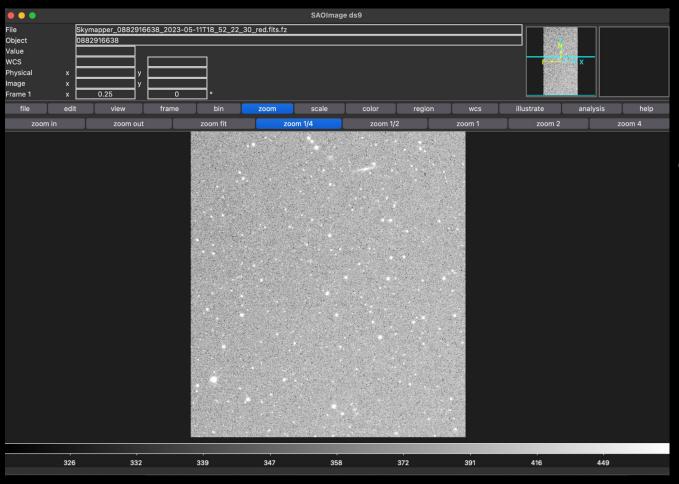
Research Updates Summer 2024

Maryann Benny Fernandes

Week 2



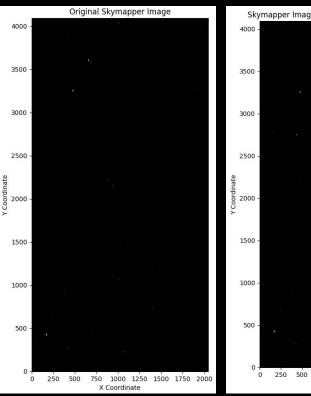
Color: Gray

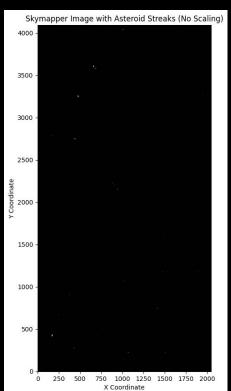
Scale: asinh, zscale

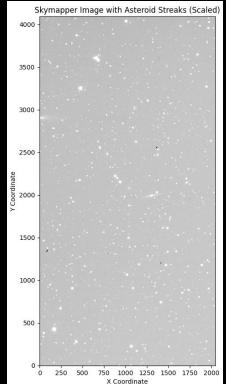
- Asinh: enhances faint and bright features simultaneously
 - zscale: Automatic contract adjustment, decreases background dust

Color: Gray, Scale: zscale, asinh

Adding asteroid streak and scaling



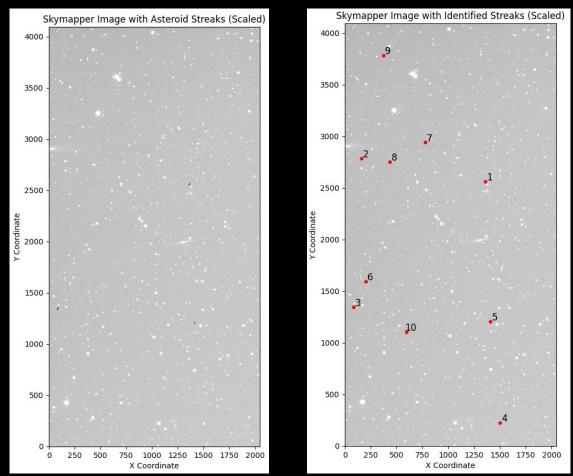




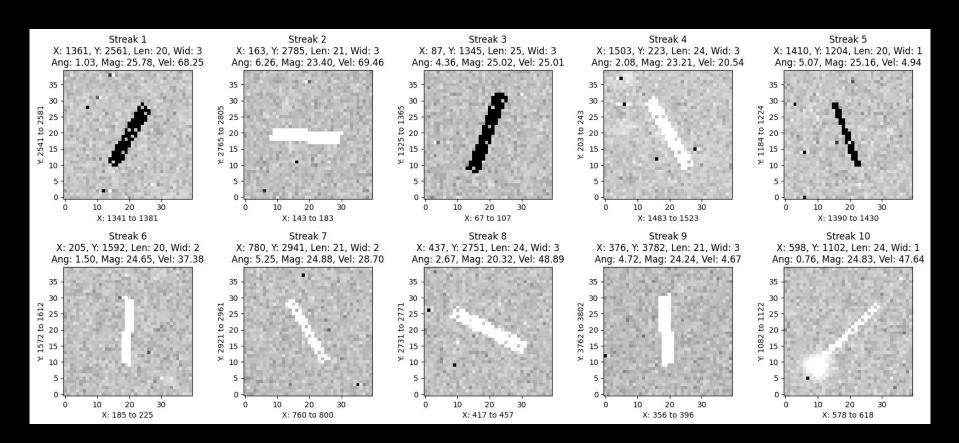
Asteroid streaks:

- 1) Random orientation (0° to 360° (clockwise from east))
- 2) Random location
- Random magnitudes from 20-26
- 4) Random velocity around 1 to 80 arcsec^-1
- 5) Random width and length (10-20 pixels)
- 6) Brightness variation not accounted

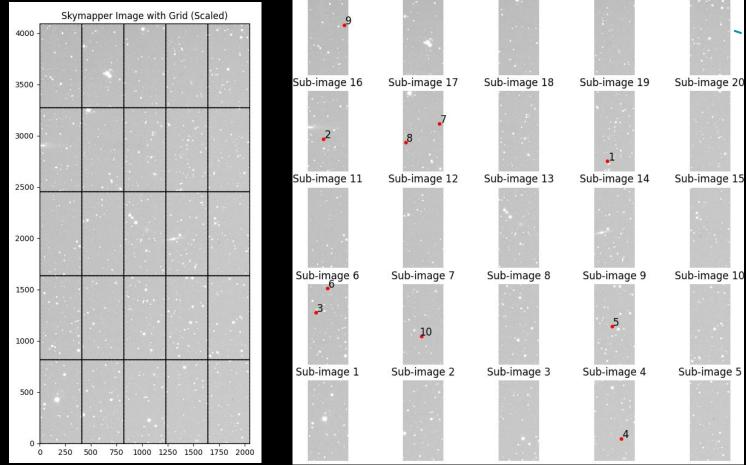
Identifying asteroid streaks



Zooming into features of asteroid streaks



Creating Sub-images and cutouts



Sub-image 21

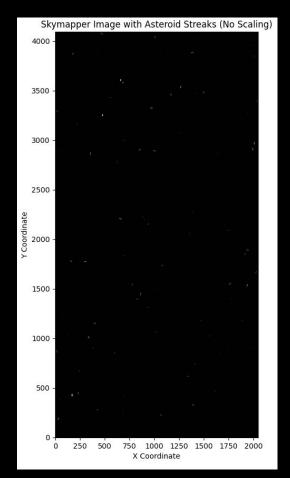
Sub-image 22

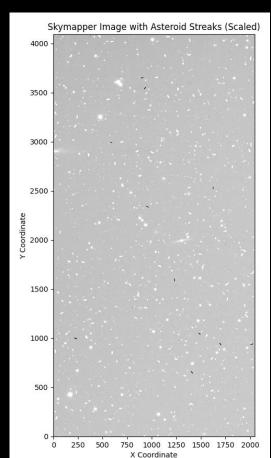
Sub-image 23

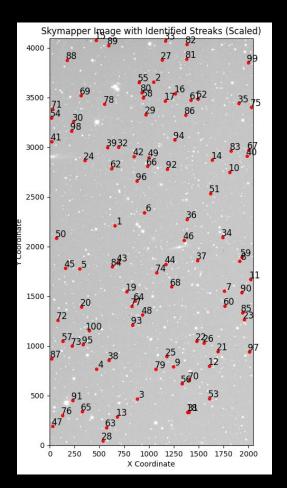
Sub-image 24

Sub-image 25

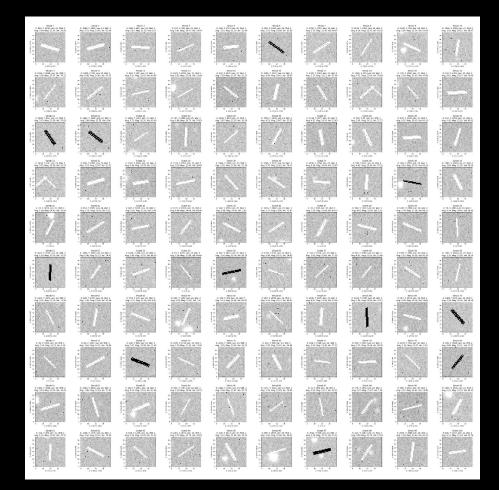
Simulating and Identifying 100 asteroid streaks

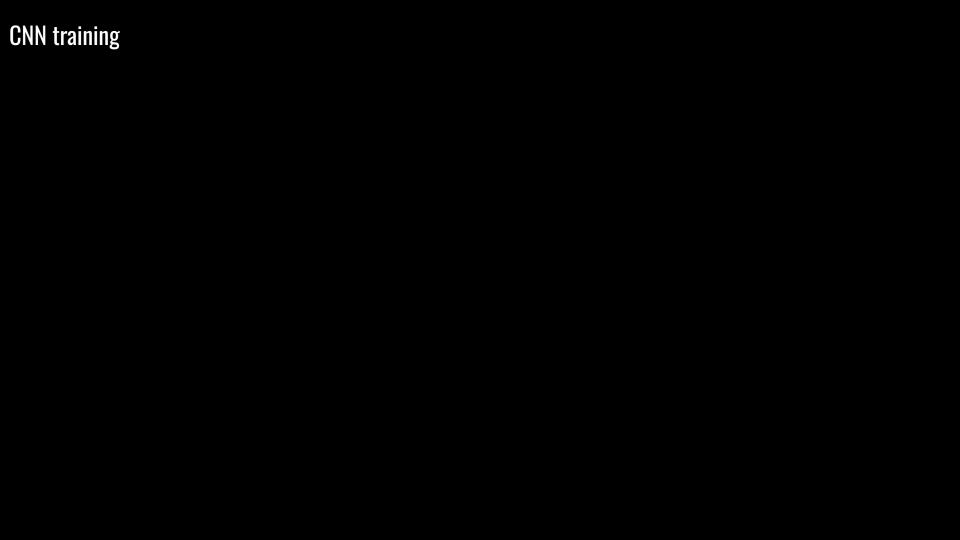


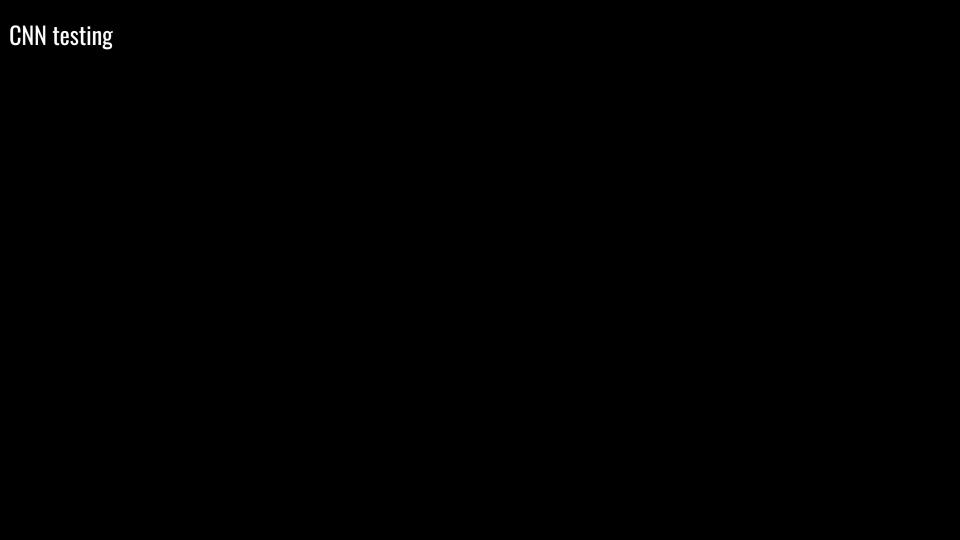




Zooming into features of asteroid streaks







Week 1

Week 1: July 19 - 29, 2024

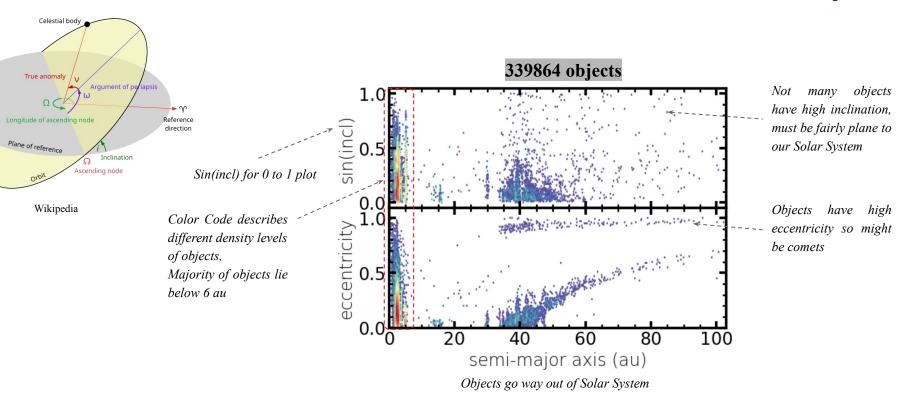
- Received access to Rubin Science Platform, Isst.data.cloud
 - Studied the notebook tutorial on 'Main Belt Asteroids'
- Studied the asteroid simulation section from the below paper
 - Euclid: Identification of asteroid streaks in simulated images using deep learning

 (DL pipeline still need to study)
- ☐ Multivariable calculus research assignment and Touch the Art final project on Simulated Asteroids

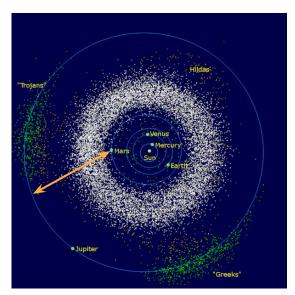
Week 2: Goals

- Explore the simulated data set on cloud and compare 'Main Belt Asteroids' tutorial with the following by also using the NEA tutorial:
 - New properties of other families in the Main Belt
 - Study objects >1.6 au to see if they make a close approach to Earth or Mars
- Euclid: Identification of asteroid streaks in simulated images using deep learning
 - Understand DL pipeline

Orbital Parameters of SSO's observed \geq 50 times in 10 years



Main Belt Asteroids (1.6 < a < 4.2 au, q > 1.6 au)



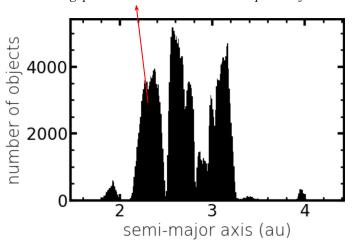
The asteroids of the inner Solar System and ^{6-J} Jupiter: The belt is located between the orbits of Jupiter and Mars.



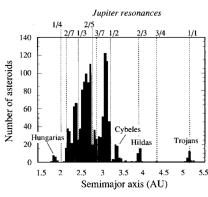
Wikipedia

Not a uniform Main Belt Asteroid

Kirkwood gaps: Resonances between orbital period of asteroids and Jupiter



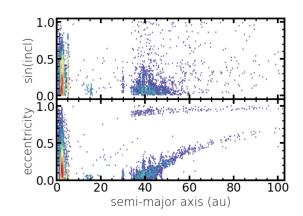
Hildas, Cybeles, and Hungarias are clearly visible in the DP0.3 dataset

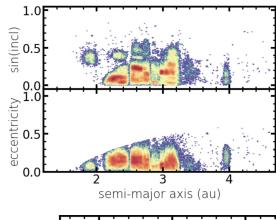


http://burro.cwru.edu/Academics/Astr221/Gravity/resonance.html

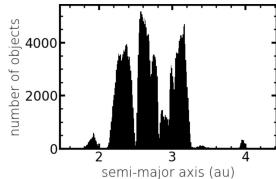
Asteroid Properties in Solar System

All objects in DP0.3 observed \geq 50 times in 10 years

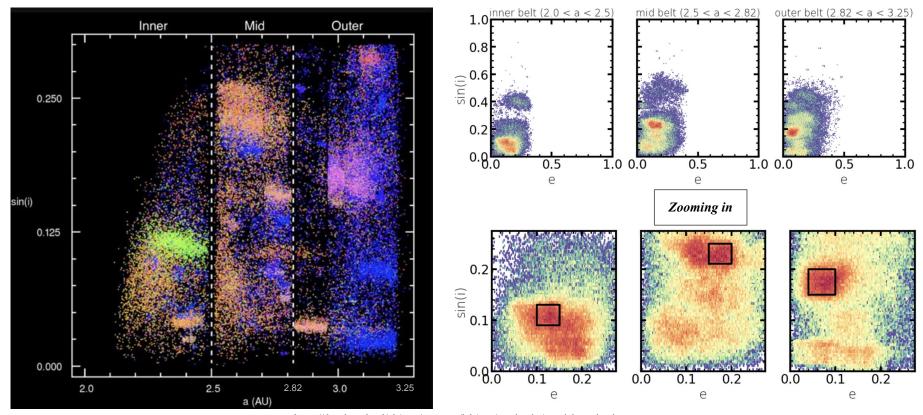




All objects in DP0.3 observed ≥ 50 times in 10 years in the Main Asteroid Belt

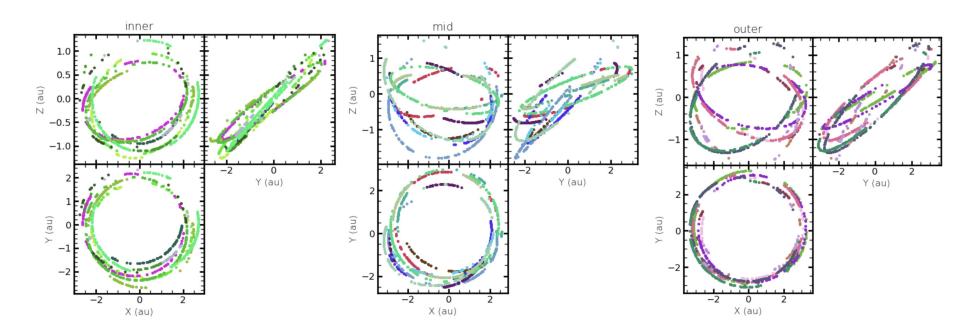


Orbit families in Main Belt Asteroids



https://data.lsst.cloud/nb/user/maryann/lab/tree/notebooks/tutorial-notebooks

Orbit eccentricity and inclination in Main Belt Asteroids



Objects in close proximity to Earth

- Earth's eccentricity: **0.0167**
- Earth Perihelion distance (closest) to Sun: **0.9832899** AU
- Earth Aphelion distance (farthest) to Sun: 1.0167103 AU

Objects close-contact with Earth?

