# PEU 416 Assignment 4

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$$\sin \theta \approx \theta - \frac{\theta^3}{6} + \frac{\theta^5}{120}$$

$$\theta - \sin \theta \approx \frac{\theta^3}{6} - \frac{\theta^5}{120}$$

$$9(\theta - \sin \theta)^2 \approx \frac{\theta^6}{4} - \frac{\theta^8}{40}$$

$$2(1 - \cos \theta)^3 \approx \frac{\theta^6}{4} - \frac{\theta^8}{16}$$

$$1 + \delta = \frac{\frac{\theta^6}{6} - \frac{\theta^8}{120}}{\frac{\theta^6}{4} - \frac{\theta^8}{16}}$$

For small  $\theta$ 

$$1 + \delta \approx \left(1 - \frac{\theta^2}{10}\right) \left(1 + \frac{\theta^2}{4}\right) \approx 1 + \frac{3}{20}\theta^2$$
 
$$\delta \approx \frac{3}{20}\theta^2$$
 
$$t = B(\theta - \sin \theta) \approx B\frac{\theta^3}{6}$$
 
$$\theta \approx \left(6\frac{t}{B}\right)^{\frac{1}{3}}$$
 
$$t_{i, \max} = B(\pi - \sin \pi) = B\pi$$
 
$$t \approx B\frac{\theta^3}{6}$$
 
$$\theta^2 = (6\pi)^{\frac{2}{3}} \left(\frac{t_i}{t_{i, \max}}\right)^{\frac{2}{3}}$$
 
$$\delta_i = \frac{3}{20}(6\pi)^{\frac{2}{3}} \left(\frac{t_i}{t_{i, \max}}\right)^{\frac{2}{3}}$$

(a) 
$$\sigma^{2}(M) = \frac{3}{R^{3}} \int_{0}^{R} \xi(r) r^{2} dr$$

$$\langle M^{2} \rangle = \langle M \rangle^{2} + \frac{\langle M \rangle^{2}}{V^{2}} \int_{V} \int_{V} \xi(\left| (x)_{1} - (x)_{2} \right|) d^{3}(x)_{1} d^{3}(x)_{2}$$

$$\sigma^{2}(M) = \frac{\langle M \rangle^{2}}{V^{2}} \int_{V} \int_{V} \xi(\left| (x)_{1} - (x)_{2} \right|) d^{3}(x)_{1} d^{3}(x)_{2}$$

$$\sigma_{R}^{2} = \frac{\sigma^{2}(M)}{\langle M \rangle^{2}} = \frac{1}{V^{2}} \int_{V} \int_{V} \xi(\left| (x)_{1} - (x)_{2} \right|) d^{3}(x)_{1} d^{3}(x)_{2}$$

$$\frac{1}{V^{2}} \int_{V} \int_{V} \xi(\left| (x)_{1} - (x)_{2} \right|) d^{3}(x)_{1} d^{3}(x)_{2} \approx \frac{1}{V} \int_{V} \xi(\left| (x) \right|) d^{3}(x)$$

$$\sigma_{R}^{2} \approx \frac{1}{\left(\frac{4}{3}\right)\pi R^{3}} \int_{0}^{R} \xi(r) (4\pi r^{2} dr)$$

$$\sigma_{R}^{2} = \frac{3}{R^{3}} \int_{0}^{R} \xi(r) r^{2} dr$$
(b) 
$$\xi(r) = \left(\frac{r}{r_{0}}\right)^{\gamma}$$
with  $r_{0} = 5h^{-1}$  Mpc and  $\gamma = -1.8$ .
$$\sigma_{R}^{2} = \frac{3}{R^{3}} \int_{0}^{R} \left(\frac{r}{r_{0}}\right)^{\gamma} r^{2} dr = \frac{3}{R^{3} r_{0}^{\gamma}} \int_{0}^{R} r^{\gamma+2} dr$$

$$\sigma_{R}^{2} = \frac{3}{R^{3} r_{0}^{\gamma}} \left[\frac{r^{\gamma+3}}{\gamma+3}\right]_{0}^{R} = \frac{3}{\gamma+3} \left(\frac{R}{r_{0}}\right)^{\gamma}$$

$$\sigma_{8}^{2} = \frac{3}{1.2} \left(\frac{8}{5}\right)^{-1.8} = 2.5 \times (1.6)^{-1.8}$$

$$\sigma_{8}^{2} \approx 2.5 \times 0.4295 \approx 1.0738$$

$$\sigma_{8} = \sqrt{1.0738} \approx 1.036$$

The implied value is  $\sigma_8\approx 1.04.$ 

(a) 
$$\sigma^2(R) = \int_0^\infty \frac{k^2 dk}{2\pi^2} P(k) \left| W.\widetilde{\text{kR}} \right|^2$$
 
$$W.\widetilde{\text{kR}} = \begin{cases} 1 \text{ if kR} \le 1\\ 0 \text{ if kR} > 1 \end{cases}$$

This sets the integral's upper limit to  $k = \frac{1}{R}$ . With  $P(k) = Ak^n$ :

$$\sigma^2(R) = \frac{A}{2\pi^2} \int_0^{\frac{1}{R}} k^{n+2} dk = \frac{A}{2\pi^2(n+3)} \left(\frac{1}{R}\right)^{n+3}$$

Thus,  $\sigma^2(R) \propto R^{-(n+3)}$ . Since mass  $M \propto R^3$ , we have  $R \propto M^{\frac{1}{3}}$ . Substituting this:

$$\sigma^2(M) \propto \left(M^{\frac{1}{3}}\right)^{-(n+3)} = M^{-\frac{n+3}{3}}$$

Comparing this to  $\sigma^2(M) \propto M^{\gamma}$ , we find the relation:

$$\gamma = -\frac{n+3}{3}$$

(b) For a **Gaussian filter**, the window function is  $W.\widetilde{kR} = e^{-\frac{(kR)^2}{2}}$ , so  $\left|W.\widetilde{kR}\right|^2 = e^{-(kR)^2}$ . The variance integral becomes:

$$\sigma^2(R) = \frac{A}{2\pi^2} \int_0^\infty k^{n+2} e^{-(\mathbf{k}\mathbf{R})^2} dk$$

We use the substitution  $t = (kR)^2$ :

$$\sigma^2(R) = \frac{A}{4\pi^2 R^{n+3}} \int_0^\infty t^{\frac{n+1}{2}} e^{-t} dt$$

The integral is the definition of the Gamma function,  $\Gamma(z)$ , with  $z = \frac{n+3}{2}$ .

$$\sigma^2(R) = \frac{A\Gamma\left(\frac{n+3}{2}\right)}{4\pi^2 R^{n+3}}$$

The variance has the same dependence on scale,  $\sigma^2(R) \propto R^{-(n+3)}$ . This leads to the **same relationship** between  $\gamma$  and n as in part a:

$$\gamma = -\frac{n+3}{3}$$

(c) We compute the ratio  $\frac{\sigma_G^2}{\sigma_S^2}$  for n=1.

Sharp k-space filter  $(\sigma_S^2)$ :

$$\sigma_S^2(R) = \frac{A}{2\pi^2(1+3)} \left(\frac{1}{R}\right)^{1+3} = \frac{A}{8\pi^2 R^4}$$

Gaussian filter  $(\sigma_G^2)$ :

$$\sigma_G^2(R) = \frac{A\Gamma(\frac{1+3}{2})}{4\pi^2 R^{1+3}} = \frac{A\Gamma(2)}{4\pi^2 R^4}$$

The ratio is:

$$\frac{\sigma_G^2(R)}{\sigma_S^2(R)} = \frac{\frac{A\Gamma(2)}{4\pi^2 R^4}}{\frac{A}{8\pi^2 R^4}}$$

The terms  $A, \pi^2$ , and  $R^4$  cancel, leaving:

Ratio = 
$$\frac{\frac{\Gamma(2)}{4}}{\frac{1}{8}} = \left(\frac{8}{4}\right)\Gamma(2) = 2\Gamma(2)$$

This expresses the answer in terms of the special function, **Gamma** ( $\Gamma$ ). To find the numerical value, we use  $\Gamma(z) = (z-1)!$  for integers:

$$\Gamma(2) = (2-1)! = 1! = 1$$

The numerical value of the ratio is:

Ratio = 
$$2 \times 1 = *2*$$

(a) 
$$M = \frac{4}{3}\pi r_i^3 \rho_i = \frac{4}{3}\pi r_i^3 \bar{\rho}_i (1+\delta_i)$$
 
$$\ddot{r} = -\frac{G \cdot \frac{4}{3}\pi r_i^3 \bar{\rho}_i (1+\delta_i)}{r^2}$$
 
$$\frac{\ddot{r}}{r} = -\frac{4\pi G \bar{\rho}_i}{3} (1+\delta_i) \left(\frac{r_i}{r}\right)^3$$
 (b) 
$$\frac{1}{2} \dot{r}^2 - \frac{GM}{r} = E$$
 
$$\frac{1}{2} \dot{r}^2 - \frac{4\pi G \bar{\rho}_i}{3} (1+\delta_i) \frac{r_i^3}{r} = E$$
 
$$E = K(r_i) + V(r_i)$$
 
$$= \frac{1}{2} \dot{r}_i^2 - \frac{4\pi G \bar{\rho}_i}{3} (1+\delta_i) r_i^2$$
 
$$= \frac{1}{2} \left(H_i r_i \left(1 - \frac{\delta_i}{3}\right)\right)^2 - \frac{4\pi G \bar{\rho}_i}{3} (1+\delta_i) r_i^2$$

At the turn-around radius  $r_{\rm ta}$ ,  $\dot{r} = 0$ ,

$$E = -\frac{4\pi G \bar{\rho}_i}{3} (1 + \delta_i) \frac{r_i^3}{r_{\text{ta}}}$$

$$\therefore -\frac{4\pi G \bar{\rho}_i}{3} (1 + \delta_i) \frac{r_i^3}{r_{\text{ta}}} = \frac{1}{2} \left( H_i r_i \left( 1 - \frac{\delta_i}{3} \right) \right)^2 - \frac{4\pi G \bar{\rho}_i}{3} (1 + \delta_i) r_i^2$$

$$-\frac{4\pi G \bar{\rho}_i}{3} (1 + \delta_i) \frac{r_i}{r_{\text{ta}}} = \frac{1}{2} H_i^2 \left( 1 - \frac{\delta_i}{3} \right)^2 - \frac{4\pi G \bar{\rho}_i}{3} (1 + \delta_i)$$

$$H_i^2 = \frac{8\pi G \bar{\rho}_i}{3}$$

$$-(1 + \delta_i) \frac{r_i}{r_{\text{ta}}} = \left( 1 - \frac{\delta_i}{3} \right)^2 - (1 + \delta_i)$$

$$-(1 + \delta_i) \frac{r_i}{r_{\text{ta}}} = -\frac{5}{3} \delta_i + \frac{1}{9} \delta_i \approx -\frac{5}{3} \delta_i \quad \text{for } \delta_i \ll 1$$

$$\therefore r_{\text{ta}} = \frac{3}{5} \left( \frac{1 + \delta_i}{\delta_i} \right) r_i$$
(c)
$$\frac{dr}{d\theta} = A \sin \theta, \quad \frac{dt}{d\theta} = B(1 - \cos \theta)$$

$$\dot{r} = \frac{dr}{dt} = \frac{\frac{dr}{d\theta}}{\frac{dt}{d\theta}} = \frac{A\sin\theta}{B(1-\cos\theta)}$$

$$\frac{d\dot{r}}{d\theta} = \frac{d}{d\theta} \left( \frac{A\sin\theta}{B(1-\cos\theta)} \right)$$

$$= \frac{A}{B} * \frac{\cos\theta(1-\cos\theta) - (\sin\theta)(-\sin\theta)}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - \cos^2\theta + \sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - (1-\sin^2\theta) + \sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - (1-\sin^2\theta) + \sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - 1 + 2\sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta(1-\cos\theta) + \sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - \cos^2\theta + 1 - \cos^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{1+\cos\theta - 2\cos^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{1+2\cos\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - \cos^2\theta + \sin^2\theta}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{\cos\theta - \cos^2\theta + (1-\cos^2\theta)}{(1-\cos\theta)^2}$$

$$= \frac{A}{B} * \frac{1+\cos\theta - 2\cos^2\theta}{(1-\cos\theta)^2}$$

$$\ddot{r} = \frac{\frac{A}{B} * \frac{1 + 2\cos\theta}{1 - \cos\theta}}{B(1 - \cos\theta)} = \frac{A}{B^2} * \frac{1 + 2\cos\theta}{(1 - \cos\theta)^2}$$

Since  $r = A(1 - \cos \theta)$ , we have  $1 - \cos \theta = \frac{r}{A}$ , and  $\cos \theta = 1 - \frac{r}{A}$ .

$$\ddot{r} = \frac{A}{B^2} * \frac{1 + 2(1 - \frac{r}{A})}{\left(\frac{r}{A}\right)^2} = \frac{A}{B^2} * \frac{3 - 2\frac{r}{A}}{\frac{r^2}{A^2}} = \frac{A^3}{B^2 r^2} * \left(3 - 2\frac{r}{A}\right) = \frac{A^2}{B^2 r^2} (3A - 2r)$$

$$\ddot{r} = \frac{\frac{A}{B} * \frac{1 + \cos \theta}{1 - \cos \theta}}{B(1 - \cos \theta)} = \frac{A}{B^2} * \frac{1 + \cos \theta}{(1 - \cos \theta)^2}$$

With  $1 - \cos \theta = \frac{r}{A}$  and  $1 + \cos \theta = 2 - (1 - \cos \theta) = 2 - \frac{r}{A}$ :

$$\ddot{r} = \frac{A}{B^2} * \frac{2 - \frac{r}{A}}{\left(\frac{r}{A}\right)^2} = \frac{A}{B^2} * \frac{\frac{2A - r}{A}}{\frac{r^2}{A^2}} = \frac{A^3}{B^2} * \frac{2A - r}{r^2}$$

#### Apply Initial Conditions

At turn-around,  $\dot{r} = 0$ , which is at  $\theta = \pi$ .

$$r_{\rm ta} = r(\theta = \pi) = A(1-\cos\pi) = 2A$$

(d) 
$$\delta(\theta) = \frac{9(\theta - \sin \theta)^2}{2(1 - \cos \theta)^3} - 1$$

At turn-around,  $\theta = \pi$ :

$$\delta(\pi) = \frac{9(\pi - \sin \pi)^2}{2(1 - \cos \pi)^3} - 1$$

$$\pi - \sin \pi = \pi - 0 = \pi$$

$$1 - \cos \pi = 1 - (-1) = 2$$

$$\delta(\pi) = \frac{9\pi^2}{2 \cdot 2^3} - 1 = \frac{9\pi^2}{16} - 1 \approx 5.55$$

(e) The virial theorem states 2T + U = 0. The potential energy is  $U = -\frac{3}{5} \frac{GM^2}{r}$ . At turn-around  $(r_{\rm ta})$ , kinetic energy is zero  $(T_{\rm ta} = 0)$ , so the total energy is:

$$E = T_{\rm ta} + U_{\rm ta} = 0 - \frac{3}{5} \frac{GM^2}{r_{\rm ta}}$$

After virialization, the system is in equilibrium.

$$E = T_{\rm vir} + U_{\rm vir}$$

$$U_{\rm vir} = -\frac{3}{5} \frac{GM^2}{r_{\rm vir}}$$

From the virial theorem,  $T_{\rm vir}=-\frac{U_{\rm vir}}{2}=\frac{1}{2}\Big(\frac{3}{5}\frac{GM^2}{r_{\rm vir}}\Big)=\Big(\frac{3}{10}\Big)\frac{GM^2}{r_{\rm vir}}$ . So, the total energy is:

$$E = \left(\frac{3}{10}\right) \frac{GM^2}{r_{\rm vir}} - \left(\frac{3}{5}\right) \frac{GM^2}{r_{\rm vir}} = -\left(\frac{3}{10}\right) \frac{GM^2}{r_{\rm vir}}$$

By conservation of energy:

$$-\frac{3}{5}\frac{GM^2}{r_{\rm ta}} = -\frac{3}{10}\frac{GM^2}{r_{\rm vir}}$$
 
$$r_{\rm vir} = \frac{r_{\rm ta}}{2}$$

The density contrast at virialization,  $\Delta_{\text{vir}}$ :

$$\Delta_{
m vir} = rac{
ho_{
m vir}}{ar
ho(t_{
m vir})}$$

We can find this by relating densities and times at turn-around and virialization. The mass density of the collapsed object changes as  $\rho \propto r^{-3}$ .

$$\frac{\rho_{\text{vir}}}{\rho_{\text{ta}}} = \left(\frac{r_{\text{ta}}}{r_{\text{vir}}}\right)^3 = \left(\frac{r_{\text{ta}}}{\frac{r_{\text{ta}}}{2}}\right)^3 = 2^3 = 8$$

The background density evolves as  $\bar{\rho} \propto t^{-2}$  (in a matter-dominated universe). The time to virialization is twice the time to turn-around,  $t_{\rm vir} = 2t_{\rm ta}$ .

$$\frac{\bar{\rho}(t_{\rm ta})}{\bar{\rho}(t_{\rm vir})} = \left(\frac{t_{\rm vir}}{t_{\rm ta}}\right)^2 = 2^2 = 4$$

Combining these results:

$$\Delta_{\rm vir} = \frac{\rho_{\rm vir}}{\bar{\rho}(t_{\rm vir})} = \frac{\rho_{\rm vir}}{\rho_{\rm ta}} \cdot \frac{\rho_{\rm ta}}{\bar{\rho}(t_{\rm ta})} \cdot \frac{\bar{\rho}(t_{\rm ta})}{\bar{\rho}(t_{\rm vir})}$$

The density contrast at turn-around is  $\frac{\rho_{\rm ta}}{\bar{\rho}(t_{\rm ta})}=1+\delta(\pi)=\frac{9\pi^2}{16}$ .

$$\Delta_{\rm vir} = 8 \cdot \frac{9\pi^2}{16} \cdot 4 = 18\pi^2 \approx 178$$

## References

[1] M. El-Deeb, "PEU-405 Assignments." [Online]. Available: https://github.com/mhdeeb/peu-assignments/tree/main/peu-405