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**PEU438-Fall2024**  
**Compact Objects and High Energy Astrophysics**

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**Due Date:** October 26, 2024, 11:30 PM  
**Assignment:** Number 2

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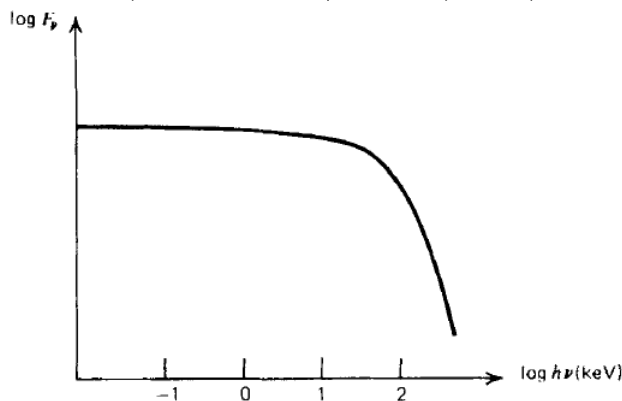
**Problem 1** Consider a sphere of ionized hydrogen plasma that is undergoing spherical gravitational collapse. The sphere is held at constant isothermal temperature  $T_0$ , uniform density and constant mass  $M_0$  during the collapse, and has decreasing radius  $R(t)$ . The sphere cools by emission of bremsstrahlung radiation in its interior. At  $t = t_0$  the sphere is optically thin.

- a. What is the total luminosity of the sphere as a function of  $M_0$ ,  $R(t)$  and  $T_0$  while the sphere is optically thin?
- b. What is the luminosity of the sphere as a function of time after it becomes optically thick?
- c. Give an implicit relation, in terms of  $R(t)$ , for the time  $t_1$  when the sphere becomes optically thick.

**Problem 2** Suppose X-rays are received from a source of known distance  $L$  with a flux  $F$  ( $\text{erg cm}^{-2} \text{ s}^{-1}$ ). The X-ray spectrum has the form of the Figure. It is conjectured that these X-rays are due to bremsstrahlung from an optically thin, hot, plasma cloud, which is in hydrostatic equilibrium around a central mass  $M$ . Assume that the cloud thickness  $\Delta R$  is roughly its radius  $R$ ,  $\Delta R \sim R$ .

- Find  $R$  and the density of the cloud,  $\rho$ , in terms of the known observations and conjectured mass  $M$ .

Hint: Hydrostatic equilibrium gives constraint on  $\rho$  and  $R$ . From the virial theorem we know that  $2 \times (\text{kinetic energy/particle}) = - (\text{gravitational energy/particle})$



### Problem 3

- (a) Determine the root-mean-square speed,  $v_{\text{rms}}$ , and corresponding kinetic energy,  $E_k$ , of an electron in an HII region of electron temperature,  $T_e = 10^4$  K.
- (b) If the electron lost all of its kinetic energy (came to a stop) during an encounter with a nucleus without recombining, what would be the frequency of the resulting emission and in what part of the spectrum does this frequency occur?
- (c) Compare the frequency of part (b) with a typical thermal Bremsstrahlung photon at a radio frequency of  $\nu = 5 \times 10^9$  Hz from an HII region. Based on this comparison, what fraction of an electron's kinetic energy has actually been lost during an encounter with a nucleus?

### Problem 4

- (a) Consider the volume emissivity of an optically thin plasma emitting thermal bremsstrahlung radiation at temperature  $T$ . Integrate the expression for  $j_\nu(\nu)$  (37) from frequency  $\nu_1$  to  $f\nu_1$  to obtain the power radiated in a fixed logarithmic frequency interval; if  $f = 10$ , your result would give the power in one decade as a function of frequency. Demonstrate that the power drops rapidly at both  $h\nu \ll kT$  and  $h\nu \gg kT$ . Let  $g = Z = 1$ .
- (b) Find the frequency at which the power in a fixed logarithmic interval is at a maximum as a function of  $\nu$ . To obtain a final solution, let the interval factor  $f$  become infinitely close to unity,  $f = 1 + \varepsilon$  for  $\varepsilon \ll 1$ . How does  $h\nu$  compare with  $kT$  at the maximum as  $\varepsilon \rightarrow 0$ ?

### Problem 5

- (a) Verify the result of the integration of  $j_\nu(\nu, T)$  to obtain  $j(T)$  for  $g = 1$ ; see (39).
- (b) The Orion nebula, an H II region, is radiating by thermal bremsstrahlung. Consider it to be spherical (radius  $R = 8\text{LY}$ ), optically thin, and at temperature  $T = 8000$  K throughout. Let  $Z = 1, g = 1$ , and  $n_e = n_i = 6 \times 10^8 \text{ m}^{-3}$ . Find the luminosity (W) of the entire nebula in terms of solar luminosities.
- (c) In what wavelength bands will the power from the Orion nebula be radiated?