

Homework #3 – ASEN 5050

Due: Thursday, 9/17/2015

Note: Use Appendix D of the book for all constants not given in the problem.

1. (40 pts) Write a computer program (any language you like, though I suggest making your life easier and using Matlab) that converts ECI position and velocity vectors into orbital elements (Algorithm 9 in the textbook). To be perfectly useful, also include the gravitational parameter, μ , as an input. The function call should look like:

$$\text{function [a, e, i, } \Omega, \omega, v] = \text{RV2COE(Rijk, Vijk, } \mu)$$

You can test if your function works by reproducing Example 2-5 in the textbook. Once you have a working function, compute orbital elements for the following ECI position/velocity:

$$\bar{\mathbf{r}} = \begin{bmatrix} -5633.9 \\ -2644.9 \\ 2834.4 \end{bmatrix} \text{ km} \quad \bar{\mathbf{v}} = \begin{bmatrix} 2.425 \\ -7.103 \\ -1.800 \end{bmatrix} \text{ km/sec}$$

The goal is to make a program that works in as many circumstances as possible, reliably and in the most useful way. Test it out on many types of orbits!

2. (40 pts) Write a computer program that converts orbital elements to ECI position and velocity vectors (Algorithm 10 in the textbook). The function call should look like:

$$\text{function [Rijk, Vijk] = COE2RV(a, e, i, } \Omega, \omega, v, \mu)$$

You can test if your function works by reproducing Example 2-6 in the textbook.

Once you have a working function, compute the ECI position and velocity for the International Space Station (ISS) using the following [NORAD Two Line Element Set \(TLE\)](http://www.celestrak.com/NORAD/documentation/tle-fmt.asp) (<http://www.celestrak.com/NORAD/documentation/tle-fmt.asp>), which I downloaded from the [CelesTrak web site](http://www.celestrak.com/) (<http://www.celestrak.com/>). Note: TLEs use mean orbital elements; treat the values as instantaneous values.

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1 25544U 98067A 01260.91843750 .00059354 00000-0 74277-3 0 4795

2 25544 51.6396 342.1053 0008148 106.9025 231.8021 15.5918272116154

3. (20 pts) Write a program that computes the eccentric anomaly (E) given the mean anomaly (M) and the eccentricity (e) (Algorithm 2 in the textbook). You can test your program using Example 2-1 in the textbook. Embed this in a program that computes the true anomaly given the time since periape passage. Use this routine and "COE2RV" from Problem 2 to predict the position and velocity of the ISS 1 hour past the epoch of the TLE.