

Week 2 and 3 discussion questions

2 types of supernovae (1,2)

#3

Type 2:

- core collapse
 - onion structure
 - iron core
 - implode / explode
 - production of neutrinos

Type 1: (most common)

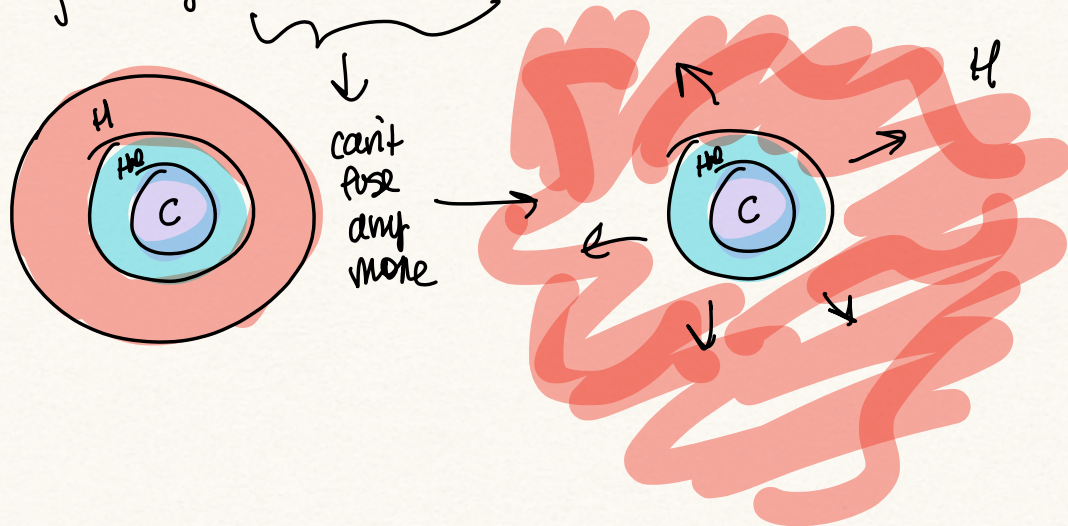
- Binary Star Systems
 - white dwarf strips energy of other star
 - one star \gg other, burns fuel faster
 - can produce nebula or supernova

Planetary Nebula?

#4

- less massive star (no supernova)
 - goes through life span

- after no fuel, No fusion
- becomes red giant?
- gravity compresses star, temp/pressure \uparrow .



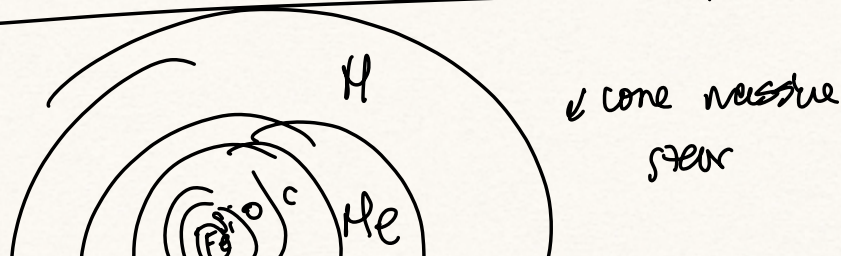
Why we detect neutrinos first. #2

- Neutrino cross section $<$ photon cross section
 \rightarrow interacts w/ less during supernova process

$$\sigma_{em} \sim 10^{-27} \text{ cm}^2 \leftarrow \text{photons interaction area}$$

$$\sigma_{weak} \sim 10^{-44} \text{ cm}^2 \leftarrow \text{neutrinos}$$

How neutrinos formed in supernovae #1





↓
Heavier elements
to center

→ When Fe is $\sim 1.4 M_{\odot}$, star implodes (~~Fusion~~)

Typical
atom:

○
nucleus

empty
space

electrons

Nuclei
compressed Fe

All layers
pressed in

$p + e \rightarrow n + \nu$



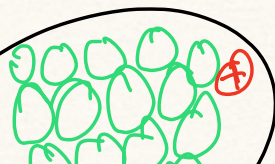
zoom
in.

(compressed
within seconds)

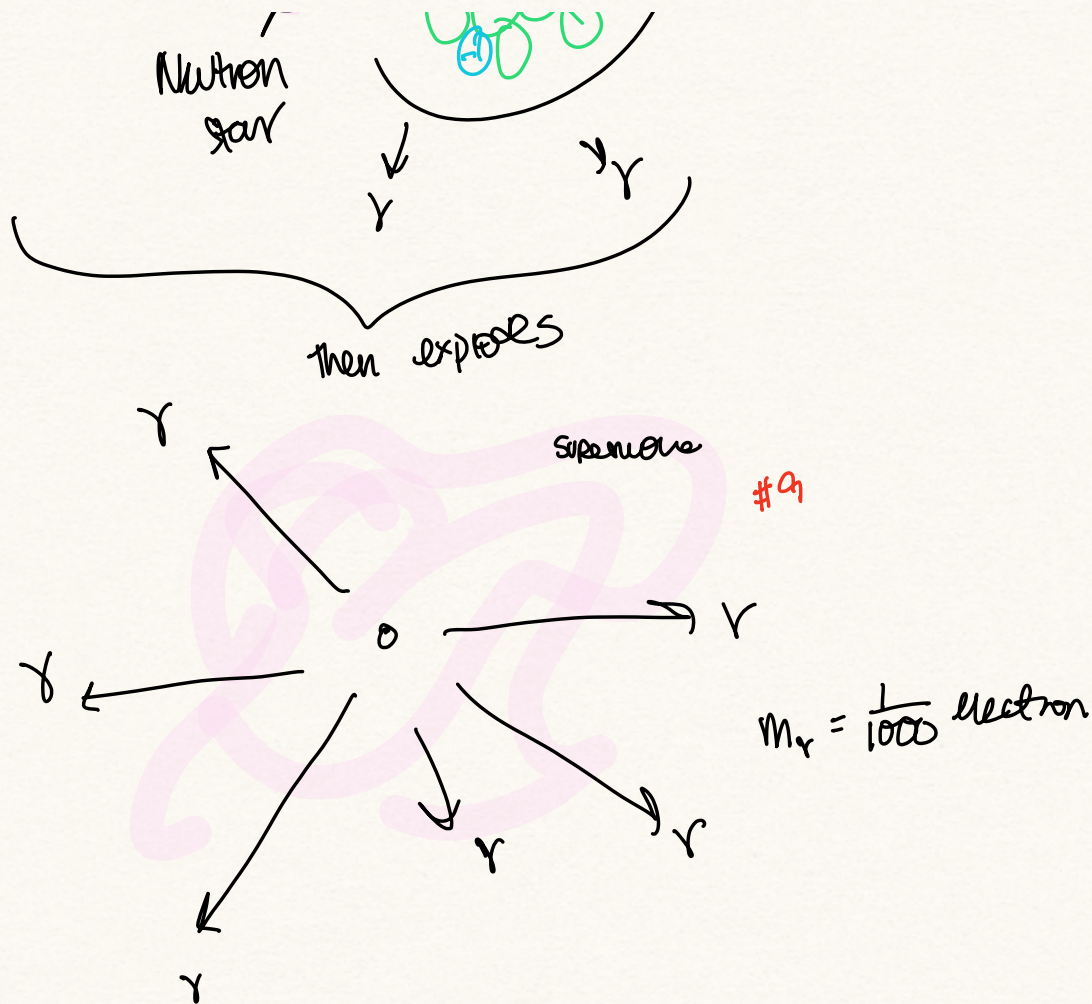
* cannot put 2 electrons within orbital *
→ electrons NEED space, cannot be near.

* Detect neutrino: $\gamma + p \rightarrow n + e^+$ *

then:



→ ν



How can neutrinos explain things #8

- # of neutrinos, time
- comparing to models to know how supernova form
- Neutrinos for high energy processes
 - neutron star mergers
 - gamma ray bursts

- only react via weak force

- extremely low mass

- have mass

- C , experience no time

- γ do oscillate

- γ flavors → change over time

- must have mass

What makes up supernova remnant. #5

- hydrogen, helium, other elements

- heavy elements in fusion that
ends around explosion

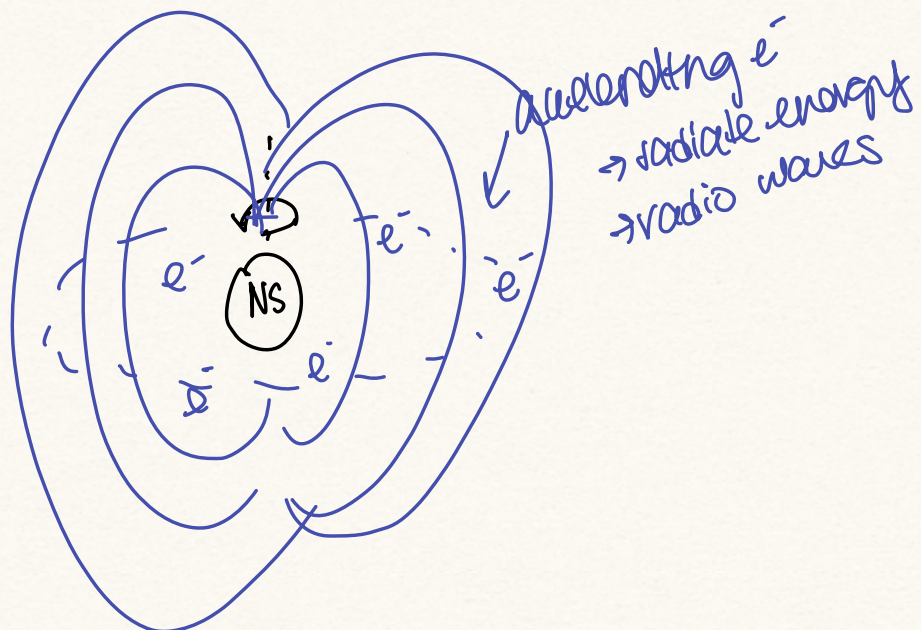
→ Temperature

- $\sim 10^4 K$, hotter than surface of sun

- Temp cools down because of getting
density dispersed into the universe

→ Energy goes to neutrinos/light and
 due to entropy transfers to the gas, in the
 universe.

Blue glow → neutron star
 → supernova



Density of Neutron Star: $\rho_{NS} = \frac{M_{NS}}{V_{NS}} = \frac{1.5 M_{\odot}}{\frac{4}{3} \pi R_{NS}^3}$

$R_{NS} \sim 10 \text{ km}$

$\rho_{NS} = 7 \times 10^{14} \text{ g cm}^{-3}$

#7

$\rho_{\text{center of } \odot} \sim 100 \text{ g cm}^{-3} \rightarrow 10^{12} \times \text{denser}$
 (trillion)

$\rho_{\oplus} \sim 5.5 \text{ g cm}^{-3}$

Earth will only react w/ $\frac{1}{161}$ neutrinos

$6.6 \times 10^{10} \frac{\text{cm}^3}{\text{s}} \sim \text{Google}$ #7

$$60 \times 10^9 \frac{1}{\text{cm}^2 \text{s}}$$