SIO 209: Signal Processing for Ocean Sciences Class 9

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Normalization Factors for Broadband Power Spectra

• Power Spectral Density

$$P_{xx}(\omega) = \sum_{m=-\infty}^{\infty} \phi_{xx}(m) e^{-j\omega m}$$

Power is obtained by integrating the power spectral density over all frequencies

$$\mu_x^2 + \sigma_x^2 = \phi_{xx}(0) = \frac{1}{2\pi} \int_{-\pi}^{\pi} P_{xx}(\omega) d\omega = \int_{-0.5}^{0.5} P_{xx}(f) df$$

see also Sections 10.2 and 10.5 in *Oppenheim & Schafer, 2009*

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Normalization Factors for Broadband Power Spectra

· Welch's method of windowed and overlapped FFTs

$$\hat{P}_{xx}^{L}(f_k) = \frac{1}{MUf_c} \overline{\left| X(k) \right|^2}$$

averaged over ${\cal L}$ potentially overlapping and windowed records

in "Power"/Hz

$$U = \frac{1}{M} \sum_{n=0}^{M-1} w^2[n]$$

normalization constant

$$X^{(l)}(k) = \sum_{n=0}^{M-1} w[n] x^{(l)}[n] e^{-j\frac{2\pi}{M}nk}$$
$$l = 1, \dots, L$$

 $\,M\,$ FFT length

 $f_k = k/M \;\; {
m frequency} \; {
m at} \; {
m FFT} \; {
m bin} \; {
m with} \;\; {
m index} \; k$

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Normalization Factors for Broadband Power Spectra

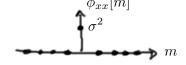
• Recall periodogram

$$I_N(\omega) = \frac{1}{N} |X(e^{j\omega})|^2 = \sum_{m=-(N-1)}^{N-1} c_{xx}[m]e^{-j\omega m}$$

$$E[I_N(\omega)] = \sum_{m=-(N-1)}^{N-1} E\{c_{xx}[m]\}e^{-j\omega m} = \sum_{m=-(N-1)}^{N-1} \frac{N-|m|}{N} \phi_{xx}[m]e^{-j\omega m}$$

- Consider uncorrelated (and thus IID) zero-mean Gaussian noise, i.e., $x[n] \sim \mathcal{N}(0,\sigma_x^2)$

$$\phi_{xx}[m] = \begin{cases} \sigma^2 & m = 0\\ 0 & m \neq 0 \end{cases}$$



$$\mathrm{E}\big[I_N(\omega)\big] = \sigma_x^2$$

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Normalization Factors for Broadband Power Spectra

• Consider the absolute value squared of the Fourier transform

$$X(e^{j\omega}) = \sum_{n=0}^{N-1} x[n]e^{-j\omega n}$$

$$|X(e^{j\omega})|^2 = X(e^{j\omega})X^*(e^{j\omega}) = \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} x[n]x[m]e^{-j\omega n}e^{j\omega m}$$

• Assuming uncorrelated zero-mean Gaussian noise, we obtain

If window w[n] had been used, this factor would be $\sum_{n=0}^{N-1} w^2[n]$

$$\mathbb{E}\{|X(e^{j\omega})|^2\} = \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} \mathbb{E}\{x[n]x[m]\}e^{-j\omega n}e^{j\omega m} = \sum_{n=0}^{N-1} \mathbb{E}\{x[n]^2\} = \sum_{n=0}^{N-1} \sigma^2 = N\sigma^2$$

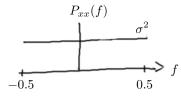
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Normalization Factors for Broadband Power Spectra

• For zero-mean Gaussian noise, we have the power spectral density

$$P_{xx}(f) = \sum_{m=-\infty}^{\infty} \phi_{xx}(m) e^{-j2\pi fm} = \phi_{xx}(0) = \sigma_x^2$$



and the corresponding power

$$\int_{-0.5}^{0.5} P_{xx}(f) \, \mathrm{d}f = \int_{-0.5}^{0.5} \sigma_x^2 \, \mathrm{d}f = \sigma_x^2$$

- Note that f is in cycles per sample or, equivalently, $f_{
m s}=1$

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Normalization Factors for Broadband Power Spectra

• In case $f_{
m s}
eq 1$, we have

$$\frac{1}{f_{s}} \int_{-f_{s}/2}^{f_{s}/2} P_{xx}(f) \, \mathrm{d}f = \frac{1}{f_{s}} \int_{-f_{s}/2}^{f_{s}/2} \sigma_{x}^{2} \, \mathrm{d}f = \sigma_{x}^{2}$$

 $P_{xx}(f)$ σ^2/f_s $-f_s/2$ $f_s/2$

since total power remains unchanged

• This motivates the normalization constant used for power spectral density estimation (see second page), e.g., for L=1 and N=M, we have

$$\hat{P}_{xx}(f_k) = \frac{1}{NUf_s} \left| X(k) \right|^2$$

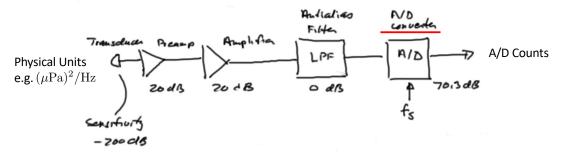
 ${\cal U}$ is the normalization that corrects for the distortion (tapering) of the segment time series by the window function

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Calibration of Power Spectral Density

System Block Diagram



Typical A/D Parameters:

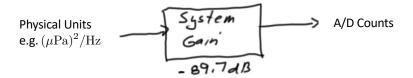
A/D Gain: $20 \log \frac{65536}{20} = 70.3 dB$

System Gain: -200dB + 20dB + 20dB + 70.3dB = -89.7dB

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Calibration of Power Spectral Density

• For calibration purposes the system block diagram can be summarized as follows



• Thus, to go from $P_{xx}(f)$ in $(A/D \ counts)^2/Hz$ to $P_{xx}(f)$ in $(\mu Pa)^2/Hz$, 89.7dB must be added

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