1. Radiative Transfer Fundamentals

• Specific Intensity I_{ν} :

$$dE = I_{\nu} dA dt d\Omega d\nu$$

• Radiative Transfer Equation (RTE):

$$\frac{dI_{\nu}}{d\tau_{\nu}} = -I_{\nu} + S_{\nu}$$

where τ_{ν} is the optical depth and S_{ν} is the source function.

• Optical Depth τ_{ν} :

$$\tau_{\nu} = \int \alpha_{\nu} \, ds$$

where α_{ν} is the absorption coefficient per unit length.

• General Solution to RTE:

$$I_{\nu}(\tau_{\nu}) = I_{\nu}(0)e^{-\tau_{\nu}} + \int_{0}^{\tau_{\nu}} S_{\nu}(t)e^{-(\tau_{\nu} - t)} dt$$

For constant S_{ν} :

$$I_{\nu} = I_{\nu}(0)e^{-\tau_{\nu}} + S_{\nu}(1 - e^{-\tau_{\nu}})$$

• Source Function S_{ν} :

$$S_{\nu} = \frac{j_{\nu}}{\alpha_{\nu}}$$

where j_{ν} is the emissivity per unit volume, and α_{ν} is the absorption coefficient per unit length.

- Limiting Cases:
 - Optically Thin $(\tau_{\nu} \ll 1)$:

$$I_{\nu} \approx I_{\nu}(0)(1-\tau_{\nu}) + S_{\nu}\tau_{\nu}$$

– Optically Thick ($\tau_{\nu} \gg 1$):

$$I_{\nu} \approx S_{\nu}$$

2. Blackbody Radiation

• Planck Function:

$$B_{\nu}(T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{h\nu/(kT)} - 1}$$

• Energy Density of Blackbody Radiation:

$$u = aT^4$$

where $a = \frac{8\pi^5 k^4}{15h^3 c^3} \approx 7.5657 \times 10^{-16} \,\mathrm{J\,m^{-3}\,K^{-4}}$ is the radiation constant.

• Radiation Pressure:

$$P_{\rm rad} = \frac{1}{3}u = \frac{a}{3}T^4$$

• Stefan-Boltzmann Law:

$$F = \sigma T^4$$

where F is the total energy flux, and $\sigma = \frac{2\pi^5 k^4}{15h^3c^2} \approx 5.6704 \times 10^{-8} \, \mathrm{W \, m^{-2} \, K^{-4}}$ is the Stefan-Boltzmann constant.

• Average Photon Energy:

$$\langle h\nu \rangle \approx 2.7 \, kT$$

- Wien's Displacement Law:
 - Wavelength Form:

$$\lambda_{\text{max}}T = 2.8978 \times 10^{-3} \,\text{m}\,\text{K}$$

- Frequency Form:

$$\nu_{\rm max} = \frac{kT}{h} \times 2.8214$$

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3. Kirchhoff's Law for Thermal Emission

• Kirchhoff's Law:

$$\frac{j_{\nu}}{\alpha_{\nu}} = B_{\nu}(T)$$

implying that "good absorbers are good emitters."

• Blackbody Radiation Condition:

$$a_{\nu} = 1$$

for an ideal blackbody, where a_{ν} is the absorptivity at frequency ν .

4. Bremsstrahlung (Free-Free Emission)

• Emissivity per Unit Volume and Frequency j_{ν} :

$$j_{\nu} = \frac{16}{3} \left(\frac{2\pi}{3}\right)^{1/2} \frac{e^6}{m_e^2 c^3} Z^2 n_e n_i \frac{1}{\sqrt{kT}} e^{-h\nu/(kT)} \overline{g}_{ff}$$

where \overline{g}_{ff} is the velocity-averaged Gaunt factor.

• Total Emissivity Integrated over Frequency:

$$j(T) = C_1 Z^2 n_e n_i T^{1/2}$$

where C_1 is a constant that depends on fundamental constants and the Gaunt factor.

• Free-Free Absorption Coefficient α_{ν}^{ff} :

$$\alpha_{\nu}^{ff} = \frac{4}{3} \left(\frac{2\pi}{3} \right)^{1/2} \frac{e^6}{m_e^2 c} Z^2 n_e n_i \frac{1}{\nu^3 \sqrt{kT}} \left(1 - e^{-h\nu/(kT)} \right) \overline{g}_{ff}$$

• Optical Depth for Free-Free Absorption:

$$\tau_{\nu} = \alpha_{\nu}^{ff} L$$

where L is the path length.

5. Larmor's Formula for Instantaneous Power

• Instantaneous Power of an Accelerating Charge:

$$P = \frac{q^2 a^2}{6\pi\varepsilon_0 c^3}$$

6. Kinetic Theory and Thermal Velocities

• Root-Mean-Square Speed of Particles:

$$v_{\rm rms} = \sqrt{\frac{3kT}{m}}$$

• Kinetic Energy per Particle:

$$E_k = \frac{1}{2}mv_{\rm rms}^2 = \frac{3}{2}kT$$

• Hydrostatic Equilibrium:

$$\frac{dP}{dr} = -\rho \frac{GM(r)}{r^2}$$

• Virial Theorem for Spherical Systems:

$$kT = \frac{GMm_p}{3R}$$

where m_p is the proton mass.

7. Photon Energy and Frequency Relations

• Energy per Photon:

$$E = h\nu$$

• Wien's Law (High-Frequency Approximation):

$$I(\nu,T) \approx \frac{2h\nu^3}{c^2} e^{-h\nu/(kT)}$$

• Rayleigh-Jeans Law (Low-Frequency Approximation):

$$I(\nu, T) \approx \frac{2\nu^2 kT}{c^2}$$

• Brightness Temperature T_b :

$$T_b = \frac{c^2 I_{\nu}}{2k\nu^2}$$

Applicable in the Rayleigh-Jeans limit $(h\nu \ll kT)$.

8. Emission and Absorption in Plasma

• Integrated Volume Emissivity:

$$j(T) = C_1 Z^2 n_e n_i T^{1/2}$$

• Optical Depth τ_{ν} :

$$\tau_{\nu} = \alpha_{\nu} L$$

• Mean Free Path λ_{mfp} :

$$\lambda_{\rm mfp} = \frac{1}{n\sigma}$$

9. Flux and Luminosity

• Flux from a Point Source (Inverse Square Law):

$$F = \frac{L}{4\pi r^2}$$

• Flux in Terms of Specific Intensity:

$$F = \int I_{\nu} \cos \theta \, d\Omega$$

For an isotropic source:

$$F = \pi I_{\nu}$$

• Luminosity of a Spherical Blackbody:

$$L = 4\pi R^2 \sigma T^4$$

• Luminosity of Optically Thin Bremsstrahlung Emission:

$$L = j(T)V = C_1 Z^2 n_e n_i T^{1/2} \times \frac{4}{3} \pi R^3$$

10. Solutions to Radiative Transfer in Specific Cases

• Optically Thin Medium $(\tau_{\nu} \ll 1)$:

$$I_{\nu} \approx I_{\nu}(0) + j_{\nu}L$$

assuming negligible absorption.

• Optically Thick Medium $(\tau_{\nu} \gg 1)$:

$$I_{\nu} \approx S_{\nu}$$

• General Solution with No Incident Intensity $(I_{\nu}(0) = 0)$ and Constant S_{ν} :

$$I_{\nu} = S_{\nu} (1 - e^{-\tau_{\nu}})$$

11. Thermal Bremsstrahlung Spectrum

• Power Radiated in a Frequency Interval $(\nu_1, f\nu_1)$:

$$P = \int_{\nu_1}^{f\nu_1} j_{\nu}(\nu) d\nu \propto n_e n_i Z^2 T^{1/2} \left(e^{-h\nu_1/(kT)} - e^{-hf\nu_1/(kT)} \right)$$

• Frequency of Maximum Emission:

$$h\nu_{\rm max} \approx kT$$

12. Units and Constants

• Constants:

• Unit Conversions:

$$1 \text{ eV} = 1.602 \times 10^{-19} \text{ J}$$

 $1 \text{ erg} = 1 \times 10^{-7} \text{ J}$
 $1 \text{ ly} = 9.461 \times 10^{15} \text{ m}$

13. Additional Relations

• Density in Terms of Mass and Volume:

$$\rho = \frac{M}{V}$$

• Particle Number Density:

$$n = \frac{\rho}{\mu m_p}$$

where μ is the mean molecular weight.

• Total Bremsstrahlung Luminosity for Optically Thin Plasma:

$$L = C_1 Z^2 n_e n_i T^{1/2} V$$

• Hydrostatic Equilibrium (Isothermal Gas Sphere):

$$\frac{dP}{dr} = -\rho \frac{GM(r)}{r^2}$$

• Virial Theorem for Gravitational Systems:

$$2E_{\rm kin} + E_{\rm grav} = 0$$

• Relation between Temperature and Gravitational Potential:

$$kT = \frac{GMm_p}{3R}$$

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(Derived from the virial theorem for spherical systems.)