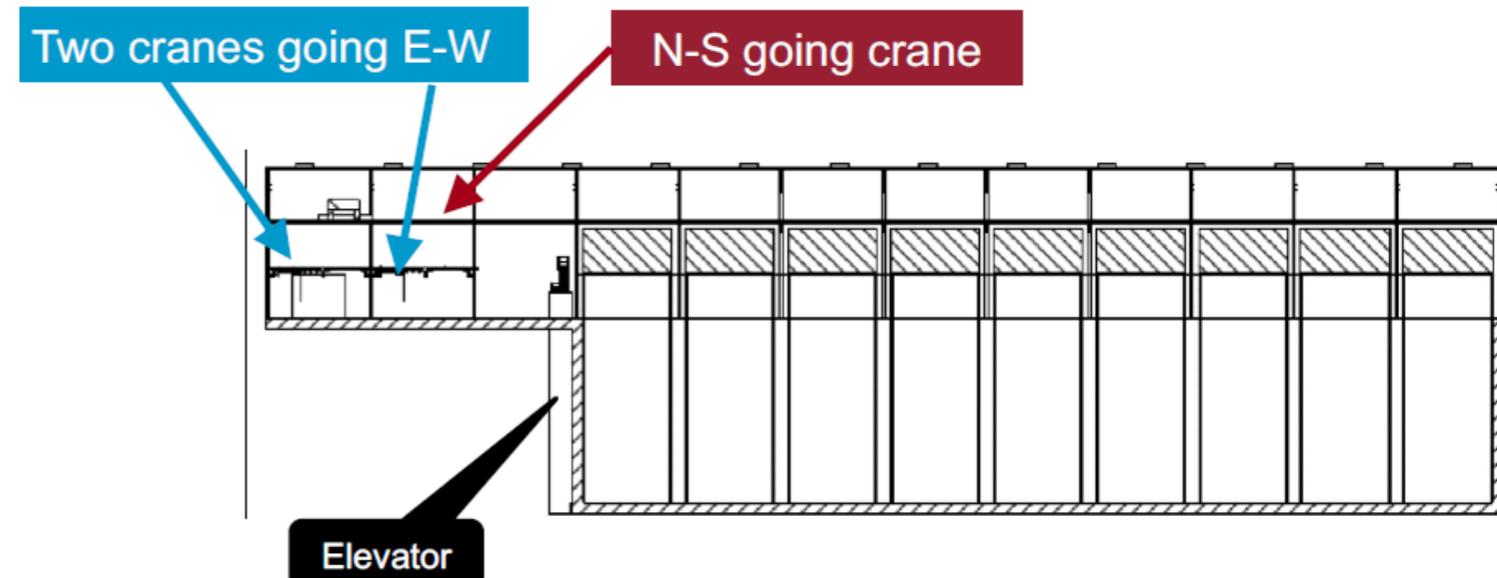
**100 Modules in  
100m × 100m  
Footprint**

**4 Detector Units  
per Module Plane**

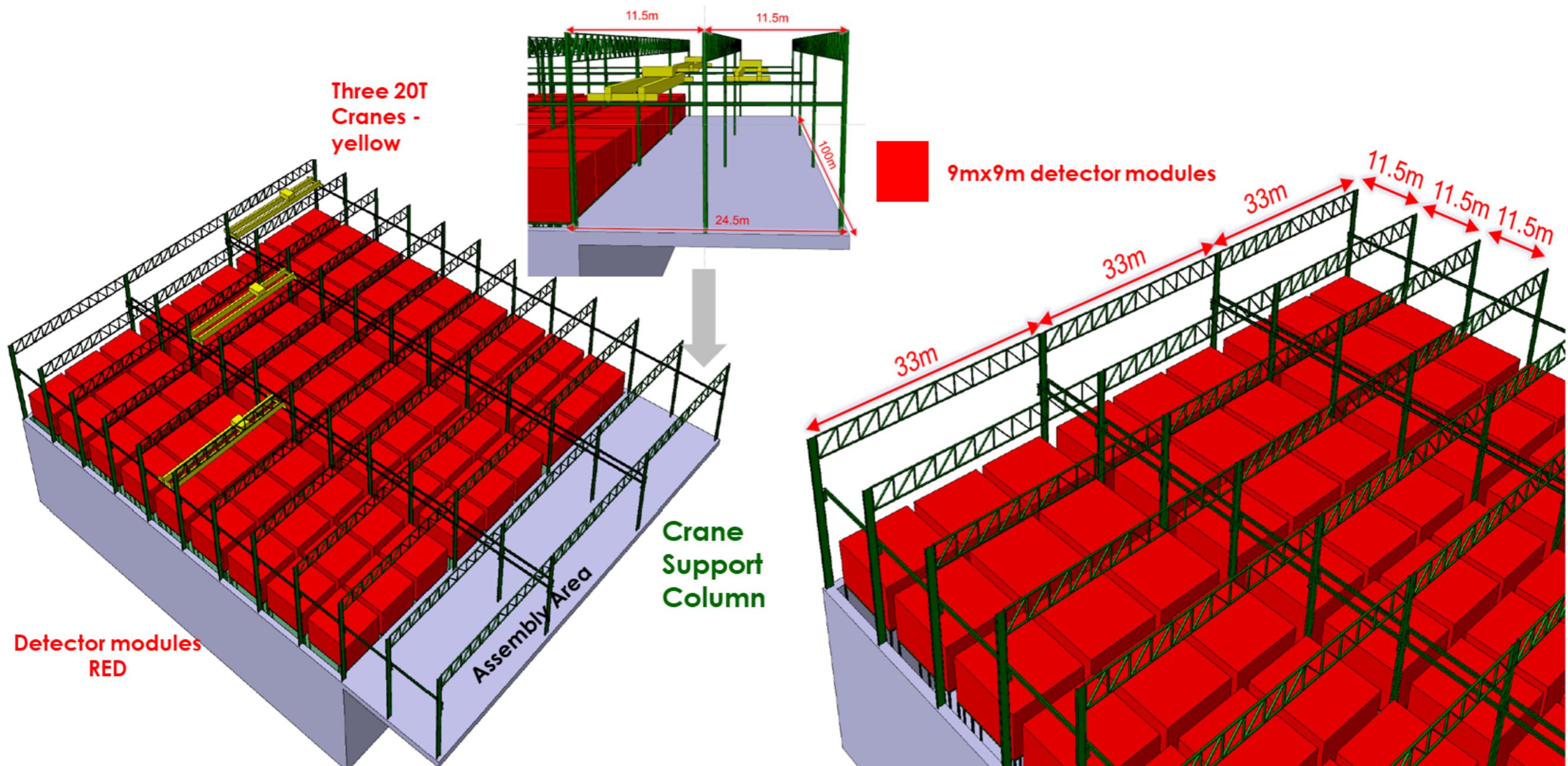


Each module has 5 tracking layers on top +  
2 floor layers + 2 mid-level layers

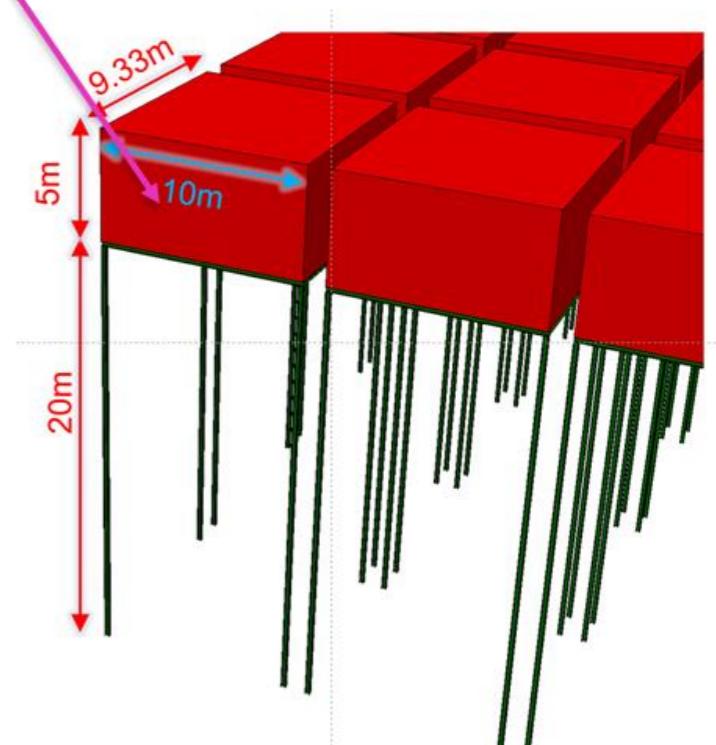
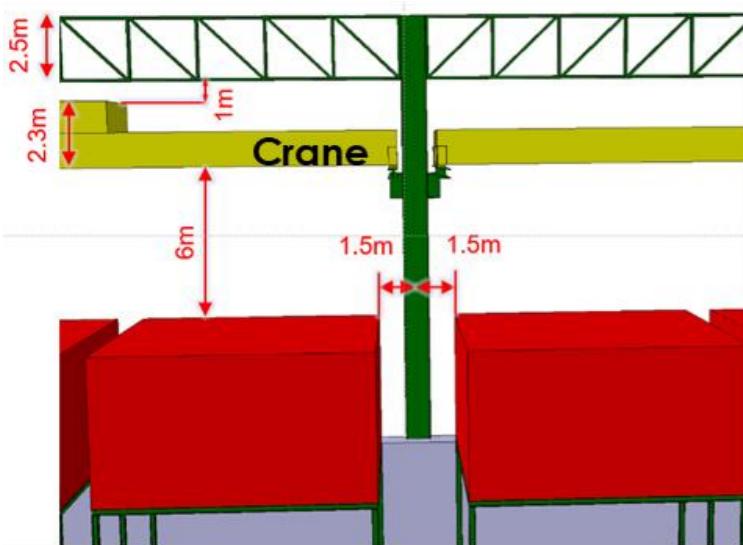
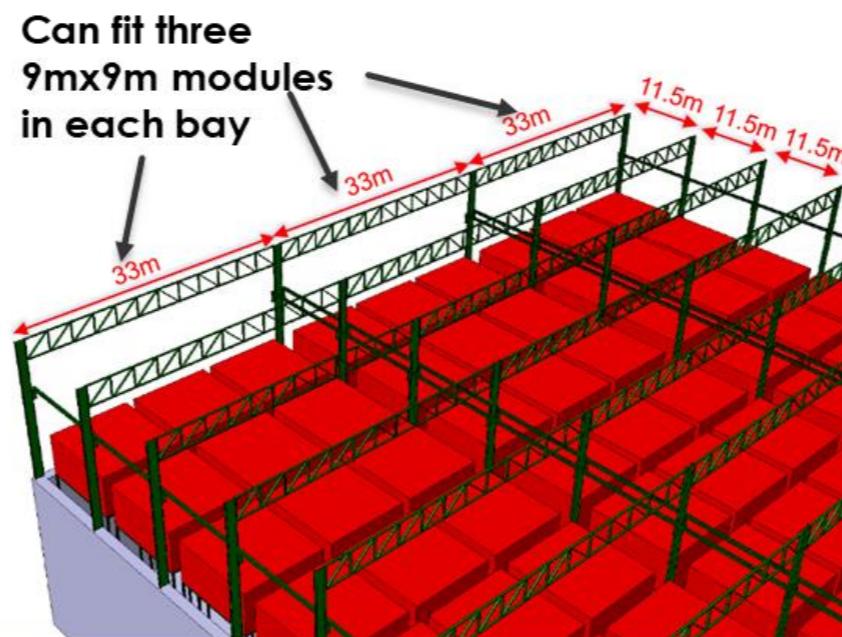
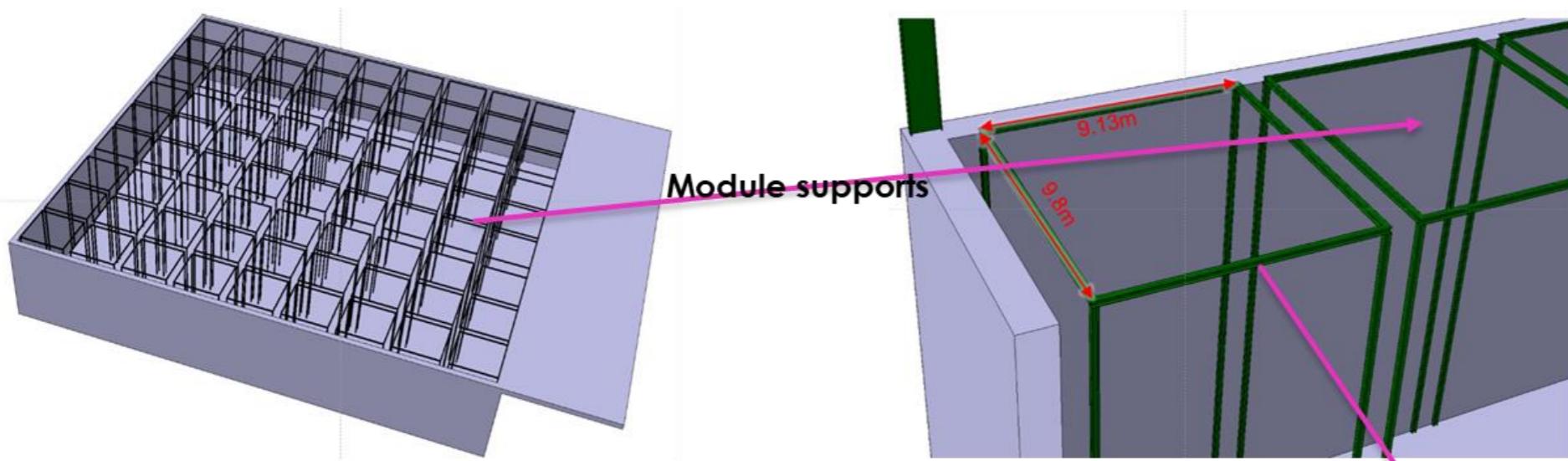
# Detector design



# Detector design



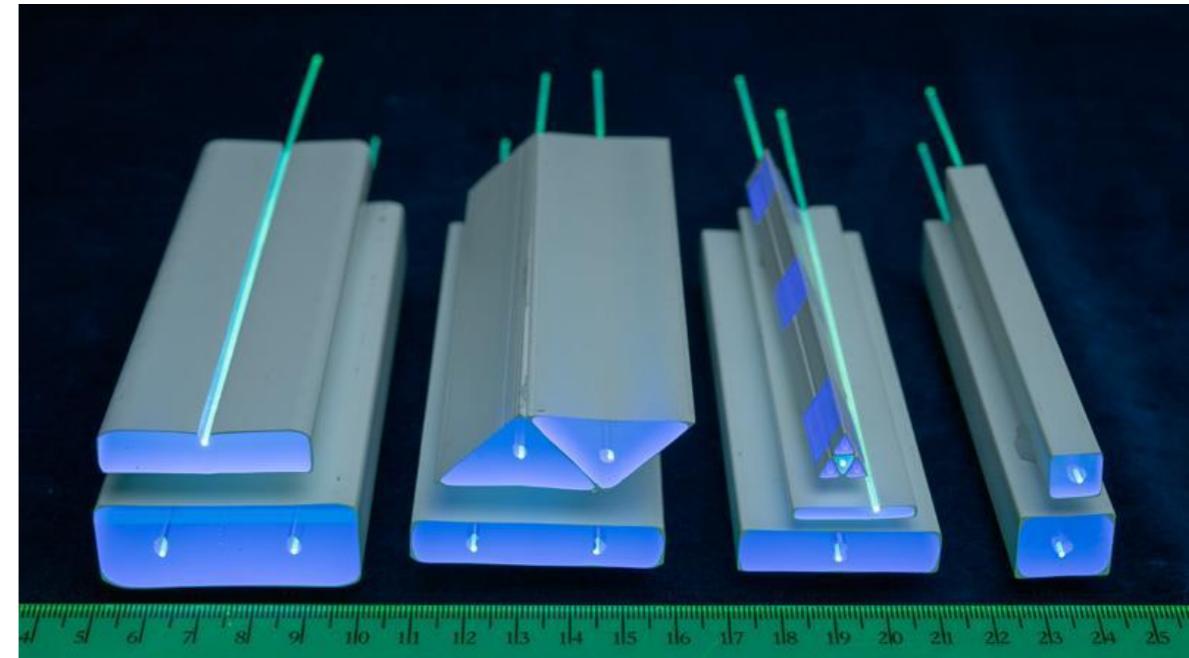
# Detector design



# Trackers, readout & trigger

Tracker layers: Composed of extruded scintillator bars with wavelength-shifting fibers coupled to Silicon Photo Multipliers

- Possibility of adding Resistive Plate Chamber layers



Each scintillator bar  $\sim 5\text{m} \times 4\text{cm} \times 2\text{cm}$ , with readout at both ends

- Transverse resolution  $\sigma \approx 1\text{ cm}$
- $\Delta t$  between two ends gives longitudinal resolution: need sub-ns precision

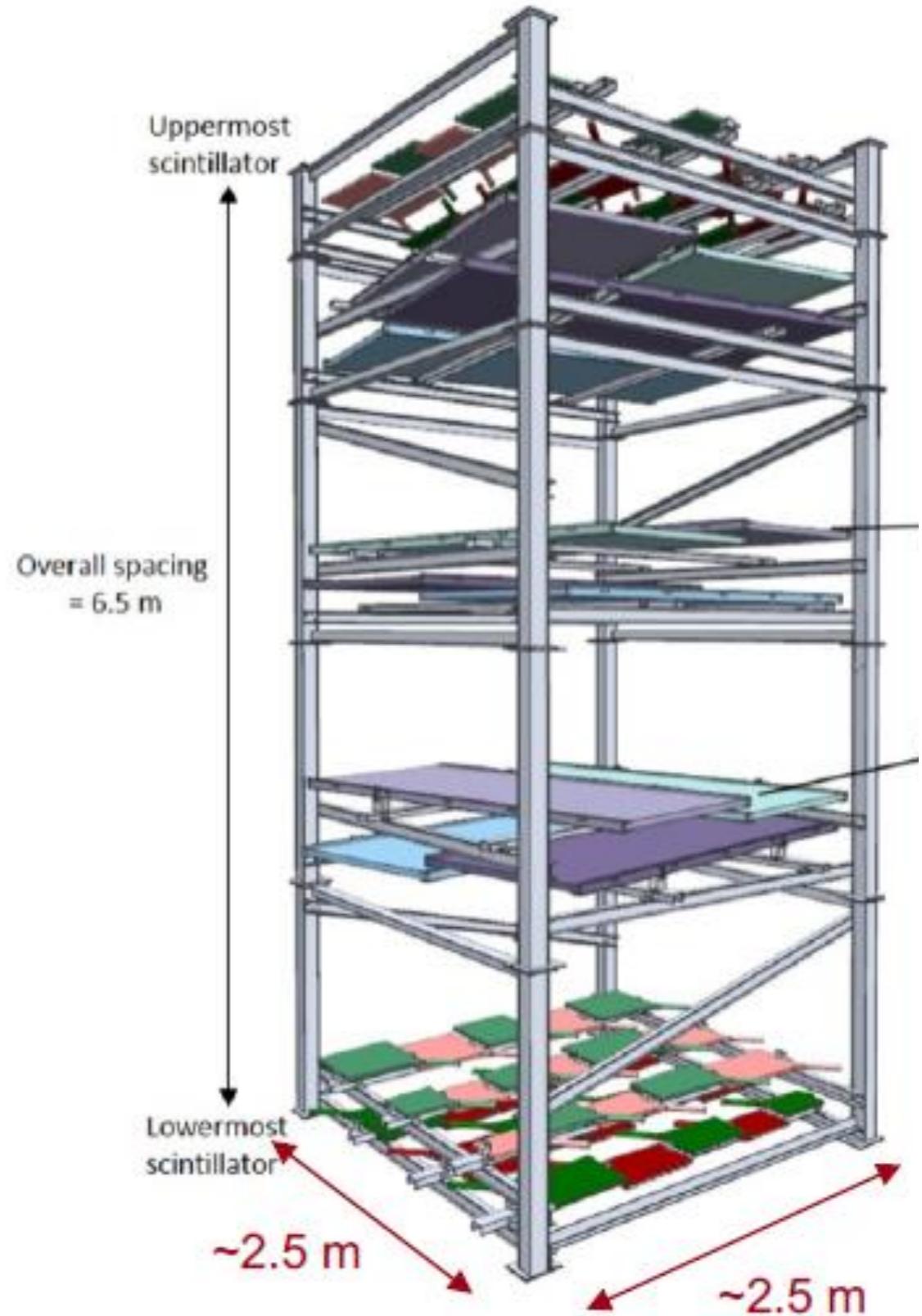
Collect all hits with no trigger selection; separately record trigger data and associate it with CMS bunch crossings

- Cosmic ray rate  $\sim 2\text{MHz}$

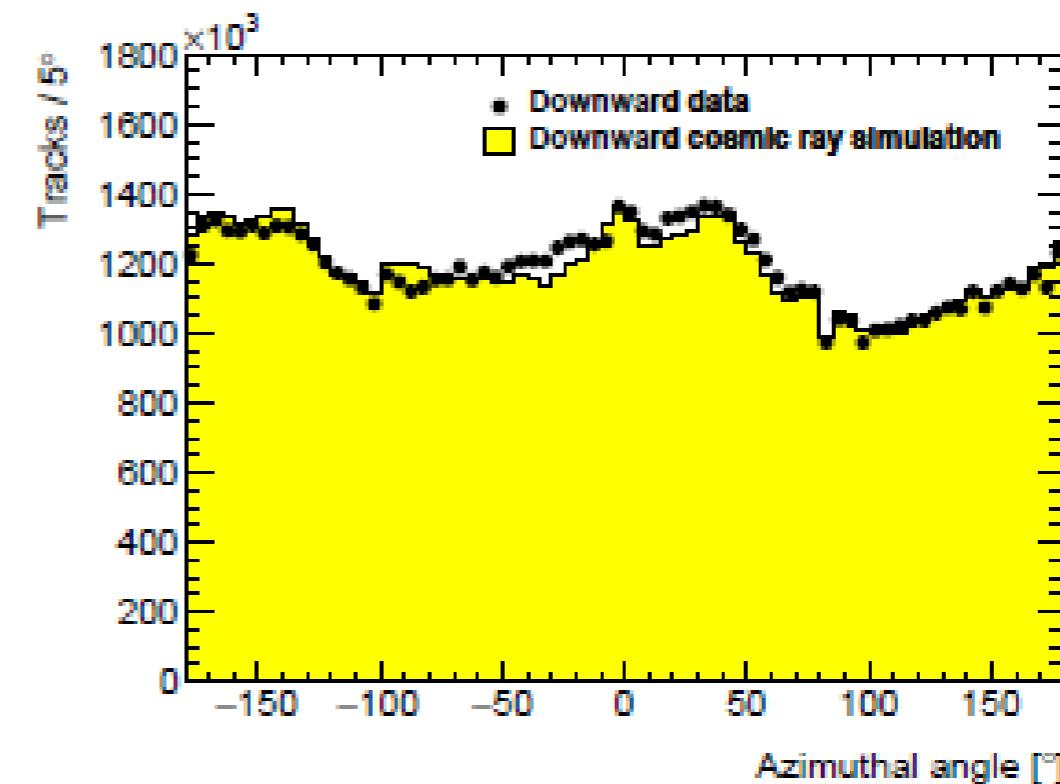
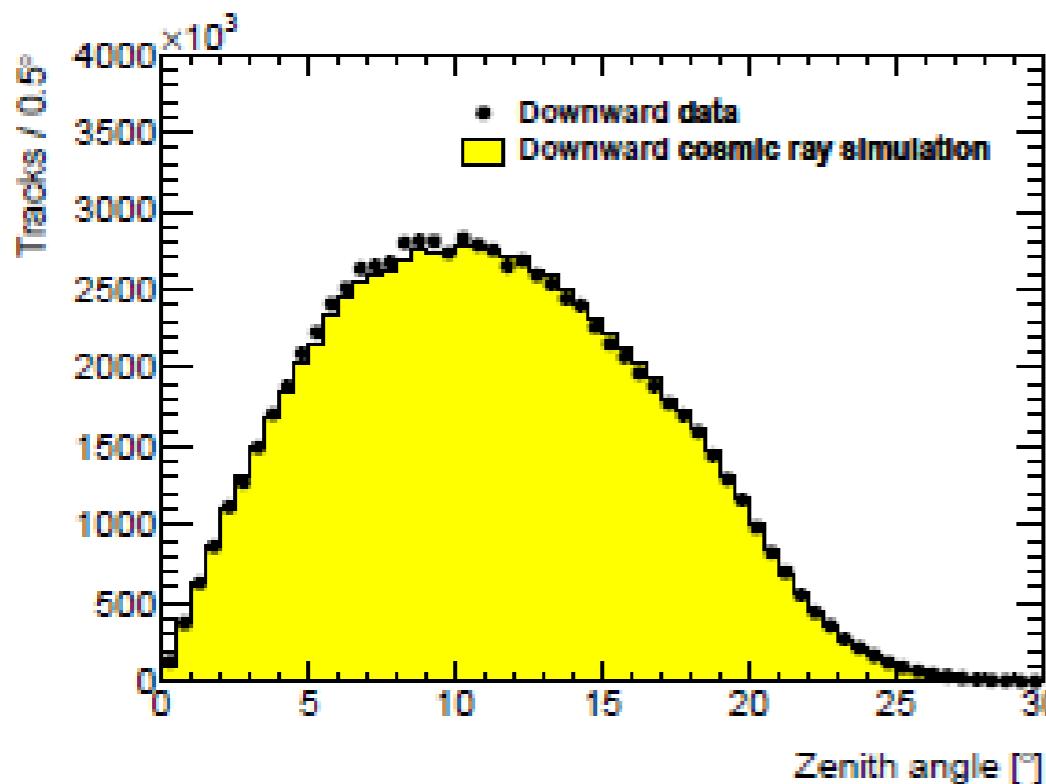
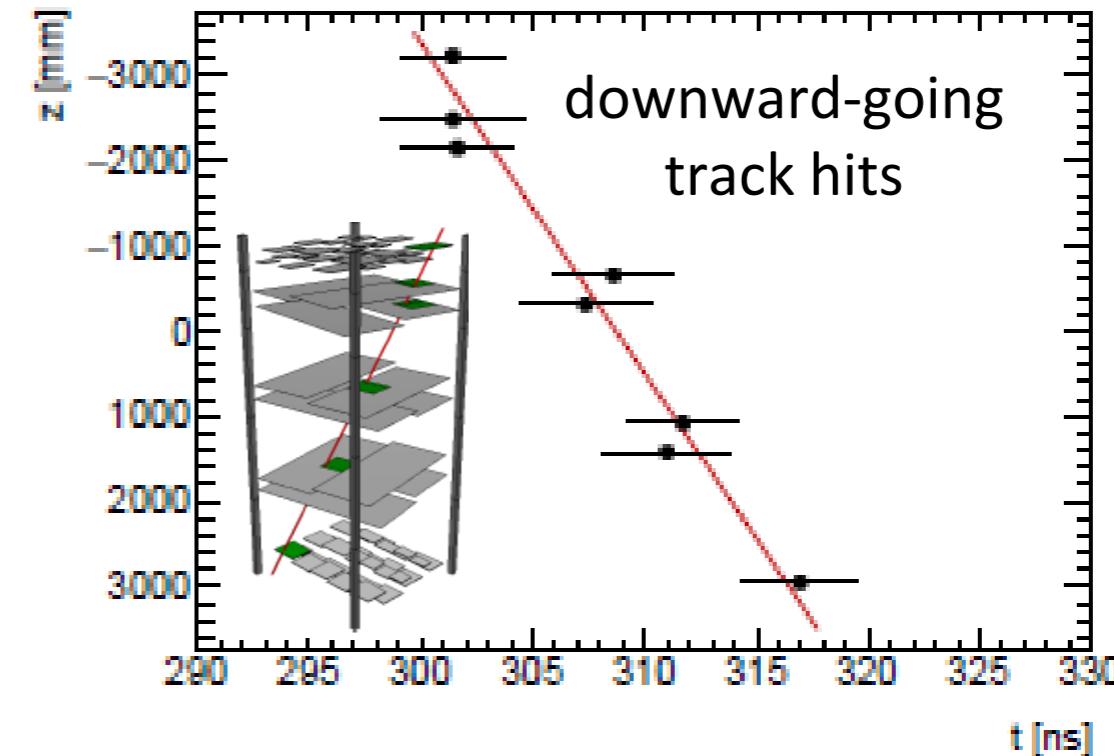
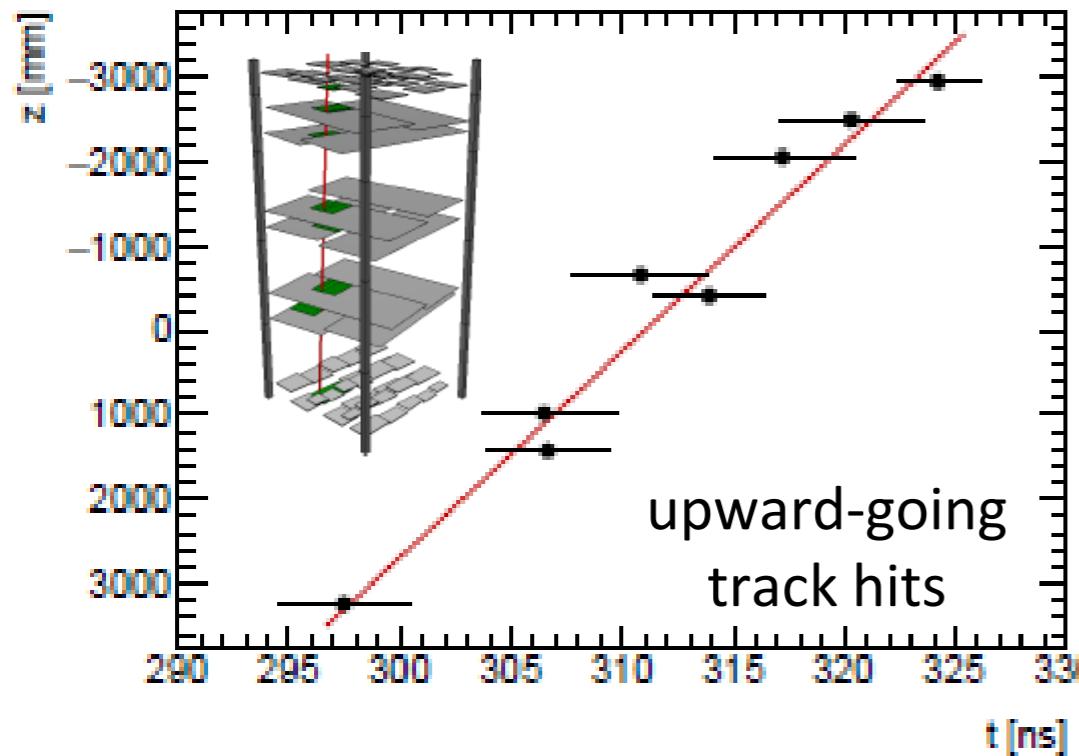
# Test Stand

- To understand LHC collision backgrounds (e.g. upward-going muons), built a test stand
- $\sim 2.5 \times 2.5 \times 6 \text{ m}^3$
- Two triggers running simultaneously:
  - Downward trigger for cosmic rays
  - Upward trigger for tracks from IP

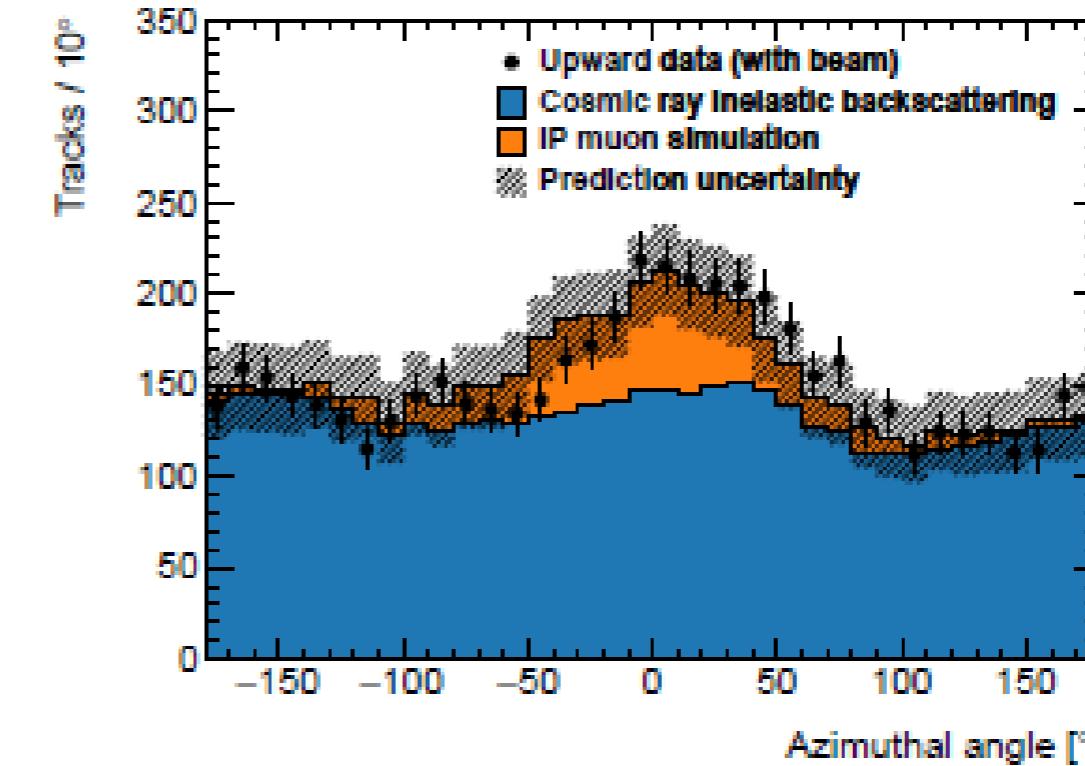
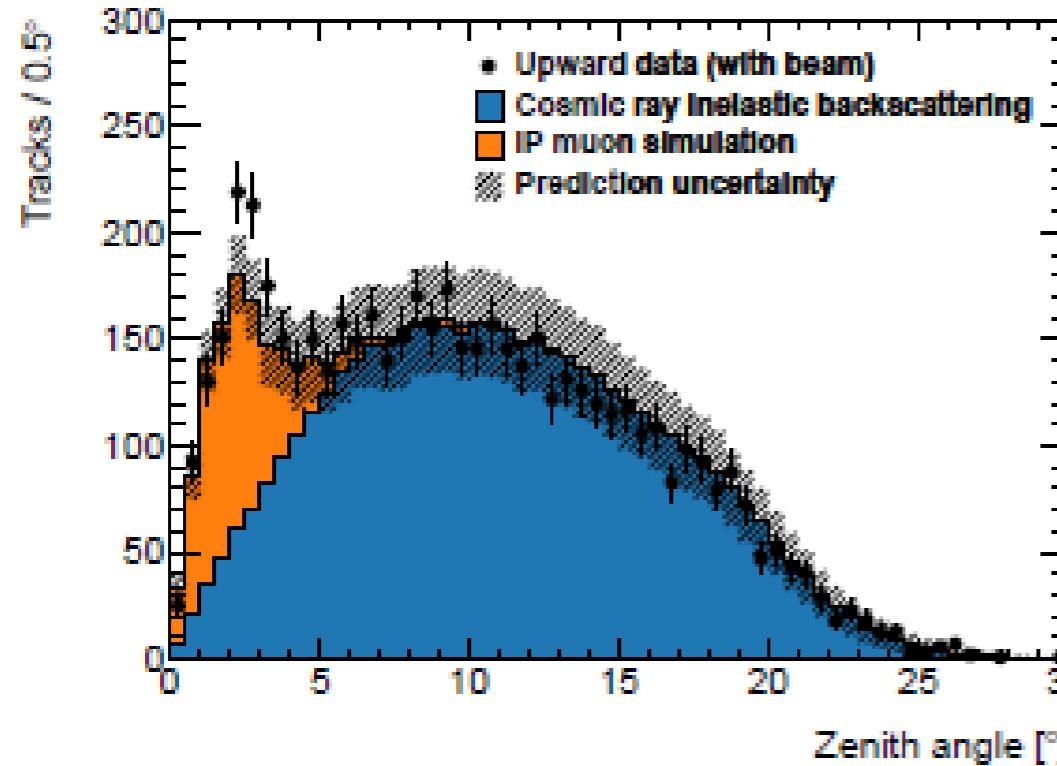
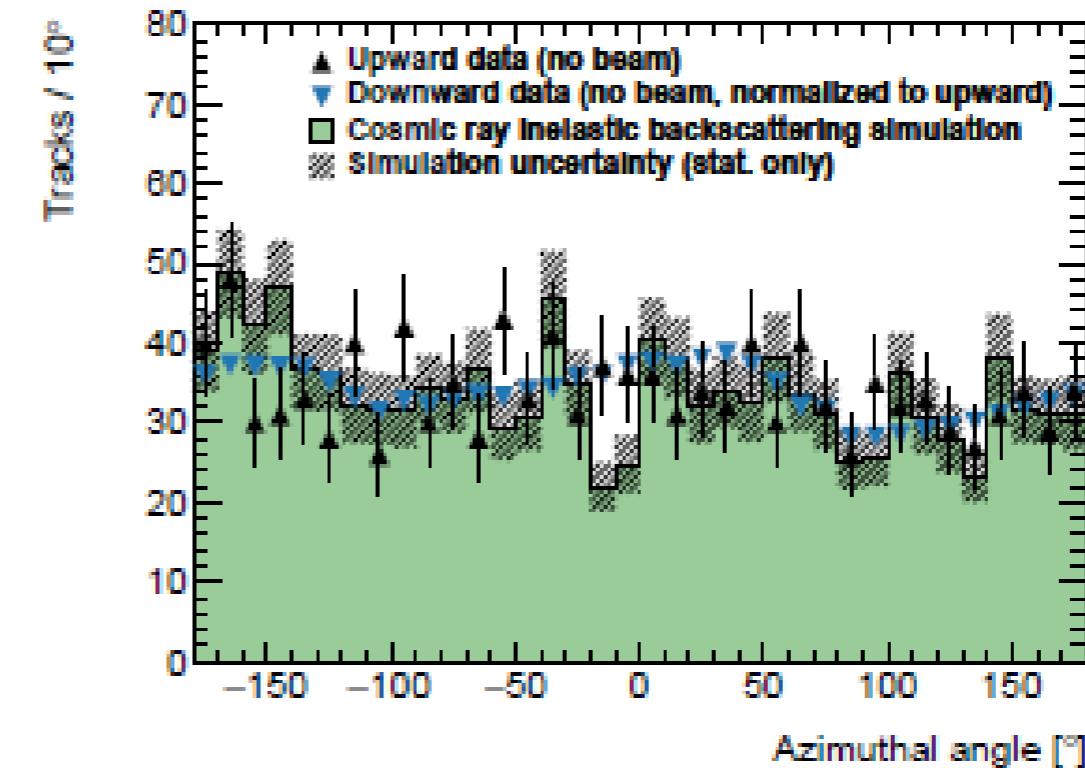
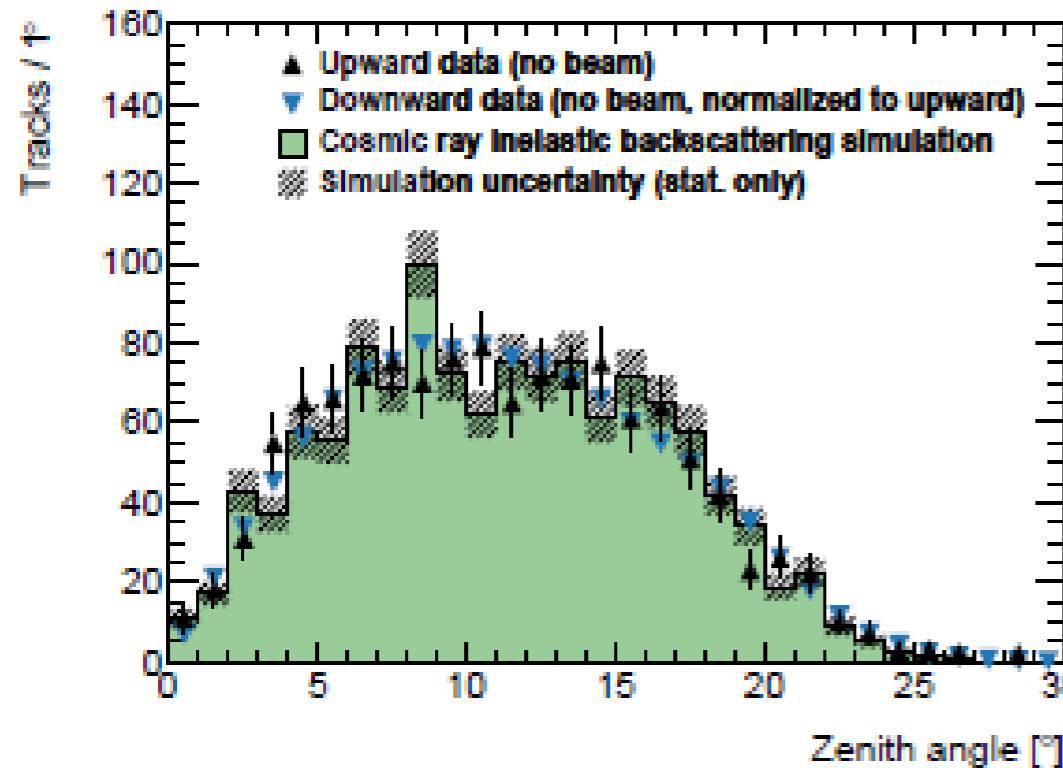
Took data above the ATLAS IP in 2018



# Test Stand



# Test Stand



MATUÍNÁ

Next Steps

# R&D plan

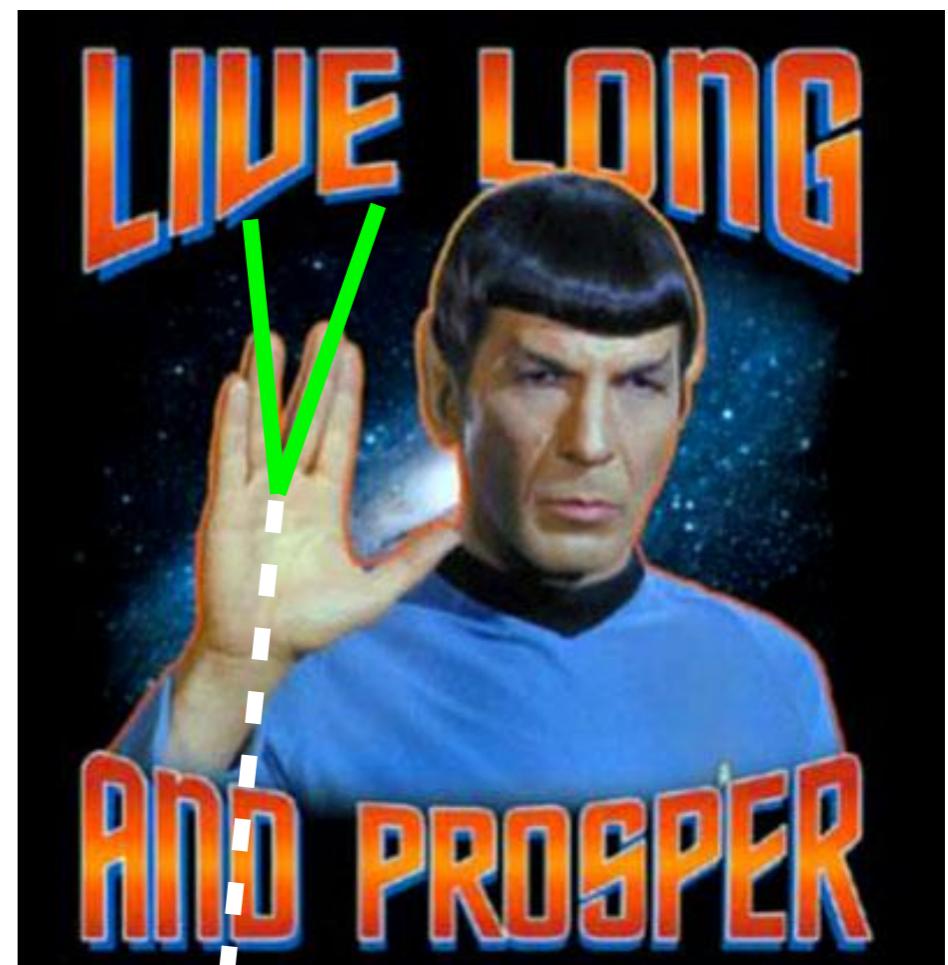
For MATHUSLA to go online ~2026, R&D required over the next few years includes:

- Front-end electronics design
- Trigger development (simple concept, nontrivial implementation)
- Detector design details (module parameters, tracking chamber support structure, installation procedures...)
- Cost optimization of detector components (scintillator, WLS, SiPMs)

# Opportunities

Exploration of the Lifetime Frontier will be central to the future of the LHC program to discover new physics

MATHUSLA is the best way to probe deep into the LLP lifetime parameter space for a wide range of masses



# References

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