

## Fulvio Tessarotto ( I.N.F.N. – Trieste )

Introduction

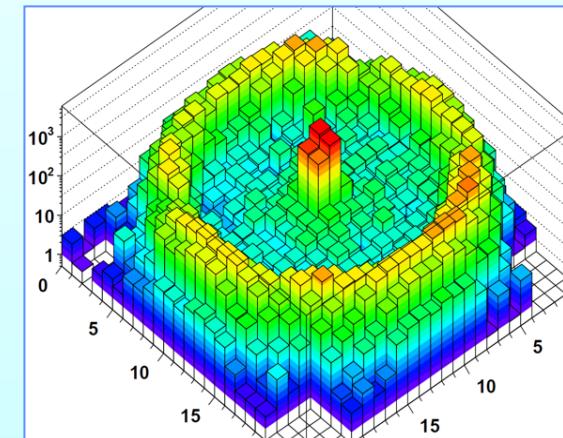
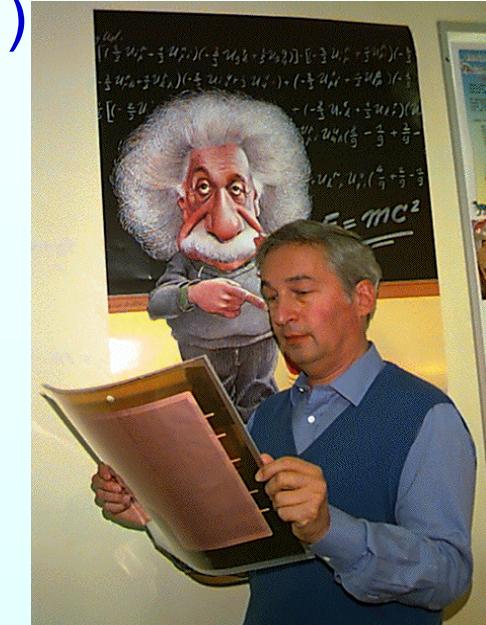
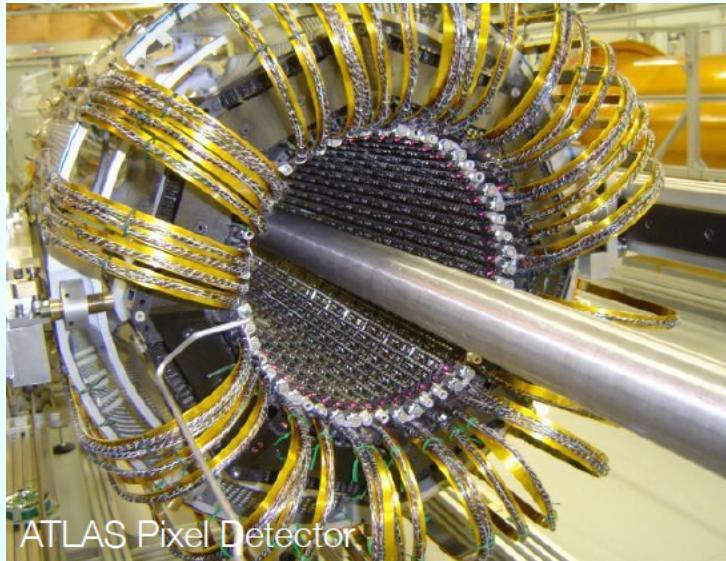
particle detectors history

CERN

Detection technologies

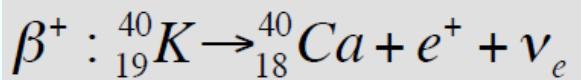
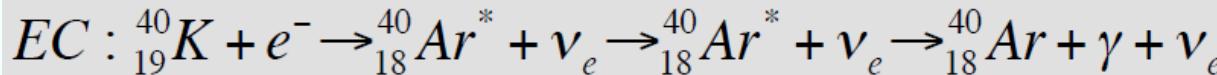
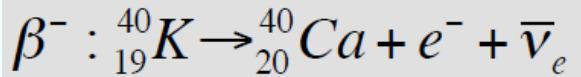
Gaseous Detectors

Particle Identification



# Radioactivity is everywhere

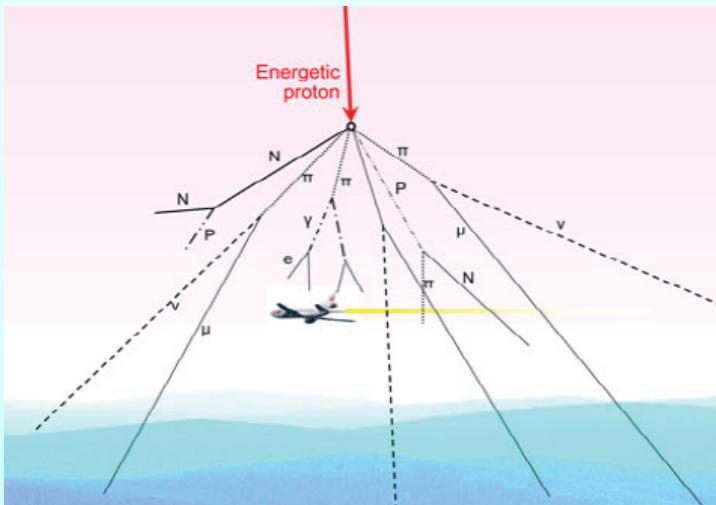
*On average a human body has ~30 mg of Potassium 40, corresponding to ~ 4 kBq*



**My activity is ~ 5000  
disintegrations per second**

*An average banana has 14 Bq*

*The average annual human exposure to natural background radiation is 2.4 mSv  
(from 0.4 to 4 mSv/y depending on the location)*



*During a flight the background dose rate increases by a factor between 10 and 30*

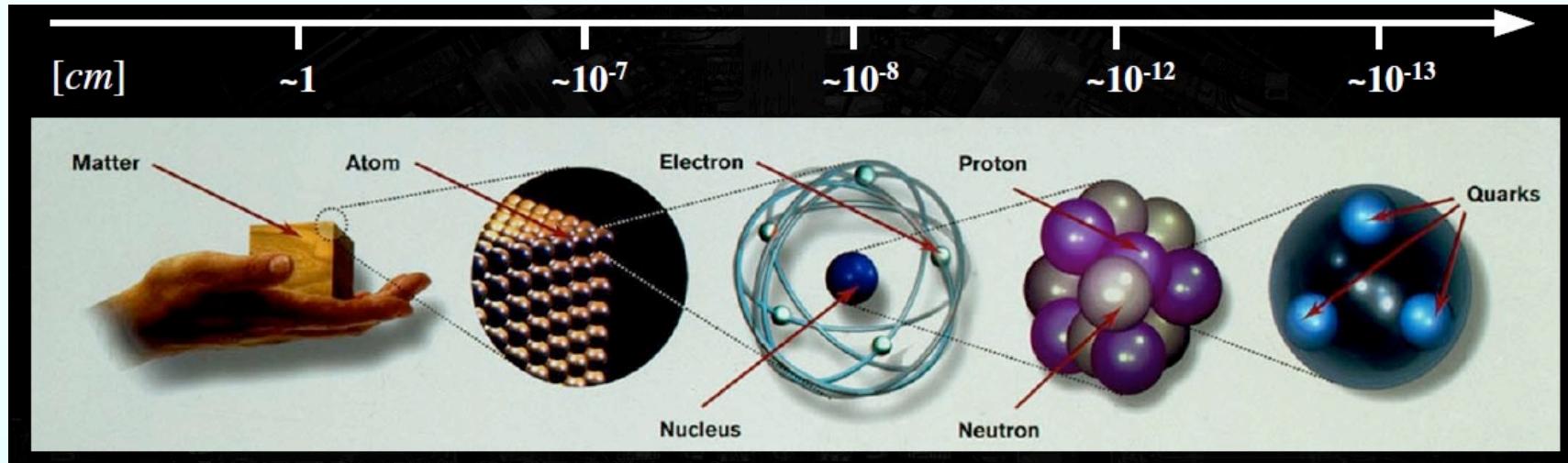
**Charged particles:**

- heavy ( $\mu$ ,  $\pi$ ,  $p$ ,  $a$ , ...)
- light (electrons and positrons)

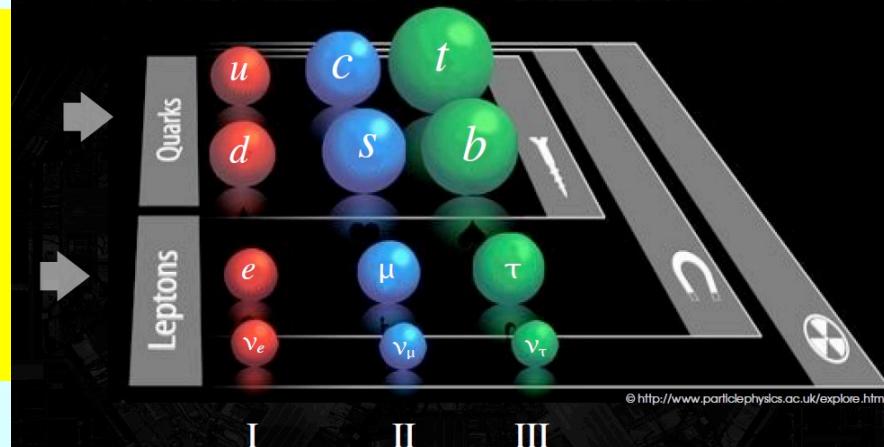
**Neutral particles:**

- photons
- neutrons
- neutrinos

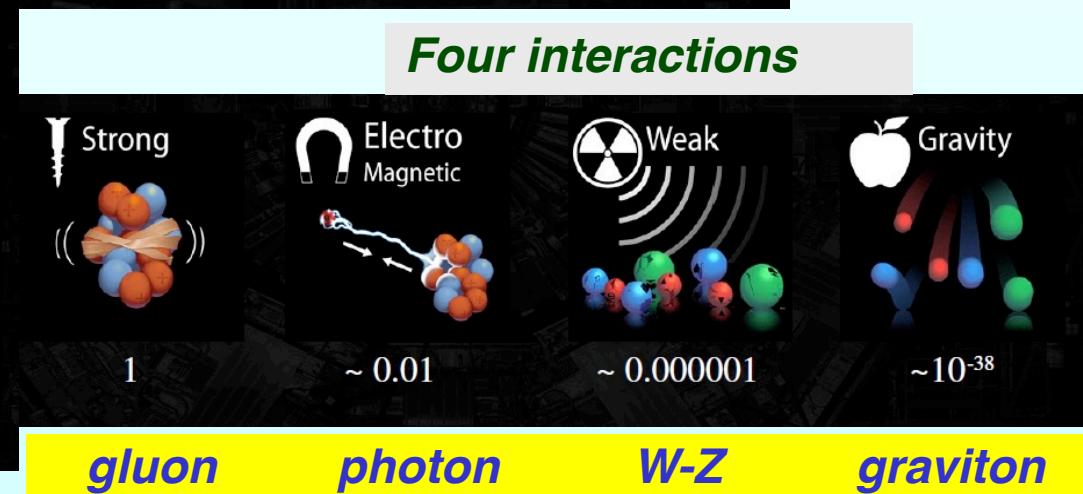
# The constituents of matter



**Two families**



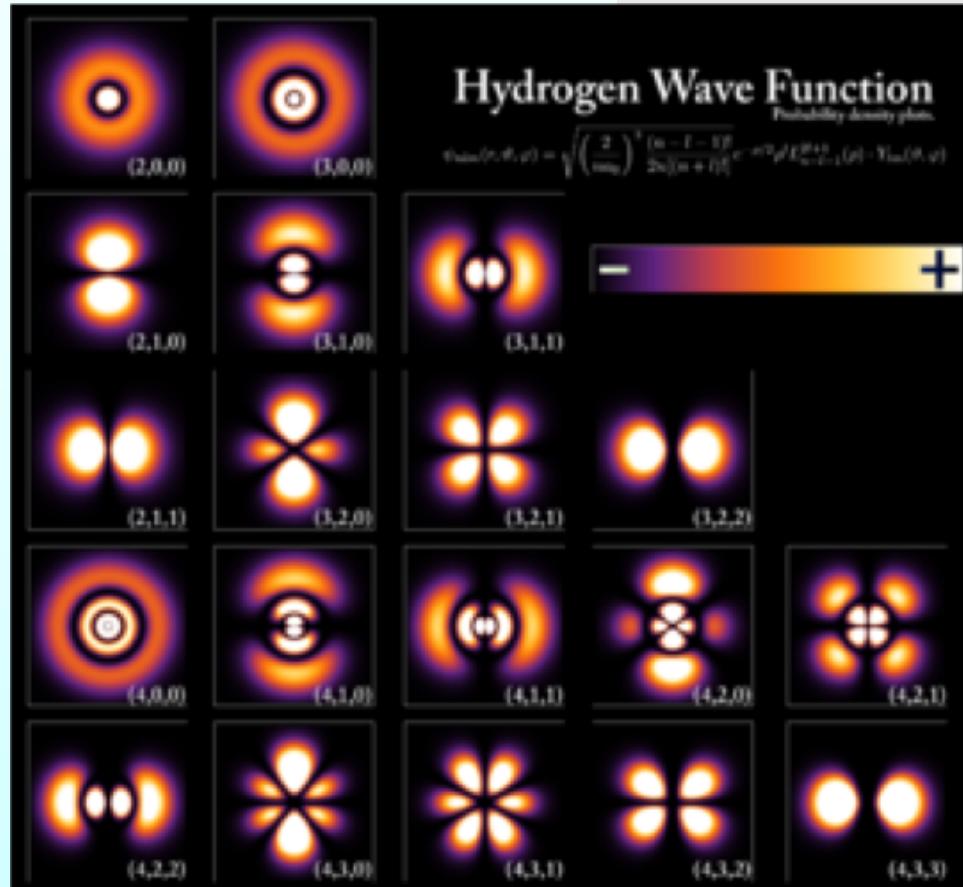
**Three generations**



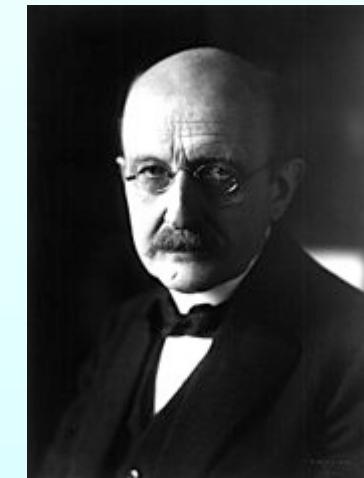
# Particles do what we cannot

**The laws which hold at the microscopic level are different from ours**

*A particle or an atom can stay at the same time in two or more different places*



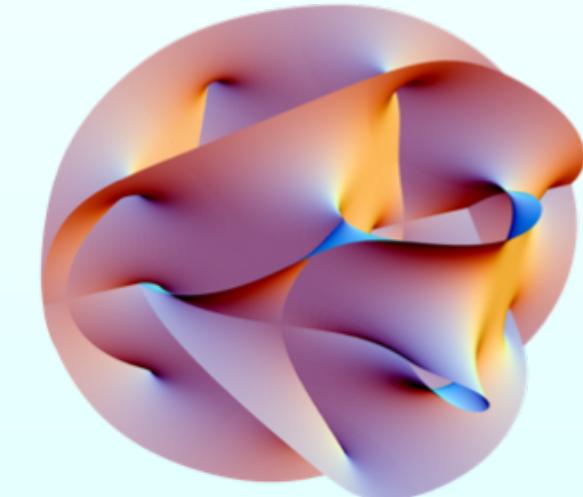
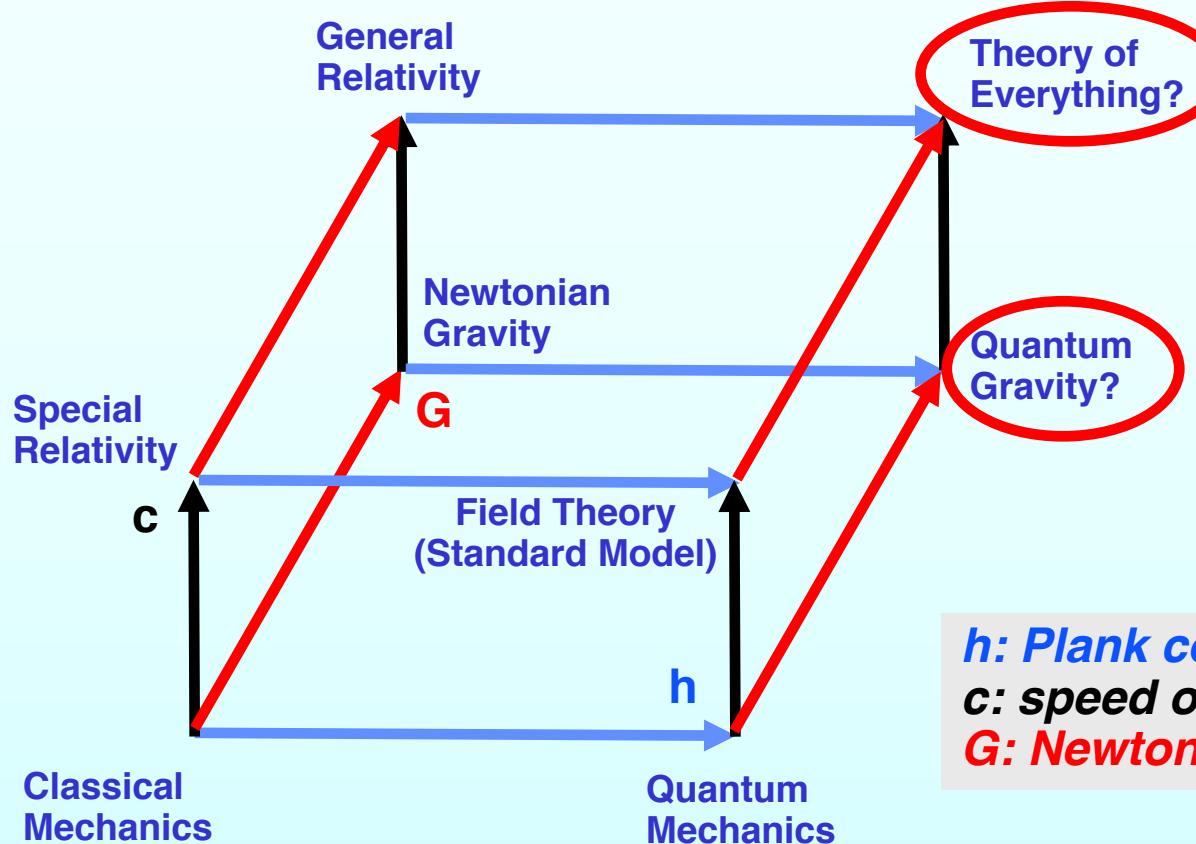
*A particle can move from a point to another in space without passing anywhere in between the two points*



**Max Planck is the father of the quantum theory**

# Theoretical panorama

## *The fundamental constants of nature and unified theories*



*h: Plank constant  
c: speed of light  
G: Newton's constant*

# The Standard Model

## Standard Model of FUNDAMENTAL PARTICLES AND INTERACTIONS

The Standard Model summarizes the current knowledge in Particle Physics. It is the quantum theory that includes the theory of strong interactions (quantum chromodynamics or QCD) and the unified theory of weak and electromagnetic interactions (electroweak). Gravity is included on this chart because it is one of the fundamental interactions even though not part of the "Standard Model."

### FERMIOS

matter constituents  
spin = 1/2, 3/2, 5/2, ...

Leptons spin = 1/2		
Flavor	Mass GeV/c <sup>2</sup>	Electric charge
$\nu_e$ neutrino	$<1 \times 10^{-8}$	0
e electron	0.000511	-1
$\nu_\mu$ muon neutrino	$<0.0002$	0
$\mu$ muon	0.106	-1
$\nu_\tau$ tau neutrino	$<0.02$	0
$\tau$ tau	1.7771	-1

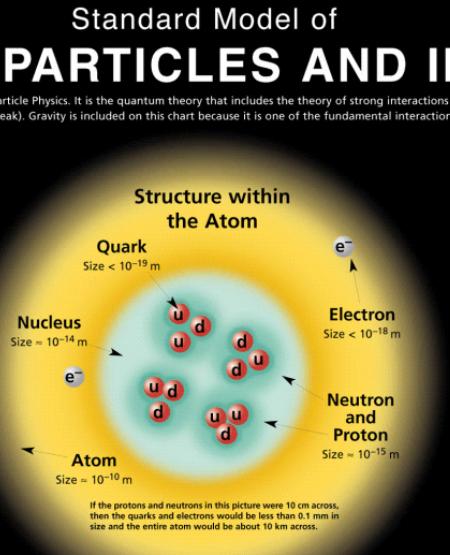
Spin is the intrinsic angular momentum of particles. Spin is given in units of  $\hbar$ , which is the quantum unit of angular momentum, where  $\hbar = h/2\pi = 6.58 \times 10^{-25}$  GeV s =  $1.05 \times 10^{-34}$  J s.

Electric charges are given in units of the proton's charge. In SI units the electric charge of the proton is  $1.60 \times 10^{-19}$  coulombs.

The energy unit of particle physics is the electronvolt (eV), the energy gained by one electron in crossing a potential difference of one volt. Masses are given in GeV/c<sup>2</sup> (remember  $E = mc^2$ ), where 1 GeV =  $10^9$  eV =  $1.60 \times 10^{-19}$  joule. The mass of the proton is 0.938 GeV/c<sup>2</sup> =  $1.67 \times 10^{-27}$  kg.

Quarks spin = 1/2

Flavor	Approx. Mass GeV/c <sup>2</sup>	Electric charge
u up	0.003	2/3
d down	0.006	-1/3
c charm	1.3	2/3
s strange	0.1	-1/3
t top	175	2/3
b bottom	4.3	-1/3



### BOSONS

force carriers  
spin = 0, 1, 2, ...

Unified Electroweak spin = 1		
Name	Mass GeV/c <sup>2</sup>	Electric charge
$\gamma$ photon	0	0
$W^-$	80.4	-1
$W^+$	80.4	+1
$Z^0$	91.187	0

Strong (color) spin = 1		
Name	Mass GeV/c <sup>2</sup>	Electric charge
g gluon	0	0

**Color Charge**  
Each quark carries one of three types of "strong charge," also called "color charge." These charges have nothing to do with the colors of visible light. There are eight possible types of color charge for gluons. Just as electrically charged particles interact by exchanging photons, in strong interactions color-charged particles interact by exchanging gluons. Leptons, photons, and  $W$  and  $Z$  bosons have no strong interactions and hence no color charge.

#### Quarks Confined in Mesons and Baryons

One cannot isolate quarks and gluons; they are confined in color-neutral particles called **hadrons**. This confinement (binding) results from multiple exchanges of gluons among the color-charged constituents. As color-charged particles (quarks and gluons) move apart, the energy in the color-force field between them increases. This energy eventually is converted into additional quark-antiquark pairs (see figure below). The quarks and antiquarks then combine into hadrons; these are the particles seen to emerge. Two types of hadrons have been observed in nature: **mesons**  $q\bar{q}$  and **baryons**  $qqq$ .

#### Residual Strong Interaction

The strong binding of color-neutral protons and neutrons to form nuclei is due to residual strong interactions between their color-charged constituents. It is similar to the residual electrical interaction that binds electrically neutral atoms to form molecules. It can also be viewed as the exchange of mesons between the hadrons.

## PROPERTIES OF THE INTERACTIONS

### Baryons qqq and Antibaryons qqq

Baryons are fermionic hadrons.  
There are about 120 types of baryons.

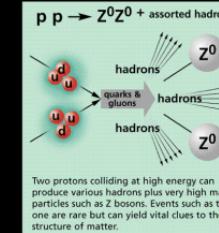
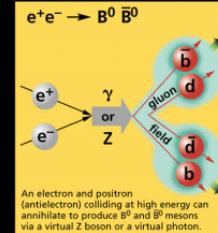
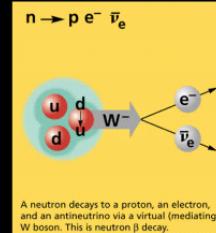
Property	Interaction	Gravitational	Weak	Electromagnetic (Electroweak)	Strong Fundamental	Residual
		Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge
Particles experiencing:		All	Quarks, Leptons	Electrically charged	Quarks, Gluons	Hadrons
Particles mediating:	Graviton (not yet observed)	$W^+$ $W^-$ $Z^0$		$\gamma$	Gluons	Mesons
Strength relative to electromag	$10^{-41}$	0.8	1	25	Not applicable to hadrons	Not applicable to quarks
for two u quarks at:	$10^{-18}$ m	$10^{-41}$	1	60		20
	$3 \times 10^{-17}$ m	$10^{-36}$	$10^{-7}$			
for two protons in nucleus						

### Matter and Antimatter

For every particle type there is a corresponding antiparticle type, denoted by a bar over the particle symbol (unless + or - charge is shown). Particle and antiparticle have identical mass and spin but opposite charges. Some electrically neutral bosons (e.g.,  $Z^0$ ,  $\gamma$ , and  $\eta_c = \bar{c}c$ , but not  $K^0 = \bar{d}s$ ) are their own antiparticles.

### Figures

These diagrams are an artist's conception of physical processes. They are not exact and have no meaningful scale. Green shaded areas represent the cloud of gluons or the gluon field, and red lines the quark paths.



### The Particle Adventure

Visit the award-winning web feature [The Particle Adventure](http://ParticleAdventure.org) at <http://ParticleAdventure.org>

This chart has been made possible by the generous support of:

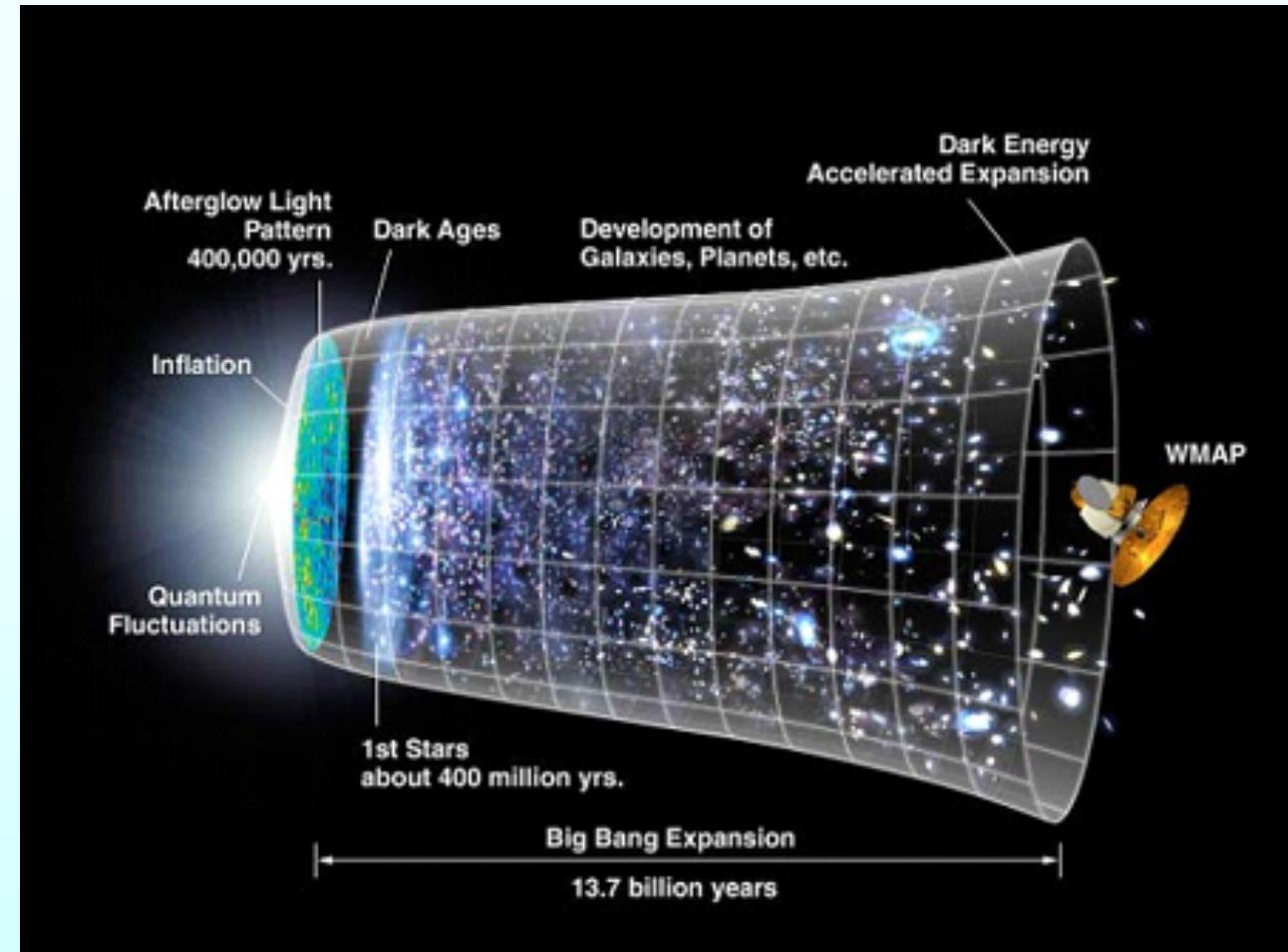
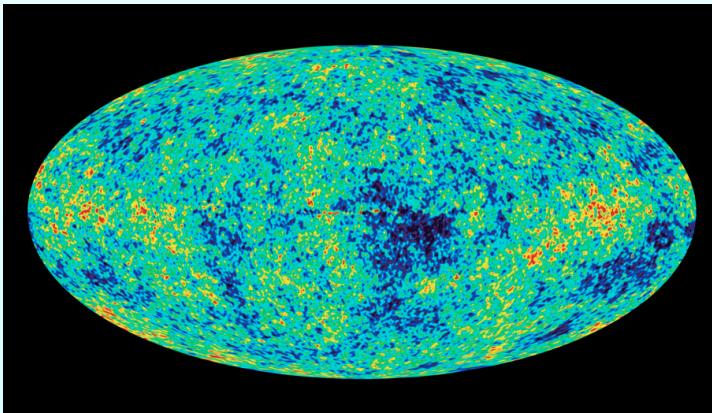
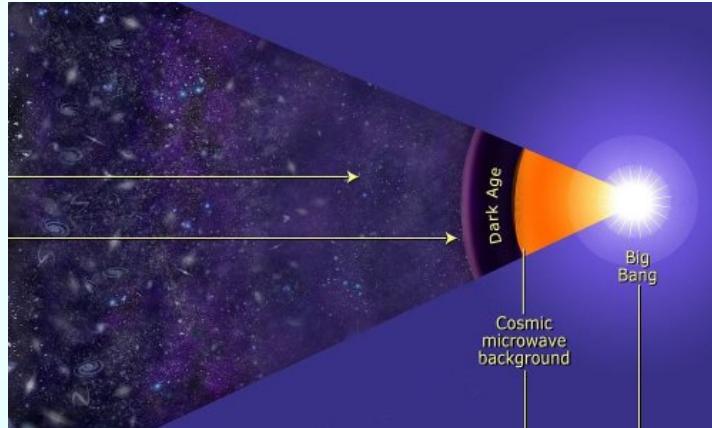
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U.S. National Science Foundation  
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<http://CPEPweb.org>

# History of the Universe

**From Cosmic Microwave Background Radiation and other measurements**



## Thanks to theoreticians and experimentalists

There is an ancient Chinese saying:

"He who labors with his mind rules over he who labors with his hand".

This kind of backward idea is very harmful to youngsters from developing countries.

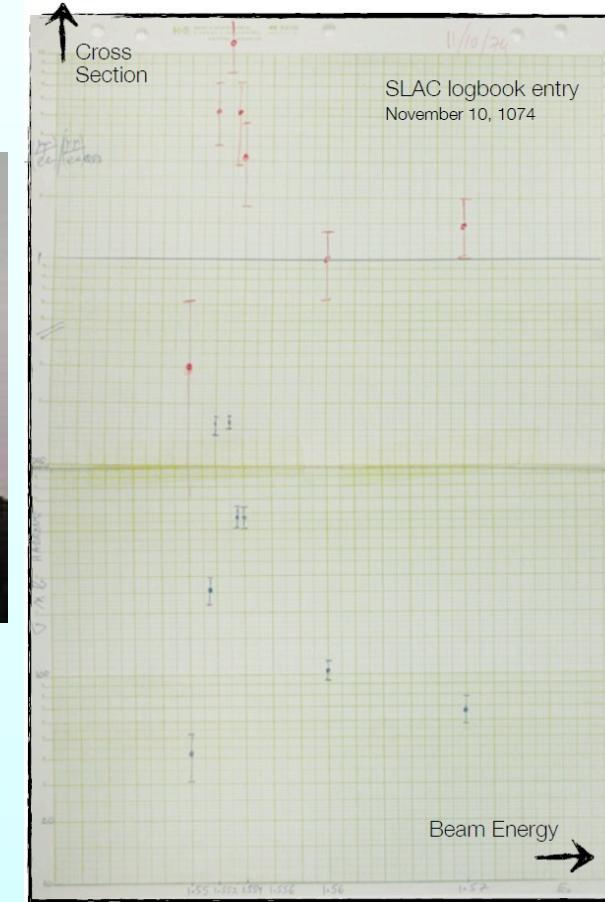
Partly because of this type of concept, many students from these countries are inclined towards theoretical studies and avoid experimental work.

In reality, a theory in natural science can not be without experimental foundations; physics, in particular, comes from experimental work ...

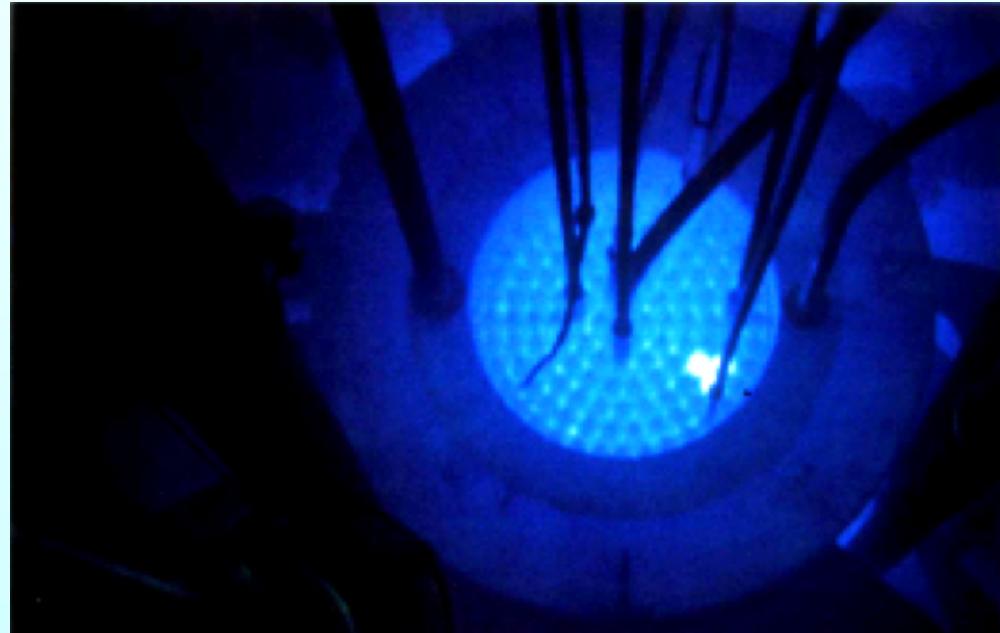
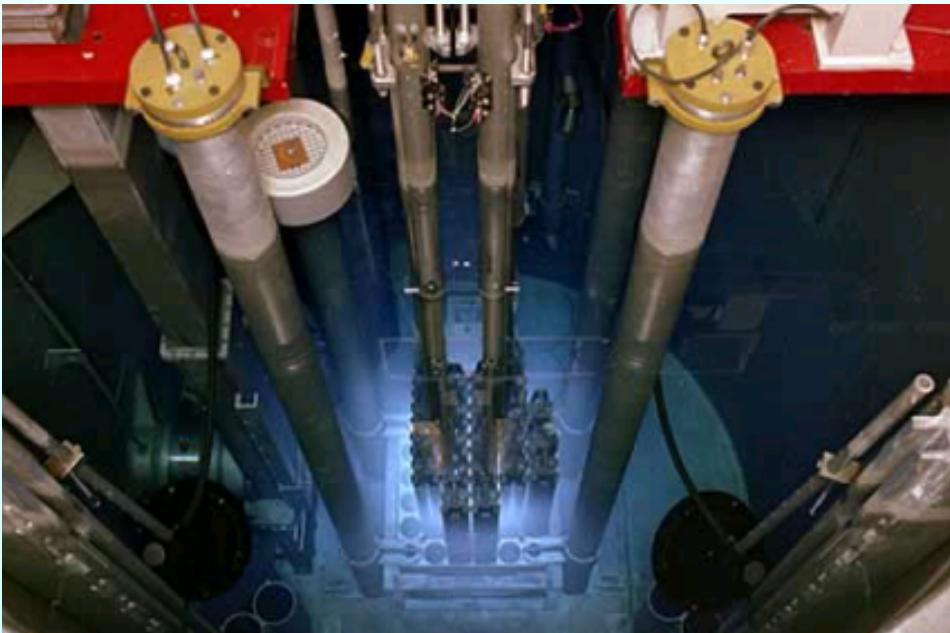
Samuel Ting  
December 10, 1976



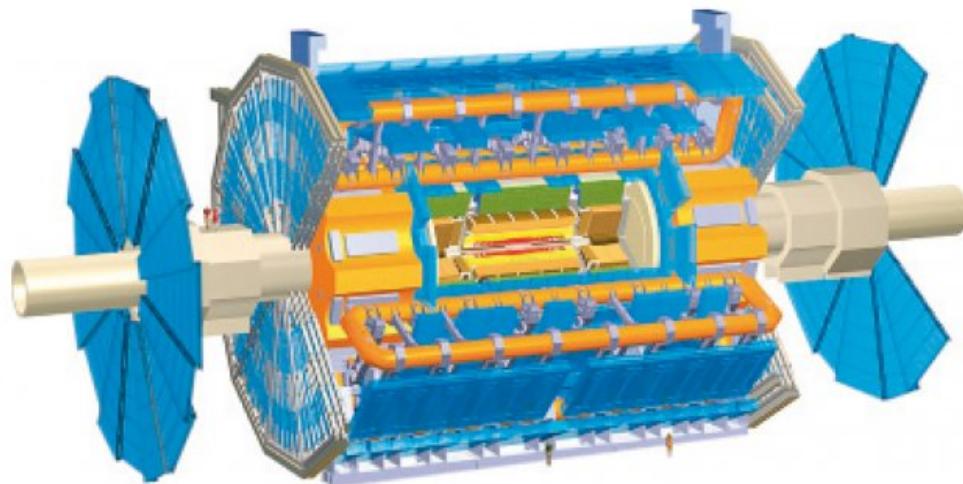
Nobel Prizes 1974  
for the J/ψ discovery



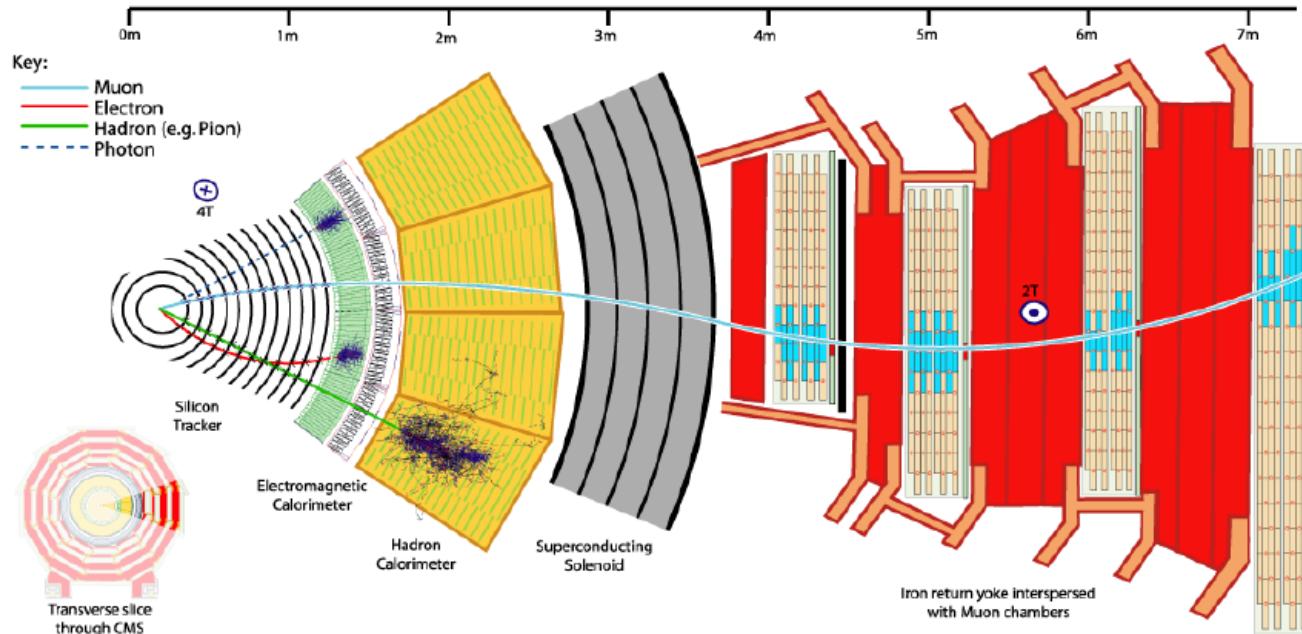
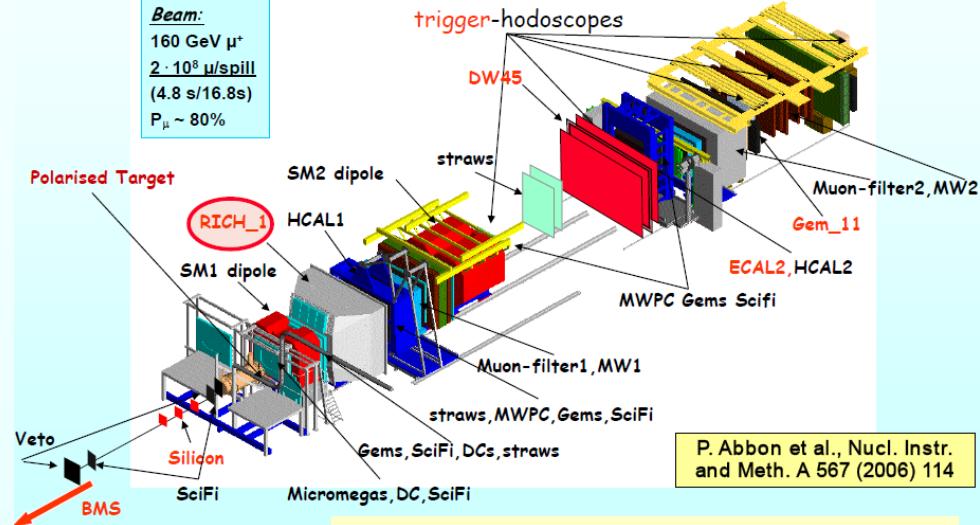
# Auroras and nuclear reactors



# A modern experiment

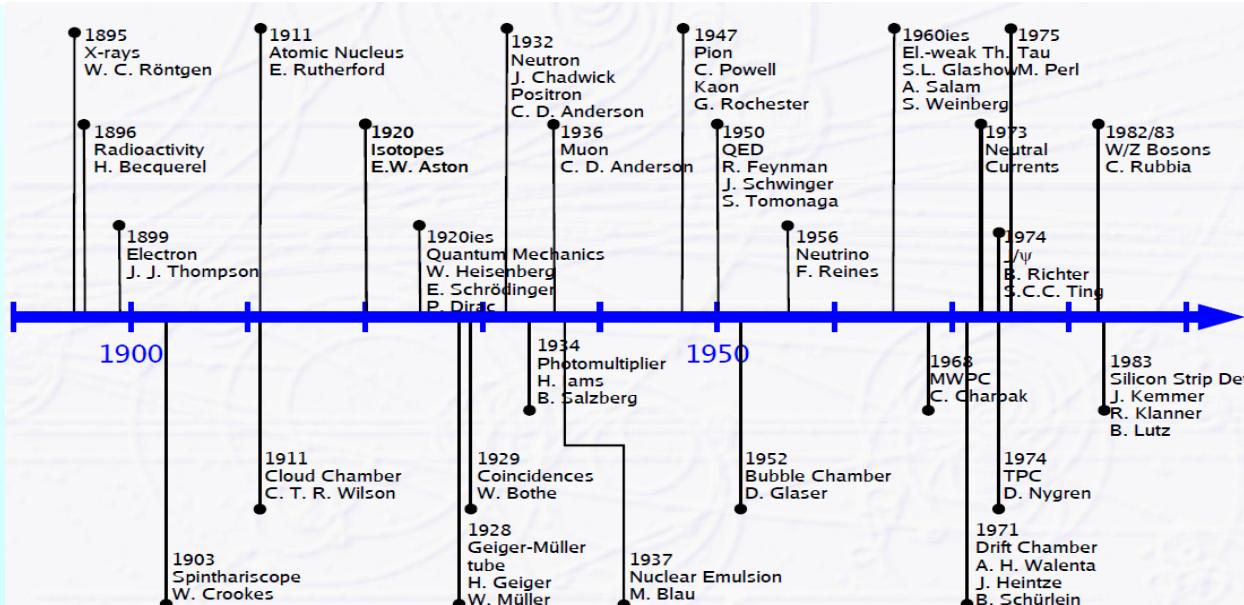
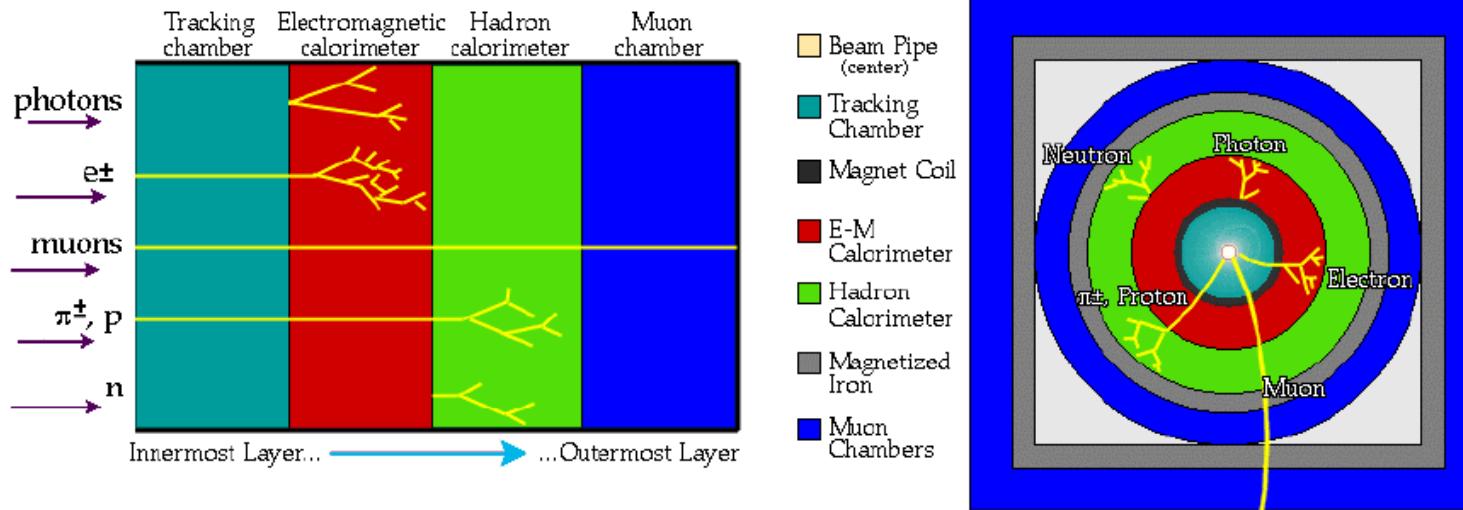


**Beam:**  
160 GeV  $\mu^+$   
 $2 \cdot 10^8 \mu\text{-spill}$   
(4.8 s/16.8 s)  
 $P_\mu \sim 80\%$



uses a combination of different detectors, combines the information from all of them and fully reconstructs the characteristics of the interesting event which took place

# Different particles are seen by different detectors



The history of discoveries and that of particle detectors are intimately interconnected

# The discovery of radiation

First

Detection of  
 $\alpha$ -,  
 $\beta$ - and  $\gamma$ -rays

1896

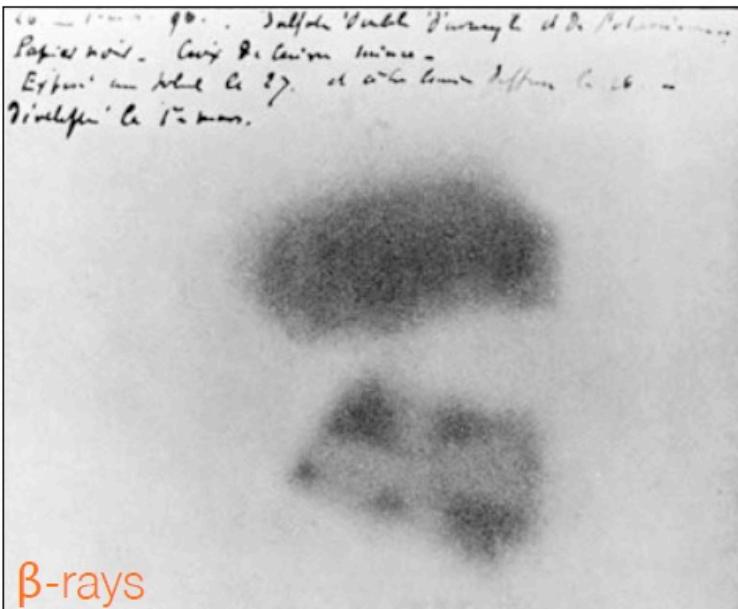


Image of Becquerel's photographic plate which has been fogged by exposure to radiation from a uranium salt.

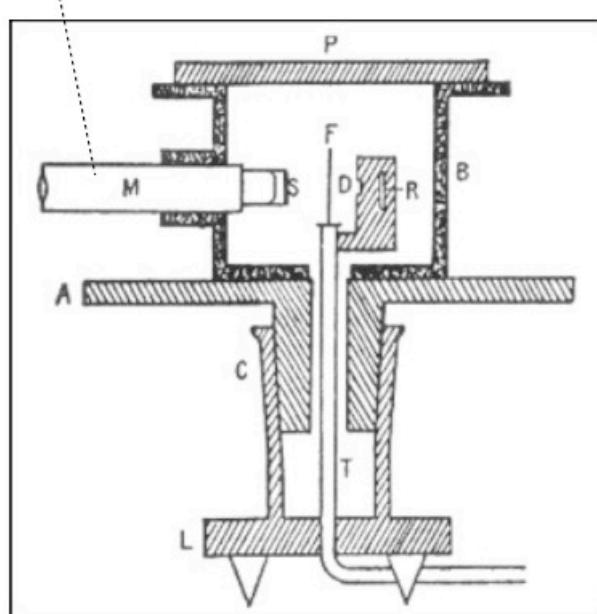
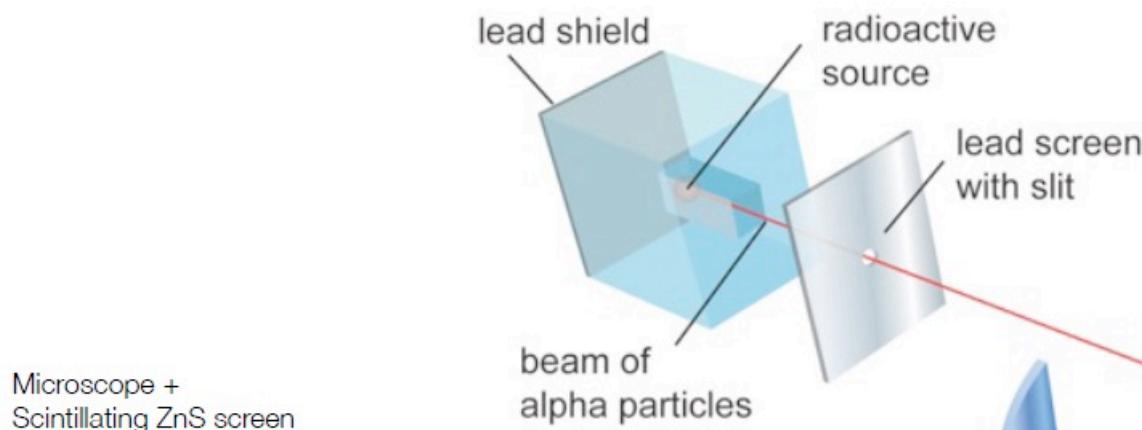
$\gamma$ -rays

1896

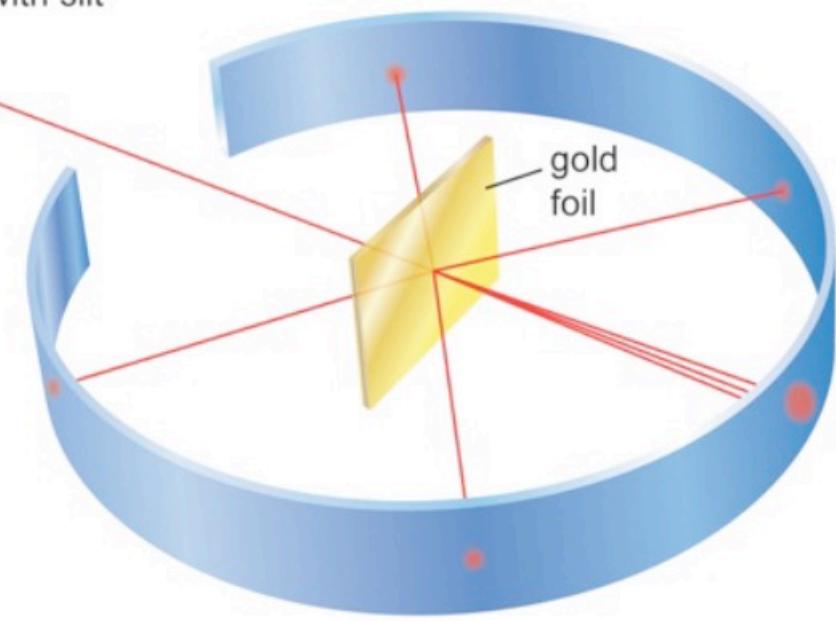


An x-ray picture taken by Wilhelm Röntgen of Albert von Kölliker's hand at a public lecture on 23 January 1896.

# The discovery of the atomic nucleus



Rutherford's original  
experimental setup



Schematic view  
of Rutherford experiment

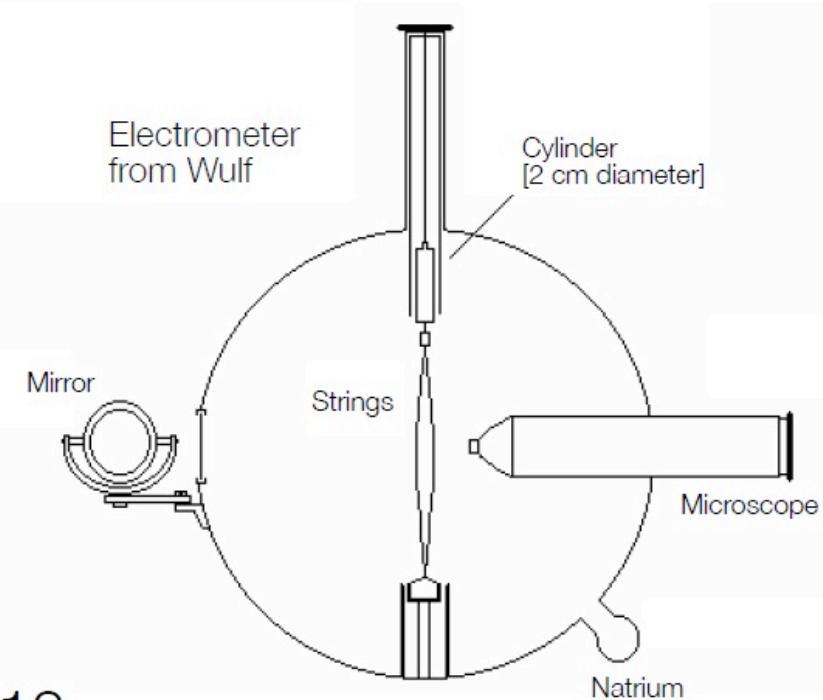
1911

# 1912: cosmic rays



## Detection of cosmic rays

[Hess 1912; Nobel prize 1936]



1912

Victor F. Hess before his 1912 balloon flight  
in Austria during which he discovered cosmic rays.

# 1911: the cloud chamber

## Cloud chamber (1911 by Charles T. R. Wilson, Noble Prize 1927)

- chamber with saturated water vapour
- charged particles leave trails of ions
  - water is condensing around ions
- visible track as line of small water droplets



UK Science Museum

## Also required

- high speed photographic methods
  - invented by Arthur M. Worthington 1908 to investigate the splash of a drop
  - ultra short flash light produced by sparks



Charles T. R. Wilson



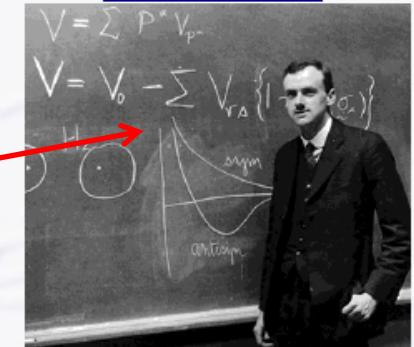
## First photographs of $\alpha$ -ray particles 1912



# 1932: antimatter, 1936: muon

Was also used for the discovery of the **positron**

- predicted by Paul Dirac 1928 (Nobel Prize 1933)
- found in cosmic rays by Carl D. Anderson 1932 (Nobel Prize 1936)



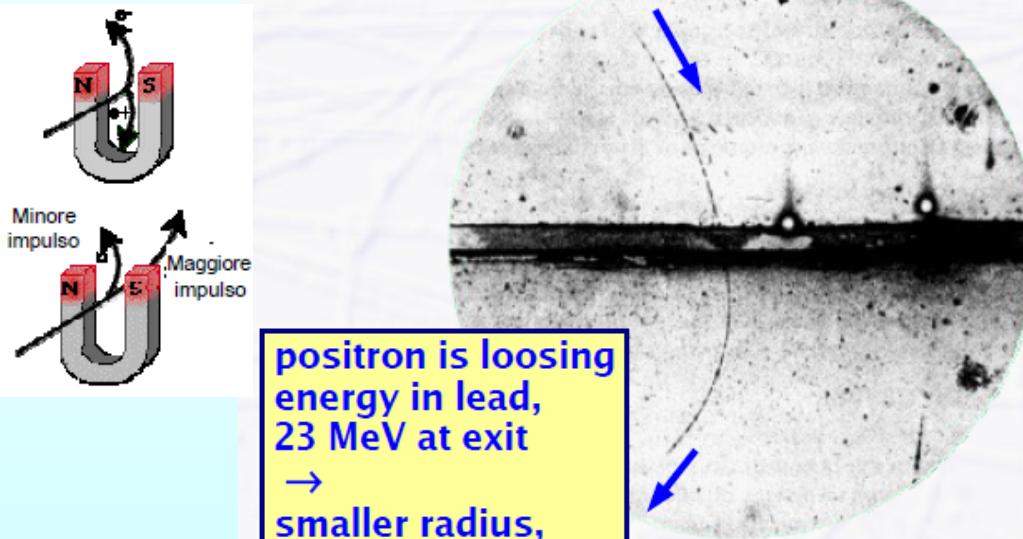
Anderson also found the **muon** in 1936, the first 2<sup>nd</sup> generation particle in the Standard Model

Isidor Isaac Rabi said:  
“Who ordered that?”

**Carl D. Anderson**



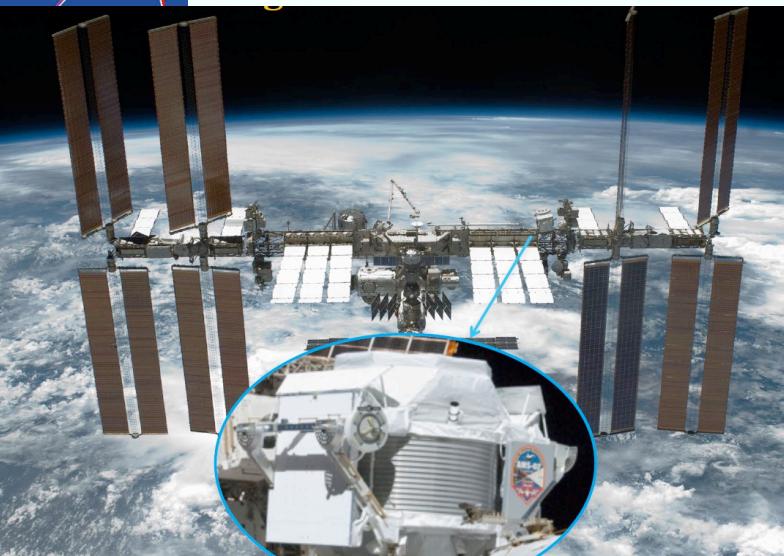
downward going positron, 63 MeV



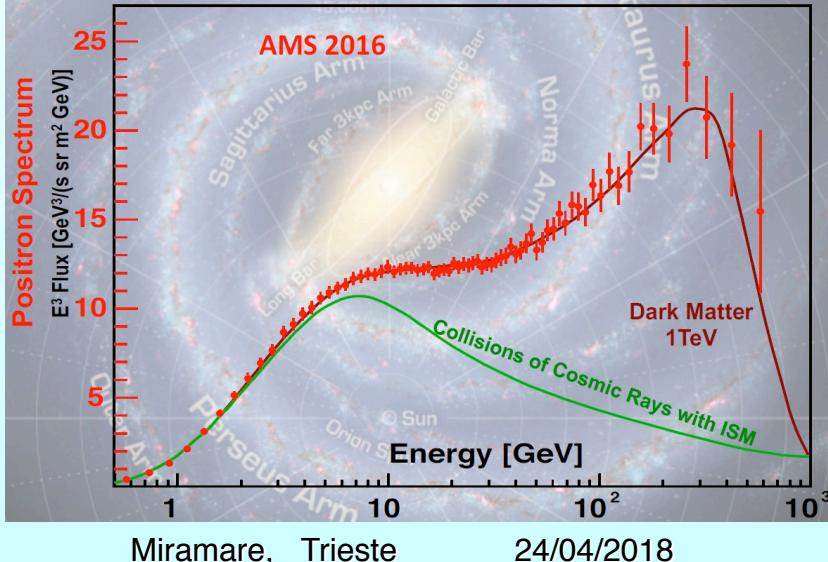
6 mm lead plate

⊗ 1.5 T magnetic field

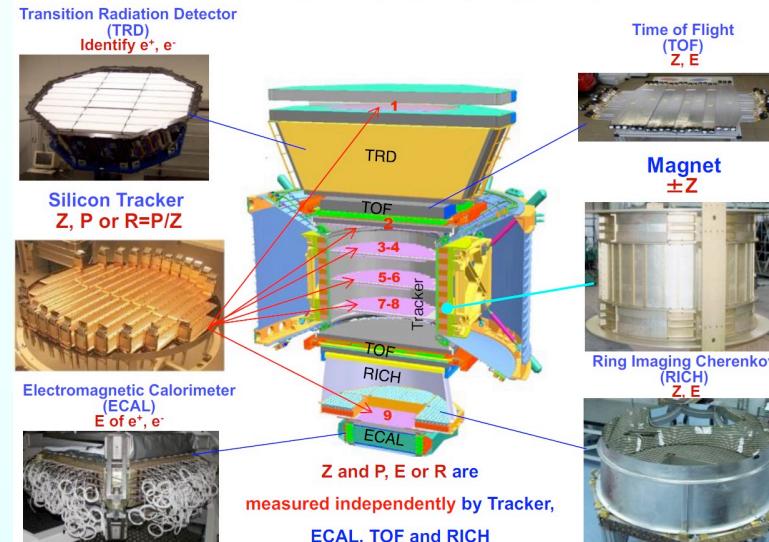
# AMS 2017: cosmic antihelium?



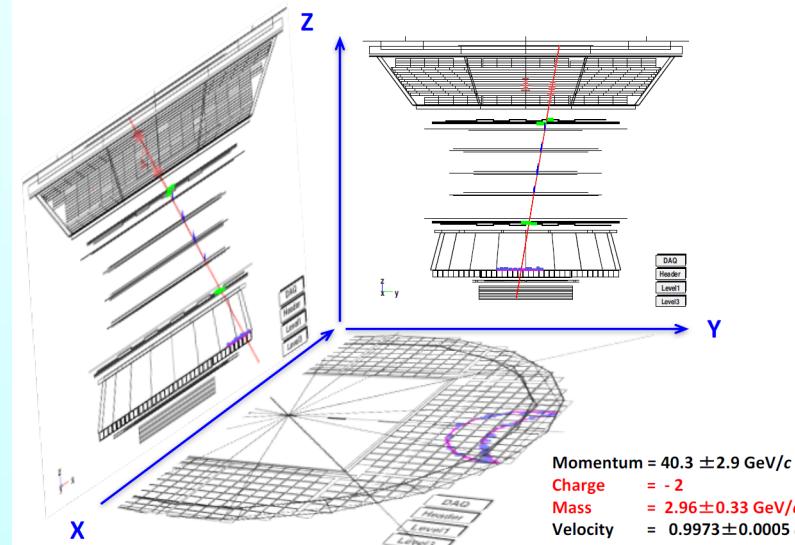
The AMS results are in excellent agreement with a Dark Matter Model



AMS: A TeV precision, multipurpose, magnetic spectrometer



An anti-Helium candidate:



# Nuclear emulsions

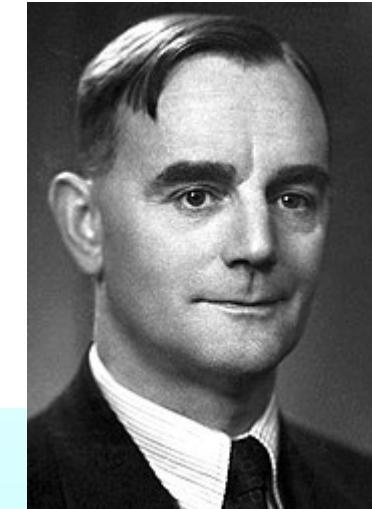
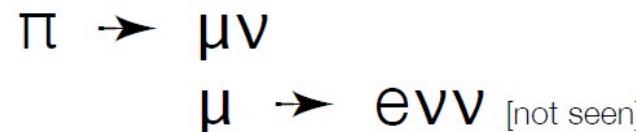
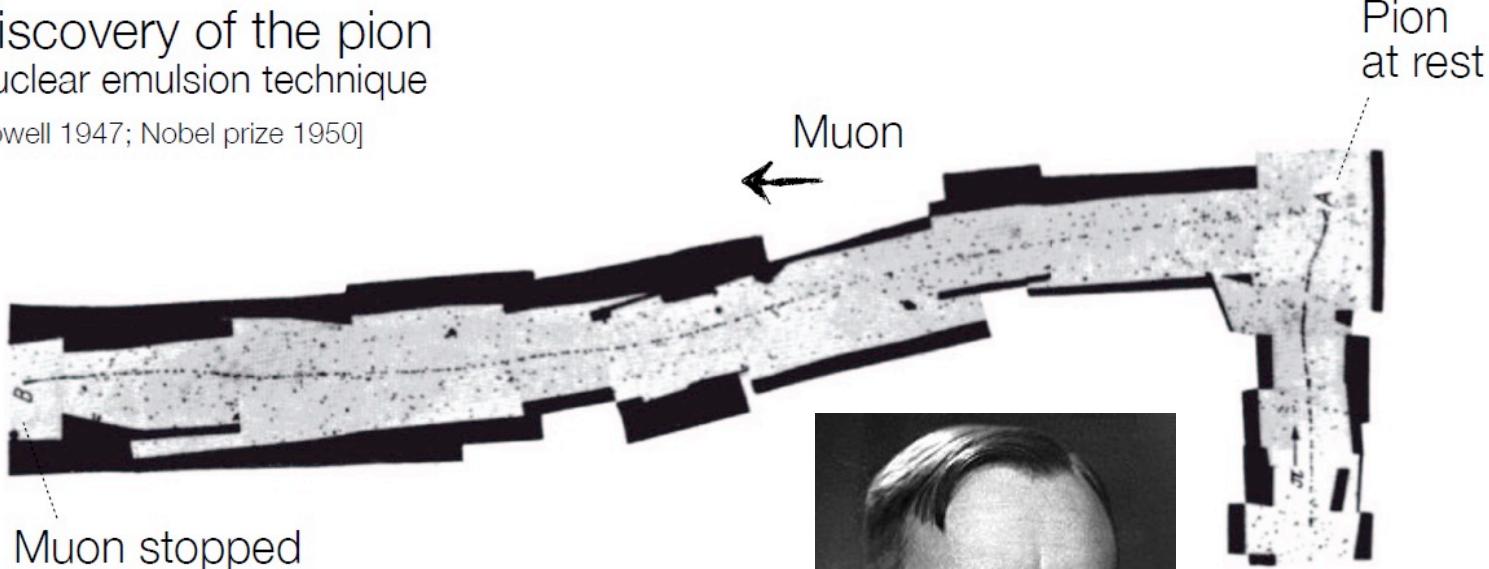


**Marietta Blau:**

she developed in Vienna the photographic nuclear emulsion technology for very accurate measurement of high energy nuclei and discovered the "disintegration stars" of spallation events

Discovery of the pion  
Nuclear emulsion technique

[Powell 1947; Nobel prize 1950]



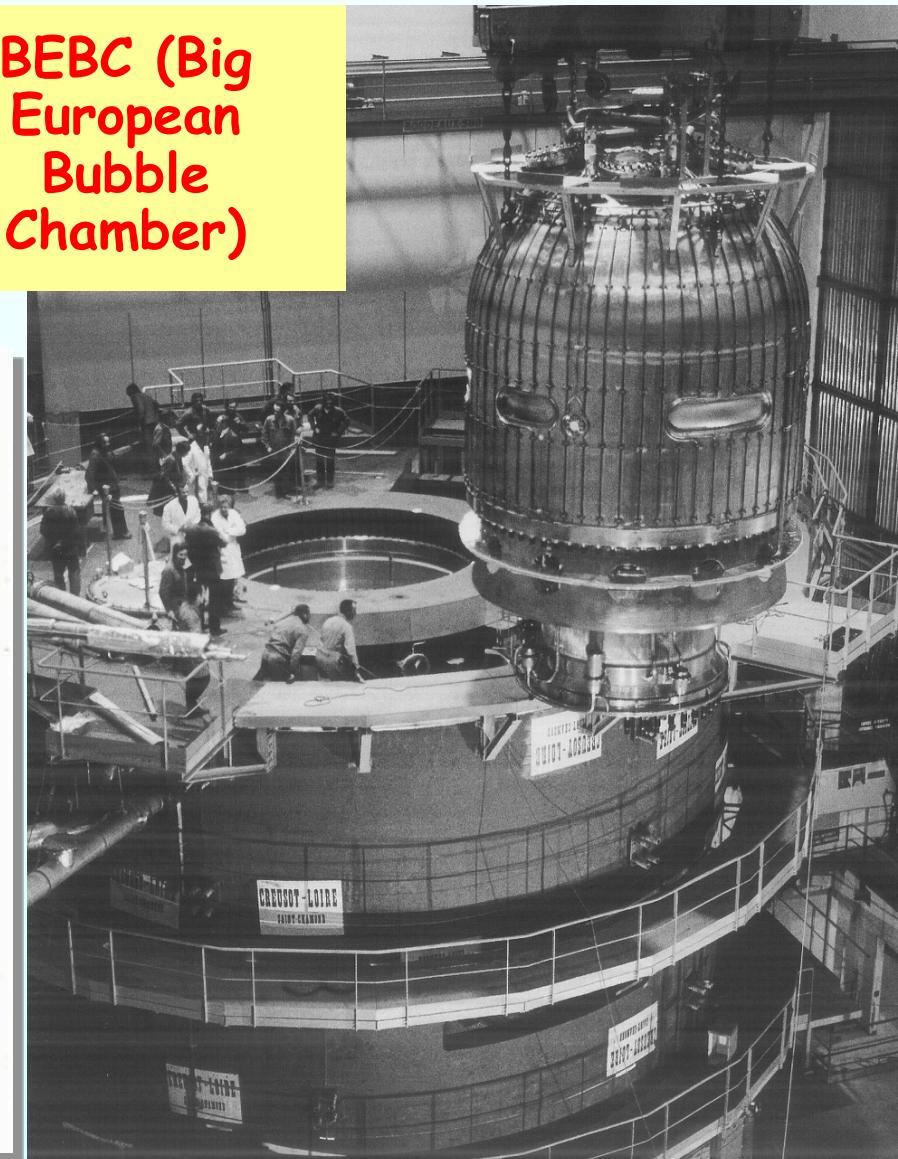
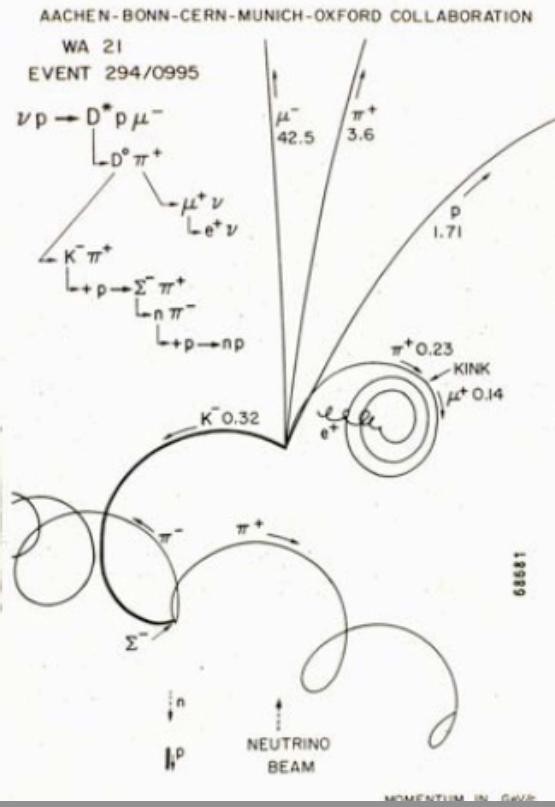
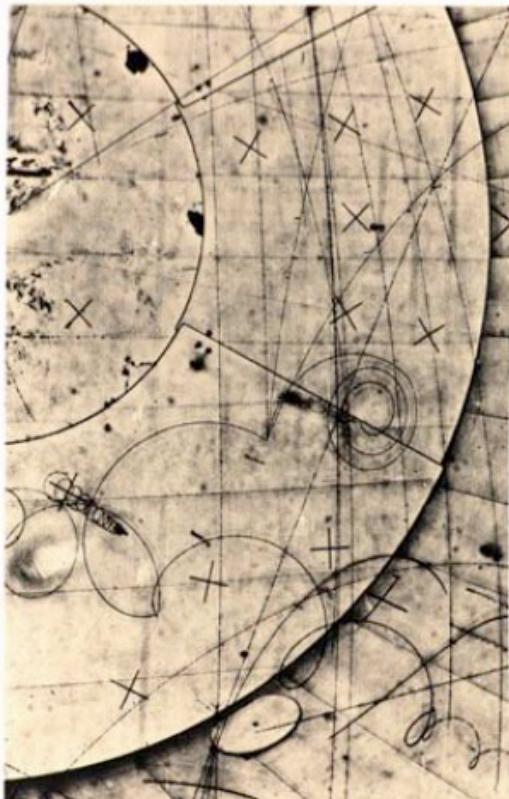
**Cecil Frank Powell**  
Nobel Prizes 1950

emulsions are still the detectors with the highest intrinsic space resolution: < 1  $\mu\text{m}$

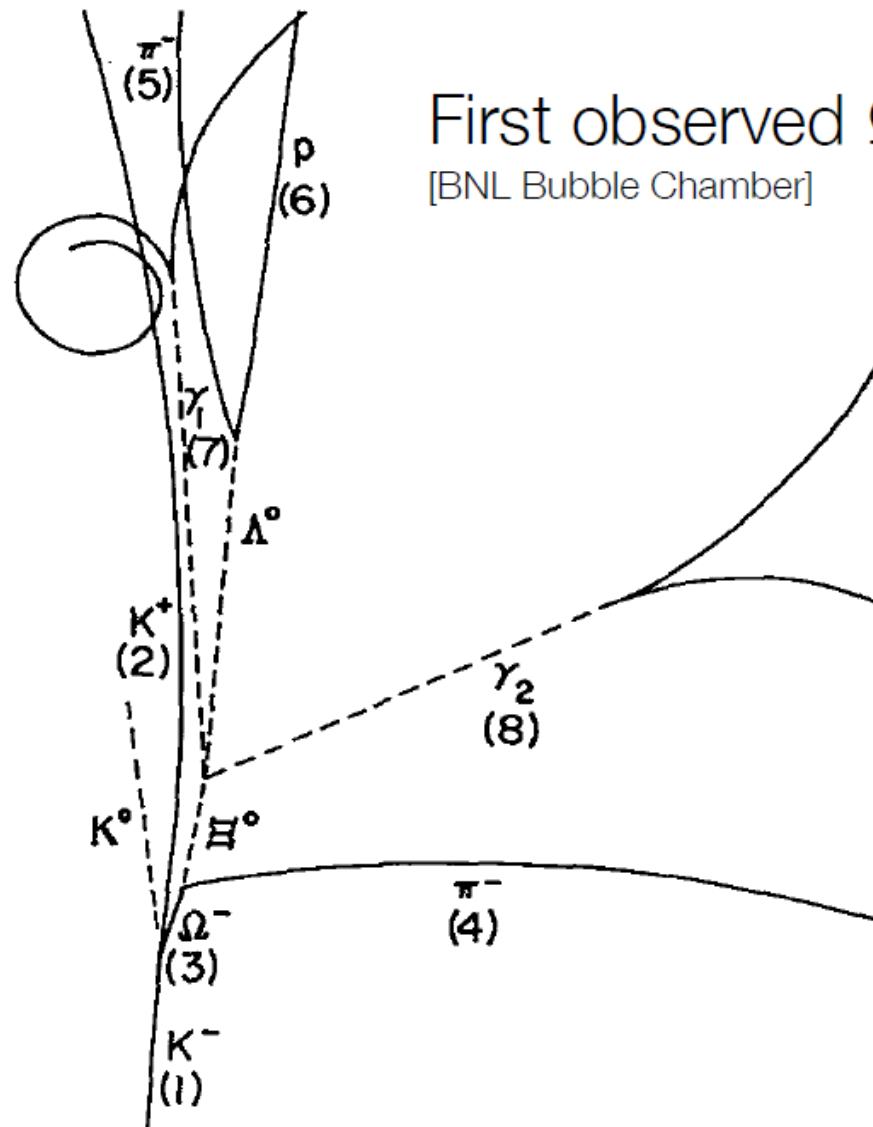
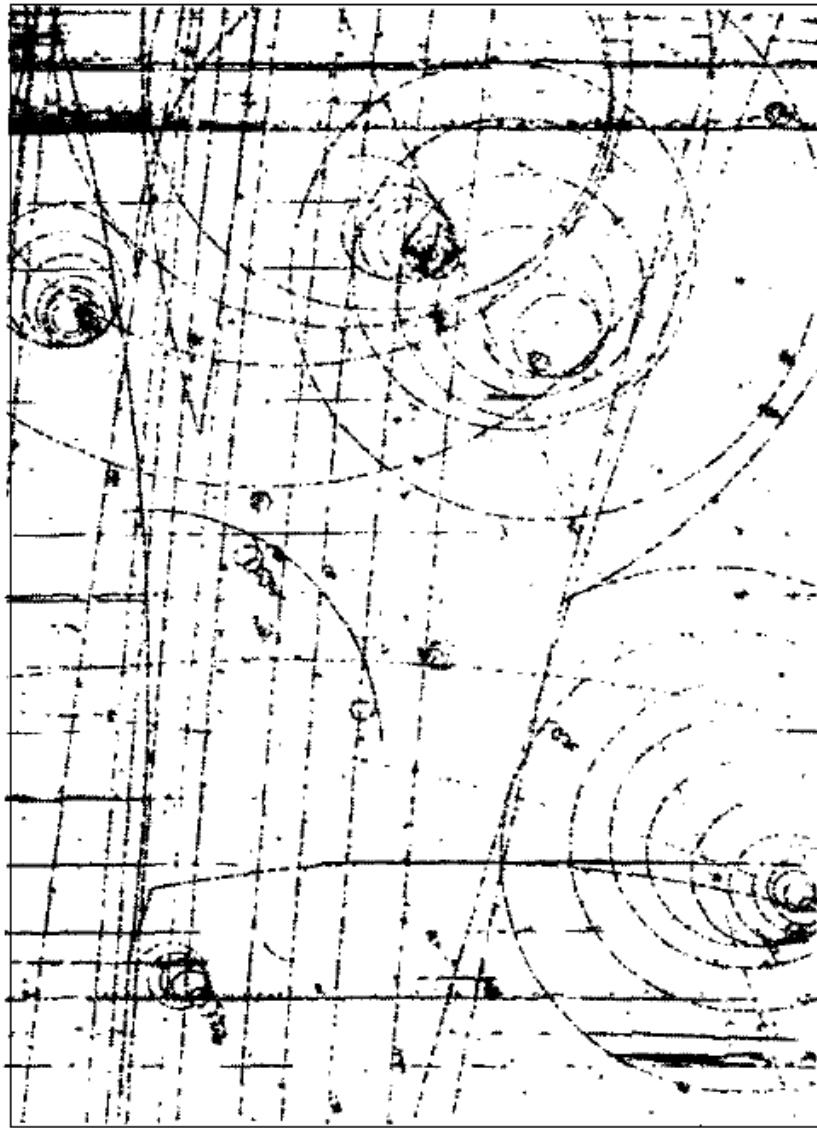
# 1952: bubble chamber

The bubble chamber, invented by Donald Arthur Glaser in 1952, has been for many years the most powerful instrument of ionizing particles investigation.

**BEBC (Big European Bubble Chamber)**

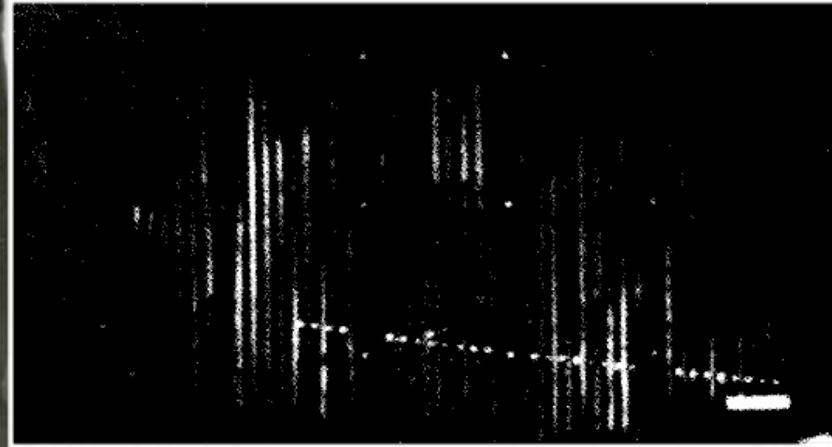
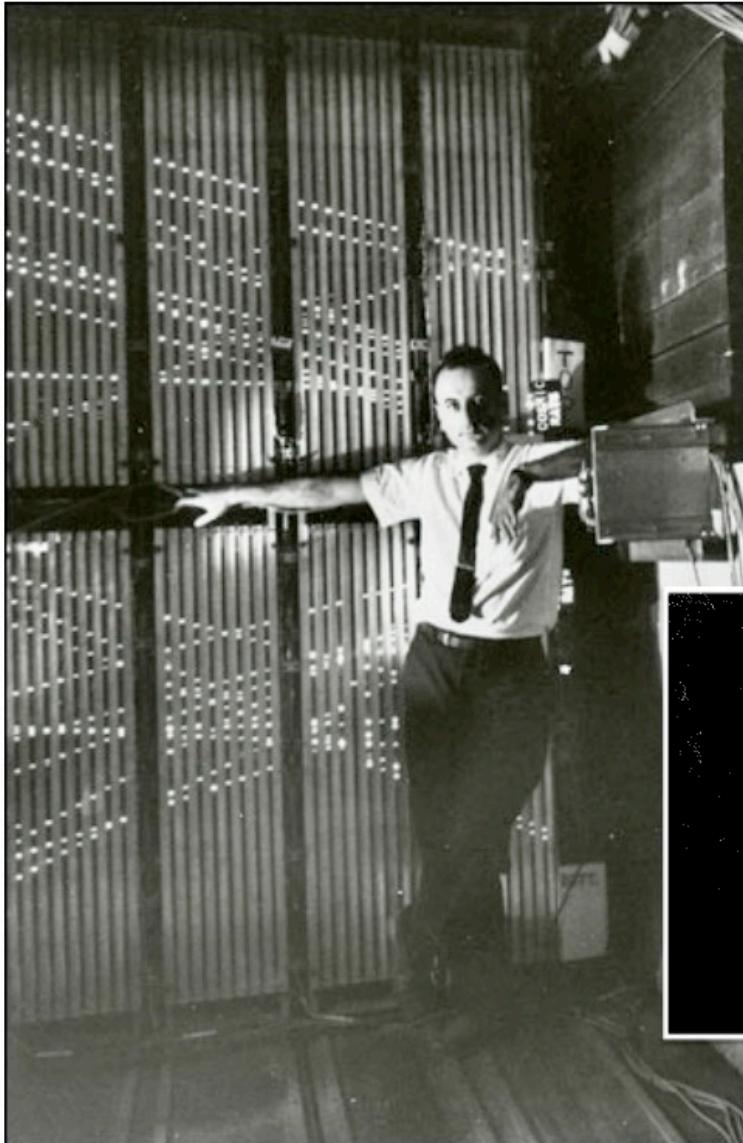


# 1964: the first predicted particle



First observed  $\Omega^-$  event  
[BNL Bubble Chamber]

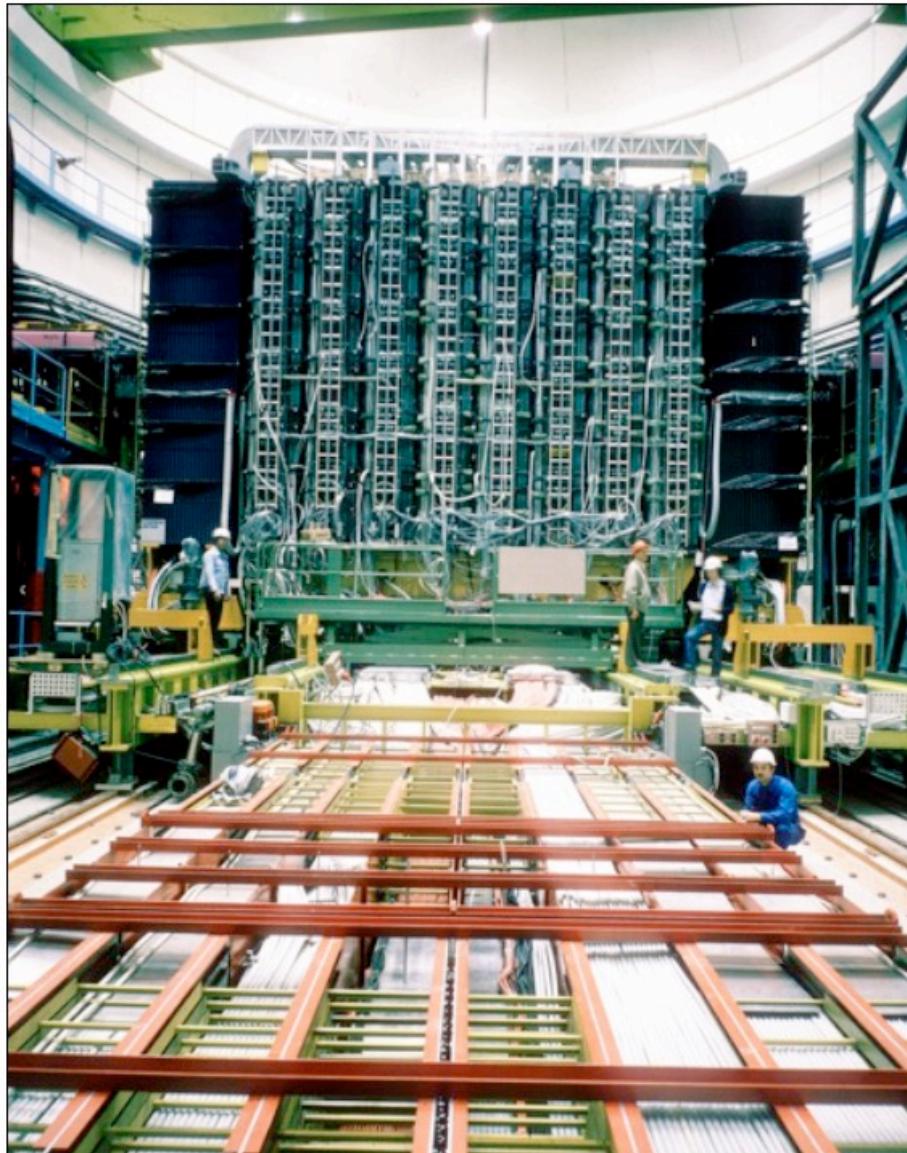
# 1962: the neutrino



Single muon event from  
original publication

Melvin Schwartz in front of the spark chamber  
used to discover the muon neutrino

# 1983: the weak bosons

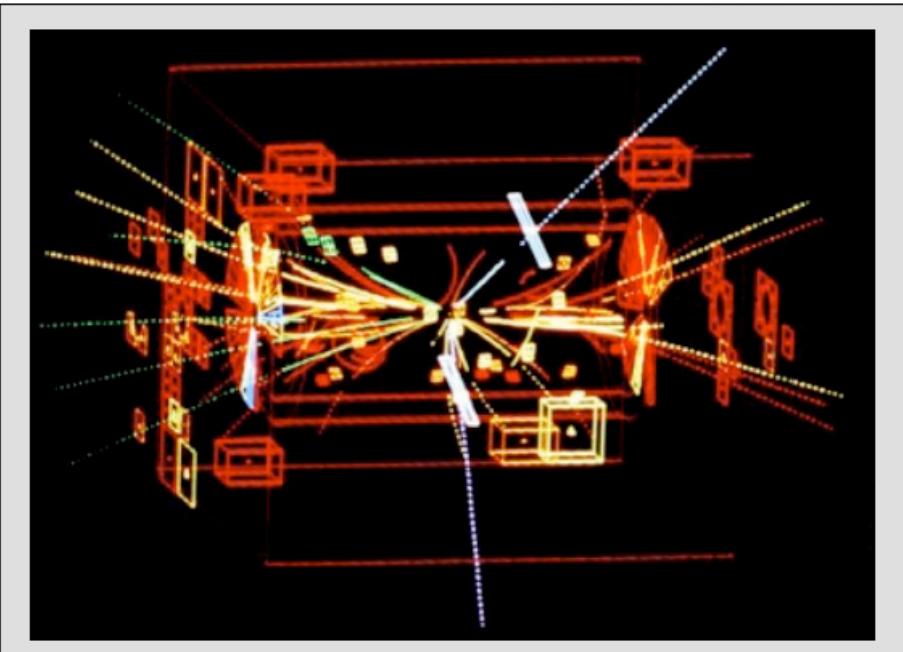


UA1  
Detector

Discovery of the  
W/Z boson (1983)

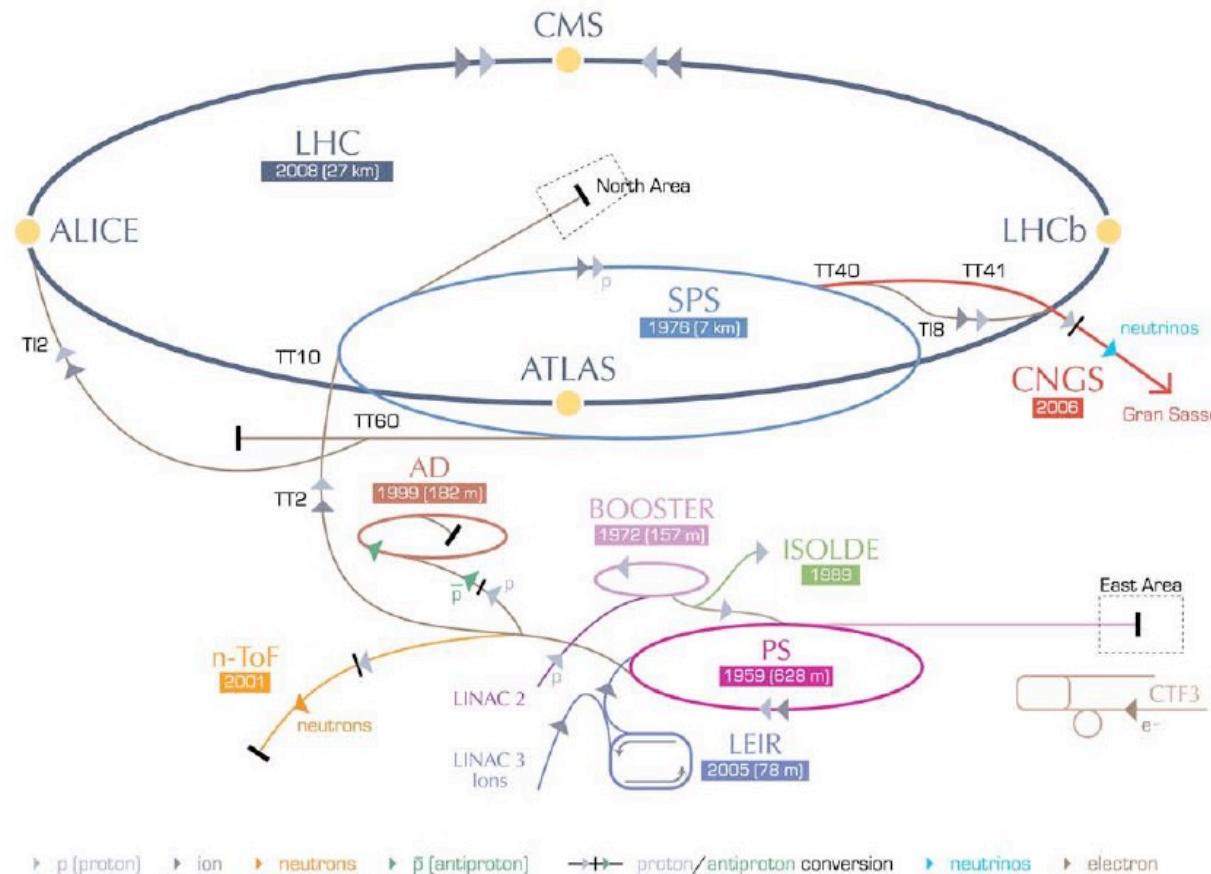
Carlo Rubbia  
Simon Van der Meer  
[Nobel prize 1984]

First  $Z^0$  particle seen by UA1





## CERN's accelerator complex



LHC Large Hadron Collider   SPS Super Proton Synchrotron   PS Proton Synchrotron

AD Antiproton Decelerator   CTF3 Clic Test Facility   CNGS Cern Neutrinos to Gran Sasso   ISOLDE Isotope Separator OnLine DEvice  
 LEIR Low Energy Ion Ring   LINAC LINear ACcelerator   n-ToF Neutrons Time Of Flight



Is the largest particle physics laboratory in the world  
It hosts ~2500 staff members and ~12000 users

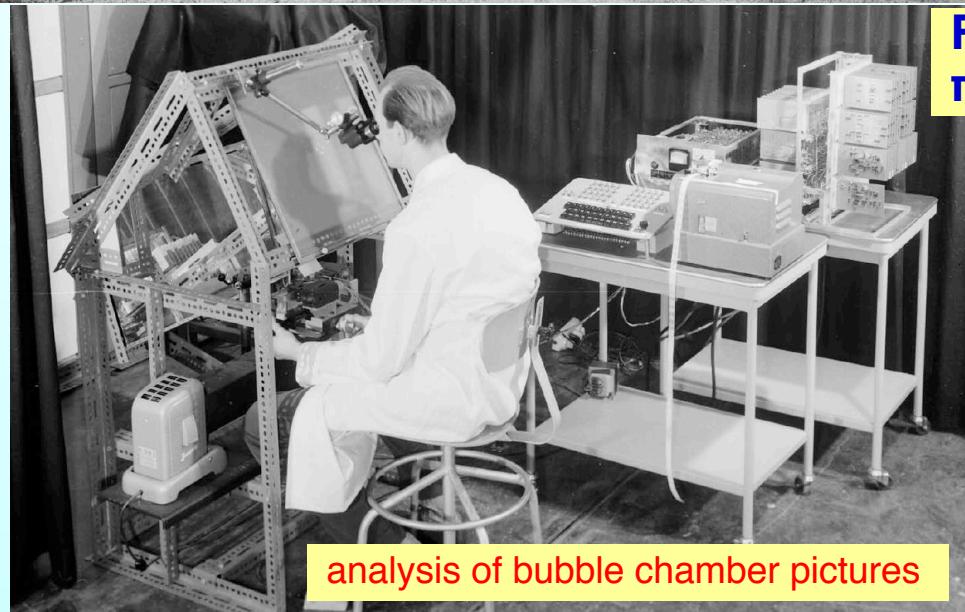
Provides particle accelerators and infrastructures for particle physics research.

Founded in 1954 as a European common project

All results are published and universally accessible



# The 600 MeV Syncro-Cyclotron



First physics result in 1958:  
 $\pi \rightarrow e \nu$  BR =  $10^{-4}$

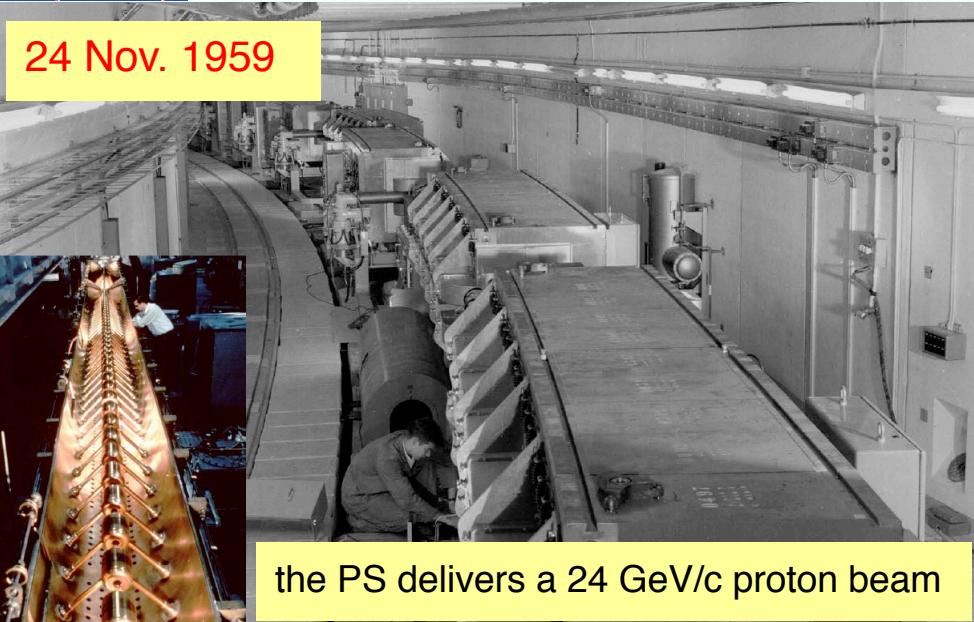


G. Fidecaro,  
(is the first INFN  
Trieste director)

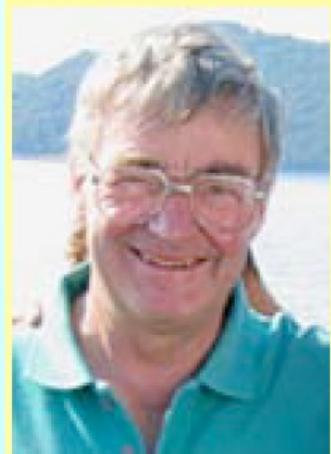


# The Proton-Synchrotron and the ISR

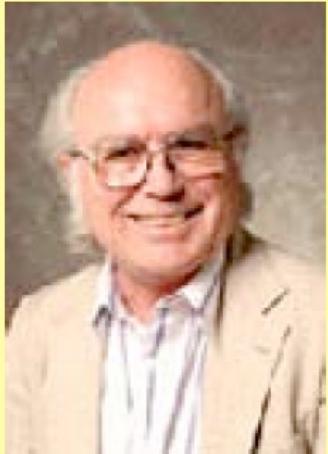
24 Nov. 1959



Julius Wess

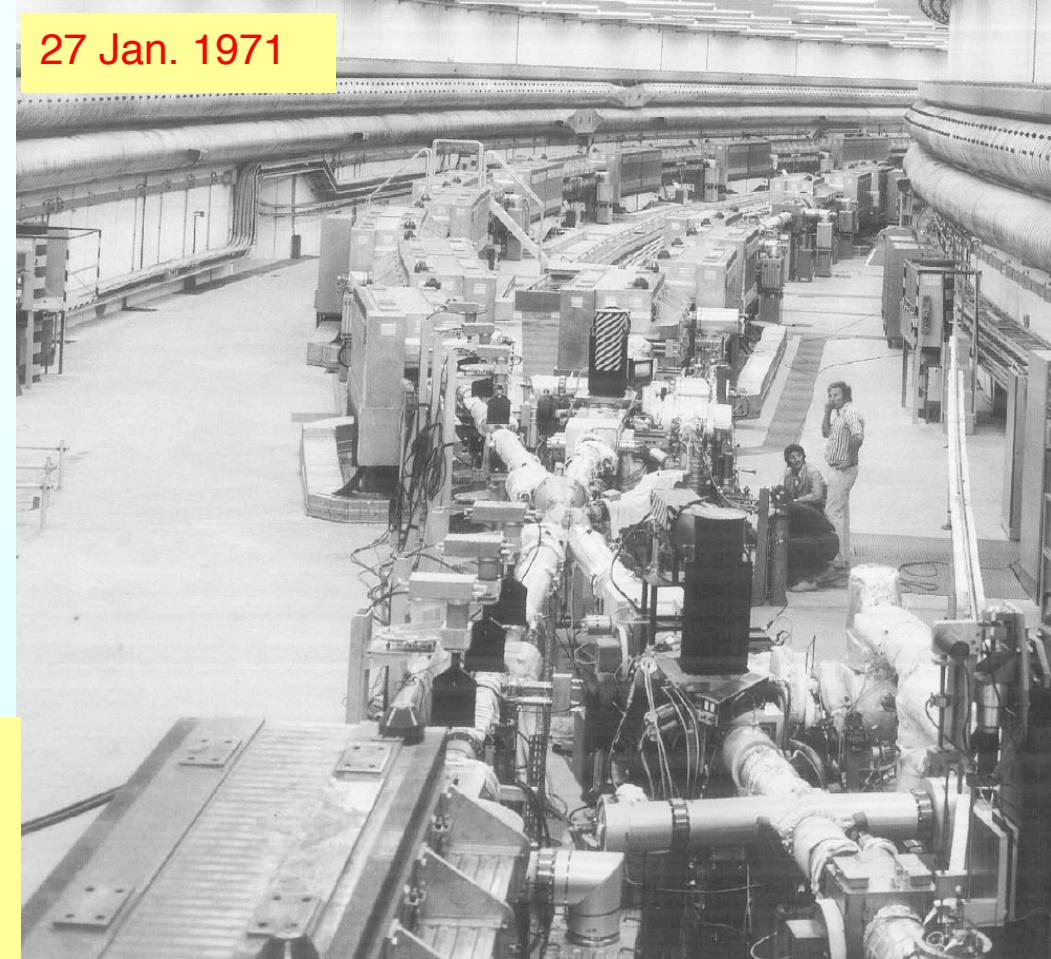


Bruno Zumino



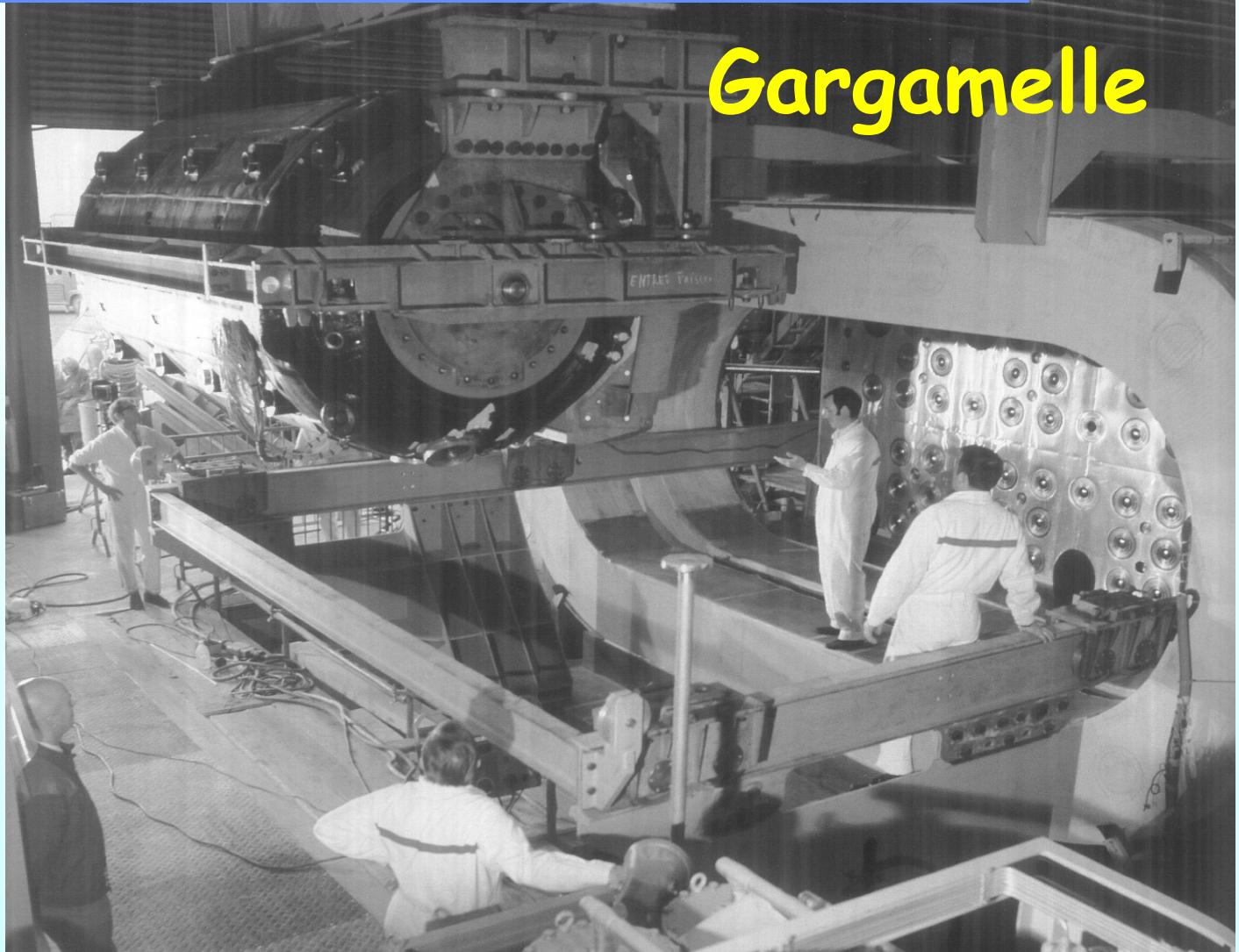
1974:  
Super-  
symmetry  
theory

27 Jan. 1971



# 1973: the e.w. theory is confirmed

1973: the first big discovery in Europe.  
**Gargamelle detects weak neutral currents**  
The electroweak theory is confirmed



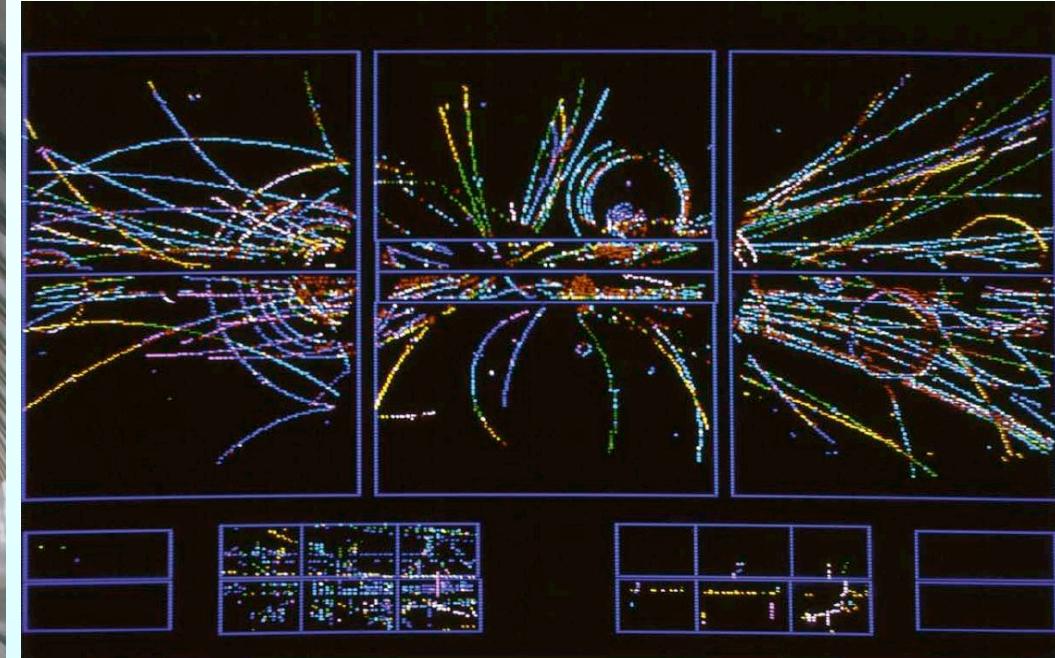
## Gargamelle

Salam receives the Nobel prize in 1979 together with Weinberg and Glashow

# The Super-Proton-Synchrotron



1981: The SPS is the first proton-antiproton collider



1983: an event in the UA1 detector

# The discovery of W and Z bosons

Nobel Prize in  
Physics 1984

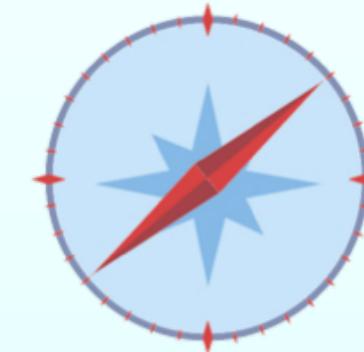
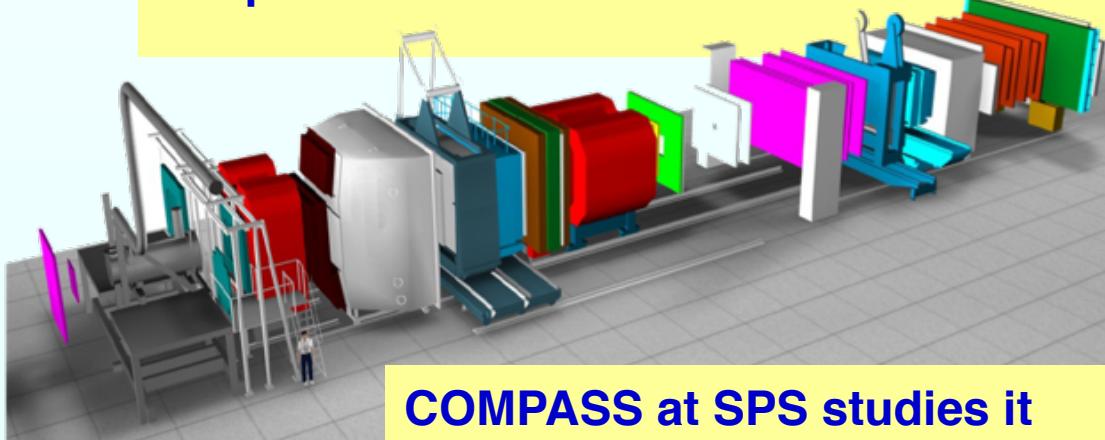
"for their decisive contributions to the large project, which led to the discovery of the field particles W and Z, communicators of weak interaction"

Carlo Rubbia and  
Simon Van der Meer

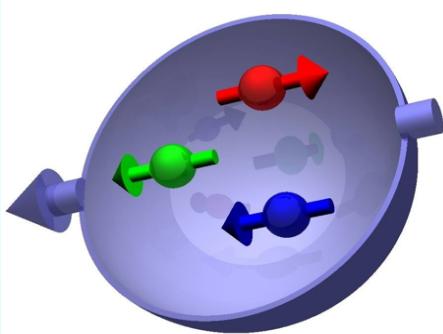


# The mystery of the proton spin

1989: The EMC at SPS measures the proton spin decomposition  
 → spin crisis

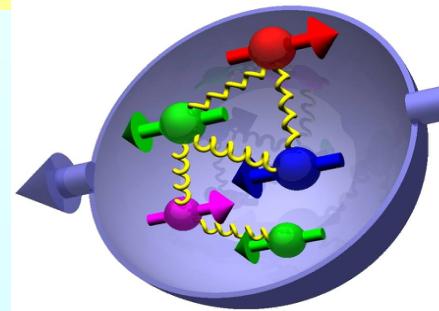


**COMPASS at SPS studies it**



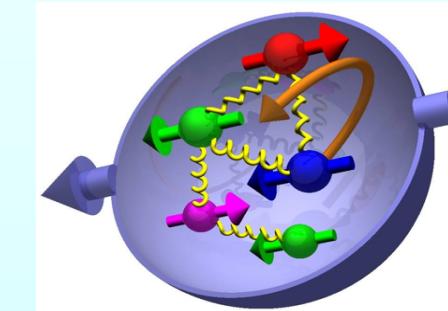
Partonmodel

1989 EMC measured  
 $\Sigma = 0.120 \pm 0.091 \pm 0.138$



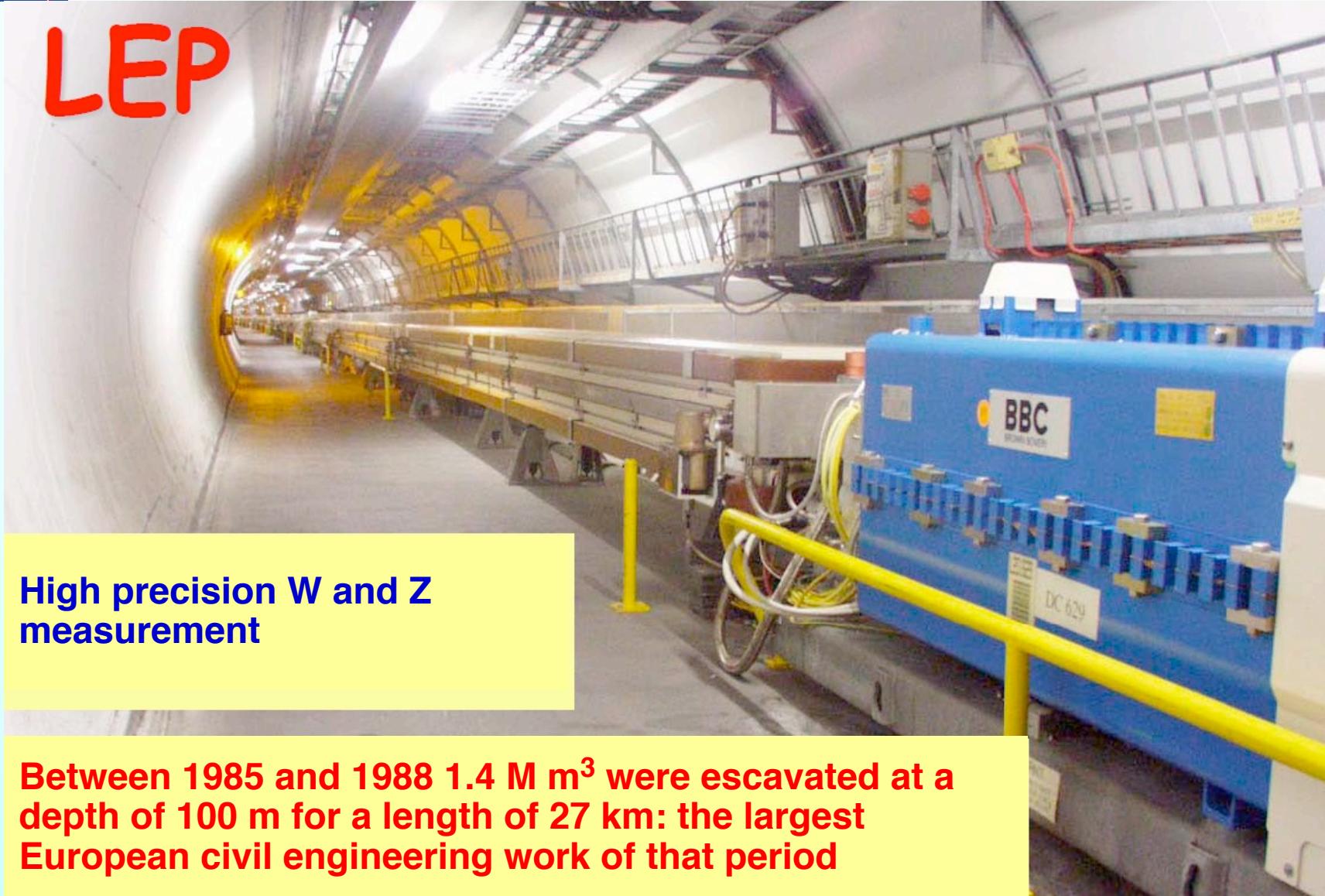
Gluons are important !

→  $\Delta g$   
 → Sea quarks  $\Delta q_s$



Full description  
 needs  
 orbital angular momentum

# LEP

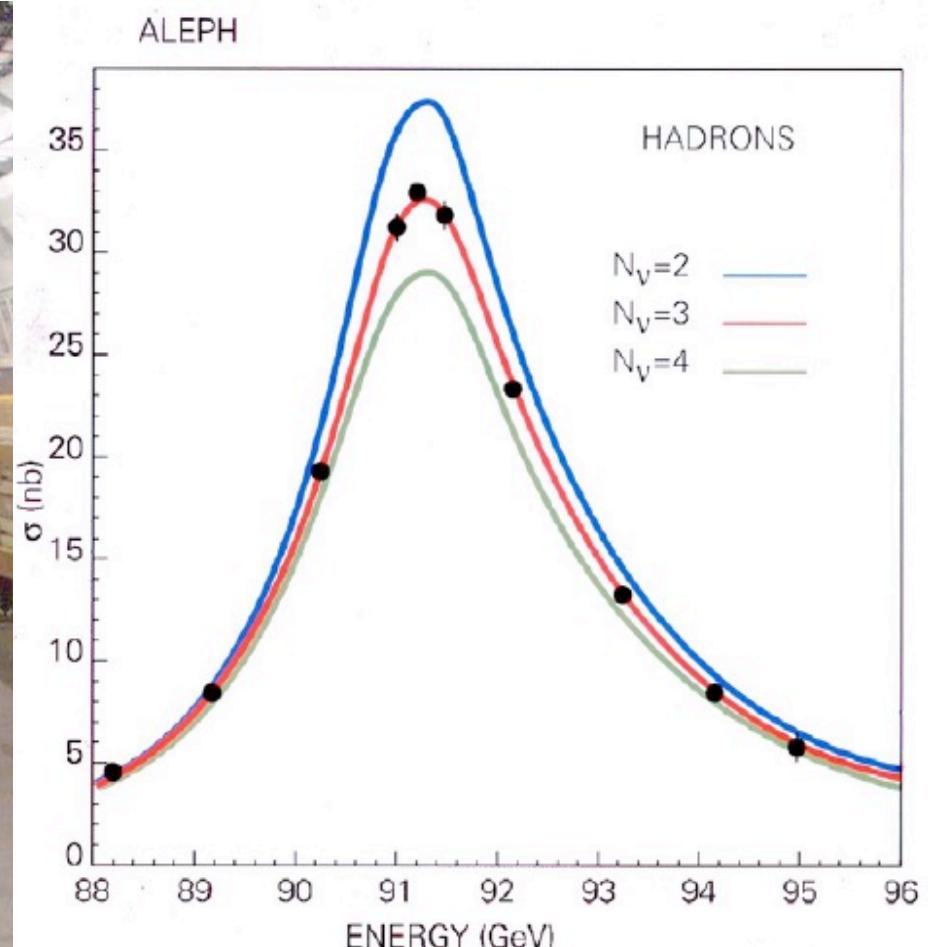


High precision W and Z measurement

Between 1985 and 1988 1.4 M m<sup>3</sup> were excavated at a depth of 100 m for a length of 27 km: the largest European civil engineering work of that period



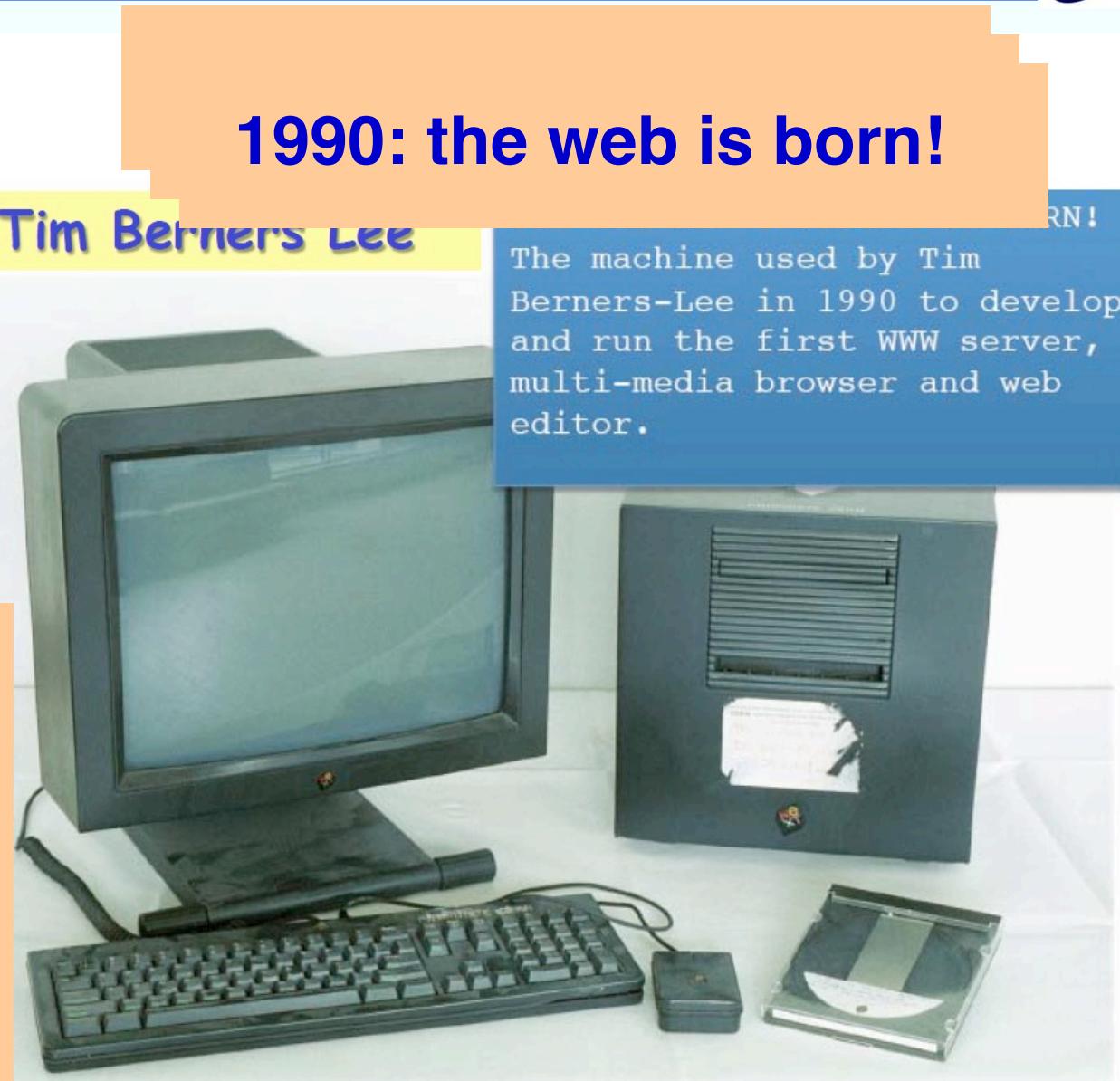
1989: there are only 3 generations of light neutrinos



# Tim Bernes-Lee invents the www



CERN gives a fundamental contribution to the informatics revolution



# The largest accelerator in the world



# 2012: discovery of the Higgs boson

## Quarks

<i>u</i>	<i>c</i>	<i>t</i>
up	charms	top

<i>d</i>	<i>s</i>	<i>b</i>
down	strange	bottom

## Forces

$Z$	$\gamma$
Z boson	photon
$W$	$g$
W boson	gluon

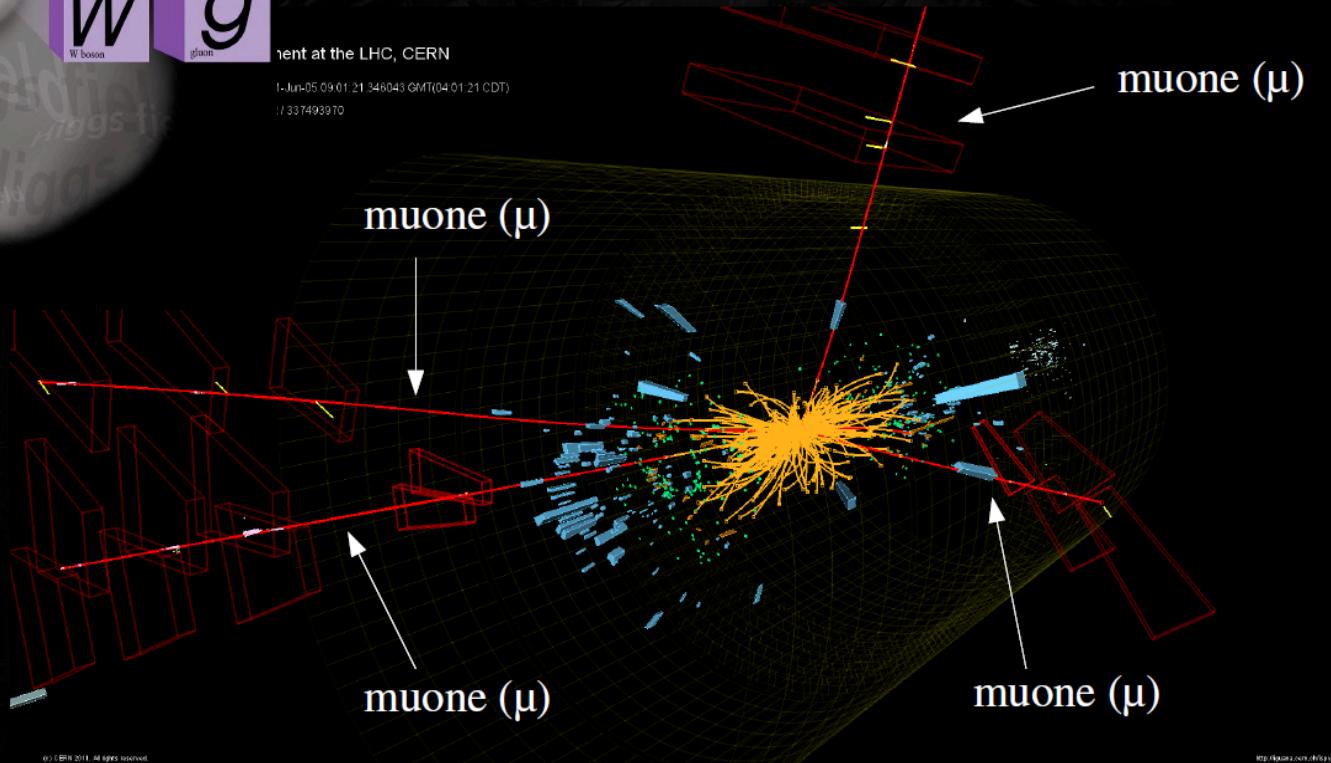
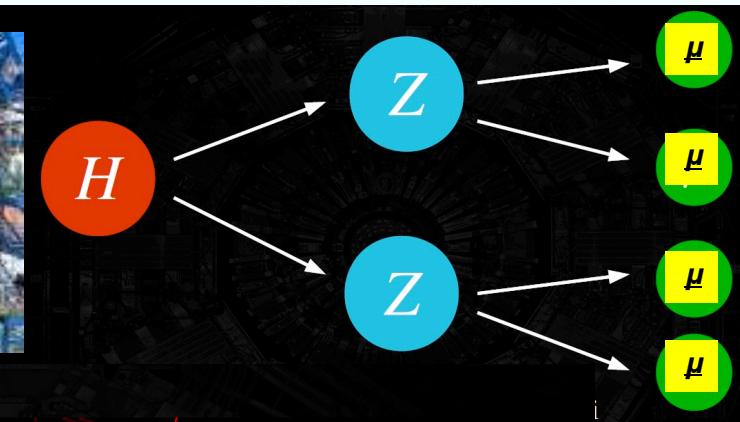
## Leptons



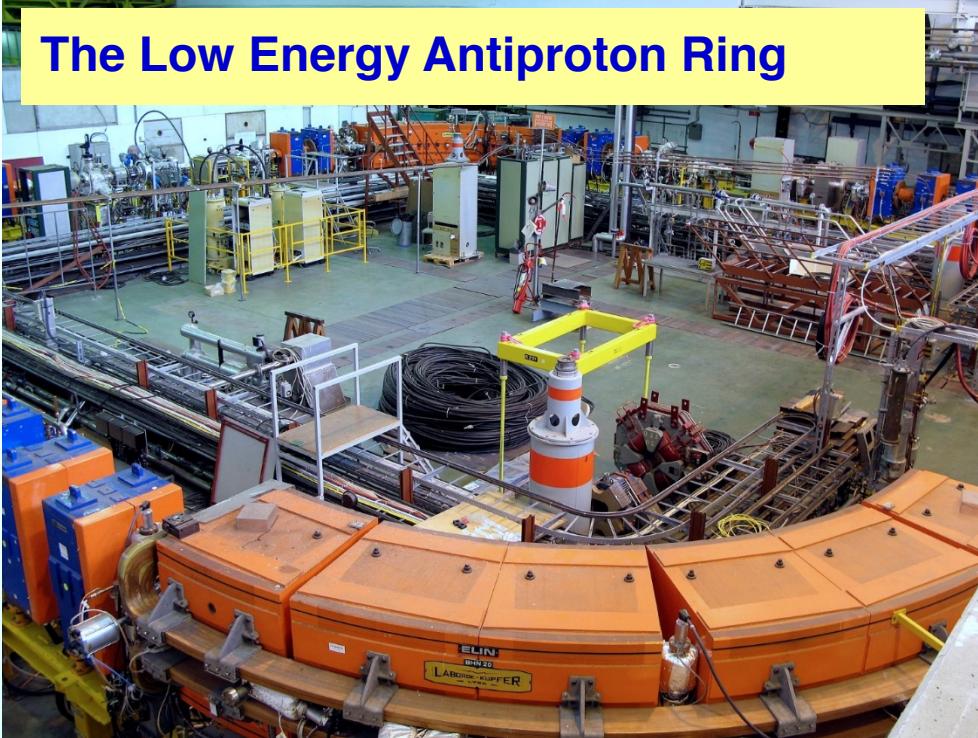
Peter Higgs,  
Nobel Prize 2013



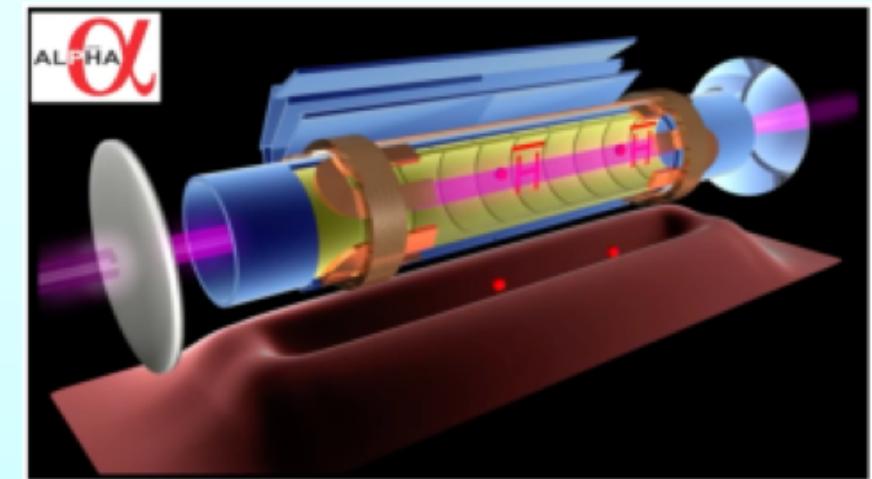
Experiment at the LHC, CERN  
1-Jun-05 09:01:21.346043 GMT(04:01:21 CDT)  
// 337493970



# The antihydrogen

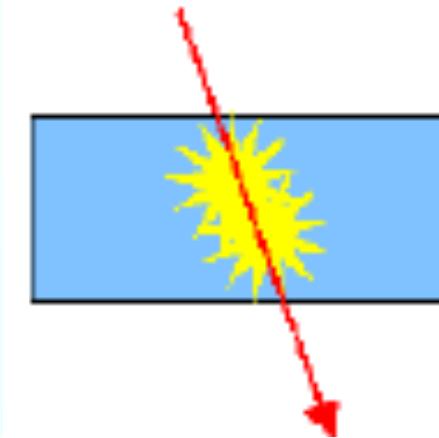


16 Dec. 2016  
first antimatter spectroscopy



# Scintillation detectors

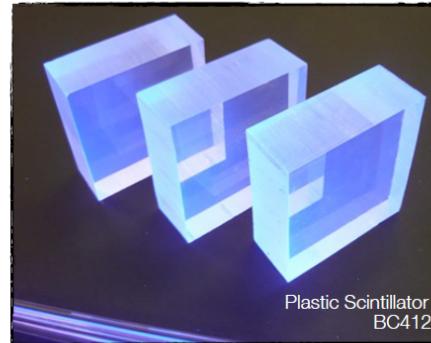
Particle Detection via Luminescence



## Scintillators – General Characteristics

### Principle:

$dE/dx$  converted into visible light  
Detection via photosensor  
[e.g. photomultiplier, human eye ...]



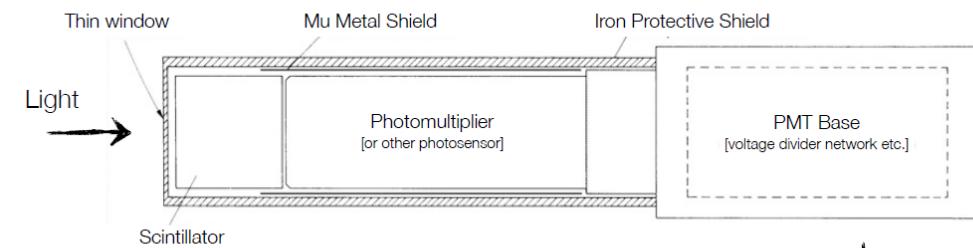
### Main Features:

Sensitivity to energy  
Fast time response  
Pulse shape discrimination

### Requirements

**High efficiency** for conversion of excitation energy to fluorescent radiation  
**Transparency** to its fluorescent radiation to allow transmission of light  
**Emission of light** in a spectral range detectable for photosensors  
**Short decay time** to allow fast response

## Scintillators – Basic Counter Setup

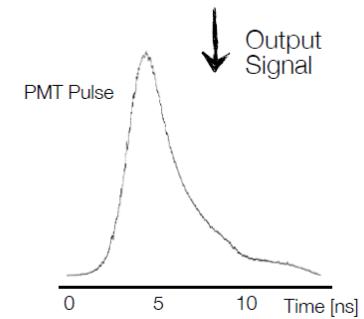


### Scintillator Types:

#### Photosensors

Photomultipliers  
Micro-Channel Plates  
Hybrid Photo Diodes  
Visible Light Photon Counter  
Silicon Photomultipliers

Organic Scintillators  
Inorganic Crystals  
Gases



# Scintillation detectors

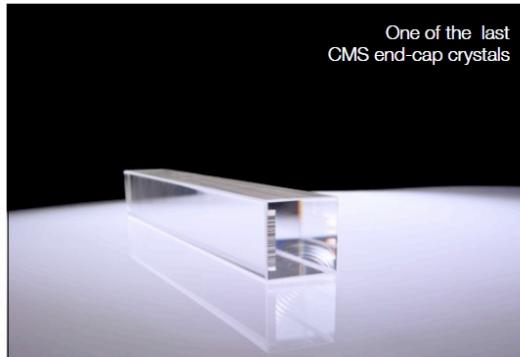
Particle Detection via Luminescence



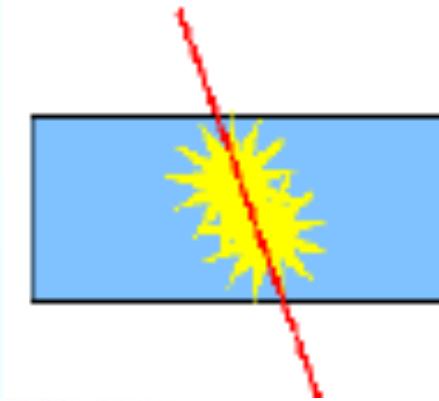
Inorganic Crystals



Example CMS  
Electromagnetic Calorimeter



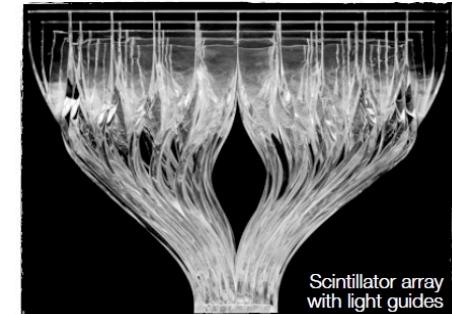
**Large light yield, good energy resolution**



Plastic and Liquid Scintillators

In practice use ...

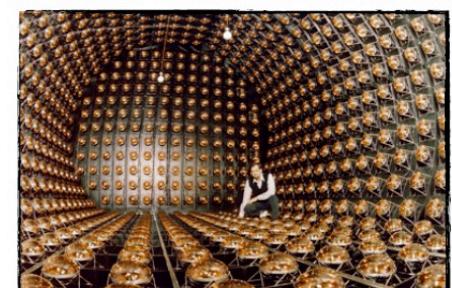
- solution of organic scintillators [solved in plastic or liquid]
- + large concentration of primary 'fluor'
- + smaller concentration of secondary 'fluor'
- + ...



Scintillator requirements:

- Solvable in base material
- High fluorescence yield
- Absorption spectrum must overlap with emission spectrum of base material

**Fast and cheaper**



## Photomultipliers

### Principle:

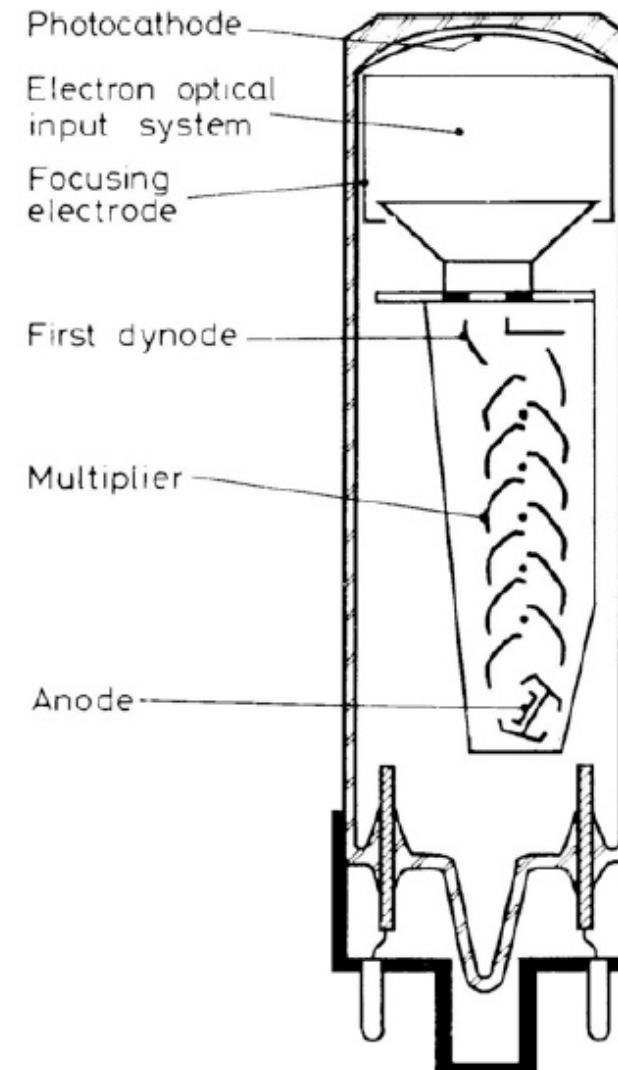
Electron emission  
from photo cathode

Secondary emission  
from dynodes; dynode gain: 3-50 [ $f(E)$ ]

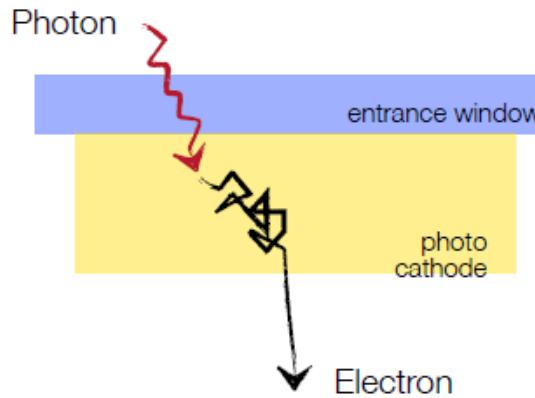
Typical PMT Gain:  $> 10^6$   
[PMT can see single photons ...]



PMT  
Collection



$\gamma$ -conversion  
via photo effect ...

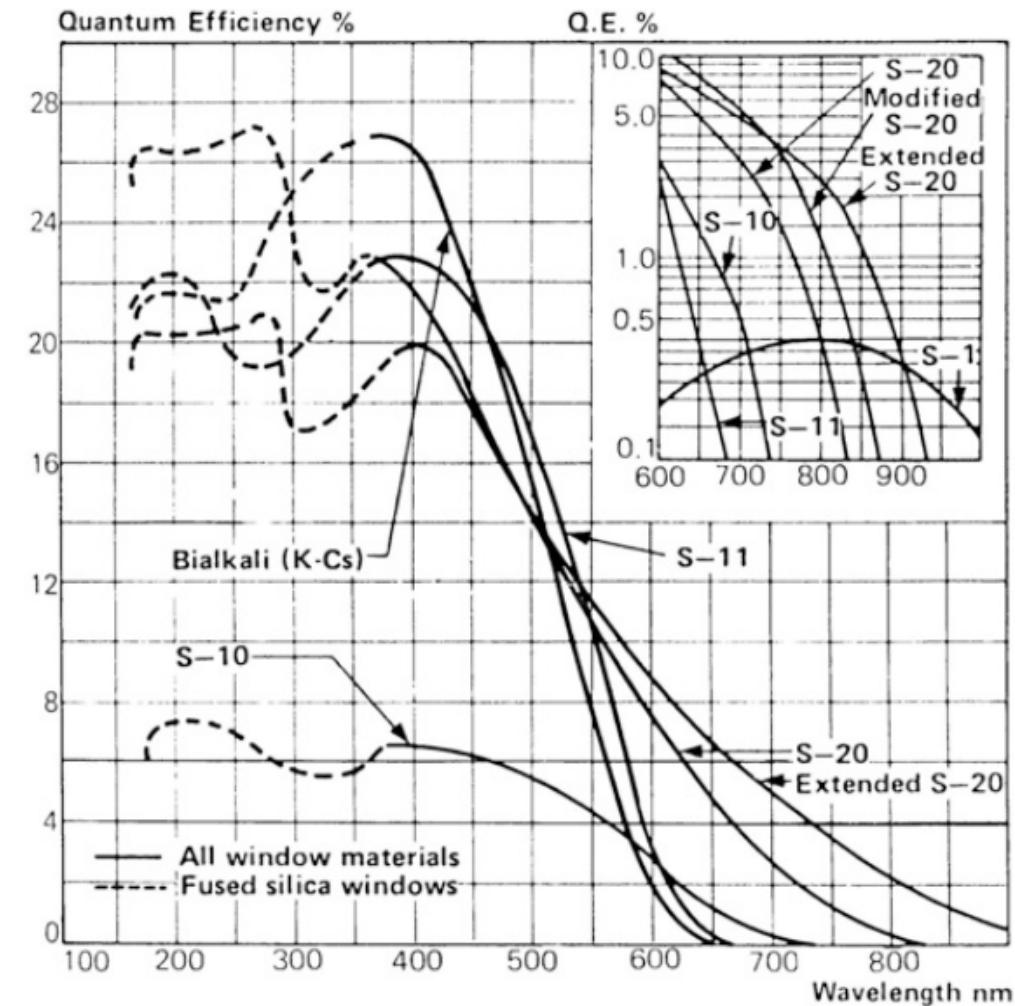


3-step process:

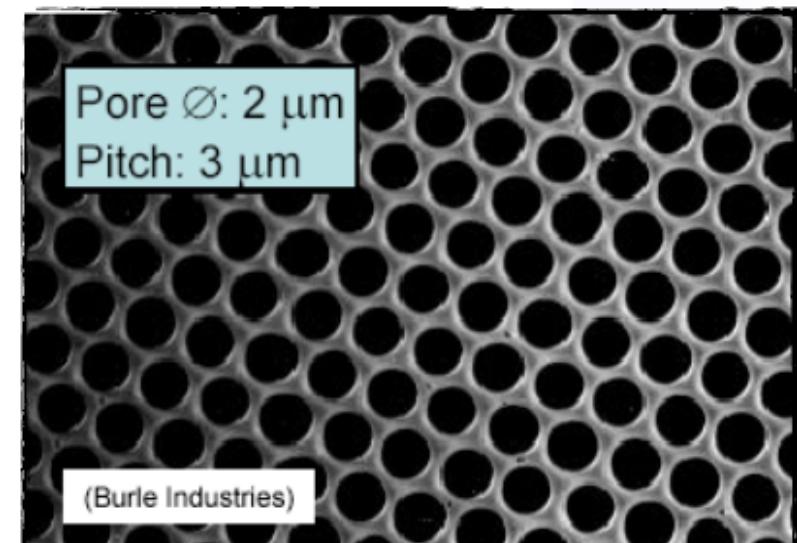
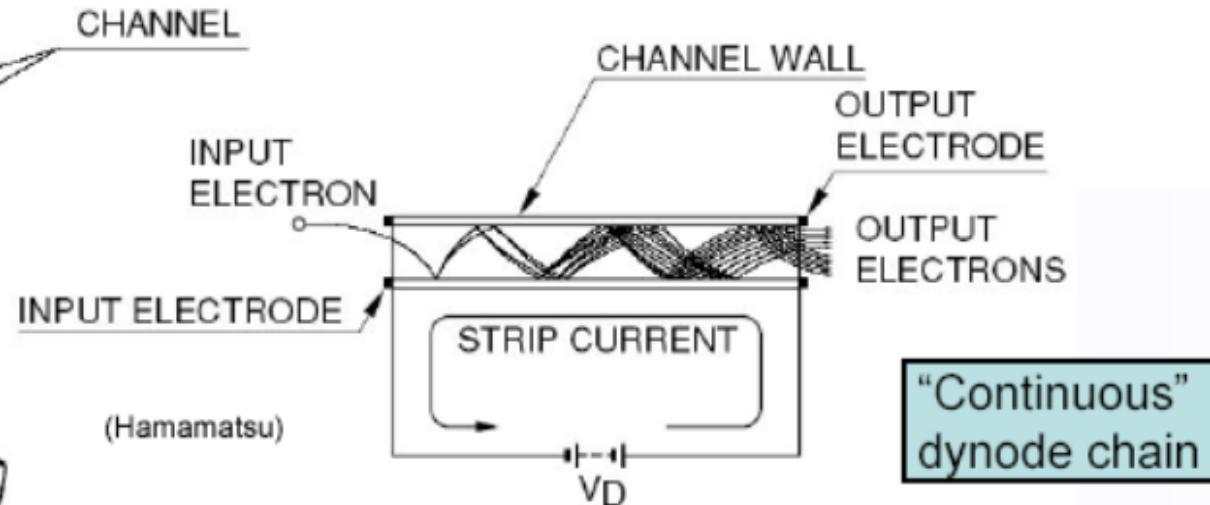
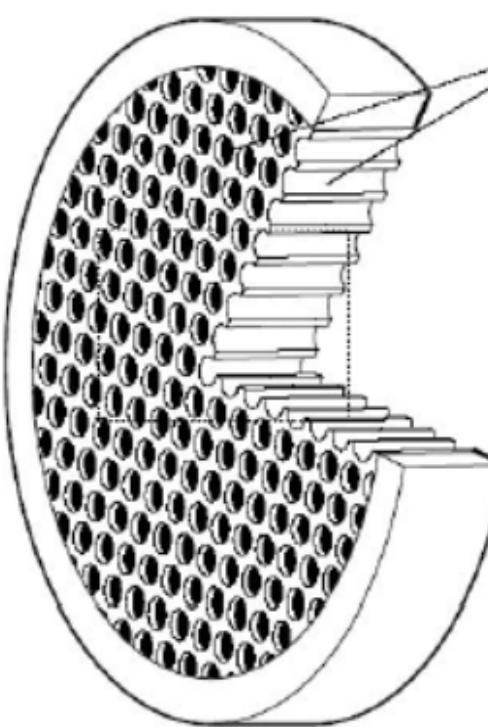
- Electron generation via ionization
- Propagation through cathode
- Escape of electron into vacuum

Q.E.  $\approx$  10-30%  
[need specifically developed alloys]

Bialkali: SbRbCs; SbK<sub>2</sub>Cs



# Micro-Channel Plates



"2D Photomultiplier"

Gain:  $5 \cdot 10^4$

Fast signal [time spread  $\sim 50$  ps]

B-Field tolerant [up to 0.1T]

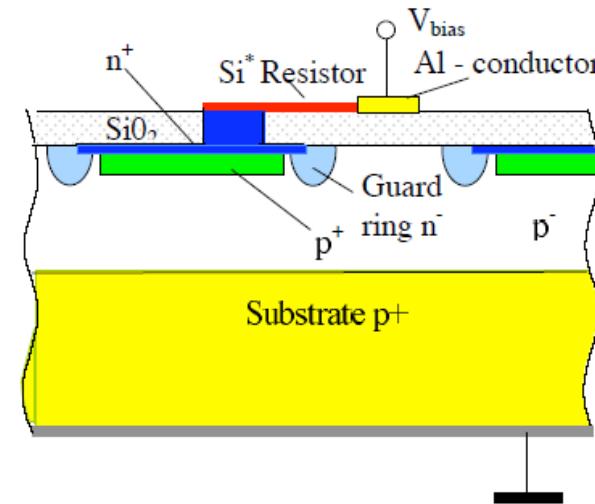
## Silicon Photomultipliers

### Principle:

Pixelized photo diodes  
operated in Geiger Mode

Single pixel works as a binary device

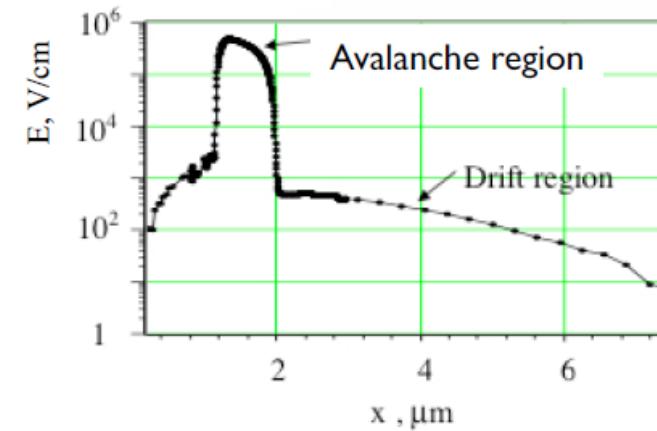
Energy = #photons seen by  
summing over all pixels



### Features:

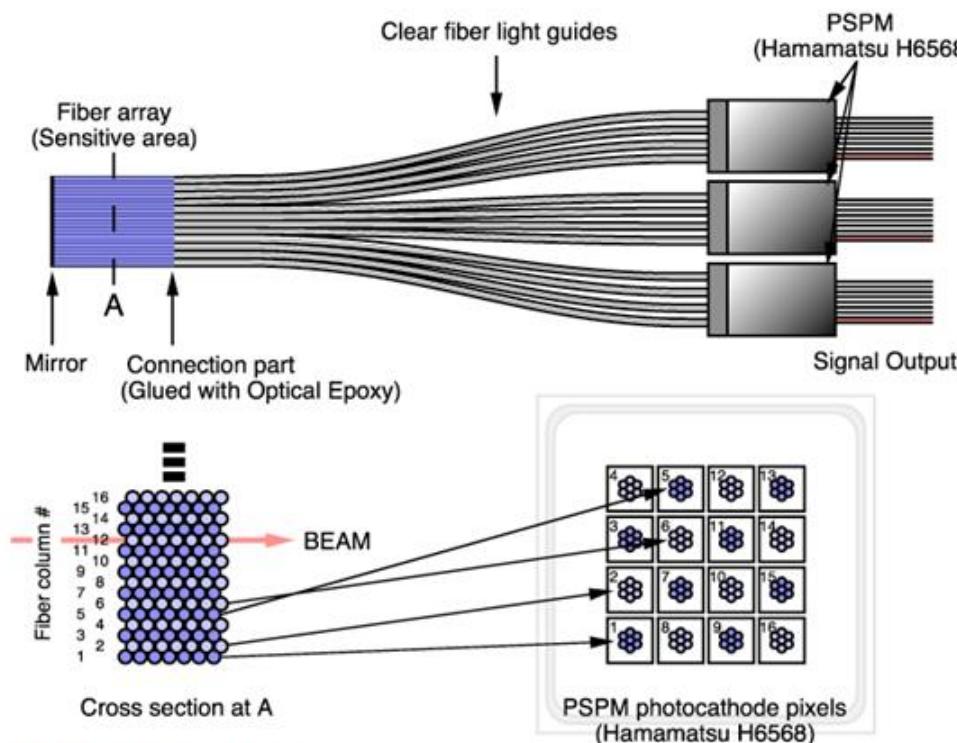
Granularity	:	$10^3$ pixels/mm <sup>2</sup>
Gain	:	$10^6$
Bias Voltage	:	< 100 V
Efficiency	:	ca. 30 %

Insensitive to magnetic fields!  
Works at room temperature ...





9 stations: 21 coordinates



Sensitive area:

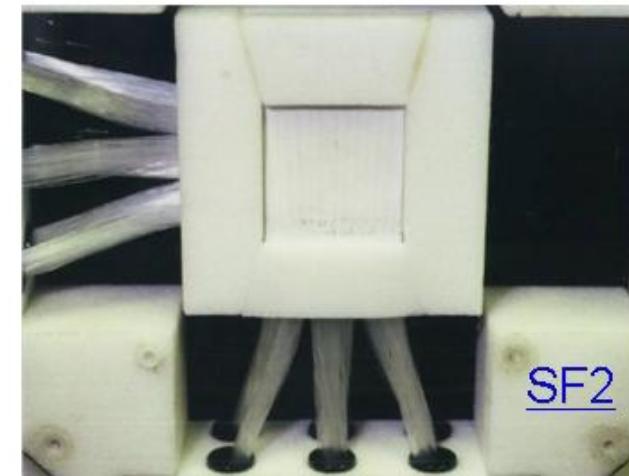
7-layers of Kuraray SCSF-78MJ 0.5 mm Ø

*Rate capability > 5 MHz  
per channel*

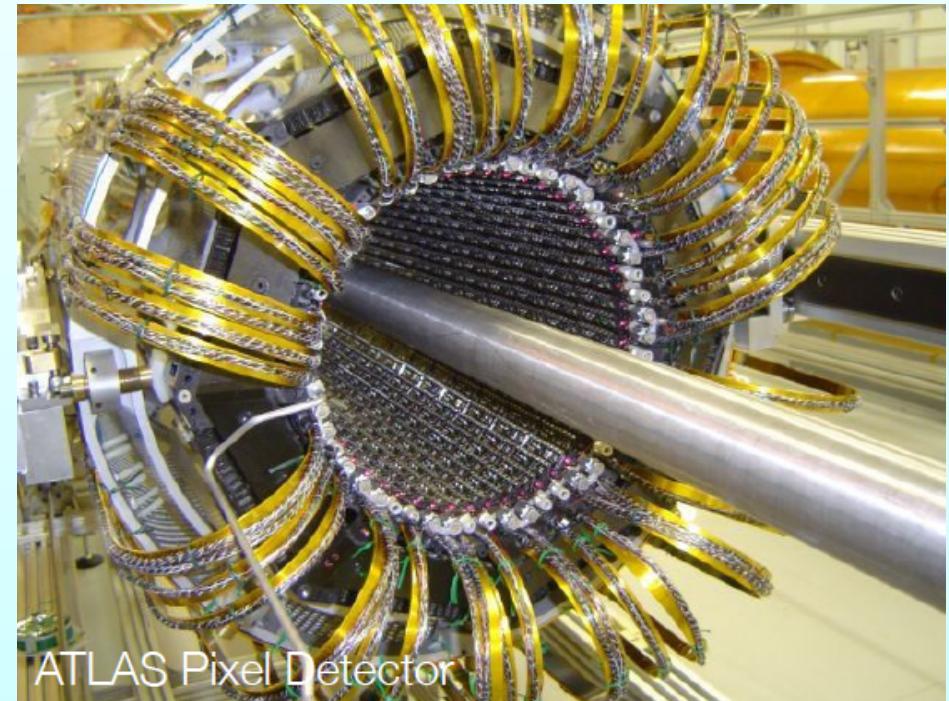
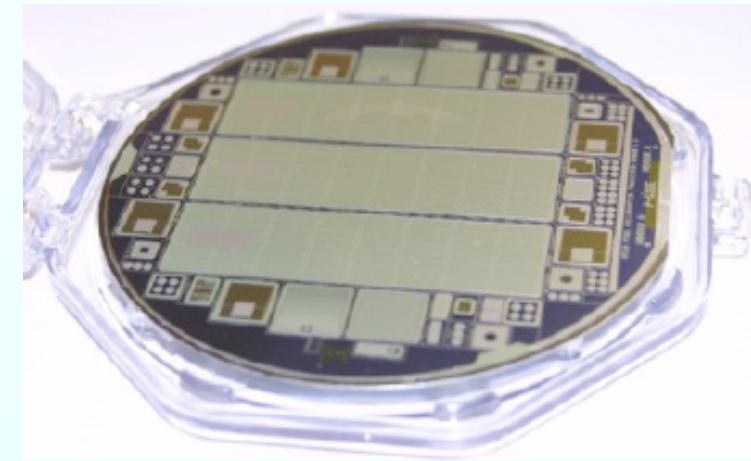
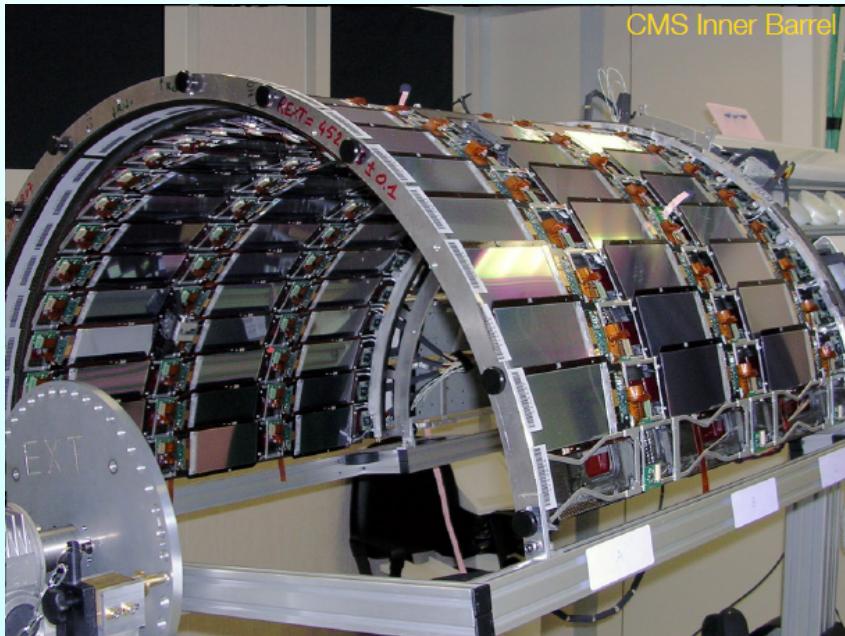
*Efficiency: 99%*

*Space resol. 130 – 250 μm*

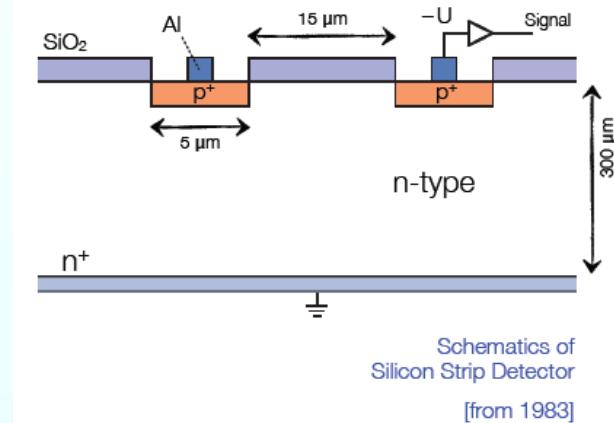
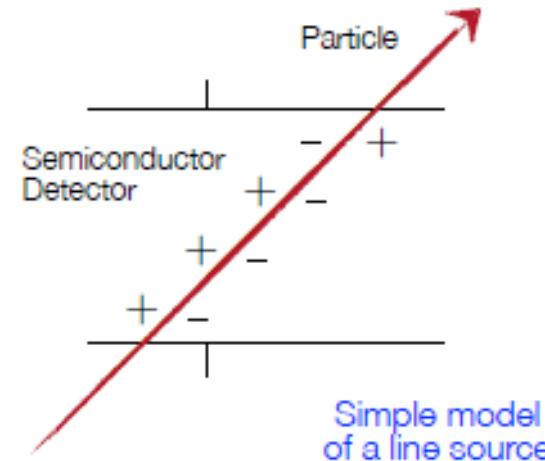
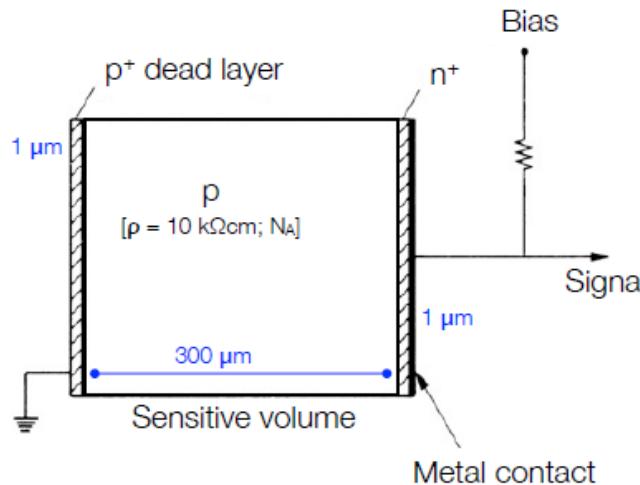
*Time resol. < 400 ps*



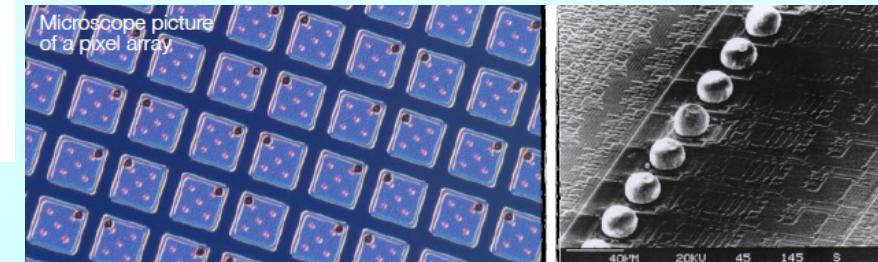
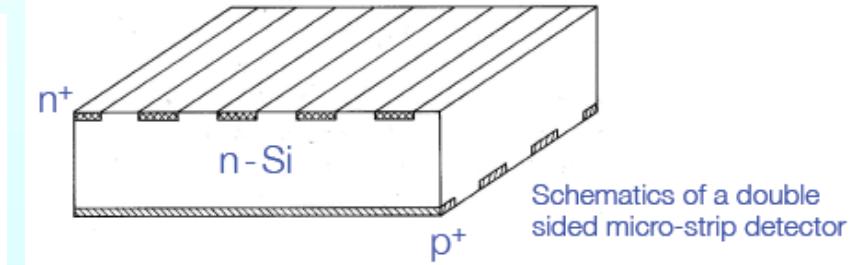
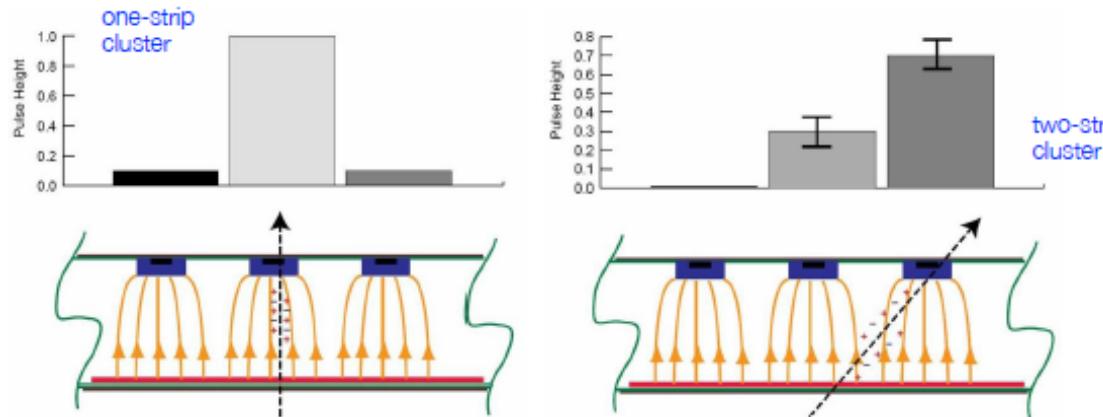
# Si detectors

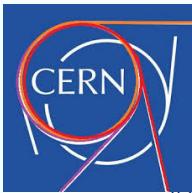


# Si microstrips

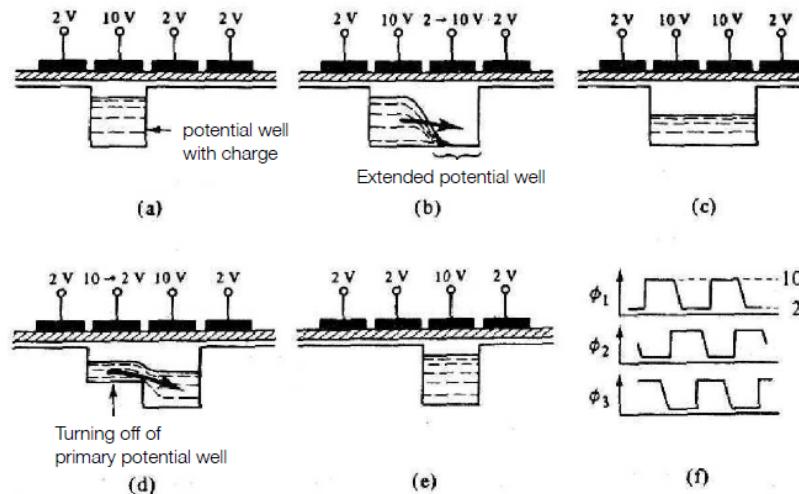
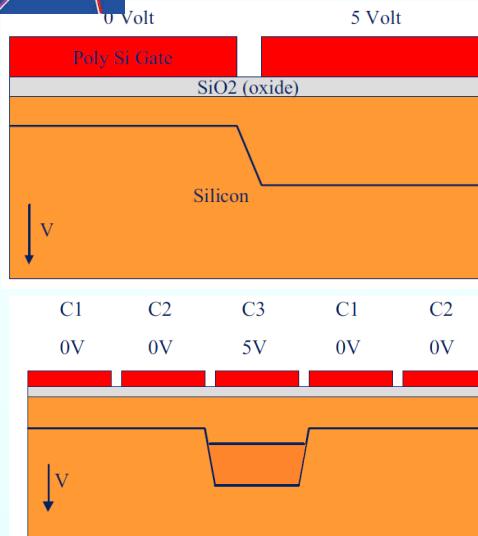


High resistive n-type silicon onto which p<sup>+</sup> diode strips with aluminum contacts are implanted





# CCD



The Nobel Prize in Physics 2009

"for groundbreaking achievements concerning the transmission of light in fibers for optical communication"

"for the invention of an imaging semiconductor circuit – the CCD sensor"



Photo: U. Montan

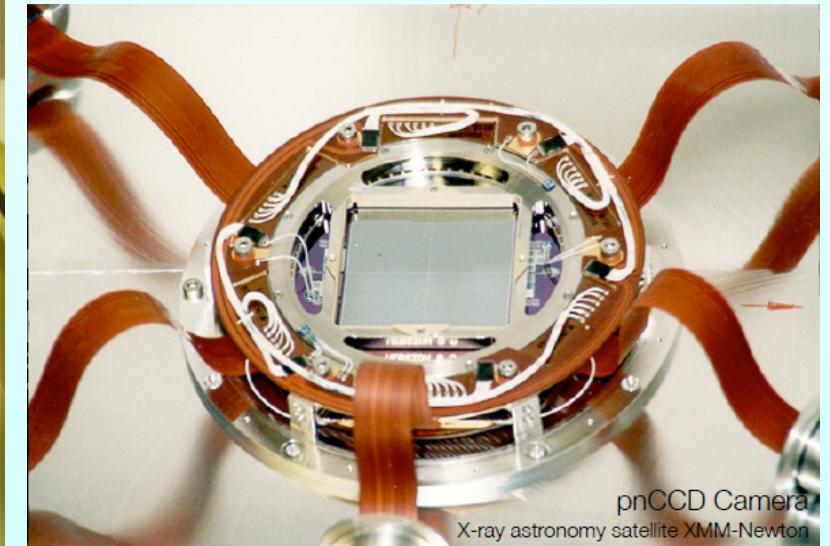
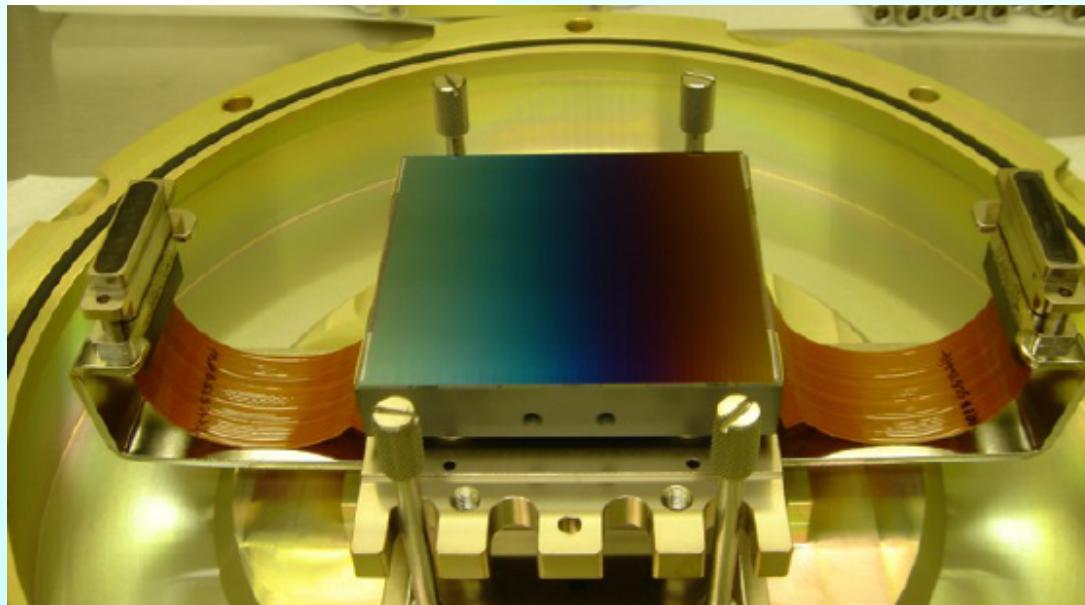
Photo: U. Montan

Photo: U. Montan

Charles K. Kao

Willard S. Boyle

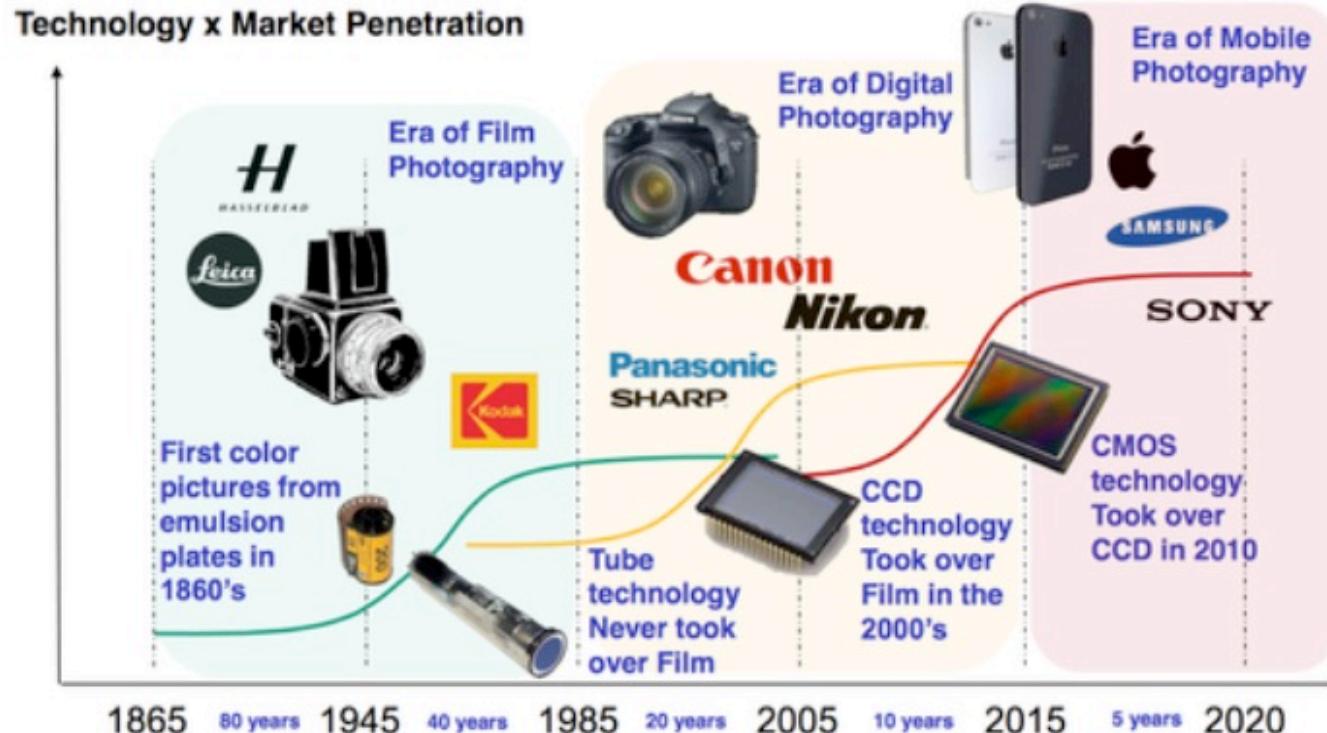
George E. Smith



pnCCD Camera  
X-ray astronomy satellite XMM-Newton

## CMOS Image Pixel Sensors

- While 1980s were dominated by CCDs (camcorder market)
- The 1990s/2000s have shown an increasing demand for CMOS imaging sensors due to the camera phone market

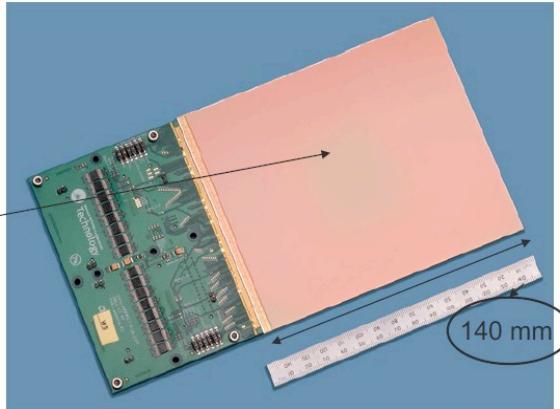


# ALICE inner tracker upgrade

## MAPS (Monolithic Active Pixel Sensors) for Imaging and More

Many developments in the field of CMOS imaging sensors and MAPS in general within the community!

Example:  
Wafers scale (8") imaging sensor developed by the RAL team (stitched)

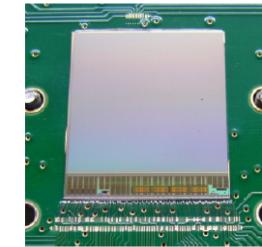
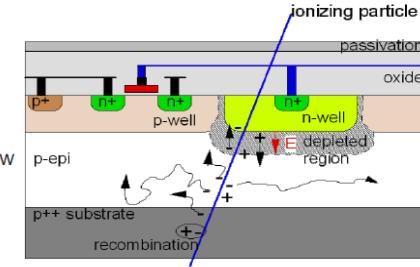


N. Guerrini, RAL, 5<sup>th</sup> school on detectors, Legnaro, April 2013

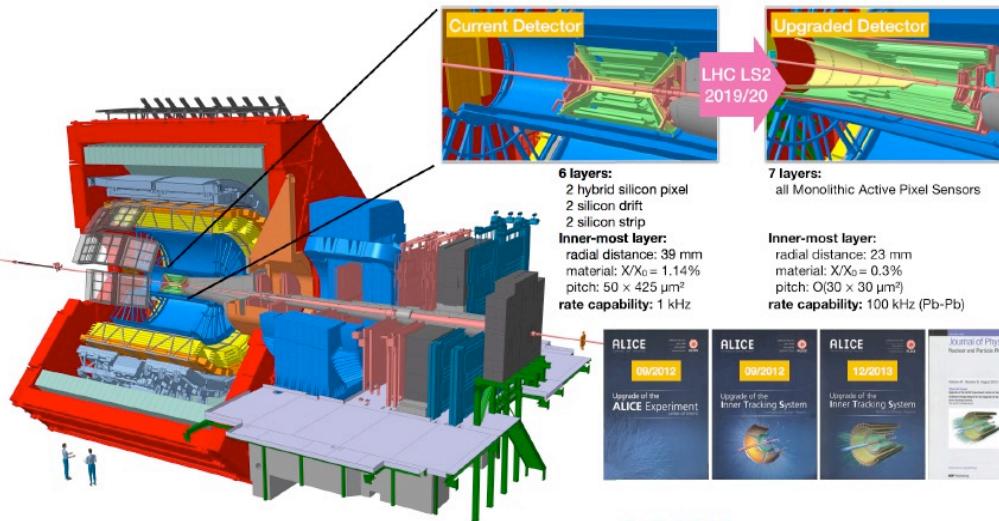


Developments lead by IPHC created a number of monolithic pixel sensors of the MIMOSA family:

- Epitaxial wafers with collection diode and few transistors per cell (size  $\sim 20 \times 20 \mu\text{m}^2$ )
- 0.35  $\mu\text{m}$  CMOS technology with only one type of transistor (NMOS)
- Rolling shutter architecture (readout time  $O(100 \mu\text{s})$ )
- **Charge collection mostly by diffusion**
- Limited radiation tolerance ( $< 10^{13} \text{n}_{\text{eq}} \text{cm}^{-2}$ )



ULTIMATE chip for STAR HFT (IPHC Strasbourg)



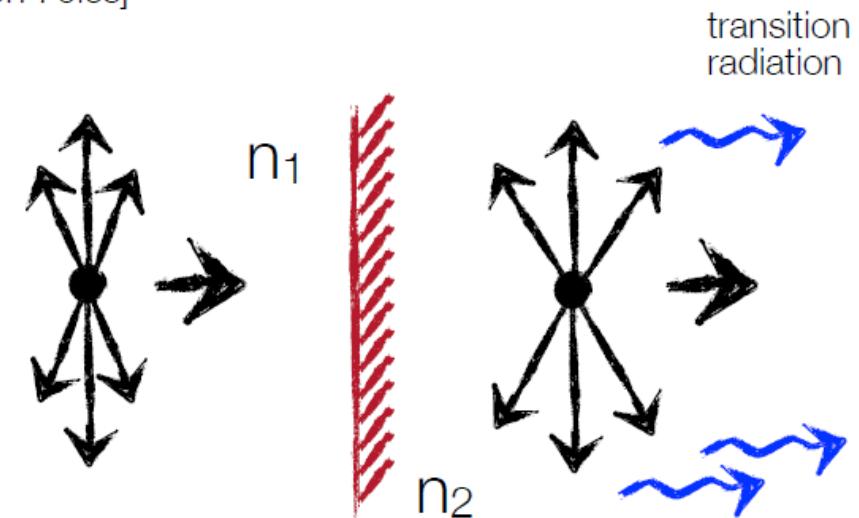
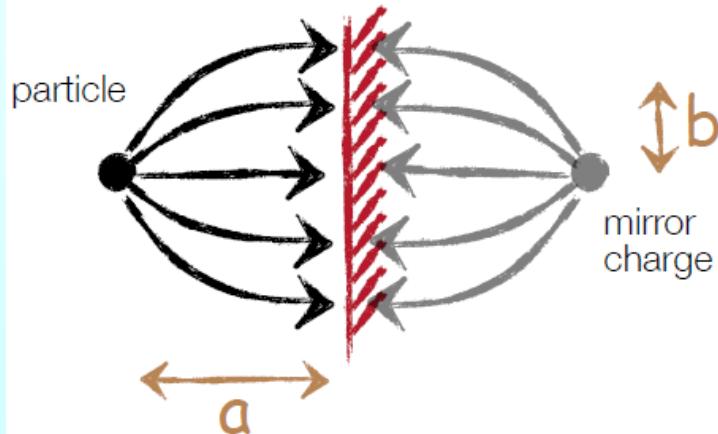
# Transition radiation detectors

Transition radiation occurs if a relativistic particle (large  $\gamma$ ) passes the boundary between two media with different refraction indices ...

[predicted by Ginzburg and Frank 1946; experimental confirmation 70ies]

Effect can be explained by re-arrangement of electric field ...

Simple model: Electron moves towards conducting plate ...



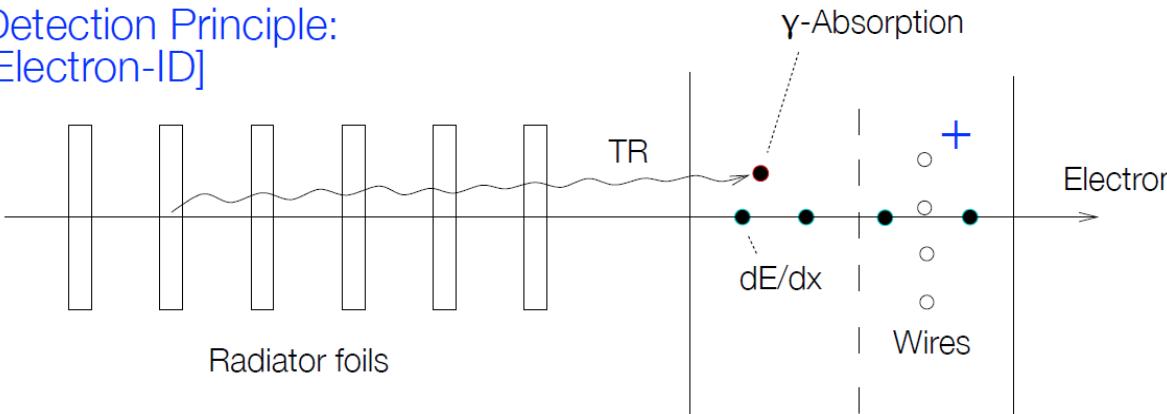
with dipole moment:

$$\vec{p} = 2e\vec{a}$$

and electric field at surface

$$|\vec{E}_n| \propto \frac{\vec{p}}{(a^2 + \rho^2)^{\frac{3}{2}}}$$

Detection Principle:  
[Electron-ID]



$$\text{CH}_2: \hbar\omega_p = 20 \text{ eV}; \gamma = 10^3$$

$$[\text{Air: } \hbar\omega_p = 0.7 \text{ eV}]$$

$$D = 10 \mu\text{m}$$

[ $d > D$ : absorption dominates]

Straw Tube Tracker  
with interspace filled with foam  
→ Tracking & transition radiation

