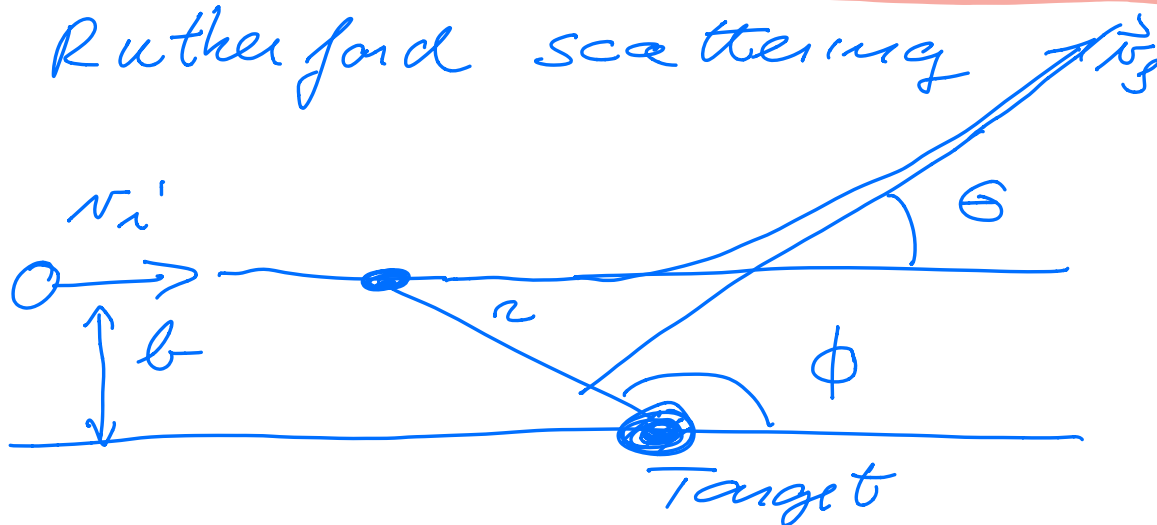


PHY 321 APRIL 7

Rutherford scattering



$$r(\phi) = \frac{c}{1 + \epsilon \cos \phi}$$

$$\frac{d\sigma}{d\Omega} = \frac{a^2}{4\pi \sin^4 \theta/2} \quad \boxed{a = \frac{\alpha}{2E}}$$

$$F(r) = -\alpha/r^2 \quad \alpha = Gm_1 m_2$$

$$\alpha = \frac{q_1 q_2}{4\pi \epsilon_0}$$

$$a = \frac{\mu \alpha}{L^2 b^2}$$

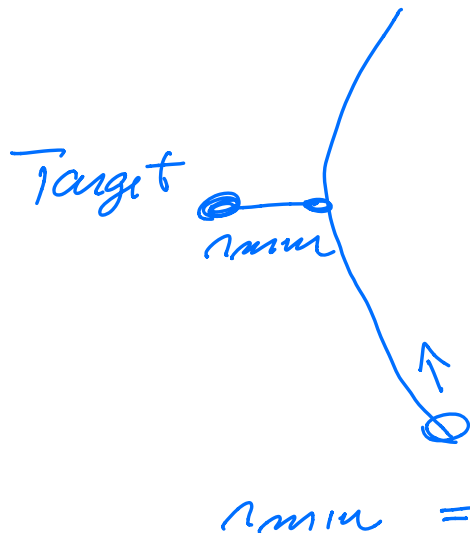
$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2}{16 E^2 \sin^4 \theta/2}$$

$$\alpha = \frac{q_1 q_2}{4\pi\epsilon_0}$$

incoming α -particles

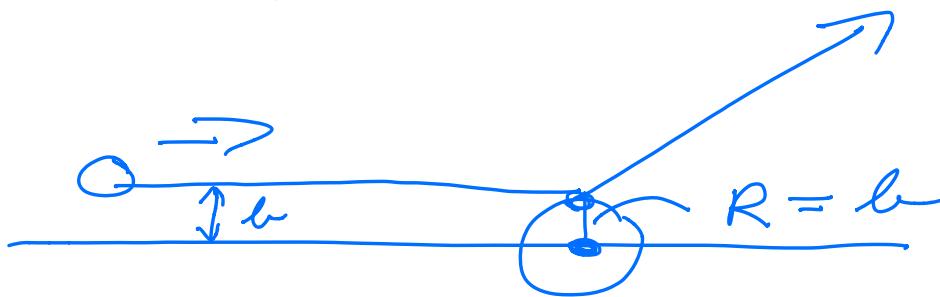
$$z=2$$

$$\text{Gold } z=79$$



$$E = \frac{L^2}{2\mu r_{\min}^2} - \frac{\alpha}{r_{\min}}$$

$$E = \frac{\frac{L^2}{2\mu R^2} - \frac{\alpha}{R}}{\frac{2\mu R^2 E + \frac{\alpha 2\mu R^2}{R}}{R}} = L^2$$



$$L^2 = 2\mu R^2 (E + \alpha/R)$$

$$L = \mu \cdot v \cdot b = b \sqrt{2\mu E}$$

$$b^2 = \frac{L^2}{2\mu E} = R^2 \left(\frac{E + \alpha/R}{E} \right)$$

$$b = R \sqrt{\frac{E + \alpha/R}{E}}$$

$$\sin \theta/2 = \frac{a}{\sqrt{a^2 + b^2}}$$

$$\Theta = 20m^{-1} \left[\frac{a}{\sqrt{a^2 + b^2}} \right]$$

$$E = 38 \text{ keV} \quad 1\text{eV} = 10^{-19} \text{ J}$$

$$\begin{aligned} \sigma &= 1400 \text{ barn} \quad 1 \text{ barn} = 10^{-28} \text{ m}^2 \\ &= \pi b^2 \quad b = R \end{aligned}$$

$$R = 7.5 \times 10^{-15} \text{ m}$$

$$Z = 2 \quad Z = 79$$

$$\sin \theta/2 = \frac{a}{\sqrt{a^2 + b^2}} \Rightarrow$$

$$\Theta = 83 \text{ degrees.}$$

===== Lagrangian formalism =====

Top down

Equations of motions from
kinetic and potential

.....

emerges

— Principle of least action

— Lagrangian

$$\mathcal{L} = K - V = \mathcal{L}(x, v, t)$$

— Euler-Lagrange equations

— variational calculus

Example 1

Harmonic oscillator

$$K = \frac{1}{2} m v^2$$

$$V = \frac{1}{2} k x^2$$

— Euler-Lagrange eqs.

$$\frac{\partial \mathcal{L}}{\partial x} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v} = 0$$

$$\frac{\partial \mathcal{L}}{\partial x} = -kx$$

$$\frac{\partial \mathcal{L}}{\partial v} = m \cdot v$$

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v} = m \frac{dv}{dt}$$

$$-kx - m \frac{dv}{dt} = 0$$

$$= \boxed{m \frac{dv}{dt} = -kx}$$

Newton's Law

$$ma = F$$

Example 2

Gravitational problem

in polar coordinates

$$\begin{aligned} K &= \frac{1}{2} \mu \dot{r}^2 + \frac{1}{2} (r \dot{\phi})^2 \mu \\ V &= -\alpha/r \end{aligned}$$

$$\mathcal{L} = \frac{1}{2} \mu \dot{r}^2 + \frac{1}{2} (r \dot{\phi})^2 \mu$$

$$+ \alpha/r = \mathcal{L}(r, \dot{r}, \phi, \dot{\phi}, t)$$

$$\frac{\partial \mathcal{L}}{\partial r} = \mu \cdot r \dot{\phi}^2 - \alpha/r^2$$

$$\frac{\partial \mathcal{L}}{\partial \dot{r}} = \mu \cdot \dot{r}$$

$$v_r$$

$$v_r = \dot{r} = \frac{dr}{dt}$$

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v_r} = \mu \cdot \frac{d^2 r}{dt^2}$$

$$\frac{\partial \mathcal{L}}{\partial r} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v_r} = 0$$

$$\mu r \dot{\phi}^2 - \alpha/r^2 - \mu \ddot{r} = 0$$

$$\boxed{\mu \ddot{r} = \mu r \dot{\phi}^2 - \alpha/r^2}$$

$$\mu a_r = \frac{L^2}{\mu r^3} - \alpha/r^2$$

$$= - \frac{dV_{\text{eff}}}{dr}$$

$$V_{\text{eff}}(r) = V(r) + \frac{L^2}{\mu r^2}$$

$$\underline{\mu a_\phi = 0}$$

$$\boxed{\dot{\phi} = \frac{d\phi}{dt} \propto L}$$