

Thermal Physics : tutorial solution

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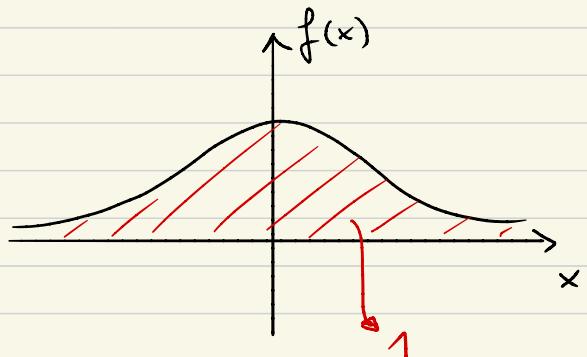
2020.03.04

1a) Show that the velocity distribution f^{h} is NORMALISED.

$$g(v_x) = \sqrt{\frac{m}{2\pi k_B T}} \exp\left[-\frac{mv_x^2}{2k_B T}\right]$$

NB condition of normalisation:

$$\int_D f(x) dx = 1$$



where D is the domain of $f(x)$

$$\begin{aligned} \int_{-\infty}^{+\infty} dv_x g(v_x) &= \int_{-\infty}^{+\infty} dv_x \sqrt{\frac{m}{2\pi k_B T}} \exp\left[-\frac{mv_x^2}{2k_B T}\right] \\ &= \sqrt{\frac{1}{2\pi \alpha}} \underbrace{\int_{-\infty}^{+\infty} dv_x \exp\left[-\frac{v_x^2}{2\alpha}\right]}_{\text{check table of integrals}} \end{aligned}$$

$$\text{where } \alpha = \frac{k_B T}{m} \geq 0$$

$$\text{Using } \int_{-\infty}^{+\infty} dx \exp\left[-\frac{x^2}{2\sigma^2}\right] = \sqrt{2\pi\sigma^{-2}},$$

$$\int_{-\infty}^{+\infty} dv_x g(v_x) = \sqrt{\frac{1}{2\pi \alpha}} \sqrt{2\pi \alpha} = 1 \quad \rightarrow \underline{\text{NORMALISED}}$$

1b) Show that $\langle v_x \rangle = 0$

NB $\langle v_x \rangle$ is called 1st moment of v_x

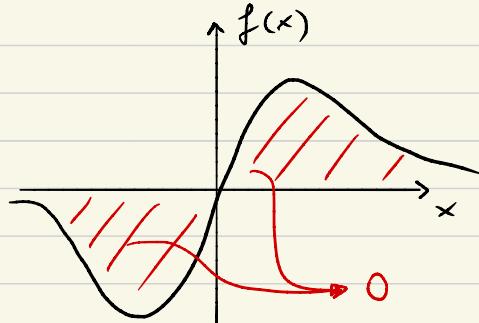
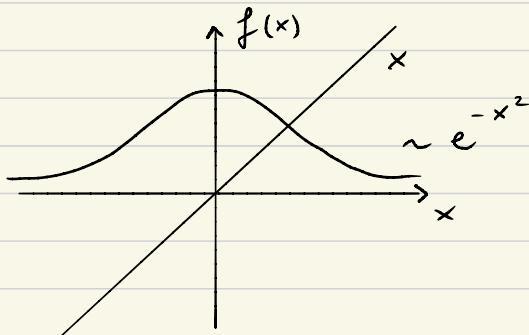
It is defined as $\langle v_x \rangle = \int_{-\infty}^{+\infty} dv_x v_x g(v_x)$

$$\int_{-\infty}^{+\infty} dv_x v_x g(v_x) = \int_{-\infty}^{+\infty} dv_x \sqrt{\frac{m}{2\pi k_B T}} v_x \exp\left[-\frac{mv_x^2}{2k_B T}\right]$$

odd
even
odd

$$= 0 \quad \underline{\text{OK}}$$

The integrand is an odd f^h on the domain of definition. The integral on the entire domain is therefore 0.



Show that $\langle v_x^2 \rangle = \frac{k_B T}{m}$

NB $\langle v_x^2 \rangle$ is called 2nd moment of v_x

It is defined as $\langle v_x^2 \rangle = \int_{-\infty}^{+\infty} dv_x v_x^2 g(v_x)$

$$\begin{aligned} \int_{-\infty}^{+\infty} dv_x v_x^2 g(v_x) &= \int dv_x \sqrt{\frac{m}{2\pi k_B T}} v_x^2 \exp\left[-\frac{mv_x^2}{2k_B T}\right] \\ &= \sqrt{\frac{1}{2\pi\alpha}} \int dv_x v_x^2 \exp\left[-\frac{v_x^2}{2\alpha}\right] \end{aligned}$$

We want to use the integral:

$$\int_0^{\infty} dx x^2 e^{-x^2} = \frac{\sqrt{\pi}}{4},$$

but before, we have to change variable and "fix" the domain of integration.

$$y^2 = \frac{v_x^2}{2\alpha}, \quad v_x^2 = 2\alpha y^2, \quad v_x = \sqrt{2\alpha} y$$

$$dv_x = \sqrt{2\alpha} dy$$

the integral becomes

$$\begin{aligned} \int_{-\infty}^{+\infty} dv_x v_x^2 g(v_x) &= \frac{1}{\sqrt{2\pi\alpha}} \int_{-\infty}^{+\infty} \sqrt{2\alpha} dy 2\alpha y^2 \exp[-y^2] \\ &= \frac{2\alpha}{\sqrt{\pi}} \int_{-\infty}^{\infty} dy y^2 \exp[-y^2] \end{aligned}$$

even

NB Since the f^n is EVEN,

$$\int_{-\infty}^{+\infty} dy \, y^2 e^{-y^2} = 2 \int_0^{\infty} dy \, y^2 e^{-y^2}$$

$$\Rightarrow \int_{-\infty}^{+\infty} dv_x \, v_x^2 g(v_x) = \frac{2\alpha}{\sqrt{\pi}} \int_{-\infty}^{+\infty} dy \, y^2 e^{-y^2}$$

$$= \frac{4\alpha}{\sqrt{\pi}} \underbrace{\int_0^{\infty} dy \, y^2 e^{-y^2}}_{\sqrt{\frac{\pi}{4}}}$$

$$= \frac{4\alpha}{\sqrt{\pi}} \frac{\sqrt{\pi}}{4}$$

$$= \alpha$$

$$= \frac{k_B T}{m} \quad \underline{\text{OK}}$$

2) the Maxwell - Boltzmann distribution for an ideal gas is:

$$p(\varepsilon) d\varepsilon = \frac{2}{\sqrt{\pi}} \beta^{3/2} \sqrt{\varepsilon} \exp[-\beta\varepsilon] d\varepsilon$$

2a) Normalisation condition: $\int_{\text{Domain}} f(x) dx = 1$

2b) Show that $p(\varepsilon)$ is properly normalised.

$$\int_0^{+\infty} d\varepsilon p(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \int_0^{+\infty} d\varepsilon \sqrt{\varepsilon} \exp[-\beta\varepsilon]$$

we want to use the integral $\int_0^{+\infty} \sqrt{x} e^{-x} dx = \frac{\sqrt{\pi}}{2}$

but before we have to change variable.

$$y = \beta\varepsilon, \sqrt{\varepsilon} = \frac{1}{\sqrt{\beta}} \sqrt{y}, dy = \beta d\varepsilon, d\varepsilon = \frac{1}{\beta} dy$$

$$\int_0^{+\infty} d\varepsilon p(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \int_0^{+\infty} \frac{1}{\beta} dy \frac{1}{\sqrt{\beta}} \sqrt{y} \exp[-y]$$

$$= \frac{2}{\sqrt{\pi}} \underbrace{\int_0^{+\infty} dy \sqrt{y} e^{-y}}_{\frac{\sqrt{\pi}}{2}}$$

$$= 1 \quad \underline{\text{OK}}$$

2c) Using $\int_0^\infty x^{3/2} e^{-x} dx = \frac{3\sqrt{\pi}}{4}$

& equipartition theorem, show that $\beta = \frac{1}{k_B T}$

The equipartition theorem relates the average energy per particle with the temperature of the gas in thermal equilibrium.

For a monoatomic gas in 3 dimensions is:

$$\langle \varepsilon \rangle = \frac{3}{2} k_B T$$

in general, $\langle \varepsilon \rangle = \underbrace{dof}_{\text{every quadratic degree of freedom}} \cdot \frac{1}{2} k_B T$

(in this case 3)

We calculate $\langle \varepsilon \rangle$, the 1st moment of ε

$$\langle \varepsilon \rangle = \int_0^\infty d\varepsilon \varepsilon p(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \int_0^\infty d\varepsilon \varepsilon^{3/2} \exp[-\beta\varepsilon],$$

we change variable

$$y = \beta \varepsilon, \quad \varepsilon = \frac{1}{\beta} y, \quad d\varepsilon = \frac{1}{\beta} dy$$

$$\int_0^\infty d\varepsilon \varepsilon p(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \int_0^\infty \frac{1}{\beta} dy \frac{1}{\beta^{3/2}} y^{3/2} e^{-y}$$

$$= \frac{2}{\sqrt{\pi}} \frac{1}{\beta} \left[\int_0^{+\infty} dy \, y^{3/2} e^{-y} \right]$$

$\frac{3}{4} \sqrt{\pi}$

$$\langle \varepsilon \rangle = \frac{3}{2} \frac{1}{\beta}$$

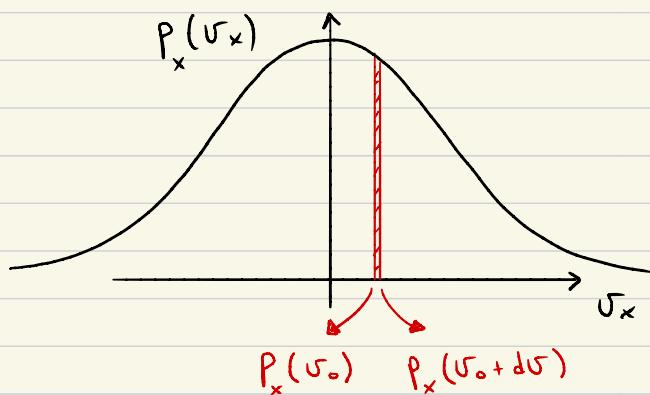
and from equipartition theorem $\langle \varepsilon \rangle = \frac{3}{2} k_B T$

$$\frac{3}{2} \frac{1}{\beta} = \frac{3}{2} k_B T \Rightarrow \beta = \frac{1}{k_B T}$$

OK

3a) State the meaning of $p_x(v_x) dv_x$.

$p_x(v_x) dv_x$ is the probability of finding a particle with velocity between $[v_x, v_x + dv_x]$.



3b) Assume that $p_x(v_x) = \exp[-\alpha v_x^2]$

Show that for IDEAL 2D gas the speed distr. is:

$$p(v) dv = f v \exp[-\alpha v^2] dv$$

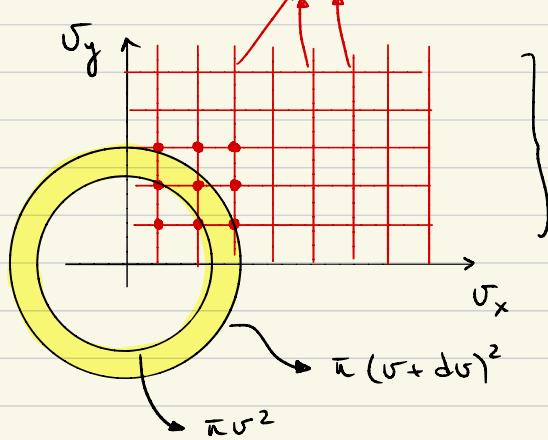
with $v = \sqrt{v_x^2 + v_y^2}$ and f is CONSTANT.

- The 2-D probability of finding a particle with velocity between $[v_x, v_x + dv_x]$ & $[v_y, v_y + dv_y]$ is :

$$\underbrace{p(v_x) p(v_y)}_{\downarrow \downarrow} \underbrace{dv_x dv_y}_{p(v) \propto v dv d\phi}$$

In more detail, in VELOCITY - space we have:

there are all the possible states of the velocity



considering this states infinitely close each other, we can consider a continuous space.

We want to find the state in an infinitesimal portion of area in the velocity space. We do this in polar coordinates.

The highlighted infinitesimal area is :

$$\underbrace{\pi(r+dr)^2}_{\text{LARGE area}} - \underbrace{\pi r^2}_{\text{small area}} = \pi(\cancel{r^2} + \cancel{dr^2} + 2rd\theta - \cancel{r^2})$$

2nd order

$$= 2\pi r dr$$

from here, we have the "r" dependence

3c) Find f by requiring $p(r)$ NORMALISED

$$f \int_0^{+\infty} dr r \exp[-ar^2]$$

NB { the integral is $\int_0^{+\infty}$ and
not $\int_{-\infty}^{+\infty}$ because $r = \sqrt{r_x^2 + r_y^2} > 0$

$$\text{changing } y = ar^2, \quad r = \frac{1}{\sqrt{a}} \sqrt{y}, \quad dr = \frac{1}{2\sqrt{a}} \frac{1}{\sqrt{y}} dy$$

$$= f \int_0^{+\infty} dy \frac{1}{2\sqrt{a}} \frac{1}{\sqrt{y}} \frac{1}{\sqrt{a}} \sqrt{y} e^{-y}$$

$$= f \frac{1}{2a} \int_0^{+\infty} e^{-y} dy$$

$$= f \frac{1}{2a} (-e^{-y}) \Big|_0^{+\infty}$$

$$= f \frac{1}{2a} (-0 - (-1)) = 1 \quad \text{for NORMALISATION}$$

$$f \frac{1}{2a} = 1 \Rightarrow f = 2a$$

OK

3d) The prefactor in 2D is " r " (see previous derivation)

while in 3D is " r^2 " because of the infinitesimal spherical surface we find deriving the probability.

$$p(r) = p_x(r_x) p_y(r_y) p_z(r_z) \frac{dr_x dr_y dr_z}{4\pi r^2 dr}$$

$$\text{where } r = \sqrt{r_x^2 + r_y^2 + r_z^2}.$$

Again the volume in probability space is:

$$\frac{4}{3}\pi (r+dr)^3 - \frac{4}{3}\pi r^3 = \frac{4}{3}\pi (\cancel{r^3} + 3r^2 dr + 3\cancel{dr^2} + \cancel{dr^3} - \cancel{r^3}) = 4\pi r^2 dr$$

2nd order 3rd order

3e) Relation particle speed & energy $E = \frac{1}{2} mv^2$

Show that:

$$P(E) dE = \frac{2a}{m} \exp\left[-\frac{2a}{m} E\right] dE$$

$$\text{We have } P(v) dv = 2a v \exp[-av^2] dv$$

$$v = \sqrt{\frac{2}{m} E}, \quad dv = \sqrt{\frac{1}{2m}} \frac{1}{\sqrt{E}} dE$$

$$\begin{aligned} p(v) dv &\rightarrow P(E) dE = 2a \sqrt{\frac{2}{m} E} \exp\left[-\frac{2a}{m} E\right] \frac{1}{\sqrt{2m}} \frac{1}{\sqrt{E}} dE \\ &= \frac{2a}{m} \exp\left[-\frac{2a}{m} E\right] dE \end{aligned}$$

OK

3f) Average energy of the particles in 2D

$$\langle E \rangle = \int_0^{+\infty} E P(E) dE \quad \underline{\text{NB}} \quad E = \frac{1}{2} mv^2 \quad E \text{ can NOT be NEGATIVE}$$

$$= \frac{2a}{m} \int_0^{+\infty} E \exp\left[-\frac{2a}{m} E\right] dE \quad \Rightarrow \int_0^{+\infty}$$

$$y = \frac{2a}{m} E, \quad E = \frac{m}{2a} y, \quad dE = \frac{m}{2a} dy$$

$$\langle E \rangle = \frac{2a}{m} \int_0^{+\infty} \frac{m}{2a} y e^{-y} \frac{m}{2a} dy$$

$$= \frac{m}{2a} \int_0^{+\infty} y e^{-y} dy$$

$$= \frac{m}{2a} \quad \underline{1} \quad \underline{\text{OK}}$$

3g) Equipartition theorem : $\langle E \rangle = \text{dof} \cdot \frac{1}{2} k_B T$

quadratic degrees of freedom

$$\text{In 2D } \langle E \rangle = k_B T .$$

$$k_B T = \frac{m}{2a} \Rightarrow a = \frac{m}{2k_B T}$$

3h) Calculate the probability that the energy of a particle randomly chosen is less than $k_B T$.

We have the probability density. To calculate the probability for a particle to be within a certain energy, we have to integrate $p(E)$ up to the energy considered.

$$\begin{aligned}
 P(E < k_B T) &= \int_0^{k_B T} p(E) dE \\
 &= \frac{2a}{m} \int_0^{k_B T} \exp\left[-\frac{2a}{m} E\right] dE \\
 &= \frac{2}{m} \frac{m}{2k_B T} \int_0^{k_B T} \exp\left[-\frac{2}{m} \frac{m}{2k_B T} E\right] dE \\
 &= \frac{1}{k_B T} \int_0^{k_B T} \exp\left[-\frac{1}{k_B T} E\right] dE \\
 &= \frac{1}{k_B T} \left(-\frac{1}{k_B T} \right) \left[e^{-\frac{1}{k_B T} E} \right]_0^{k_B T} \\
 &= -\left(e^{-1} - e^0 \right) = 1 - \frac{1}{e} \approx 63\%
 \end{aligned}$$

OK