



Searching for similarities and anomalies in a pool of galaxy images

Matias Carrasco Kind

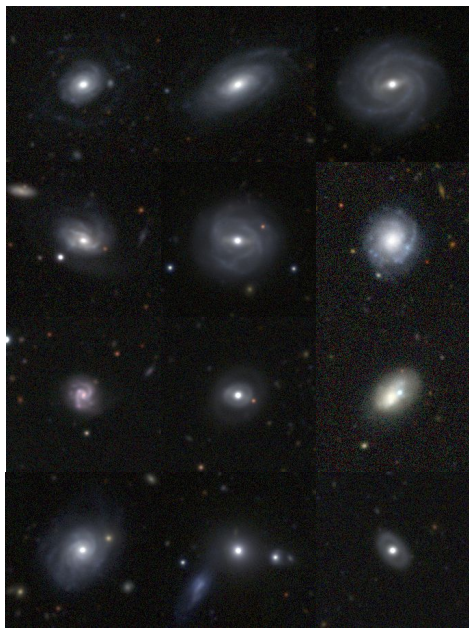
Senior Research Scientist, National Center for Supercomputing Applications

Assistant Research Professor, Astronomy

University of Illinois at Urbana-Champaign

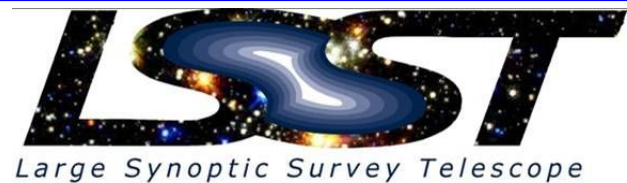
Artificial Intelligence for Data Discovery and Reuse, May 13-15 2019
Carnegie Mellon University, Pittsburgh, PA

Motivation



Astronomy is just one example where image exploration needs to be automated.

Large catalogs, Large number of images, many unexpected objects/problems → Anomaly detection



- In operations 2021
- Every night for 10 years
 - 15 TB per night
- 18 billions objects (first year), ~40 billions by the end of survey
- ~1500 images per night
- Stream and static data
- Target to capture new physics (moving and variable objects)



- More than 500 nights of observation over 5 years, 2TB per night
- 500 millions cataloged galaxies and 100 millions stars
- Many open problems: Systematics, new objects, new physics, etc.
- Almost completed

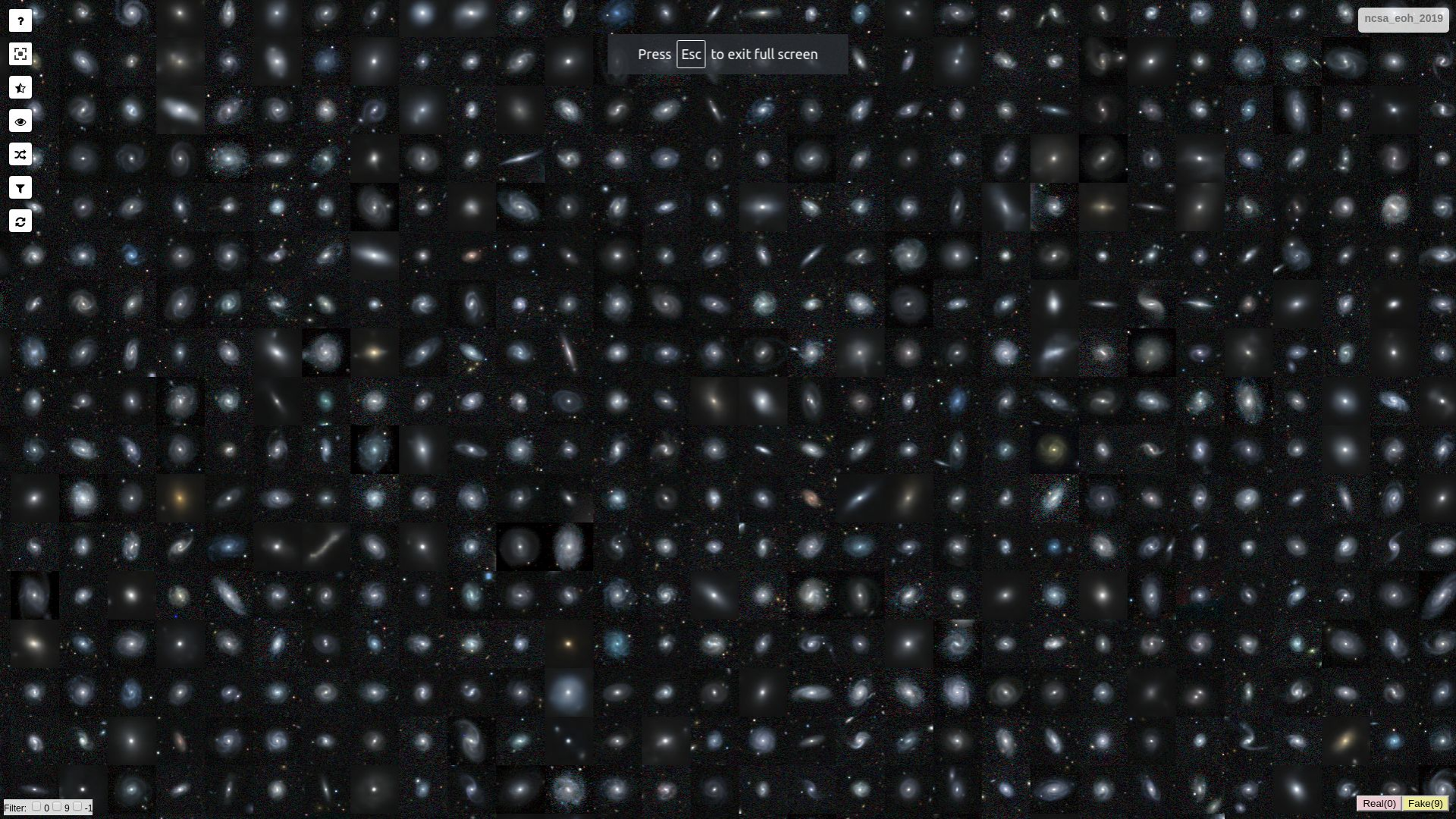
Current (personal) data discovery challenges

- Visualize large set of galaxy (or other) images
- Quickly classify images for AI using multiple experts
- Compress the important information in an efficient way
- Quickly search images by (dis)similarity (several science cases)
- Find anomalies in an image data set (new phenomena, errors, unrepresentative samples)

Not covered here

- Generate and sample realistic fake images based on a training for modeling and Monte-Carlo Sampling
- Generate and sample realistic fake images based on a training in a controlled manner (with a prior)

Press **Esc** to exit full screen

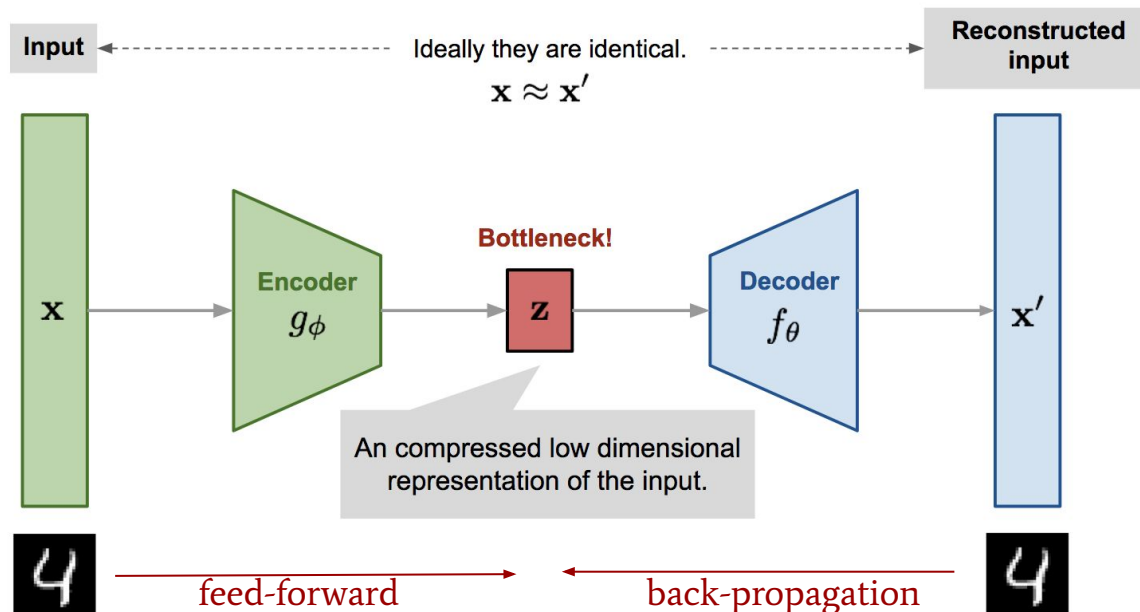


Galaxy Image Exploration and Classification

- Image Exploration
- Resize is done dynamically
- Quick Classification/Label
- Works fine with 10,000 images
- Individual classifications are saved and aggregated
- Keyboard control

<https://github.com/mgckind/cutouts-explorer>

Autoencoders review



- Around since the 80's
- Data compression
- Anomaly detection
- Denoising
- Regular Machine Learning
- PCA

$$z = g_\phi(x)$$

$$x' = f_\theta(g_\phi(x))$$

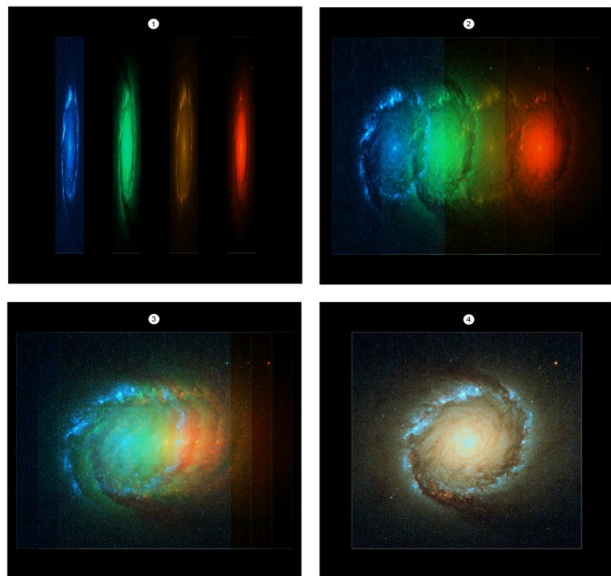
$$\mathcal{L}(x, x') + reg.$$

$$\ell_1 = \lambda \sum_i |a_i^{(h)}|$$

$$\mathcal{L}(x, x') = ||x - x'||^2$$

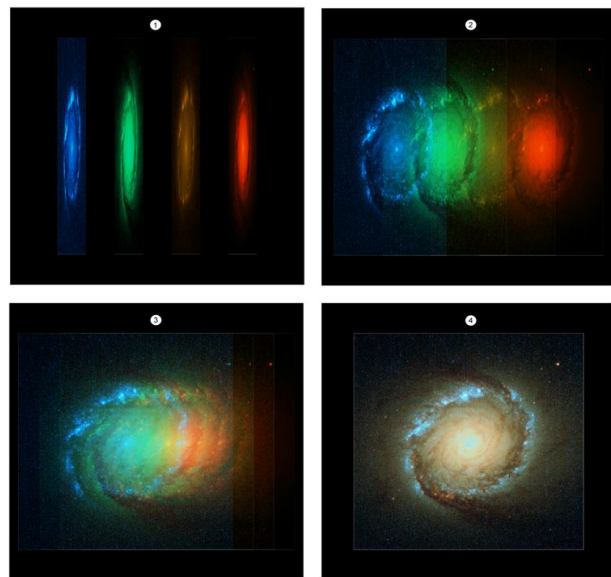
Many variants: Sparse AE, Contractive AE, Stacked AE, etc...

The latent space

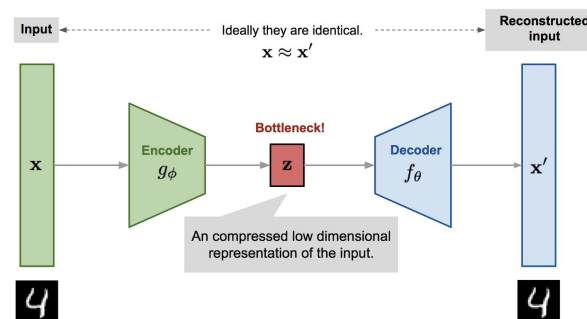


Credit: esa

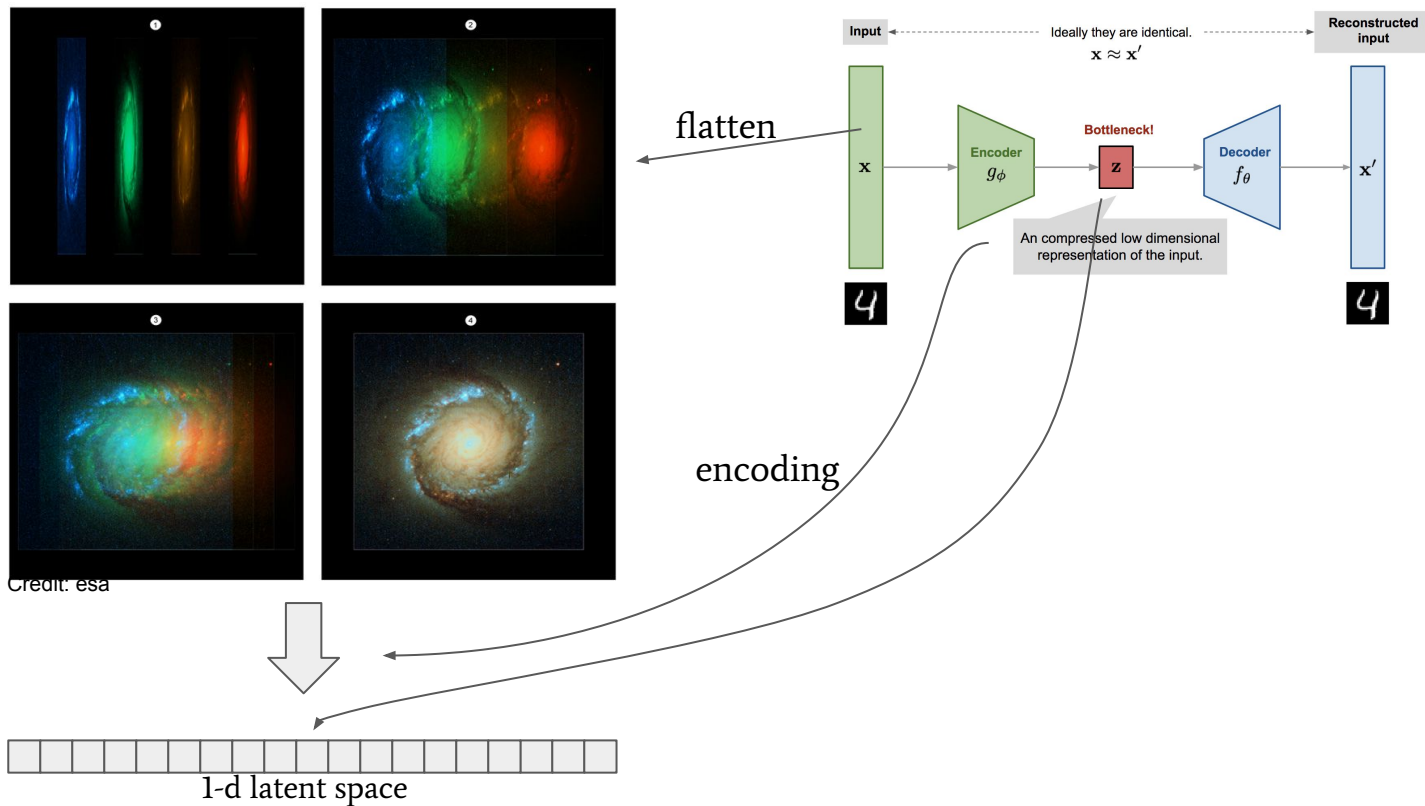
I The latent space



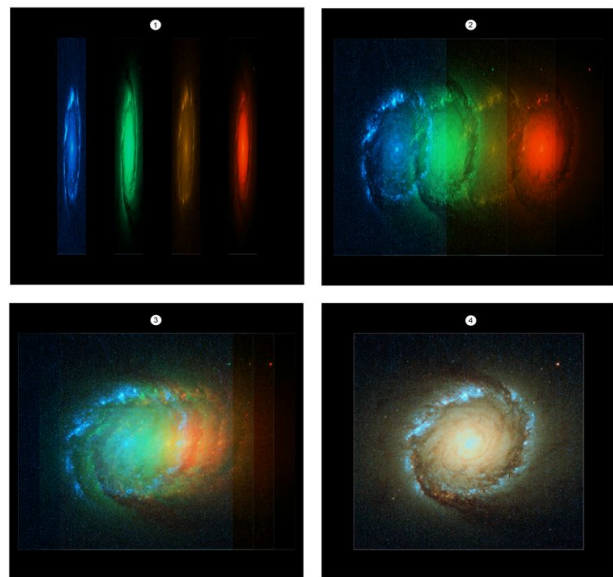
Credit: esa



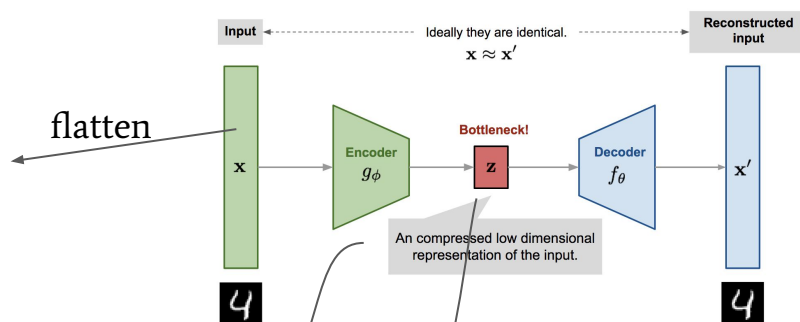
The latent space



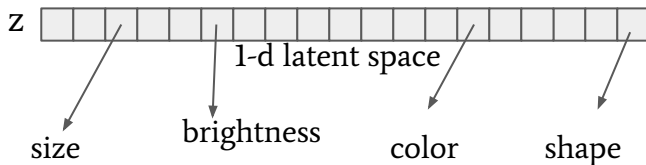
I The latent space



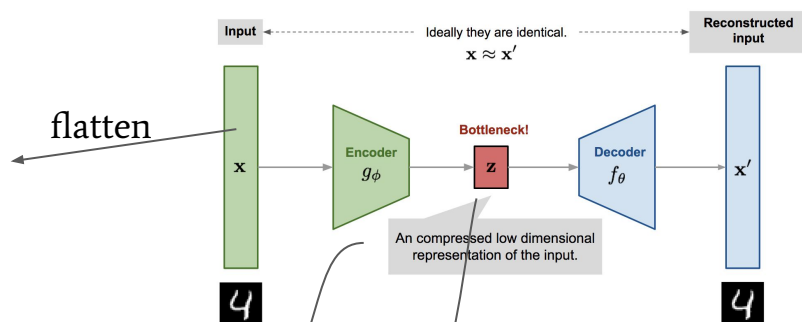
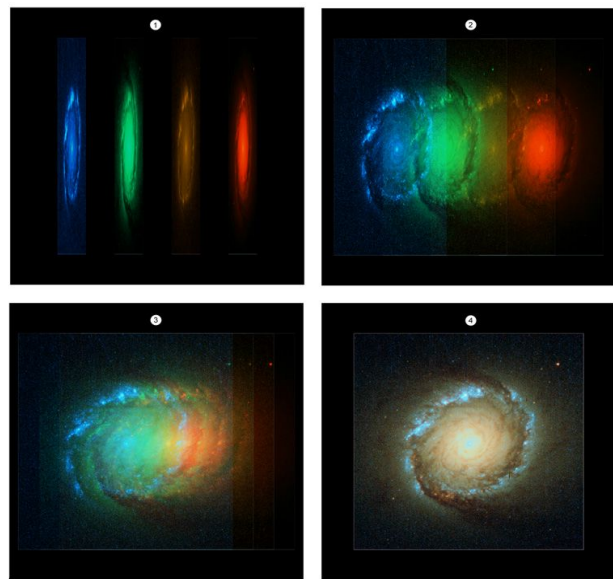
Credit: esa



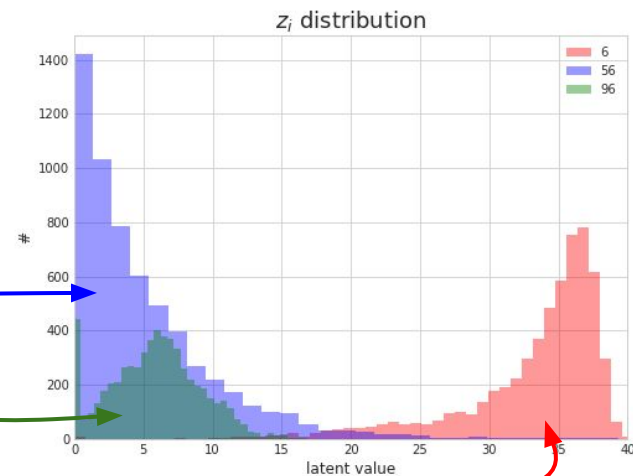
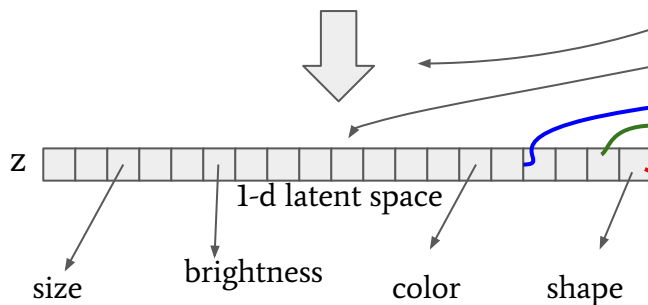
encoding



The latent space



Each galaxy image is encoded to a unique 1-d vector

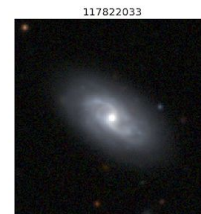
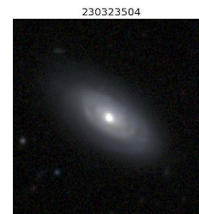
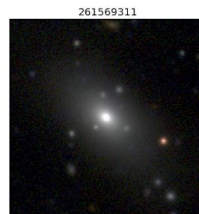
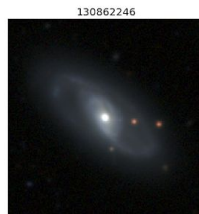
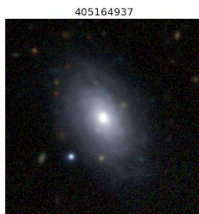
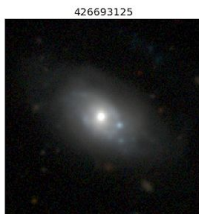


Similarity ranking

Using standard ML in latent space to look for neighbors, outliers, etc.

Compress images from $220 \times 220 \times 3$ pixels to 100-vector (2000x), for fast similarity search, anomaly detection, etc...

No need decoder (only for Loss)

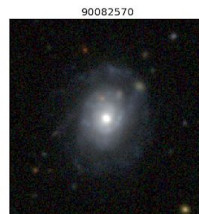
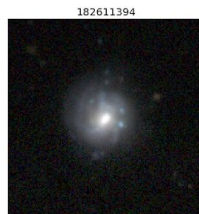
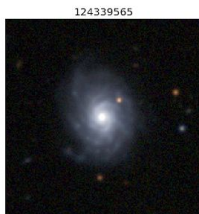
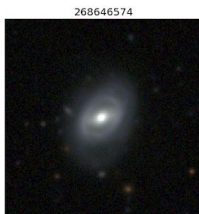


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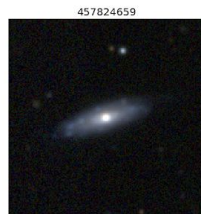
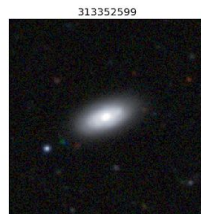
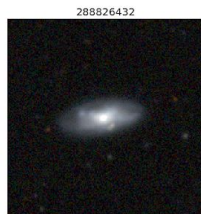
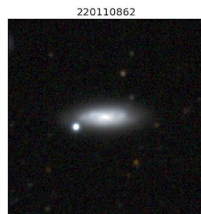
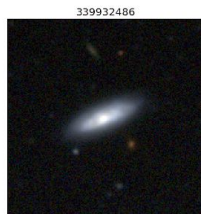
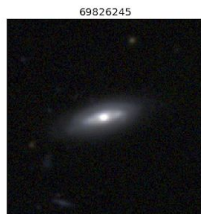
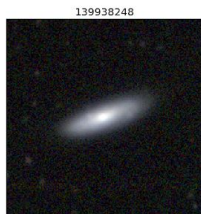
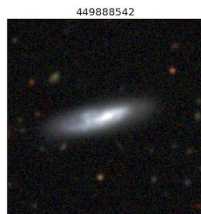
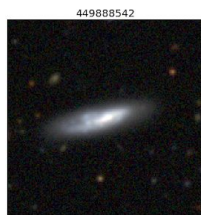


I Similarity ranking

Using standard ML in latent space to look for neighbors, outliers, etc.

Compress images from $220 \times 220 \times 3$ pixels to 100-vector (2000x), for fast similarity search, anomaly detection, etc...

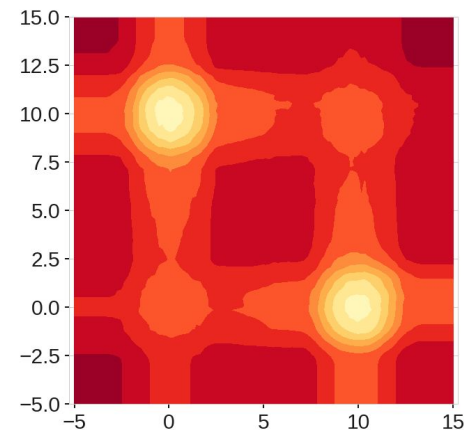
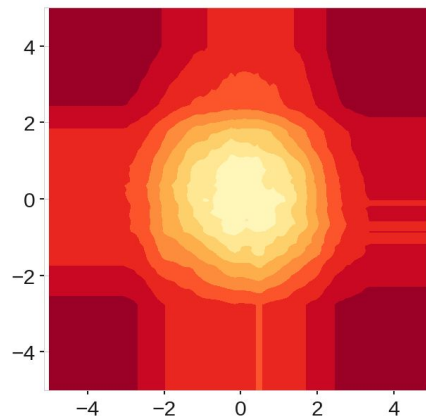
No need decoder (only for Loss)



Anomaly detection with Extended Isolation Forest

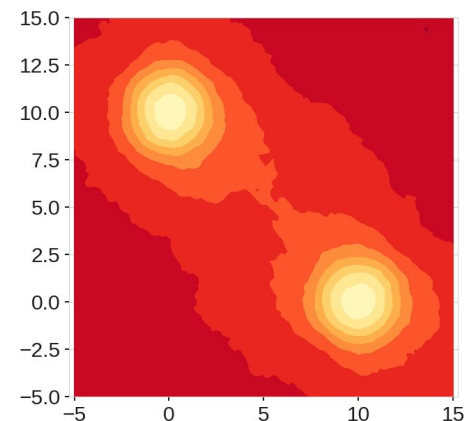
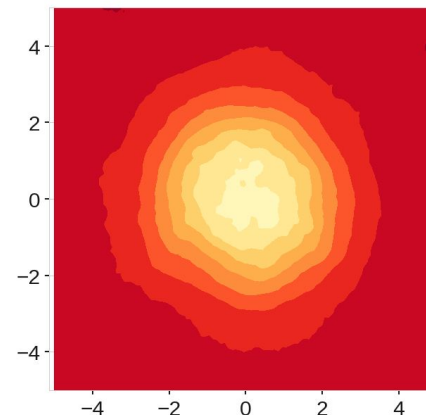
Isolation Forest:

- ✓ Model free
- ✓ Computationally efficient
- ✓ Readily application to high dimensional data
- ✗ Inconsistent scoring seen in score maps



Extended Isolation Forest:

- ✓ Model free
- ✓ Computationally efficient
- ✓ Readily application to high dimensional data
- ✓ Consistent scoring

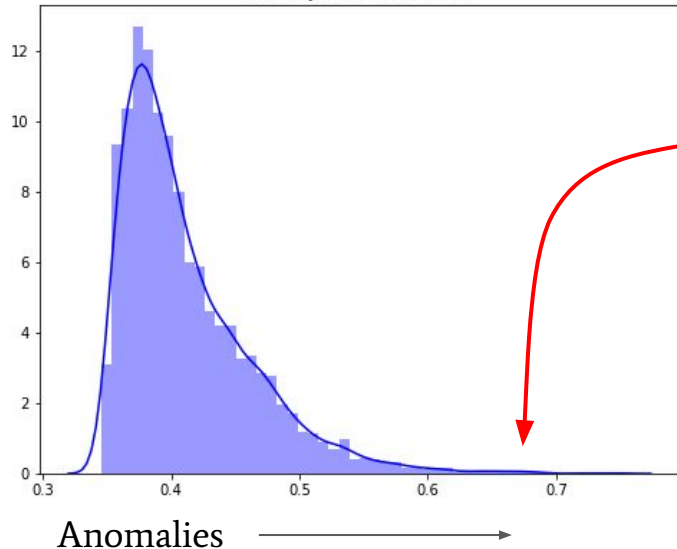


Hariri, Carrasco-Kind, Brunner, 2019, arXiv: 1811.02141

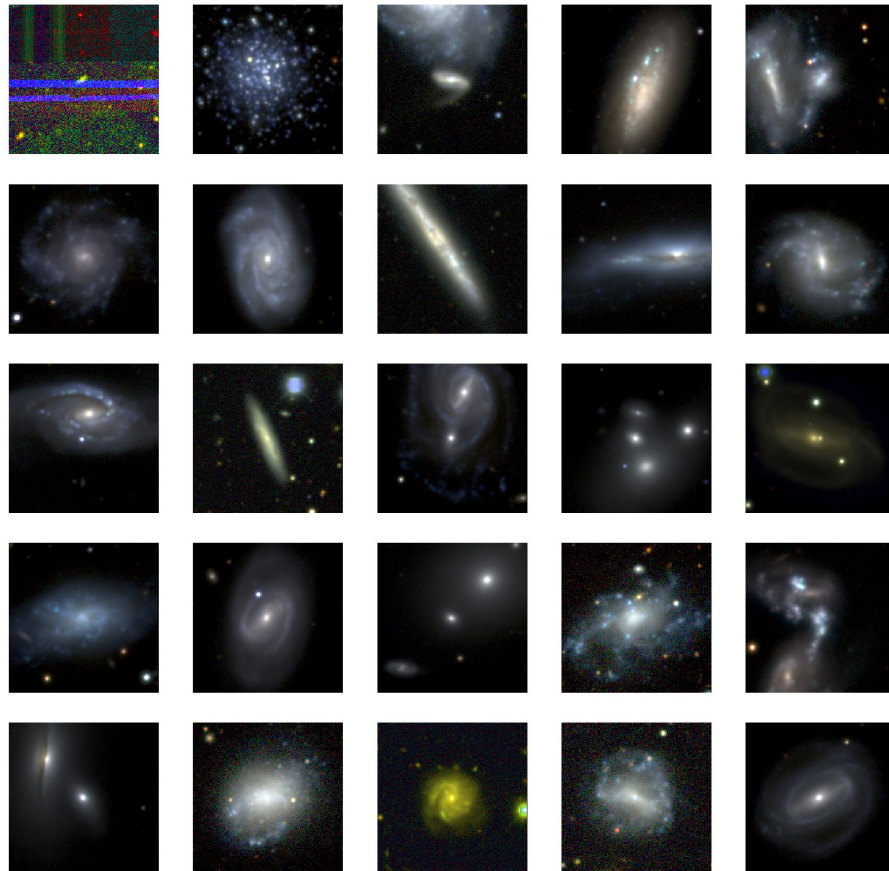
<https://github.com/sahandha/eif>

Anomaly detection with EIF in high dimensional latent space

Anomaly Score Distribution



Top 25
anomalies

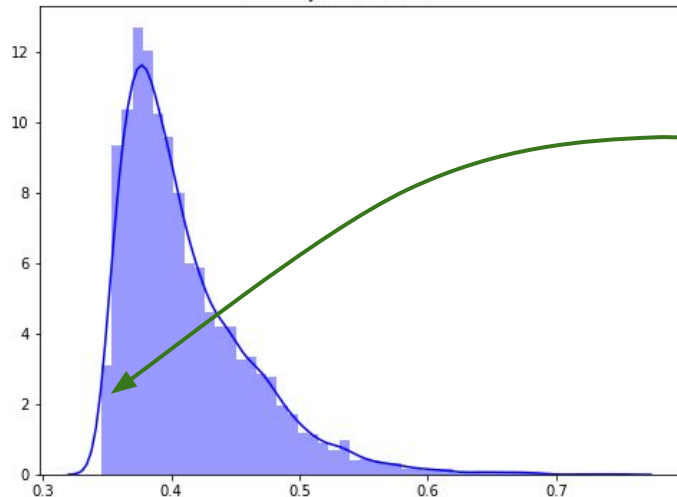


The EIF algorithm produces an anomaly score that can be used to select outlier galaxies.

Errors, special cases, unrepresented galaxies

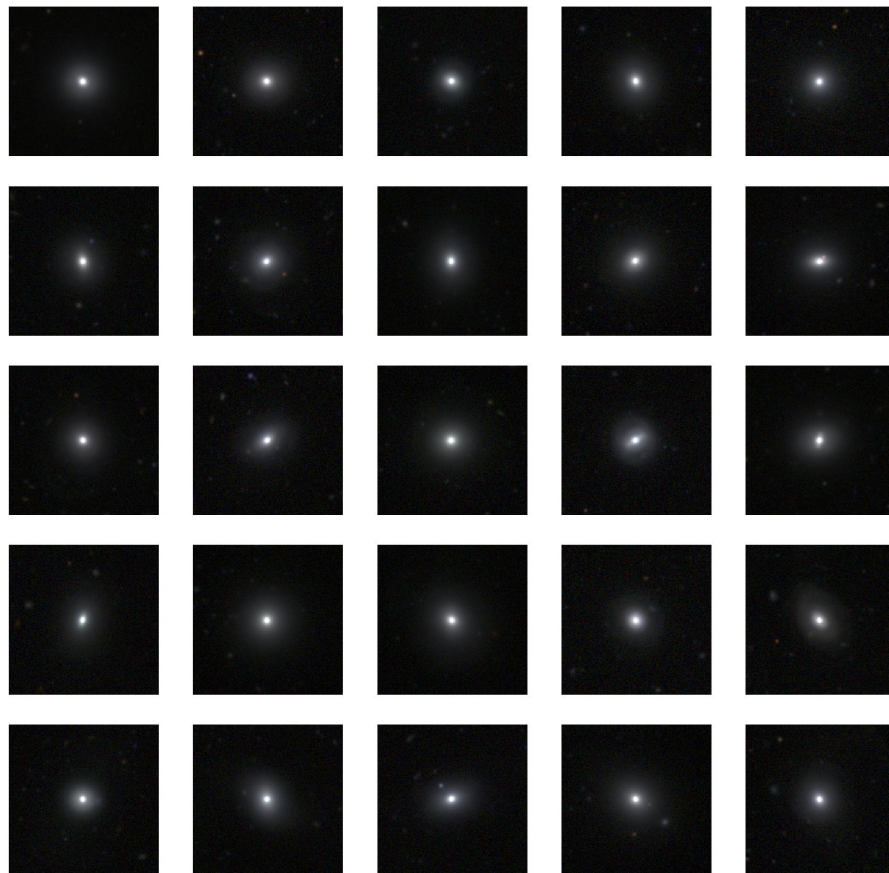
Anomaly detection with ELF in high dimensional latent space

Anomaly Score Distribution



Anomalies

Top 25
nominals

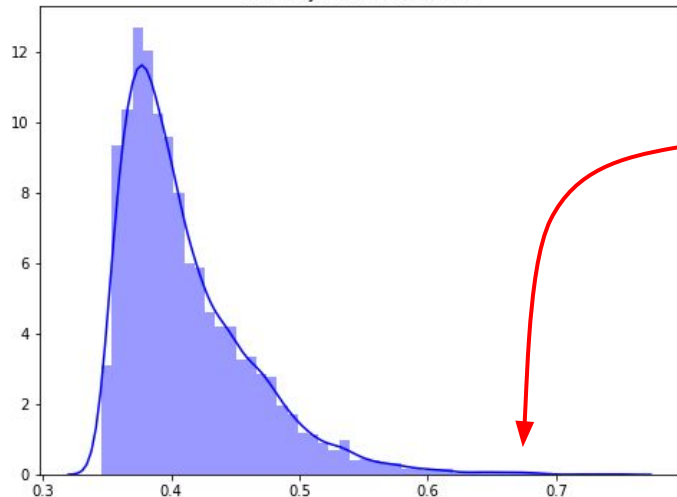


But also can tell us about the repeated and common cases as shown here

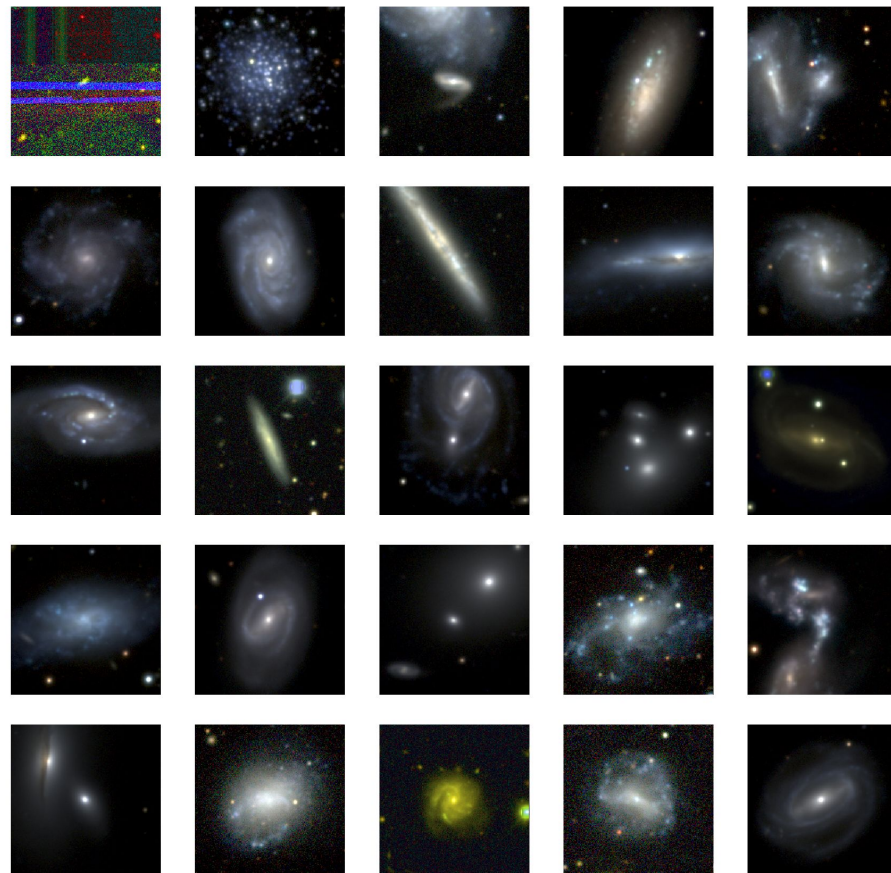
These are all different sources!

Anomaly detection with ELF in high dimensional latent space

Anomaly Score Distribution



Top 25
anomalies

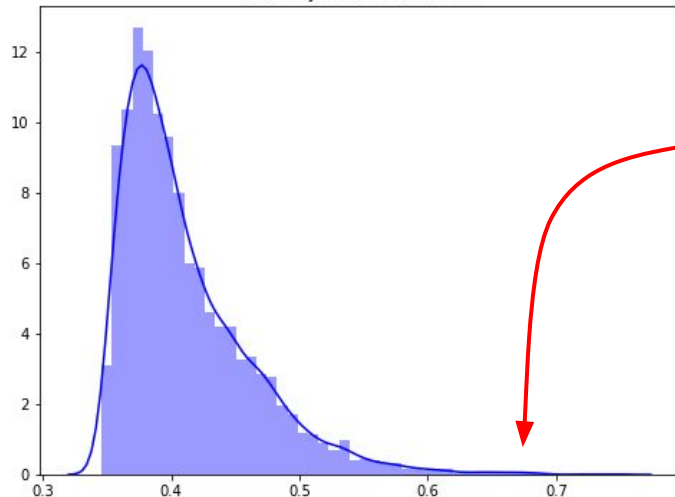


Anomalies →

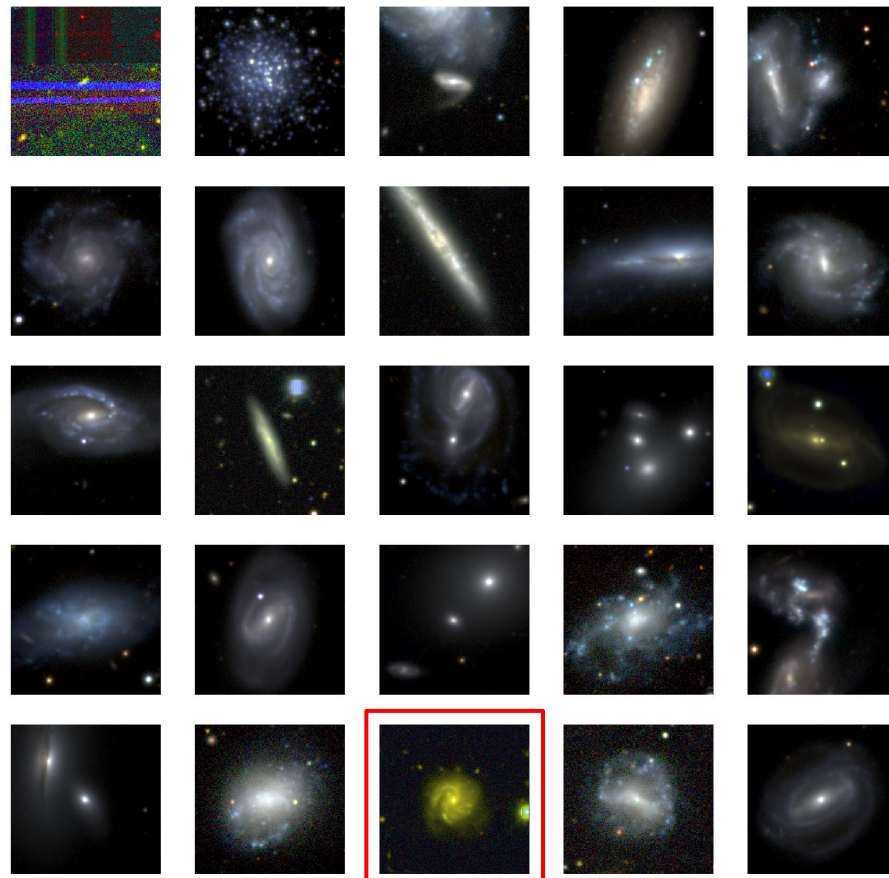
We can find other errors by
looking at similarities

Anomaly detection with ELF in high dimensional latent space

Anomaly Score Distribution



Top 25
anomalies

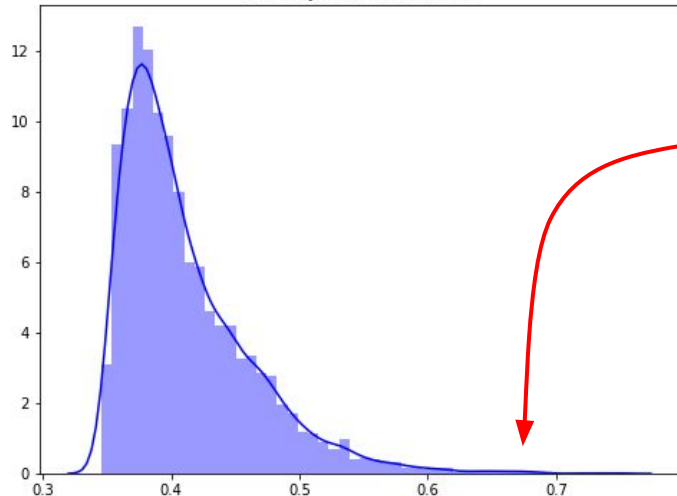


Anomalies →

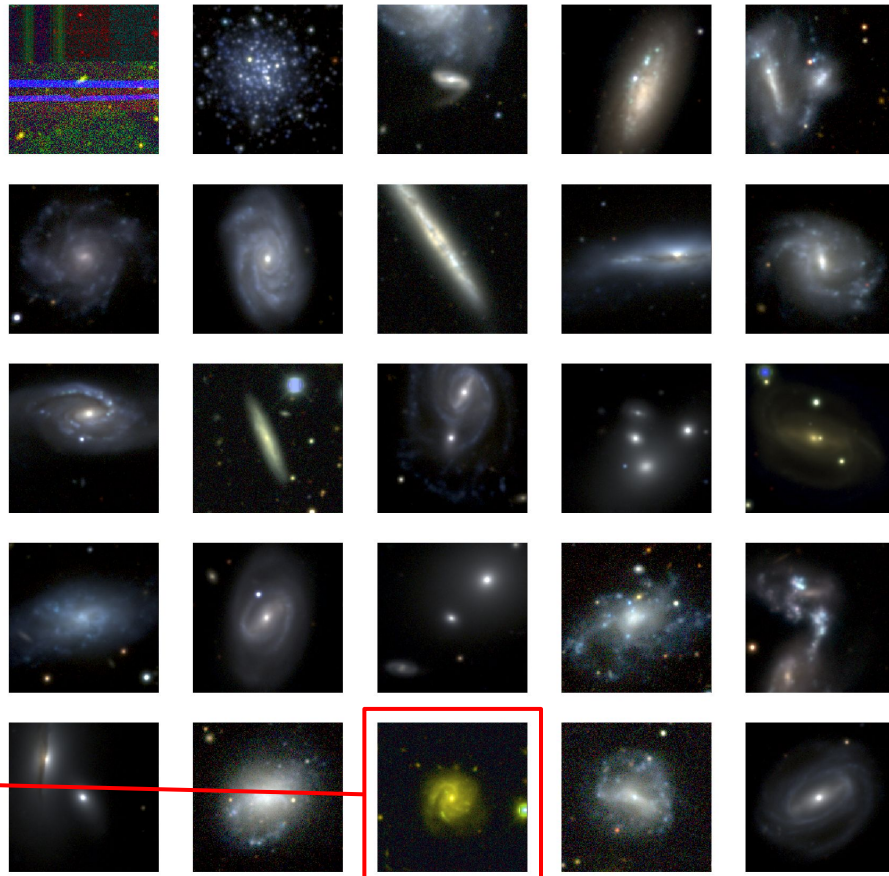
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Anomaly detection with EIF in high dimensional latent space

Anomaly Score Distribution

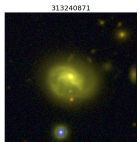
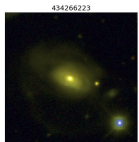
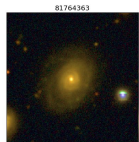
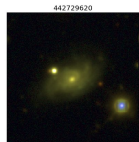
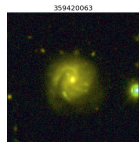
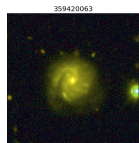


Top 25
anomalies



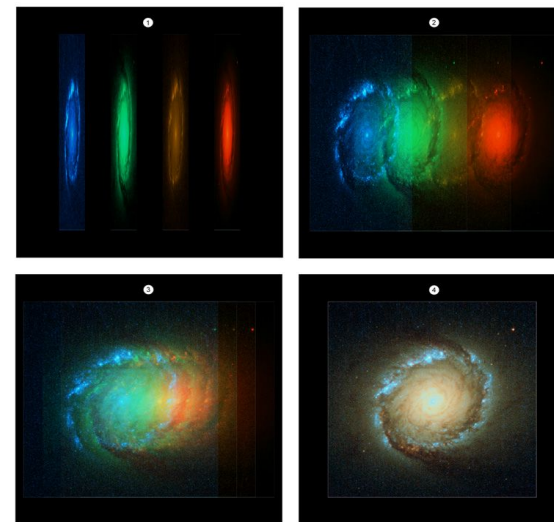
Anomalies →

We can find other errors by
looking at similarities

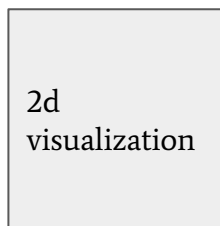


Can we keep reducing dimensions?

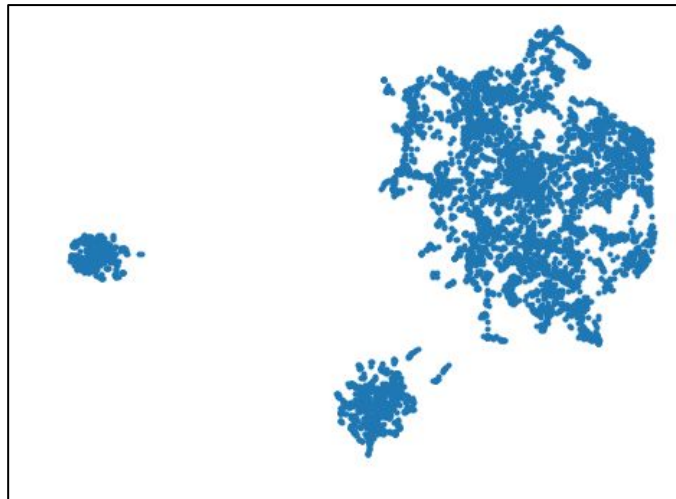
- Apply clustering and unsupervised techniques to latent space to find patterns
- Self Organizing Maps and T-SNE are perfect candidates
- We have used Uniform Manifold Approximation and Projection (UMAP)



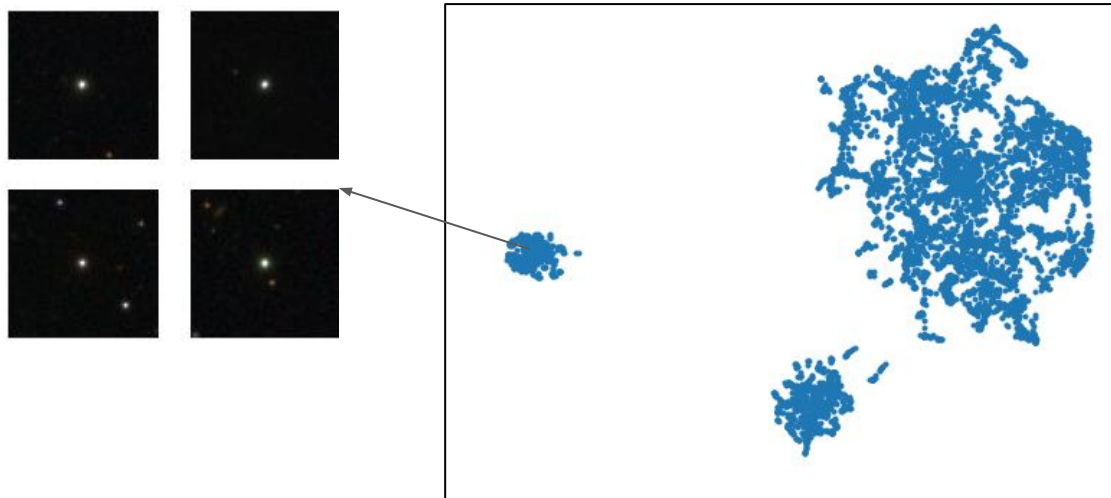
Credit: esa



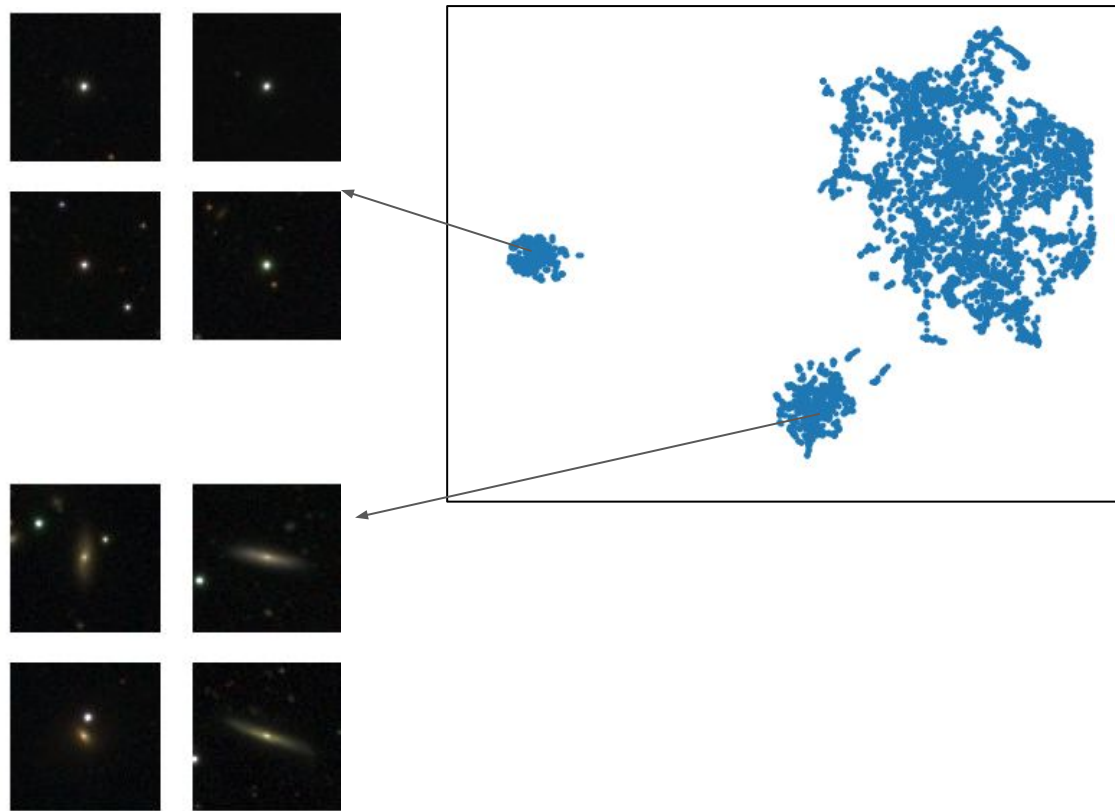
UMAP Representation (6000 galaxies)



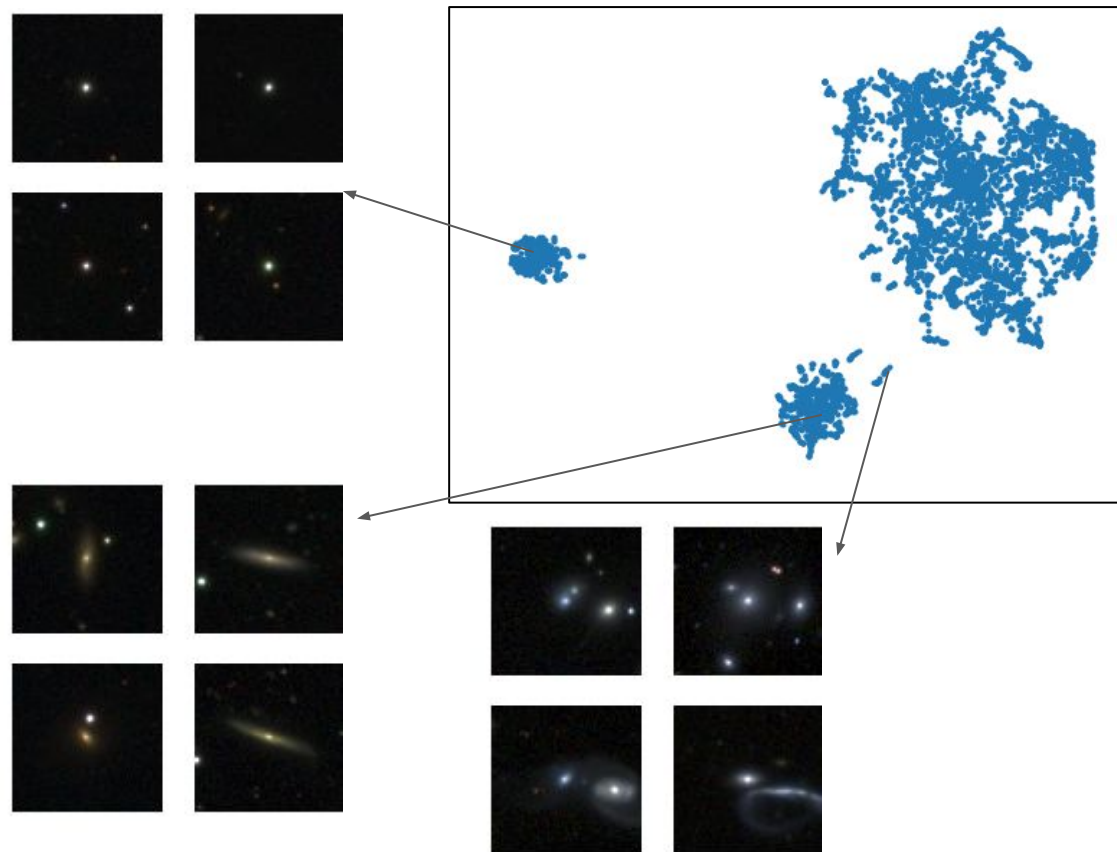
UMAP Representation (6000 galaxies)



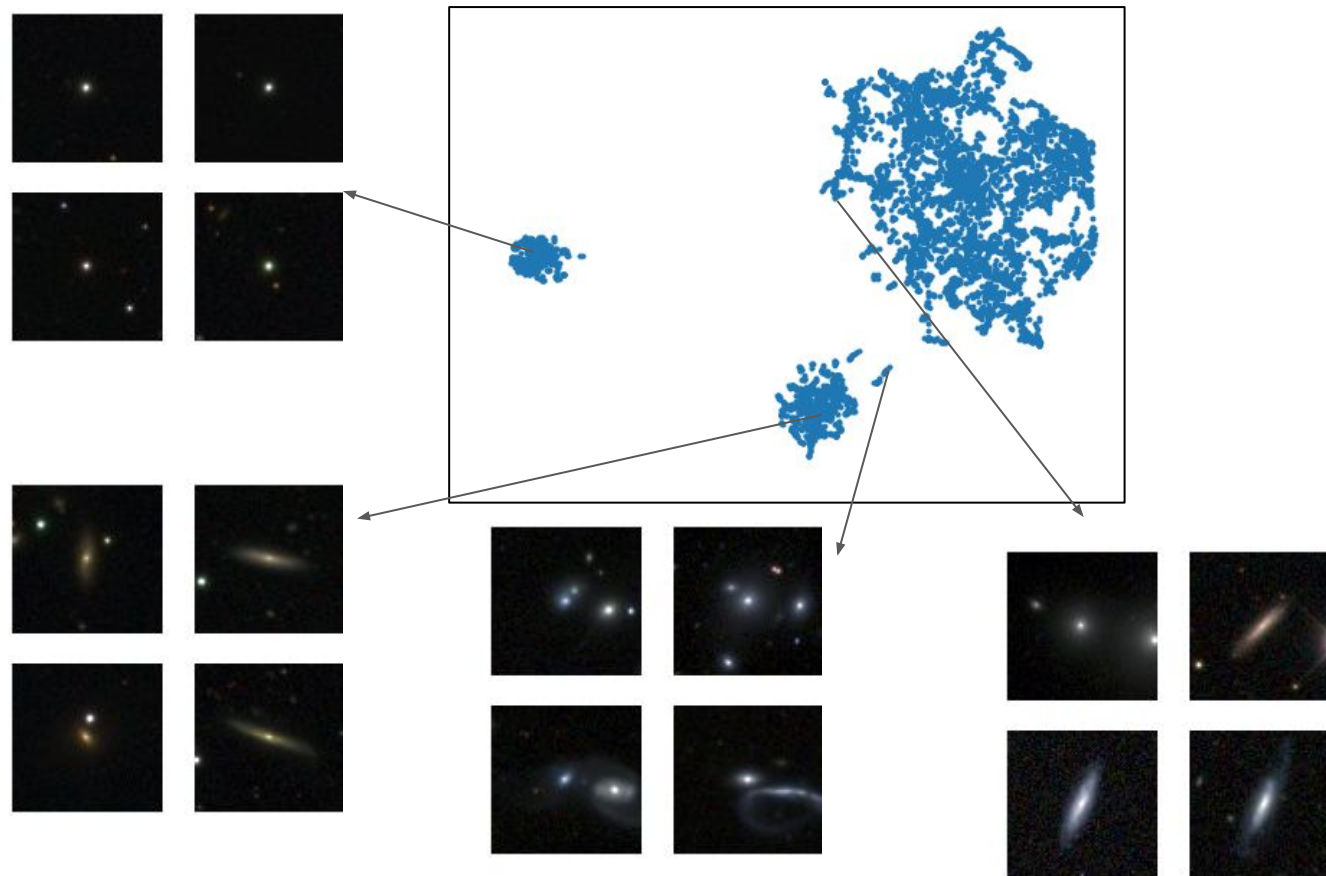
UMAP Representation (6000 galaxies)



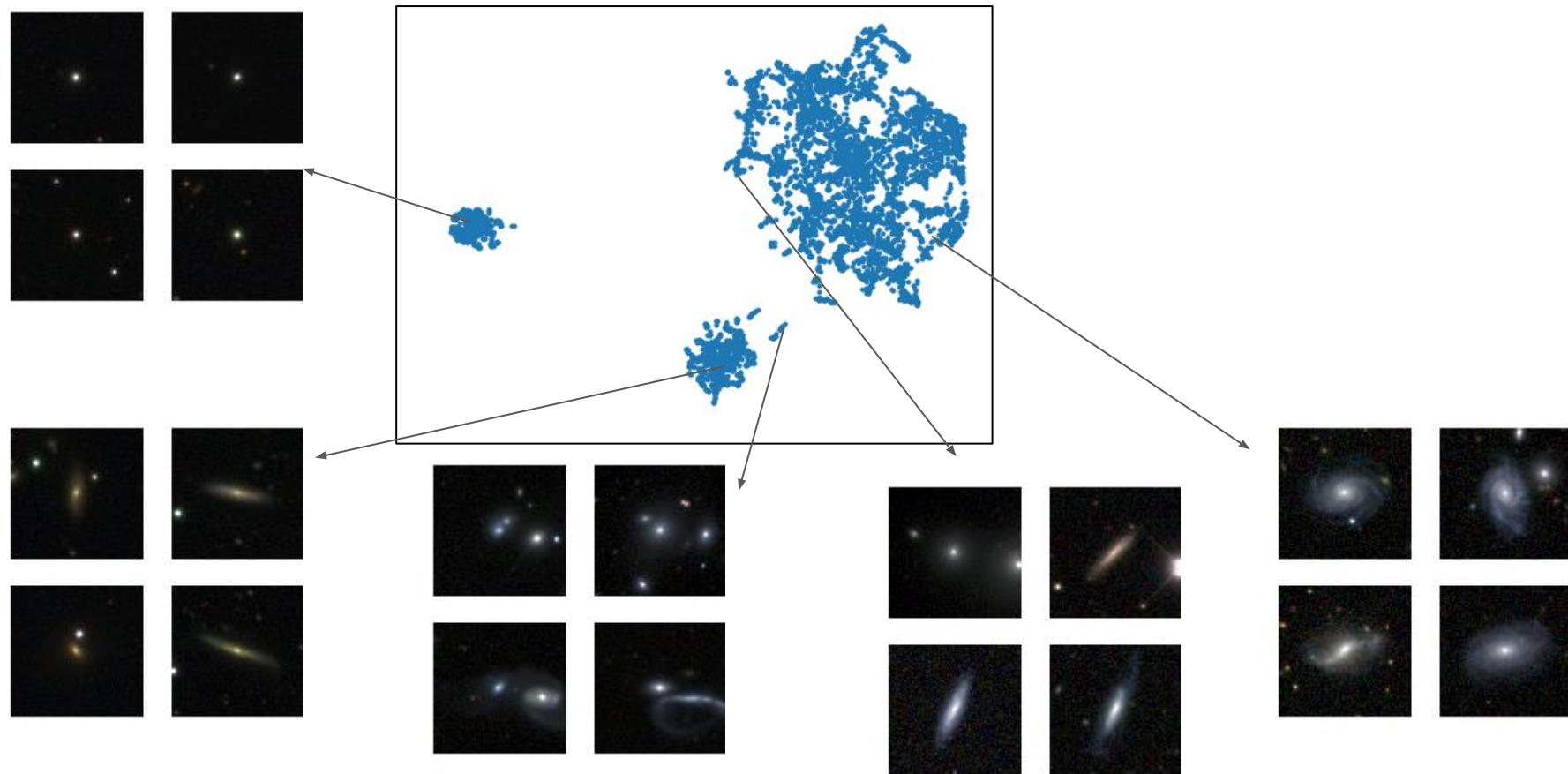
UMAP Representation (6000 galaxies)



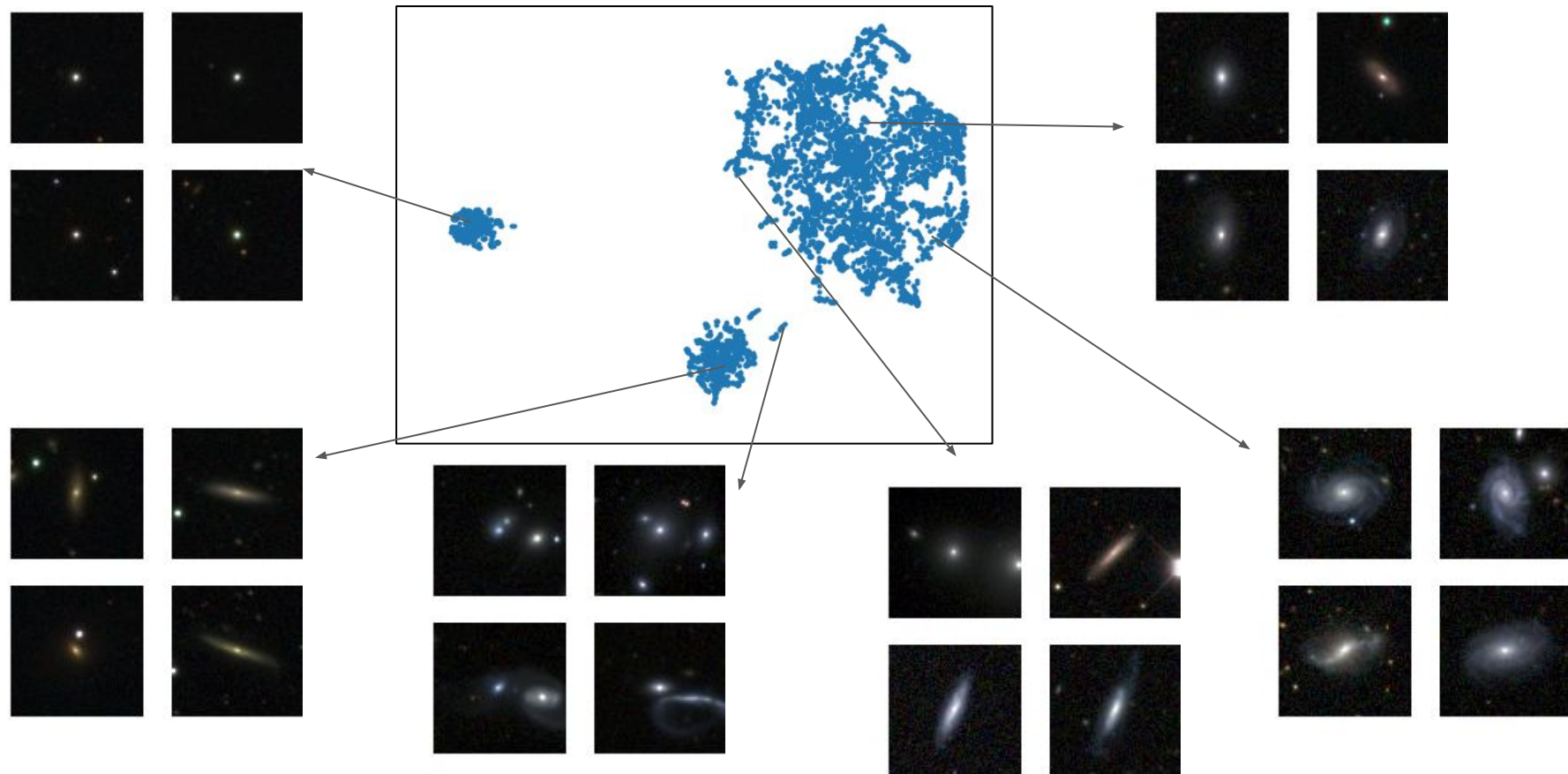
UMAP Representation (6000 galaxies)



UMAP Representation (6000 galaxies)



UMAP Representation (6000 galaxies)



Conclusions

- We developed a visualization and classification tool for multiple images
- Using Autoencoders we can compress images to small (but high-n) latent space
- Look for similarities and anomalies in that space
- Represent even more in a 2d graph using t-SNE, SOM or UMAP
- Scientific driven cases
- State-of-the-art models allows a bayesian manipulation of the latent space

Thank you!

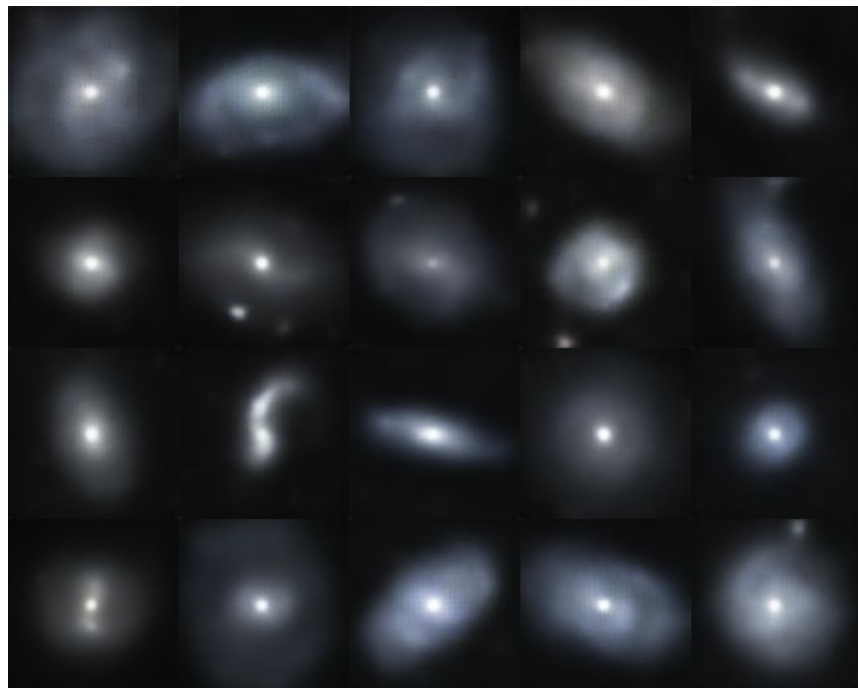
Questions?

Matias Carrasco Kind -- NCSA

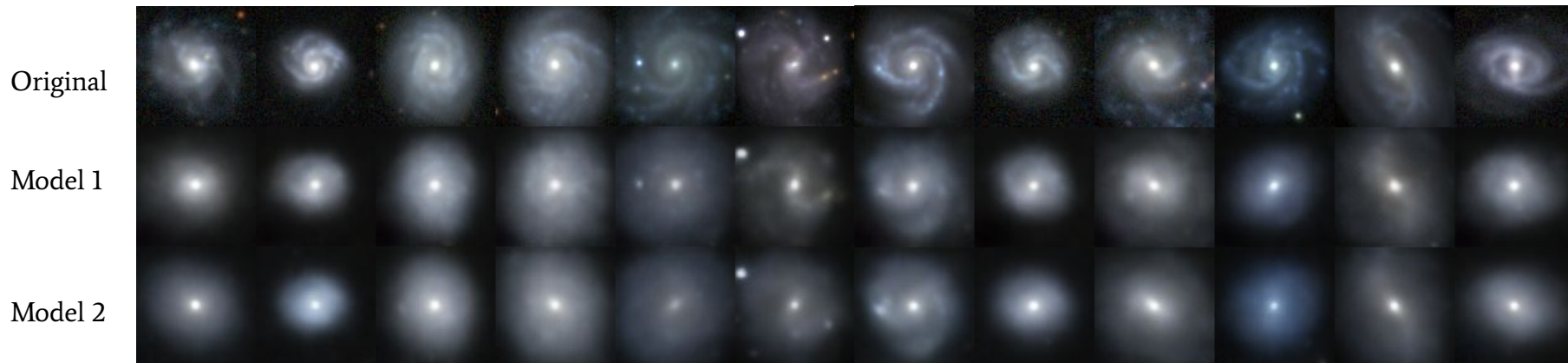
mcarras2@illinois.edu

github.com/mgkind

matias-ck.com



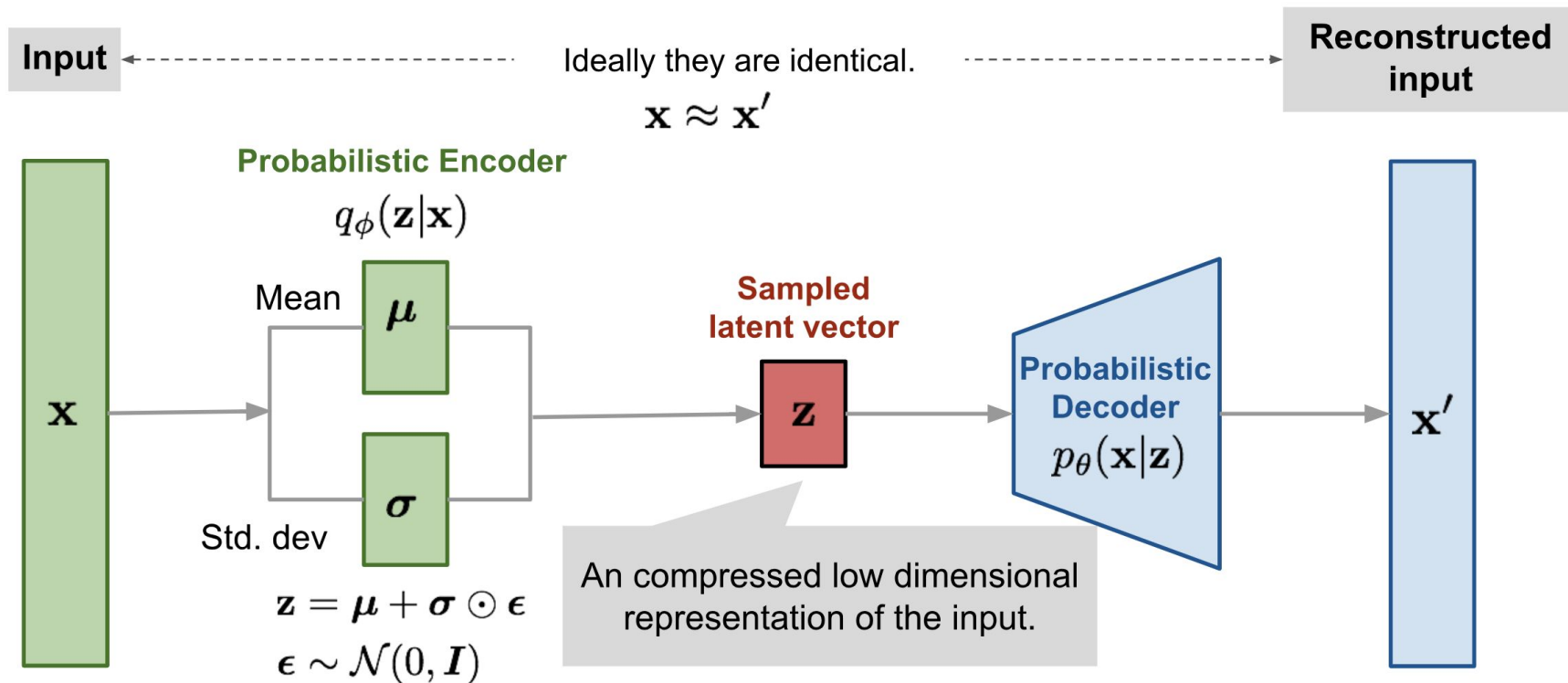
Autoencoders Applications: Reconstruction



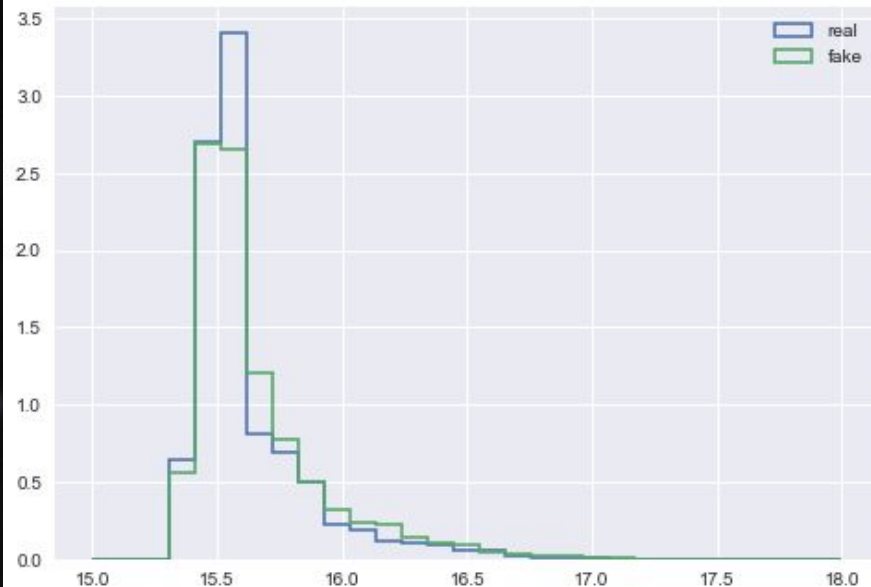
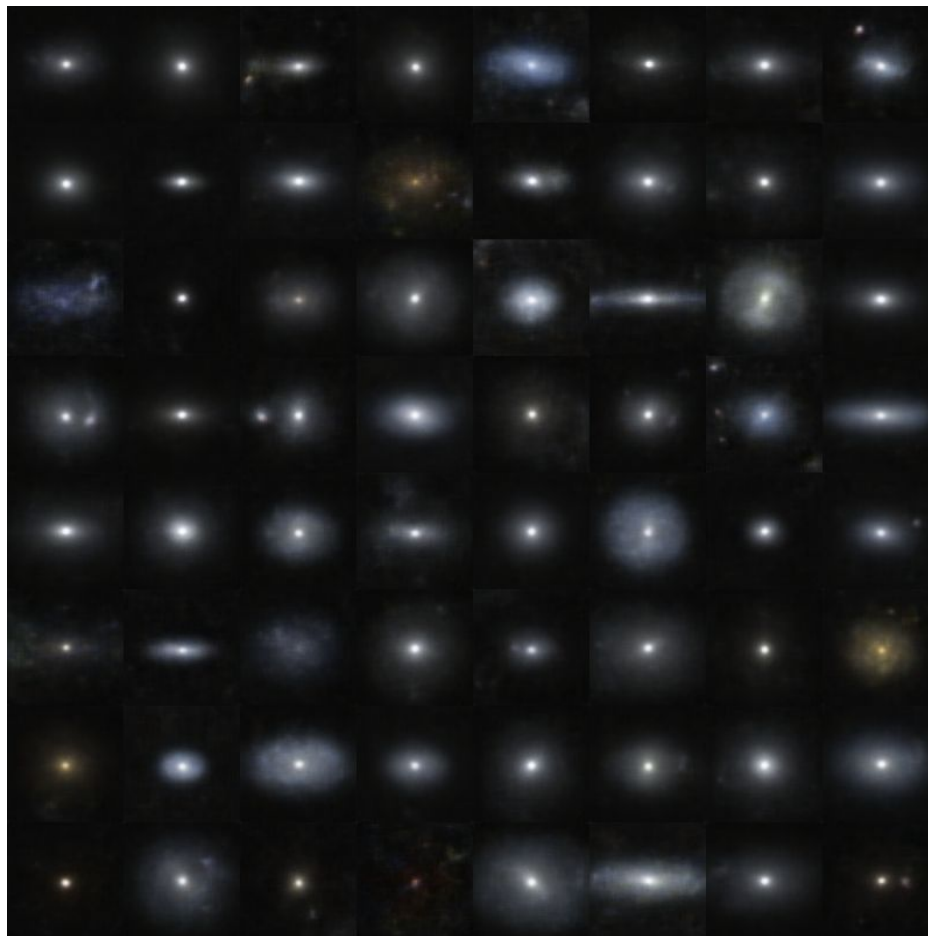
Blurry images and structure is lost, but angular sizes, radial profiles and brightness are a match.

What if we can make the model learn properties at the same time as images.? What if can sample from the latent space?

Summarizing VAE basics, more complex models built on top



VAE Sampling example (no reconstruction)



We can generate samples from z , next step is can we constrain what's being sampled?

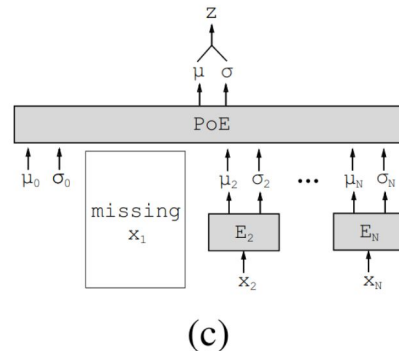
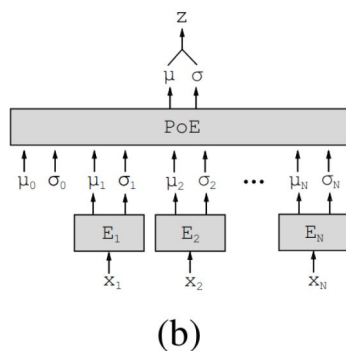
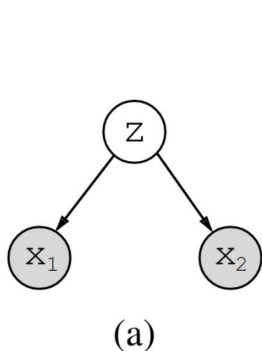
Multimodal VAE: Training modalities

Multimodal Generative Models for Scalable Weakly-Supervised Learning

Mike Wu
Department of Computer Science
Stanford University
Stanford, CA 94025
wumike@stanford.edu

Noah Goodman
Departments of Computer Science and Psychology
Stanford University
Stanford, CA 94025
ngoodman@stanford.edu

[multimodal-generative-models-for-scalable-weakly-supervised-learning](#)

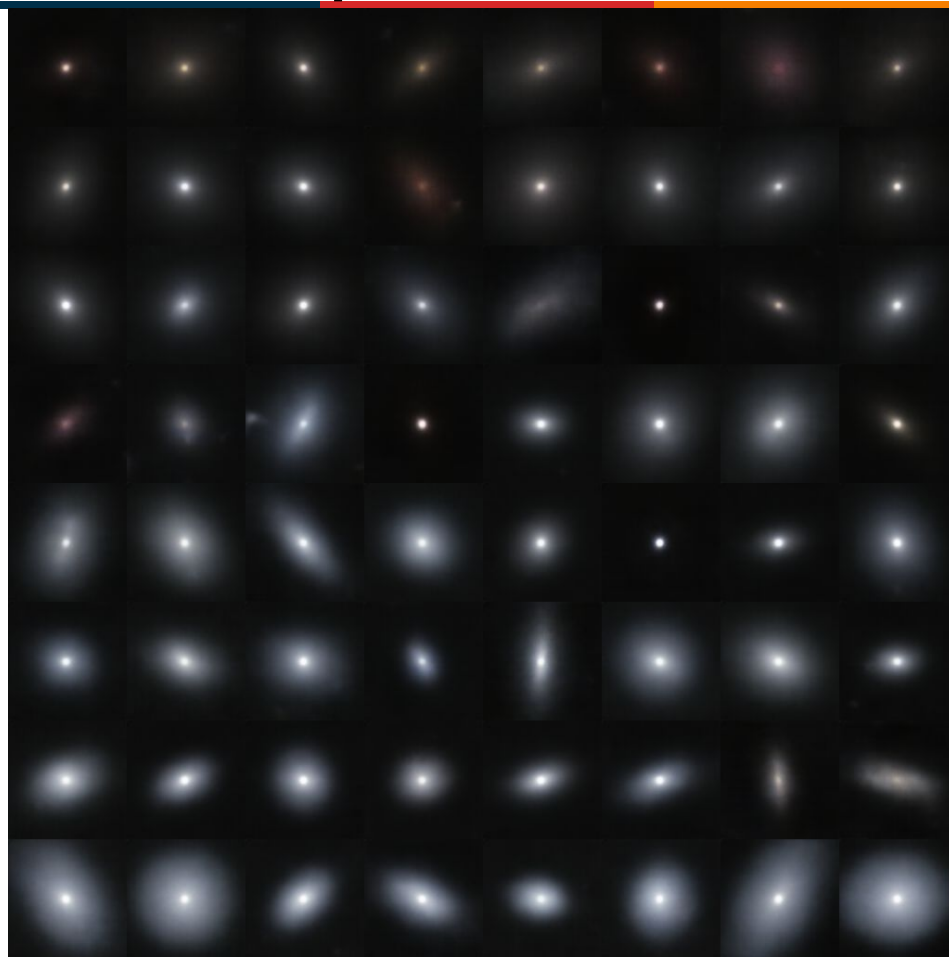


Learning joint representation of conditionally independent modalities using product of experts.

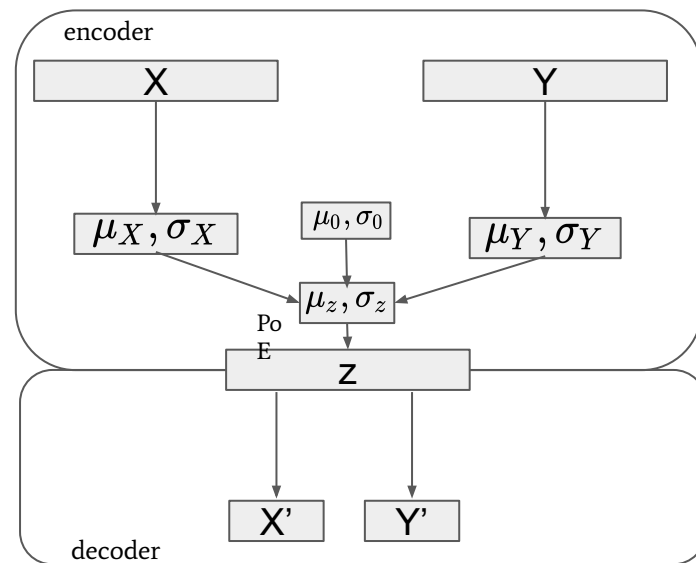
We can:

- Conditional sample with certain attributes
- Sample without any limitations
- Change the attribute of an existing input data
- Similarity search and anomaly detection
- Predict one modality from the others
- Sample and train with missing modalities

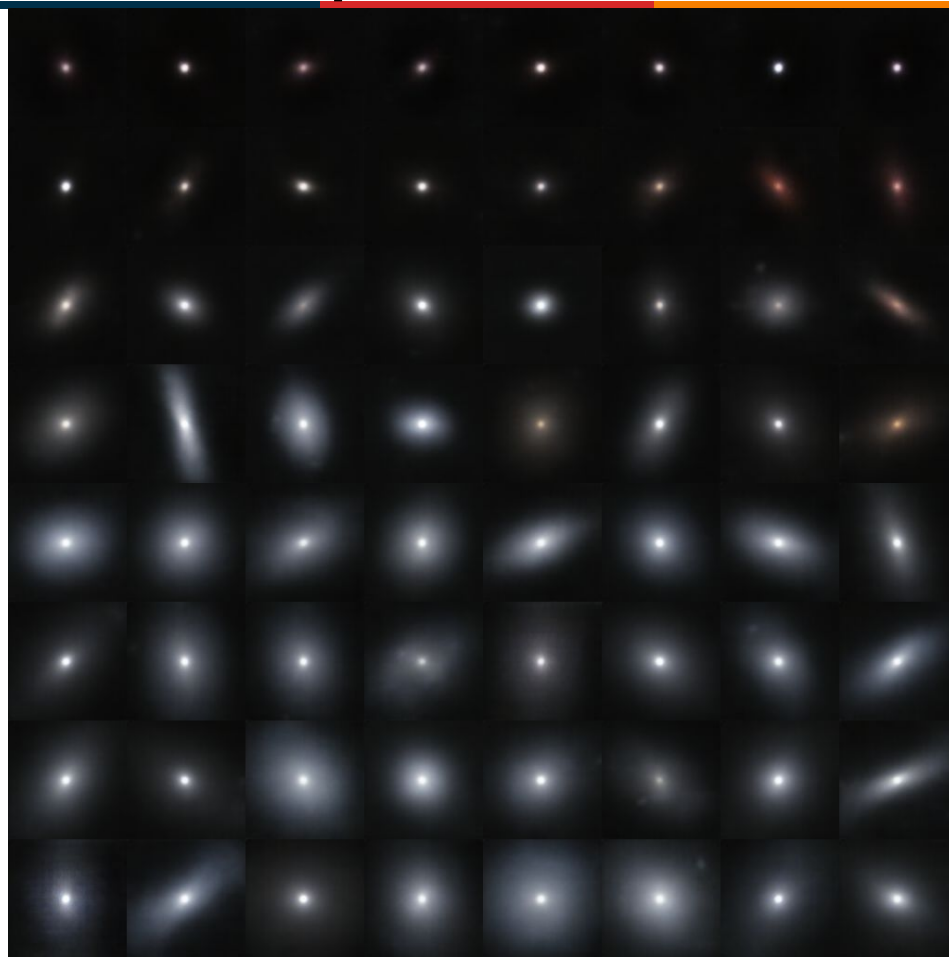
MVAE: Examples



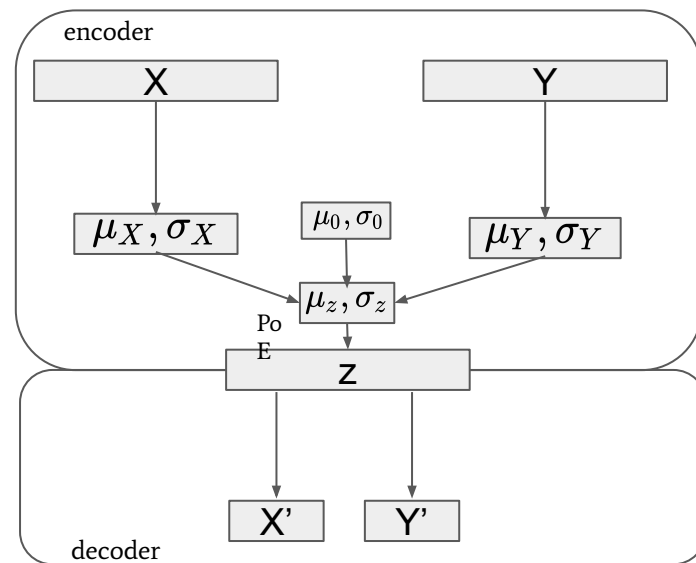
Samples with changing brightness
(increasing downwards)



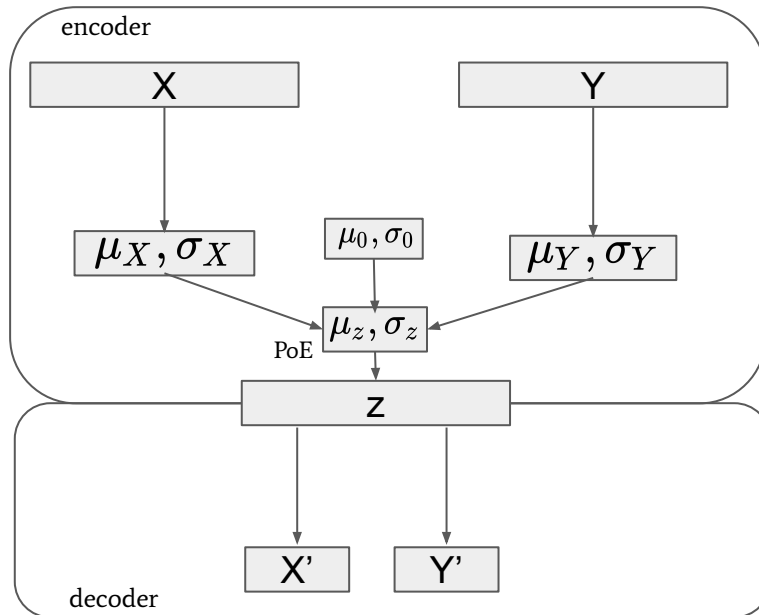
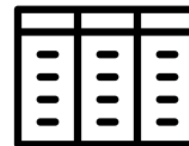
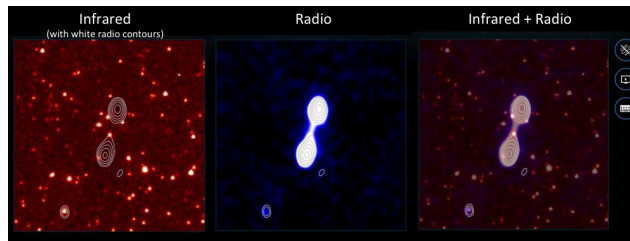
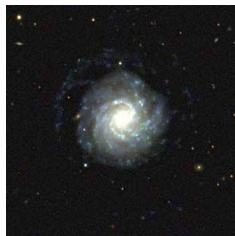
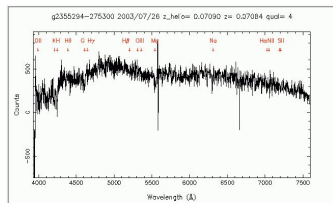
MVAE: Examples



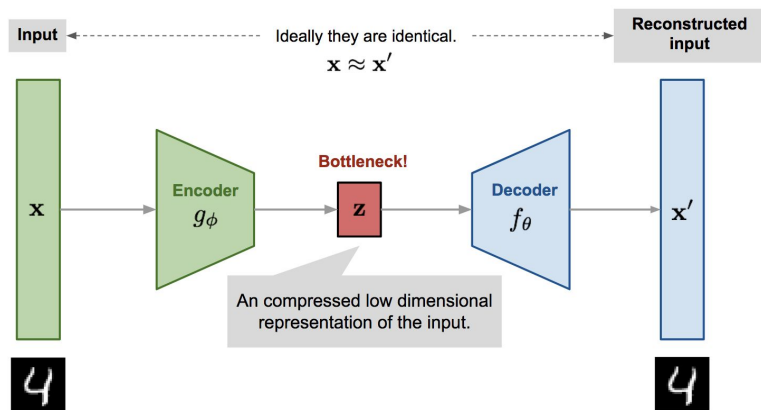
Samples with changing area
(increasing downwards)



MVAE: Opens very interesting options



Can we sample z to generate fake data?



Exist θ for max the likelihood

$$p(x) = \int p(x|z, \theta)p(z)dz$$

Too expensive

$$p(z|x) = \frac{p(x|z)p(z)}{p(x)}$$

We need an approximate posterior (prob. encoder)

$$q_\lambda(z|x) \approx p(z|x)$$

And we can use q to be Gaussian (there are other alternatives)

$$q_\lambda(z|x) = \mathcal{N}(z; \mu_\lambda(x), \sigma_\lambda(x))$$

$$p(z) = \mathcal{N}(0, I)$$

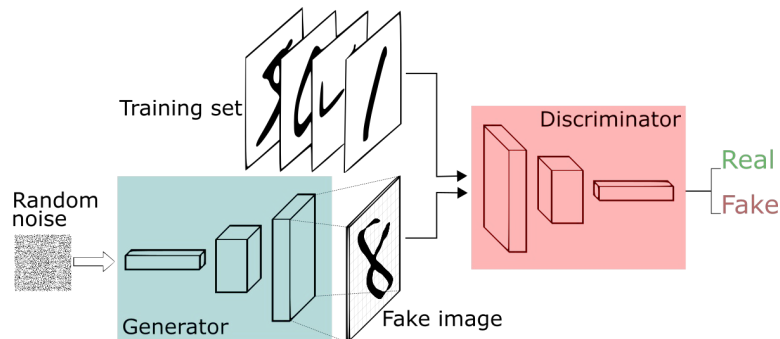
- Map x to a distribution $p(z|x)$
- Sample from distribution $z_i \sim p(z)$
- Generate fake data $x'_i \sim p_\theta(x'|z)$
- Probabilistic approach

$$p(x, z) = p(x|z)p(z)$$

Variational “Autoencoder”

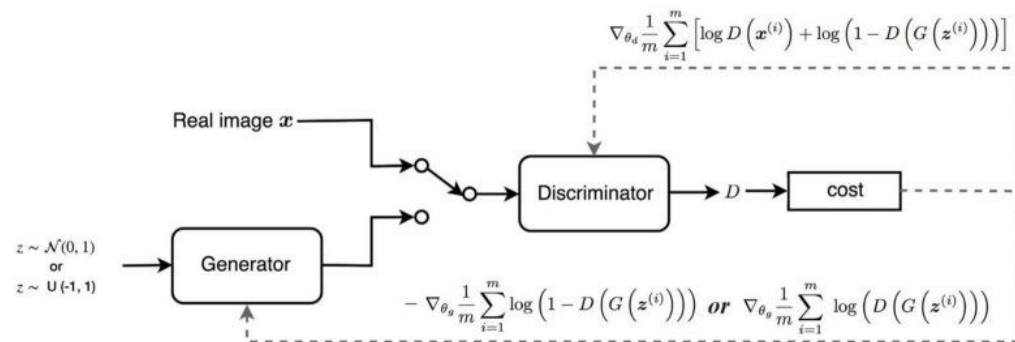
Generative Adversarial Networks (GAN)

Gan can be VERY good to specific image generation and create realistic images. Very powerful discriminator



But:

- Very hard to train (unstable)
- Not really sampling methods
- Hard to evaluate likelihood of data $p(x)$
- Tend to underfit data distribution
- Main goal is to fool the discriminator



Very powerful if combined with VAE