

Stratified flows: Equilibrium, instability and turbulence

Marco Toffolon
(University of Trento)



(Marco Adami, *Scirocco*, 2008)

References

Benoit Cushman-Roisin and Jean-Marie Beckers, **Introduction to geophysical fluid dynamics : physical and numerical aspects**, 2nd ed., Academic Press, 2011.

<https://webapps.unitn.it/Biblioteca/it/Web/LibriElettroniciDettaglio/117963>

- 11. Stratification
 - 11.1 Introduction
 - 11.2 Static Stability
- 14. Turbulence in Stratified Fluids
 - 14.1 Mixing in Stratified Fluids

Pijush K. Kundu, Ira M. Cohen, David R. Dowling, **Fluid mechanics**, 5th ed., Academic Press, Elsevier, 2012.

<https://www.sciencedirect-com.ezp.biblio.unitn.it/book/9780123821003/fluid-mechanics>

- 1. Introduction
 - 1.10. Stability of Stratified Fluid Media
- 11. Instability
 - 11.3. Kelvin-Helmholtz Instability
- 12. Turbulence
 - 12.11. Turbulence in a Stratified Medium

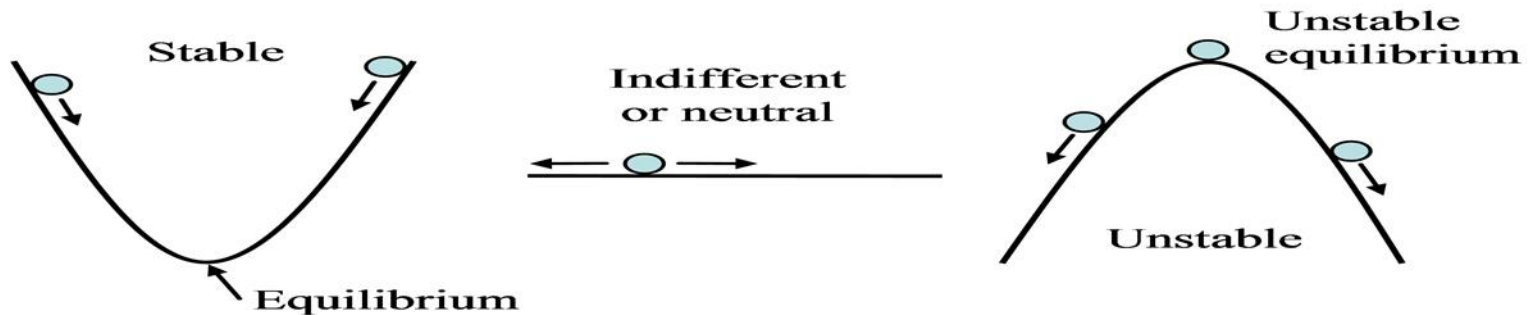
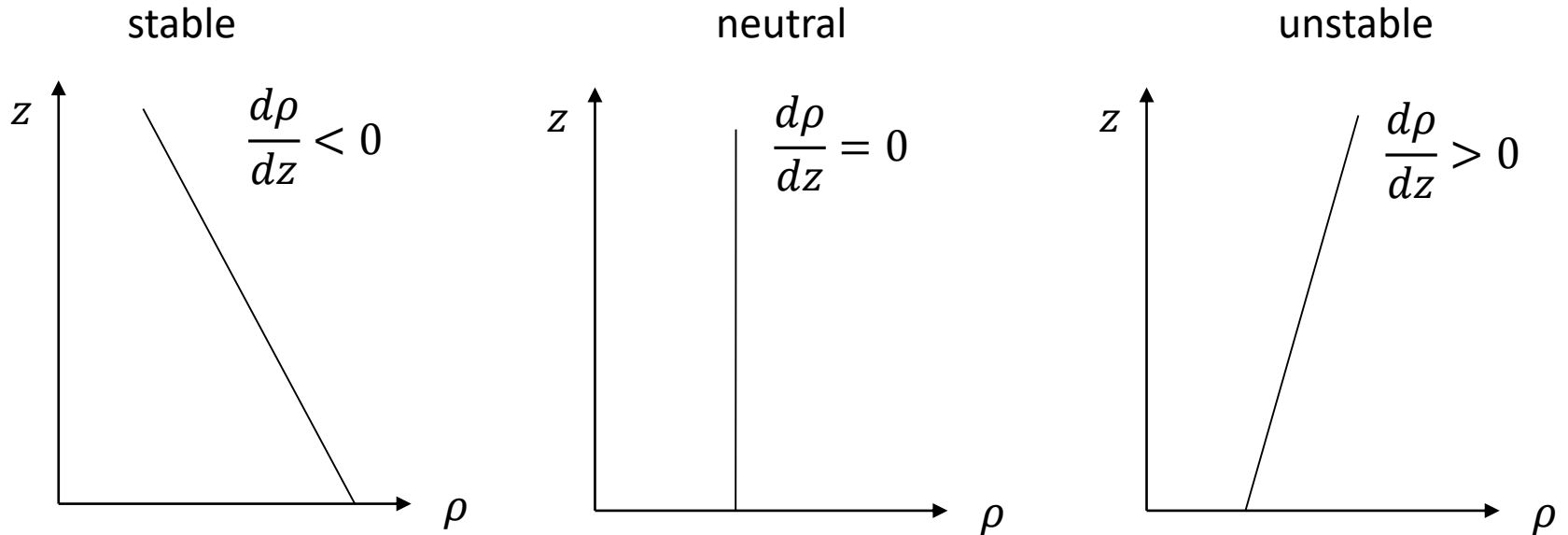
Scott A. Socolofsky, **Mixing and Transport Processes in the Environment**, Lecture notes – **part 2**.

- ch. 8 – Concepts, Definitions, and Governing Equations
- ch. 11 – Instability

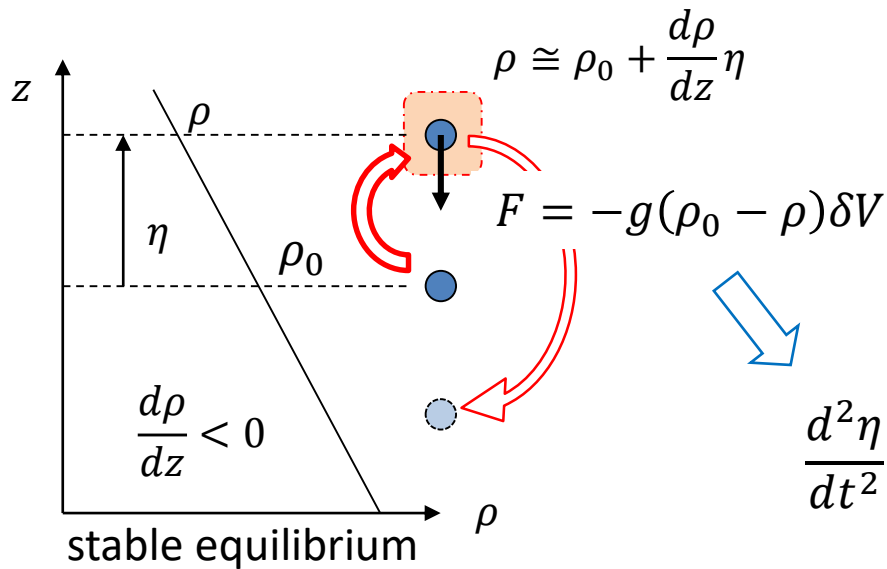
DENSITY STRATIFICATION AND GRAVITATIONAL EQUILIBRIUM

Equilibrium

(hp. neglecting the effect of pressure on density)



Oscillations around stability: a simple model



second Newton's law

$$F = ma = \rho_0 \delta V \frac{d^2 \eta}{dt^2}$$

$$\frac{d^2 \eta}{dt^2} - \frac{g}{\rho_0} \frac{d\rho}{dz} \eta = 0$$

$$\frac{d^2 \eta}{dt^2} + N^2 \eta = 0$$

$$N = \sqrt{-\frac{g}{\rho} \frac{d\rho}{dz}}$$

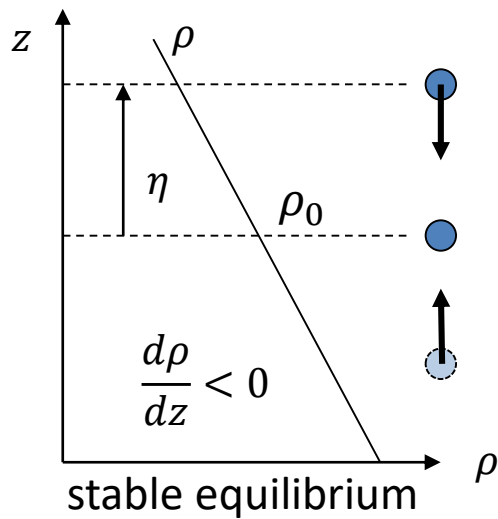
(buoyancy frequency)

solution: oscillating motion
around the equilibrium position

$$\eta = \eta_0 \cos(Nt)$$

oscillation period $T = \frac{2\pi}{N}$

Oscillations around stability: analogy



$$N = \sqrt{-\frac{g}{\rho} \frac{d\rho}{dz}}$$



David Brunt
(1886-1965)



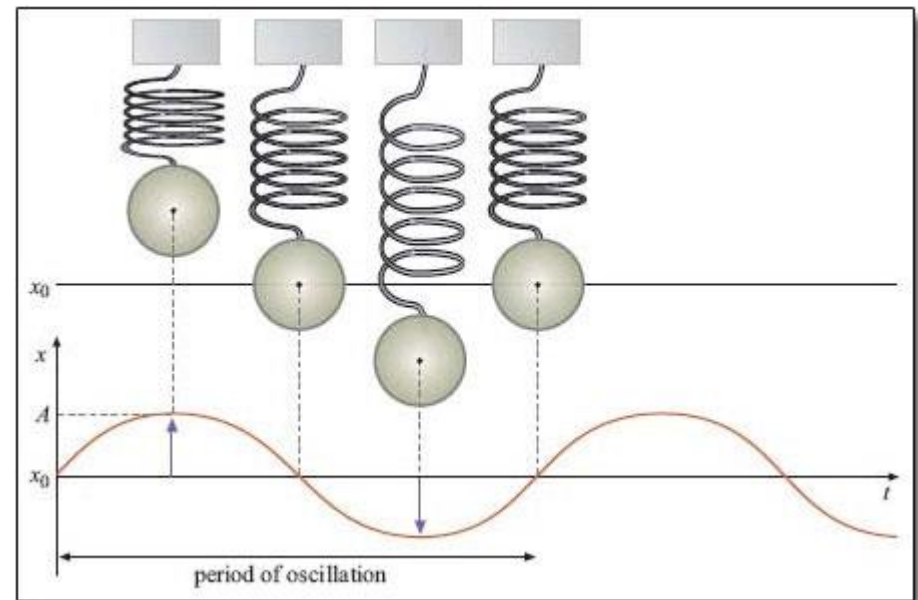
Vilho Väisälä
(1889-1969)

Brunt-Väisälä buoyancy frequency N [s^{-1}]
(physical interpretation: "spring oscillations")

$$\frac{d^2\eta}{dt^2} + N^2\eta = 0$$

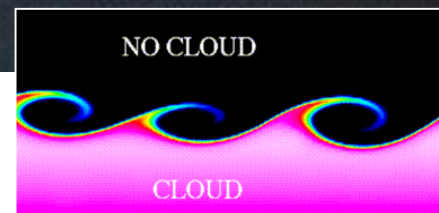
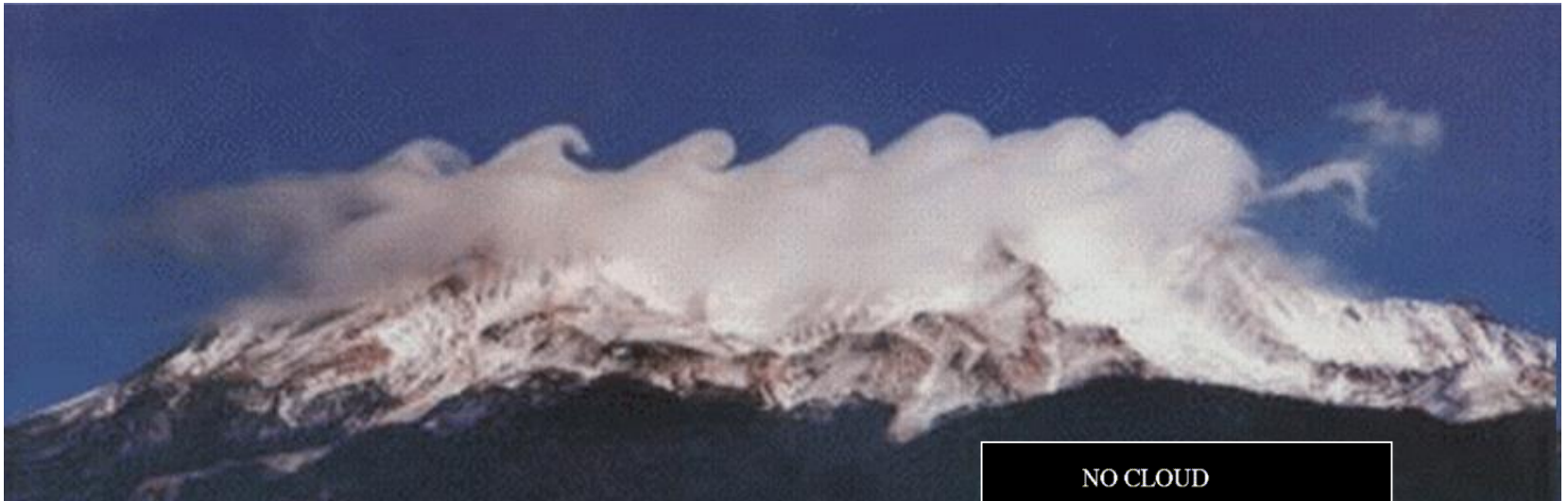
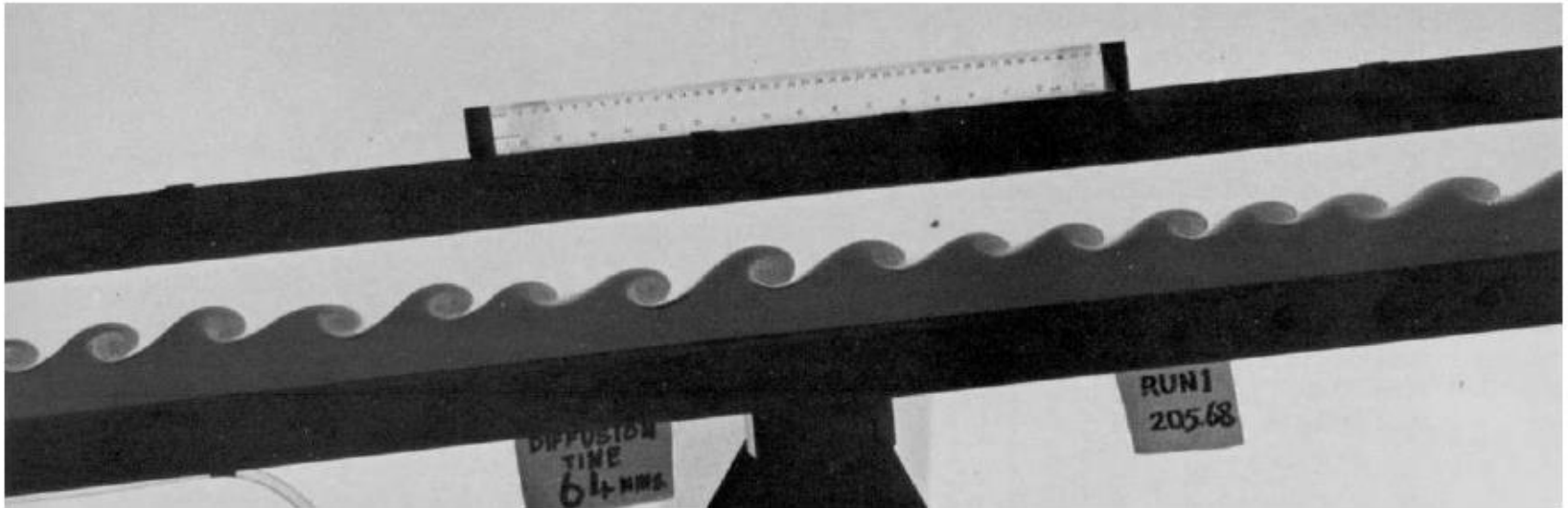
$$\eta = \eta_0 \cos(Nt)$$

$$T = \frac{2\pi}{N}$$



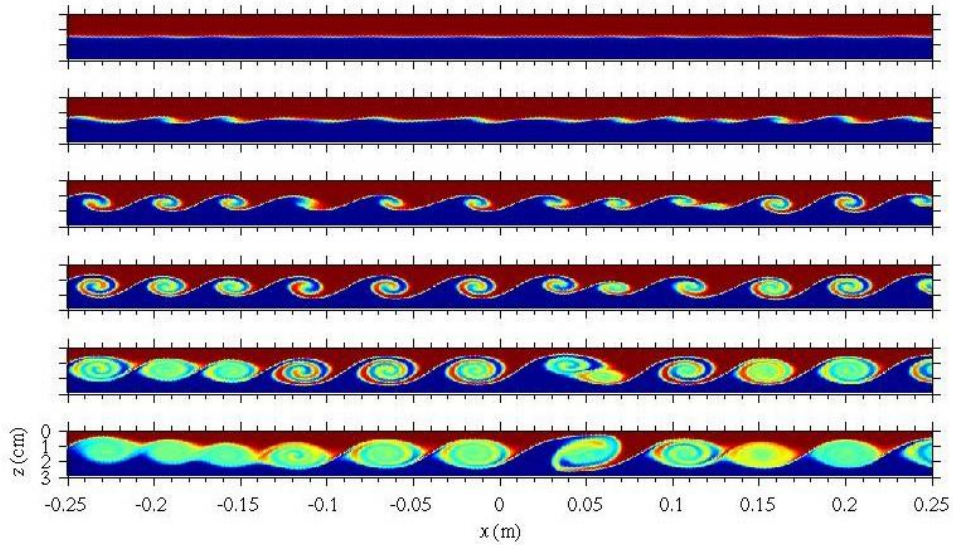
INSTABILITY ACROSS DENSITY INTERFACES

Examples of the Kelvin-Helmholtz instability





http://www.efluids.com/efluids/gallery/gallery_pages/cloud_instability_2.jsp



<http://www.physics.mun.ca/~danielb/research/aestus/aestus.html>

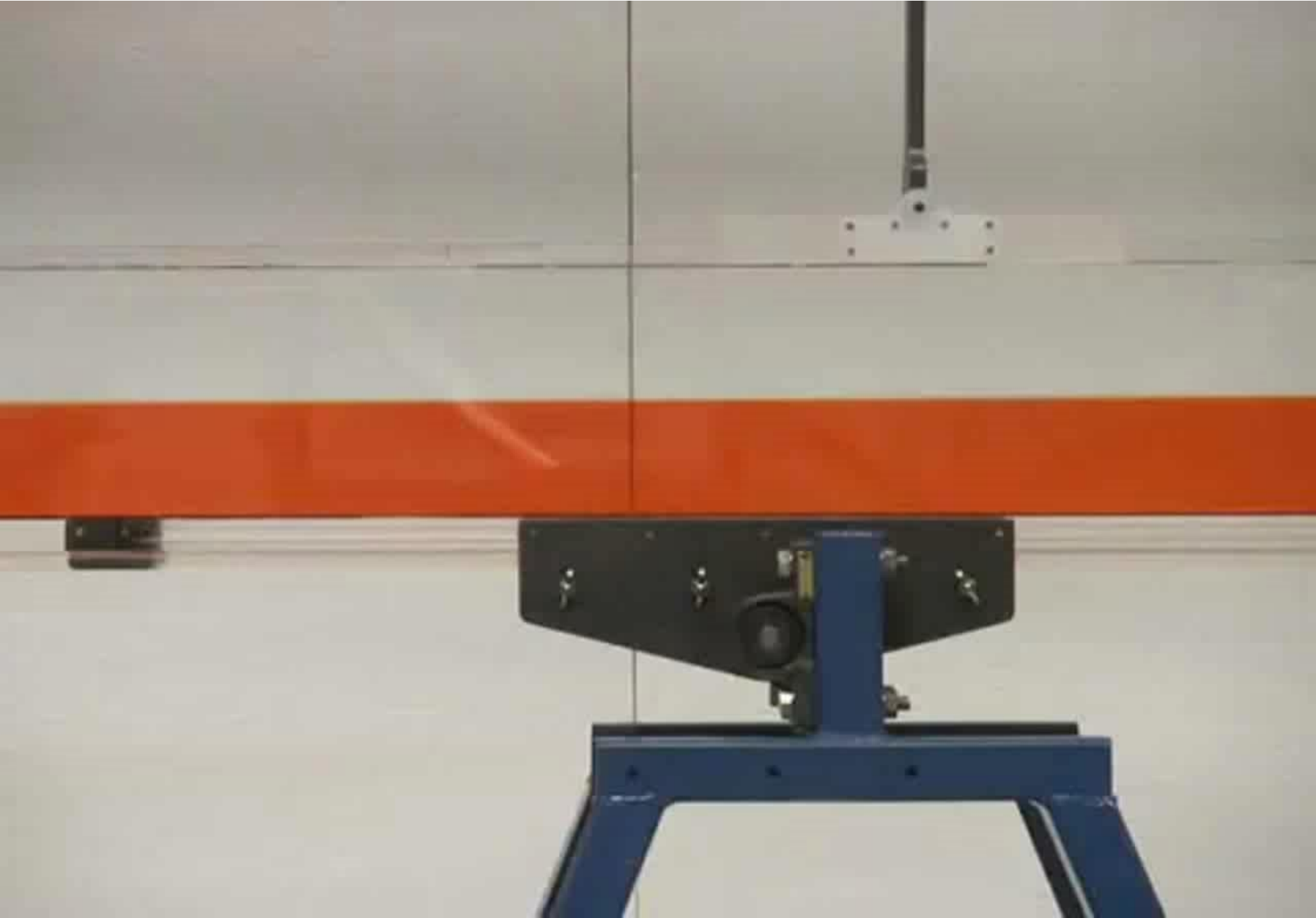


http://www.efluids.com/efluids/gallery/gallery_pages/kelvin_helm_page.jsp

Examples of the Kelvin-Helmholtz instability – in Trentino



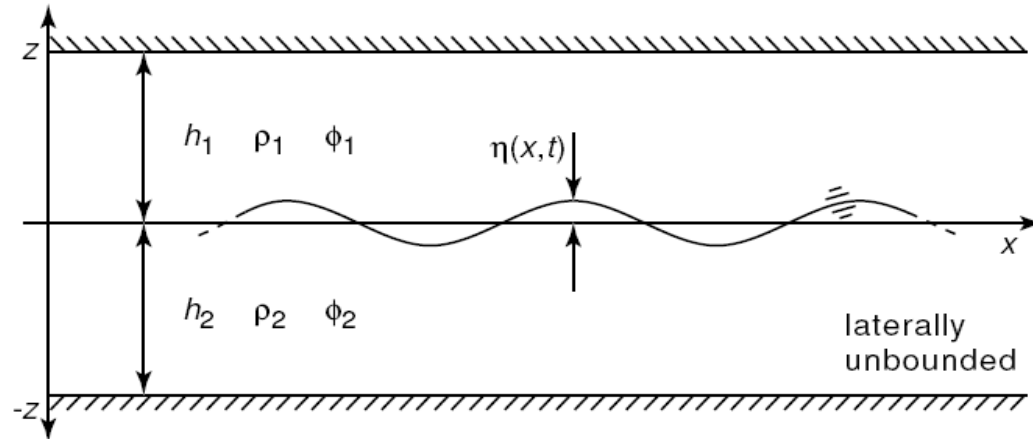
Photo: Sebastiano Piccolroaz



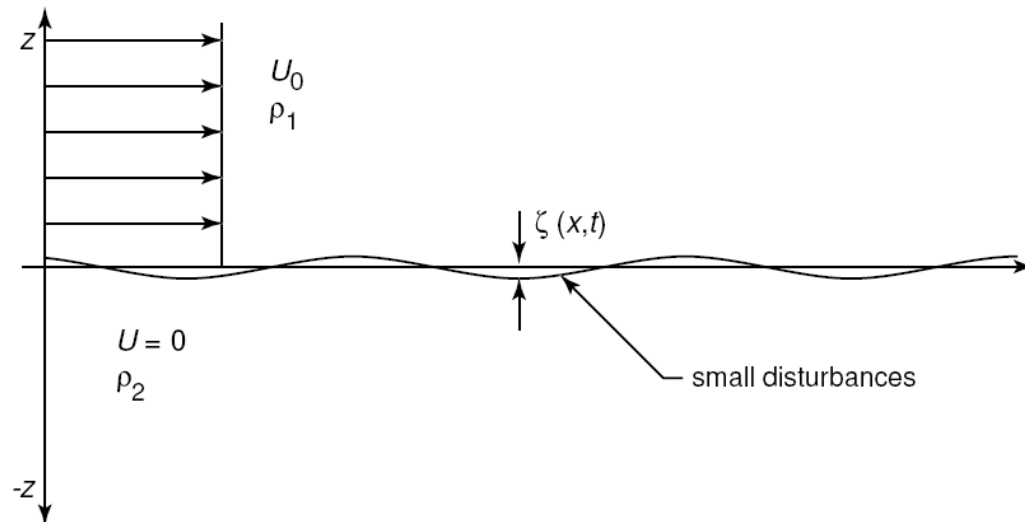
Instability in stratified flows: the Kelvin-Helmholtz instability

Hp: two layers with different densities; immiscible fluids, inviscid motion.

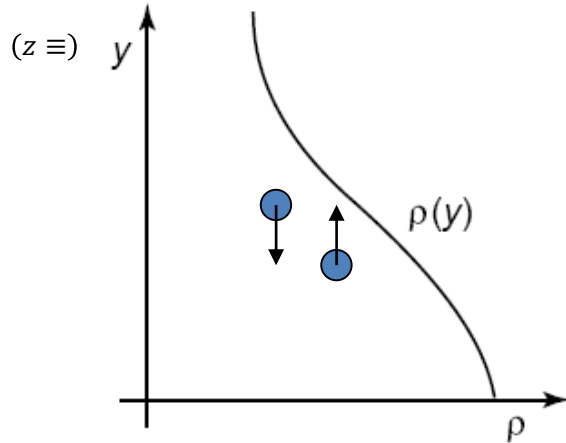
No overall motion:
interfacial waves



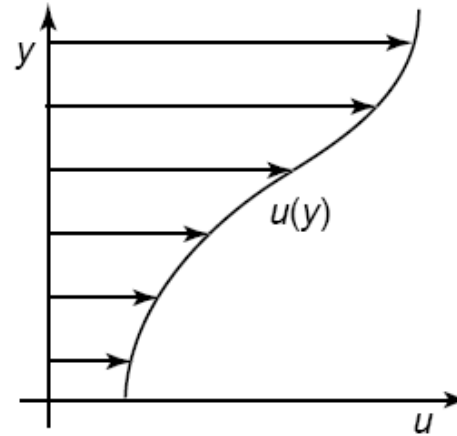
Relative motion
between layers:
disturbances
grow or decay?



Heuristic explanation



vertical gradient of density

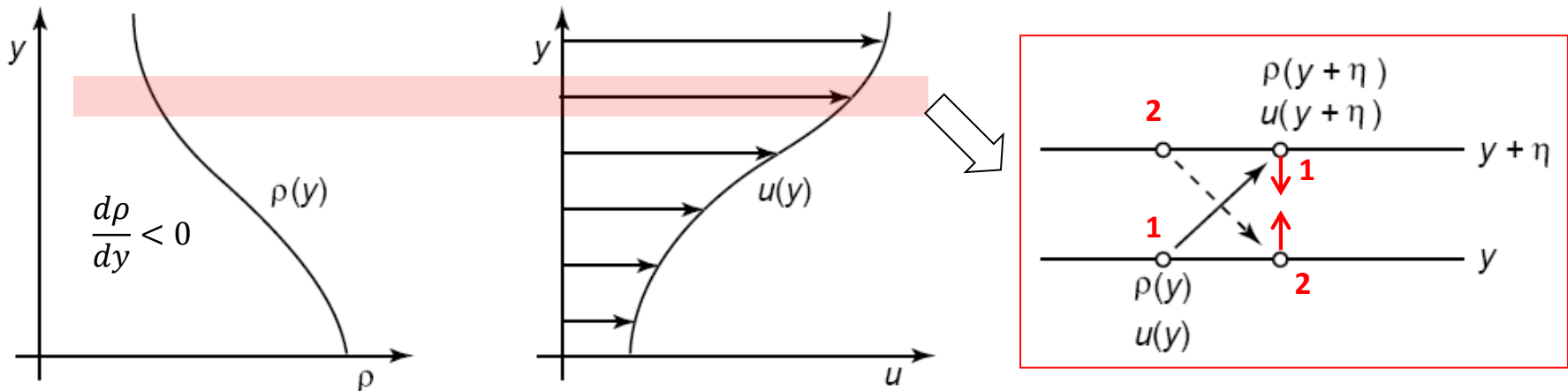


vertical gradient of velocity

instability: when the dissipated kinetic energy is larger than the work required by the buoyancy forces to move fluid particles

(without viscosity)

Mechanical work of the buoyancy forces



buoyancy force

$$\Delta B = -g\rho(y) + g\rho(y + \eta) = -g\rho(y) + g\left[\rho(y) + \frac{d\rho}{dy}\eta + O(\eta^2)\right] \cong g \frac{d\rho}{dy}\eta \quad \text{work}$$

water particle **1**

$$W_{B1} = \int_0^{\delta y} g \frac{d\rho}{dy} \eta d\eta = g \frac{d\rho}{dy} \frac{(\delta y)^2}{2}$$

$$\Delta B = -g\rho(y + \eta) + g\rho(y) \cong -g \frac{d\rho}{dy} \eta$$

water particle **2**

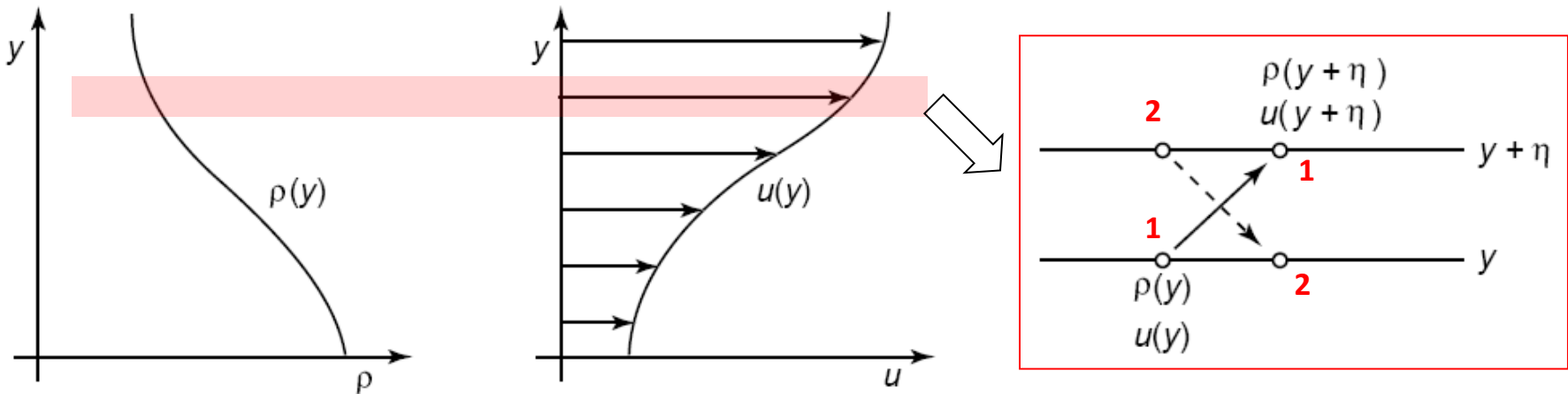
$$W_{B2} = \int_{\delta y}^0 -g \frac{d\rho}{dy} \eta d\eta = g \frac{d\rho}{dy} \frac{(\delta y)^2}{2}$$

total work

$$W_B = W_{B1} + W_{B2} = g \frac{d\rho}{dy} (\delta y)^2 < 0$$

(energy is required to exchange the particles)

Variation of kinetic energy



before $E(t_a) = \frac{\rho_0 u^2}{2} + \frac{\rho_0 (u + \delta u)^2}{2}$

(1) (2)

variation of kinetic energy

$$\Delta E = E(t_b) - E(t_a) = -\frac{\rho_0}{4} (\delta u)^2$$

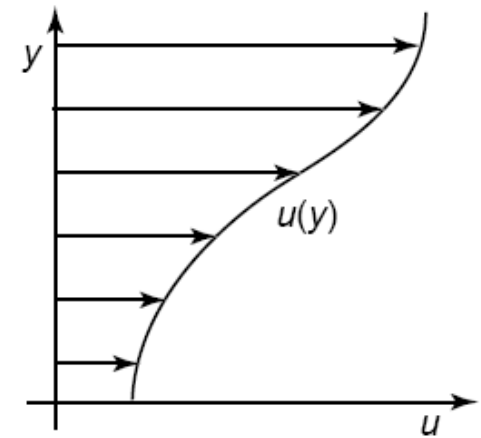
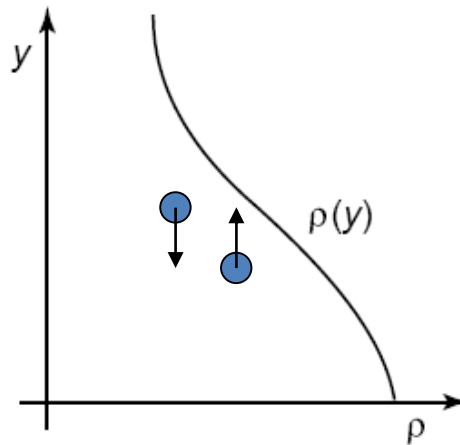
(after the exchange, the kinetic energy is smaller)

after $E(t_b) \cong 2 \frac{\rho_0}{2} \left(\frac{u + (u + \delta u)}{2} \right)^2$

hp. average velocity

Threshold for instability

instability: when the dissipated kinetic energy is larger than the work required by the buoyancy forces to move fluid particles



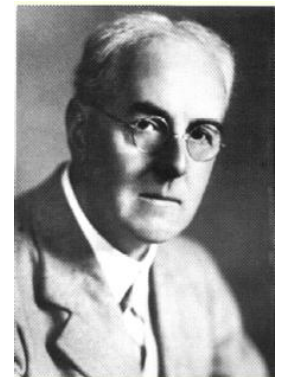
$$-\Delta E > -W_B$$

$$\frac{\rho_0}{4} (\delta u)^2 > -g \frac{d\rho}{dy} (\delta y)^2$$

(without viscosity)

Richardson number

$$\Rightarrow Ri = -\frac{g}{\rho_0} \frac{d\rho}{dy} \left(\frac{du}{dy} \right)^{-2} < \frac{1}{4}$$

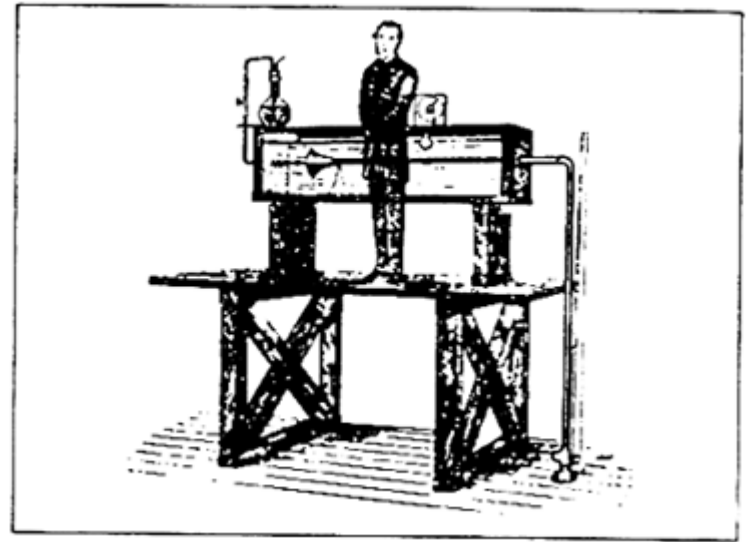
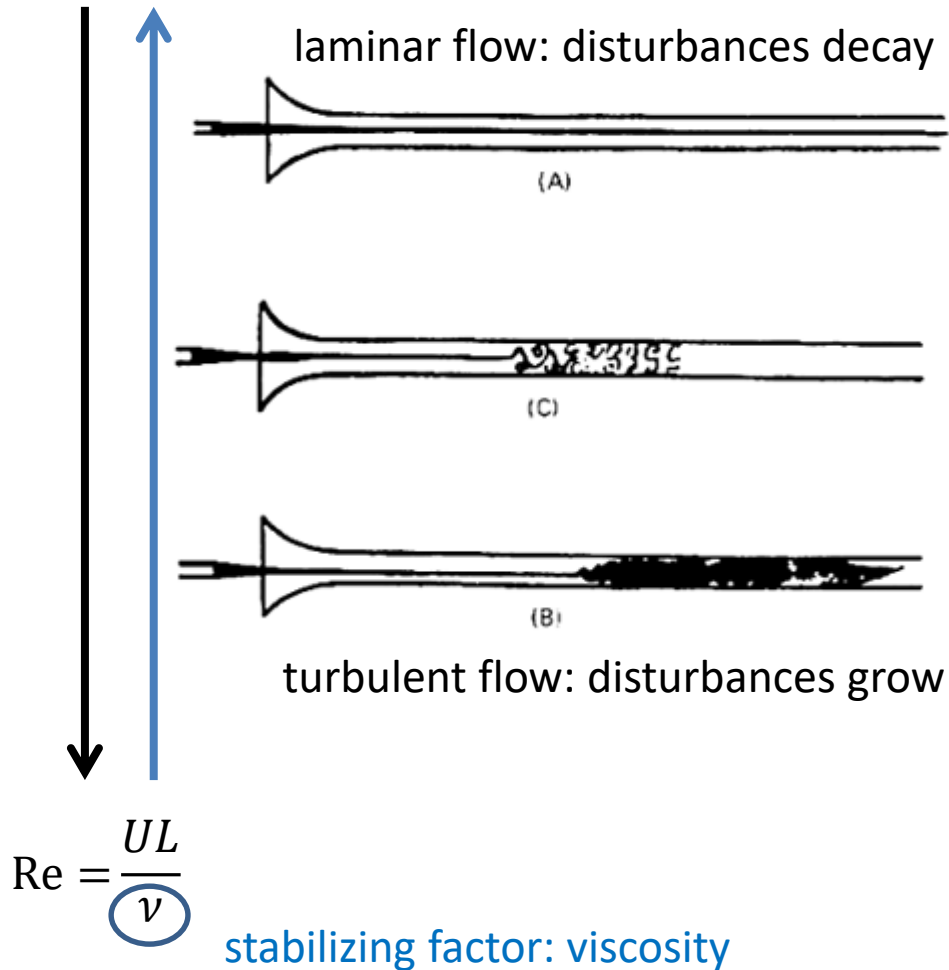


Lewis Fry Richardson
(1881-1953)

$$N^2 = -\frac{g}{\rho_0} \frac{d\rho}{dy} \quad \text{square of Brunt-Väisälä frequency [T}^{-2}\text{]}$$

Analogies with instability in turbulence

“Classical” turbulence:
Reynolds experiment



Stratified flows

stabilizing factor: density difference

$$Ri = \frac{-\frac{g}{\rho} \frac{d\rho}{dz}}{\left(\frac{du}{dz}\right)^2}$$

Relevant dimensionless number

Richardson number:

- gradient
$$Ri = N^2 \left(\frac{du}{dz} \right)^{-2} = \frac{-\frac{g}{\rho} \frac{d\rho}{dz}}{\left(\frac{du}{dz} \right)^2}$$

- “bulk”
$$Ri_b = \frac{g'H}{U^2} = \frac{1}{F_d^2}$$

$$g' = \frac{\rho'}{\rho_0} g$$

- flux
$$Ri_f = \frac{g \langle \tilde{\rho} \tilde{w} \rangle}{\bar{\rho} u_*^2 \frac{du}{dz}}$$

(turbulence) ratio between energy removal rate due to buoyancy forces and production due to shear

densimetric Froude number

$$F_d = \frac{U}{\sqrt{g'H}} = \frac{1}{\sqrt{Ri_b}}$$

Reynolds number

$$Re = \frac{UL}{\nu}$$

Prandtl number

$$Pr = \frac{\nu}{\kappa}$$

**TURBULENT MIXING
SUPPRESSION
IN STRATIFIED FLOWS**

Turbulence reduction due to stratification

Simple empirical laws

Effect of stratification
(depending on Richardson number)

$$D_{z, strat}^T = D_{z, 0}^T (1 + a Ri)^b$$

$$Ri = -\frac{g}{\rho} \frac{d\rho}{dz} \left(\frac{du}{dz} \right)^{-2}$$

$$Ri_b = -\frac{g \Delta \rho H}{\bar{\rho} \Delta U^2} \quad (\text{bulk})$$

Autore	a	b
(1) Munk and Anderson (1948)	10/3	-3/2
(2) Officer (1976)	1	-2
(3) Schiller and Sayre (1975)	10/3	-5/4

Some references:

Munk, WH, ER Anderson (1948), Notes on a theory of the thermocline, *J. Marine Res.*, 7:276–95.
Large, WG, JC McWilliams, SC Doney (1994), Oceanic vertical mixing: A review and a model with a nonlocal boundary layer parameterization, *Reviews of Geophysics*, 32:363–403.

Otherwise:

- specific single point second order **turbulence models** (k - ε , Mellor-Yamada, etc.)
- LES (Large Eddy Simulation)
- DNS (Direct Numerical Simulation) eddy-resolving methods

Examples of empirical laws

Pacanowski and Philander (JPO1981)

$$\nu_z = \nu_0(1 + \alpha Ri)^{-n} + \nu_{bg} \quad \text{viscosity}$$

$$D_z = \nu_z(1 + \alpha Ri)^{-1} + D_{bg} \quad \text{diffusivity}$$

background
viscosity/diffusivity

Large and Gent (JPO1999)

$$\nu_z = \nu_s + \nu_w + \nu_d$$

viscosity, diffusivity
(for velocity, temperature,
salinity, etc.)

resolved
vertical
shear

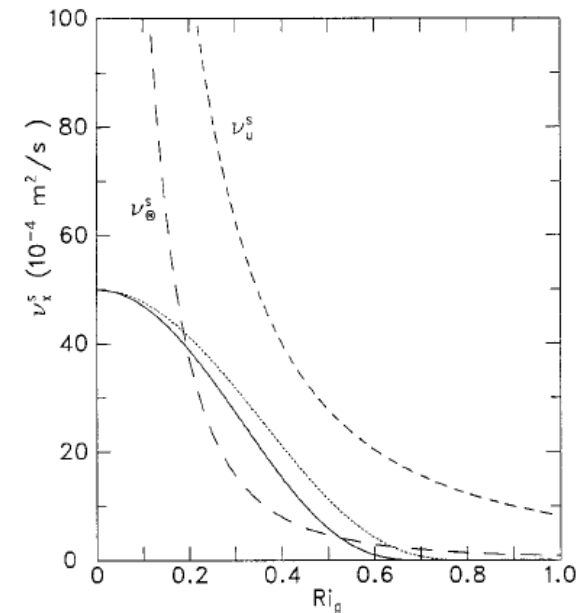
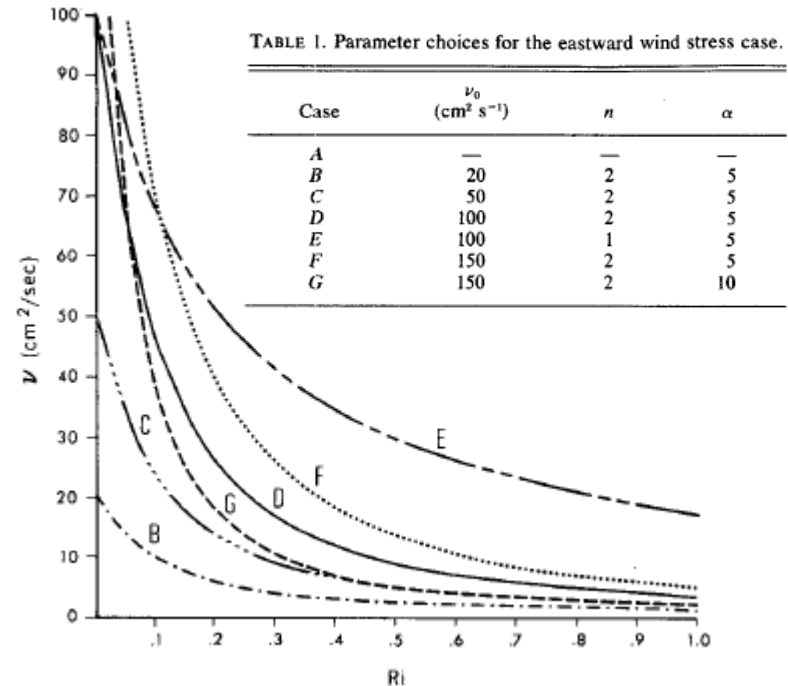
unresolved
internal
wave
breaking

double
diffusion

$$\nu_{w,u} = 10^{-4} \text{ m}^2 \text{ s}^{-1}$$

$$\nu_{w,T} = 10^{-5} \text{ m}^2 \text{ s}^{-1}$$

$$\frac{\nu_s}{\nu_0} = \begin{cases} 1 & Ri < 0 \\ \left[1 - \left(\frac{Ri}{Ri_0} \right)^2 \right]^3 & 0 < Ri < Ri_0 \\ 0 & Ri > Ri_0 \end{cases}$$



Estimating the parameters

Elements for the dynamic reconstruction of the reduced vertical eddy diffusivity:

- reference viscosity/diffusivity (possibly vertical profile)
- background viscosity/diffusivity (unresolved processes)
- gradient Richardson number:
 - buoyancy frequency (depends on temperature profile)
 - shear frequency (depends on flow field)

$$\left. \begin{array}{l} \bullet \text{ buoyancy frequency (depends on temperature profile)} \\ \bullet \text{ shear frequency (depends on flow field)} \end{array} \right\} Ri = \frac{N^2}{S^2}$$

$$N^2 = -\frac{g}{\rho} \frac{\partial \rho}{\partial z} = \beta g \frac{\partial T}{\partial z}$$

β thermal expansibility of water

$$\rho = \rho_o[1 - \beta(T - T_o)]$$

$$S^2 = \left(\frac{\partial U}{\partial z} \right)^2 + \left(\frac{\partial V}{\partial z} \right)^2$$

Estimates for models that do not resolve hydrodynamics

$$S^2 = S_{lw}^2 + S_{iw}^2 + S_{bg}^2$$

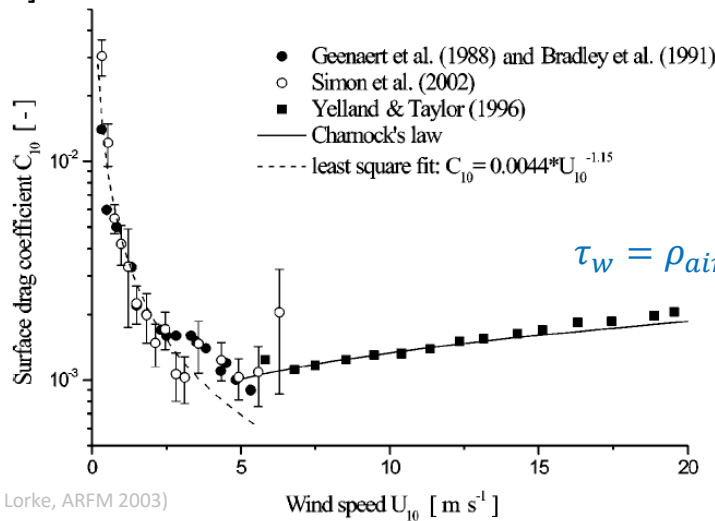
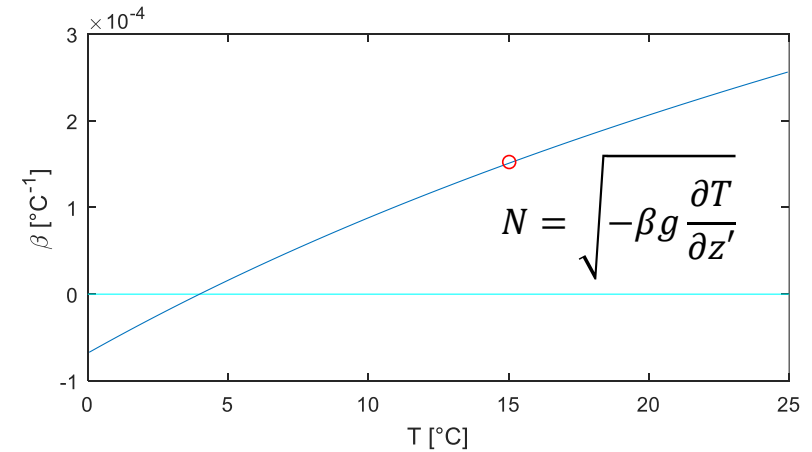
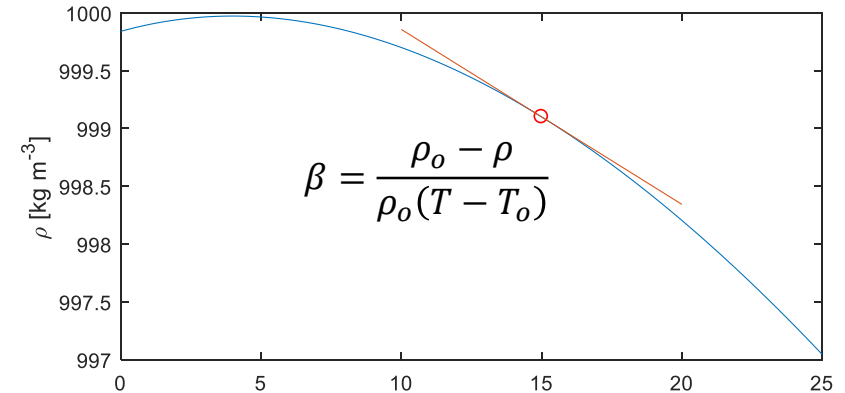
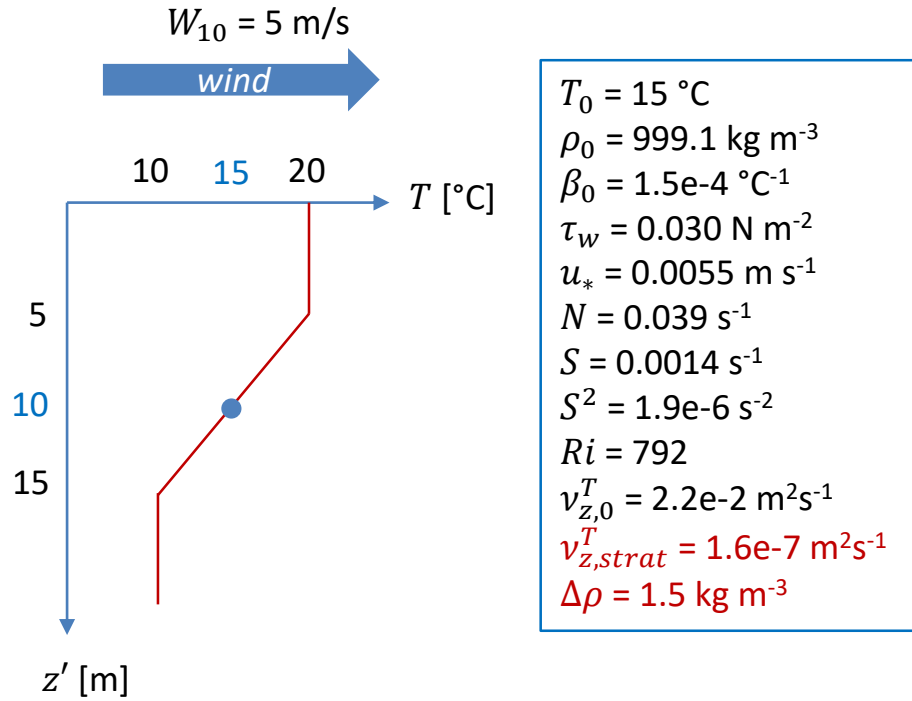
wind stress (τ_w) contribution
~ law of the wall

$$S_{lw} = \frac{\partial U_0}{\partial z} \sim \frac{u_*}{\kappa z} = \frac{\sqrt{\tau_w / \rho}}{\kappa z}$$

background
 $2 \cdot 10^{-7} \text{ m}^2 \text{ s}^{-1}$ (Lawrence et al., 2002)

internal wave contribution
 $\gamma = 0.7$ (Mellor et al., 1989)

An example of an estimate of Ri for a wind-driven flow in a lake



$$\tau_w = \rho_{air} C_D W_{10}^2$$

$$u_* = \sqrt{\tau_w / \rho}$$

$$S \cong \frac{u_*}{\kappa z}$$

$$v_{z, strat}^T = v_{z, 0}^T \left(1 + \frac{10}{3} Ri \right)^{-3/2}$$

$$Ri = \frac{N^2}{S^2}$$