# Computational Physics Lectures: Ten algorithms for quantum mechanical systems

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#### Overview

In this contribution we outline central algorithms for studying quantum mechanical systems, with an emphasis on both computational and pedagogical aspects. Using simple systems that allow for analytical solutions, we show how one can move from linear algebra and eigenvalue algorithms using for example full configuration interaction theory, to stochastic methods like variational and Diffusion Monte Carlo approaches and finally, how we can use Monte Carlo methods together with deep learning algorithms. Along this journey we will present ten central algorithms which have changed considerably the way we study interacting many-particle systems. These algorithms span from Householder's famous transformation of matrices, via iterative eigenvalue solvers to neural networks and automatic differentiation for optimizing multidimensional functions. Codes and jupyter-notebooks are provided, allowing thereby the reader to experiment and practice the various methods.

#### Algorithms

- Linear algebra and eigenvalue problems
  - 1. Houselholder's transformation
  - 2. Jacobi/Givens rotations
  - 3. Iterative methods, Lanczsos' method
- Monte Carlo methods

- 1. Variational Monte Carlo
- 2. Metropolis-Hastings algorithm
- Optimization problems
  - 1. Gradient descent and steepest gradient descent
  - 2. Adaptive methods
- Automatic differentiation
- Deep learning
  - 1. Neural Networks
  - 2. Reduced Boltzmann machines

#### Eigenvalue problems, basic definitions

Let us consider the matrix  $\mathbf{A}$  of dimension n. The eigenvalues of  $\mathbf{A}$  are defined through the matrix equation

$$\mathbf{A}\mathbf{x}^{(\nu)} = \lambda^{(\nu)}\mathbf{x}^{(\nu)},$$

where  $\lambda^{(\nu)}$  are the eigenvalues and  $\mathbf{x}^{(\nu)}$  the corresponding eigenvectors. Unless otherwise stated, when we use the wording eigenvector we mean the right eigenvector. The left eigenvalue problem is defined as

$$\mathbf{x}_L^{(\nu)}\mathbf{A} = \lambda^{(\nu)}\mathbf{x}_L^{(\nu)}$$

The above right eigenvector problem is equivalent to a set of n equations with n unknowns  $x_i$ .

#### Eigenvalue problems, basic definitions

The eigenvalue problem can be rewritten as

$$\left(\mathbf{A} - \lambda^{(\nu)} \mathbf{I}\right) \mathbf{x}^{(\nu)} = 0,$$

with  $\mathbf{I}$  being the unity matrix. This equation provides a solution to the problem if and only if the determinant is zero, namely

$$\left|\mathbf{A} - \lambda^{(\nu)} \mathbf{I}\right| = 0,$$

which in turn means that the determinant is a polynomial of degree n in  $\lambda$  and in general we will have n distinct zeros.

The eigenvalues of a matrix  $\mathbf{A} \in \mathbb{C}^{n \times n}$  are thus the n roots of its characteristic polynomial

$$P(\lambda) = \det(\lambda \mathbf{I} - \mathbf{A}),$$

or

$$P(\lambda) = \prod_{i=1}^{n} (\lambda_i - \lambda).$$

The set of these roots is called the spectrum and is denoted as  $\lambda(\mathbf{A})$ . If  $\lambda(\mathbf{A}) = \{\lambda_1, \lambda_2, \dots, \lambda_n\}$  then we have

$$det(\mathbf{A}) = \lambda_1 \lambda_2 \dots \lambda_n,$$

and if we define the trace of  $\bf A$  as

$$Tr(\mathbf{A}) = \sum_{i=1}^{n} a_{ii}$$

then

$$Tr(\mathbf{A}) = \lambda_1 + \lambda_2 + \dots + \lambda_n.$$

#### Abel-Ruffini Impossibility Theorem

The *Abel-Ruffini* theorem (also known as Abel's impossibility theorem) states that there is no general solution in radicals to polynomial equations of degree five or higher.

The content of this theorem is frequently misunderstood. It does not assert that higher-degree polynomial equations are unsolvable. In fact, if the polynomial has real or complex coefficients, and we allow complex solutions, then every polynomial equation has solutions; this is the fundamental theorem of algebra. Although these solutions cannot always be computed exactly with radicals, they can be computed to any desired degree of accuracy using numerical methods such as the Newton-Raphson method or Laguerre method, and in this way they are no different from solutions to polynomial equations of the second, third, or fourth degrees.

The theorem only concerns the form that such a solution must take. The content of the theorem is that the solution of a higher-degree equation cannot in all cases be expressed in terms of the polynomial coefficients with a finite number of operations of addition, subtraction, multiplication, division and root extraction. Some polynomials of arbitrary degree, of which the simplest nontrivial example is the monomial equation  $ax^n = b$ , are always solvable with a radical.

The Abel-Ruffini theorem says that there are some fifth-degree equations whose solution cannot be so expressed. The equation  $x^5 - x + 1 = 0$  is an example. Some other fifth degree equations can be solved by radicals, for example  $x^5 - x^4 - x + 1 = 0$ . The precise criterion that distinguishes between those equations that can be solved by radicals and those that cannot was given by Galois and is now part of Galois theory: a polynomial equation can be solved by radicals if and only if its Galois group is a solvable group.

Today, in the modern algebraic context, we say that second, third and fourth degree polynomial equations can always be solved by radicals because the symmetric groups  $S_2$ ,  $S_3$  and  $S_4$  are solvable groups, whereas  $S_n$  is not solvable for  $n \geq 5$ .

#### Eigenvalue problems, basic definitions

In the present discussion we assume that our matrix is real and symmetric, that is  $\mathbf{A} \in \mathbb{R}^{n \times n}$ . The matrix  $\mathbf{A}$  has n eigenvalues  $\lambda_1 \dots \lambda_n$  (distinct or not). Let  $\mathbf{D}$  be the diagonal matrix with the eigenvalues on the diagonal

$$\mathbf{D} = \begin{pmatrix} \lambda_1 & 0 & 0 & 0 & \dots & 0 & 0 \\ 0 & \lambda_2 & 0 & 0 & \dots & 0 & 0 \\ 0 & 0 & \lambda_3 & 0 & 0 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & \dots & \dots & \dots & \dots & \lambda_{n-1} \\ 0 & \dots & \dots & \dots & \dots & 0 & \lambda_n \end{pmatrix}.$$

If A is real and symmetric then there exists a real orthogonal matrix S such that

$$\mathbf{S}^T \mathbf{A} \mathbf{S} = \operatorname{diag}(\lambda_1, \lambda_2, \dots, \lambda_n),$$

and for j = 1 : n we have  $\mathbf{AS}(:, j) = \lambda_j \mathbf{S}(:, j)$ .

To obtain the eigenvalues of  $\mathbf{A} \in \mathbb{R}^{n \times n}$ , the strategy is to perform a series of similarity transformations on the original matrix  $\mathbf{A}$ , in order to reduce it either into a diagonal form as above or into a tridiagonal form.

We say that a matrix  $\mathbf{B}$  is a similarity transform of  $\mathbf{A}$  if

$$\mathbf{B} = \mathbf{S}^T \mathbf{A} \mathbf{S}, \quad \text{where} \quad \mathbf{S}^T \mathbf{S} = \mathbf{S}^{-1} \mathbf{S} = \mathbf{I}.$$

The importance of a similarity transformation lies in the fact that the resulting matrix has the same eigenvalues, but the eigenvectors are in general different.

To prove this we start with the eigenvalue problem and a similarity transformed matrix  ${\bf B}.$ 

$$\mathbf{A}\mathbf{x} = \lambda \mathbf{x}$$
 and  $\mathbf{B} = \mathbf{S}^T \mathbf{A} \mathbf{S}$ .

We multiply the first equation on the left by  $\mathbf{S}^T$  and insert  $\mathbf{S}^T\mathbf{S} = \mathbf{I}$  between  $\mathbf{A}$  and  $\mathbf{x}$ . Then we get

$$(\mathbf{S}^T \mathbf{A} \mathbf{S})(\mathbf{S}^T \mathbf{x}) = \lambda \mathbf{S}^T \mathbf{x},\tag{1}$$

which is the same as

$$\mathbf{B}\left(\mathbf{S}^{T}\mathbf{x}\right) = \lambda\left(\mathbf{S}^{T}\mathbf{x}\right).$$

The variable  $\lambda$  is an eigenvalue of **B** as well, but with eigenvector  $\mathbf{S}^T \mathbf{x}$ . The basic philosophy is to

• Either apply subsequent similarity transformations (direct method) so that

$$\mathbf{S}_{N}^{T} \dots \mathbf{S}_{1}^{T} \mathbf{A} \mathbf{S}_{1} \dots \mathbf{S}_{N} = \mathbf{D}, \tag{2}$$

- Or apply subsequent similarity transformations so that **A** becomes tridiagonal (Householder) or upper/lower triangular (the *QR* method to be discussed later).
- Thereafter, techniques for obtaining eigenvalues from tridiagonal matrices can be used.
- Or use so-called power methods
- Or use iterative methods (Krylov, Lanczos, Arnoldi). These methods are popular for huge matrix problems.

One speaks normally of two main approaches to solving the eigenvalue problem.

- The first is the formal method, involving determinants and the characteristic polynomial. This proves how many eigenvalues there are, and is the way most of you learned about how to solve the eigenvalue problem, but for matrices of dimensions greater than 2 or 3, it is rather impractical.
- The other general approach is to use similarity or unitary tranformations to reduce a matrix to diagonal form. This is normally done in two steps: first reduce to for example a *tridiagonal* form, and then to diagonal form. The main algorithms we will discuss in detail, Jacobi's and Householder's (so-called direct method) and Lanczos algorithms (an iterative method), follow this methodology.

Direct or non-iterative methods require for matrices of dimensionality  $n \times n$  typically  $O(n^3)$  operations. These methods are normally called standard methods and are used for dimensionalities  $n \sim 10^5$  or smaller. A brief historical overview

Year	n	
1950	n = 20	(Wilkinson)
1965	n = 200	(Forsythe et al.)
1980	n = 2000	Linpack
1995	n = 20000	Lapack
2017	$n\sim 10^5$	Lapack

shows that in the course of 60 years the dimension that direct diagonalization methods can handle has increased by almost a factor of  $10^4$ . However, it pales beside the progress achieved by computer hardware, from flops to petaflops, a factor of almost  $10^{15}$ . We see clearly played out in history the  $O(n^3)$  bottleneck of direct matrix algorithms.

Sloppily speaking, when  $n \sim 10^4$  is cubed we have  $O(10^{12})$  operations, which is smaller than the  $10^{15}$  increase in flops.

If the matrix to diagonalize is large and sparse, direct methods simply become impractical, also because many of the direct methods tend to destroy sparsity. As a result large dense matrices may arise during the diagonalization procedure. The idea behind iterative methods is to project the n-dimensional problem in smaller spaces, so-called Krylov subspaces. Given a matrix  $\bf A$  and a vector  $\bf v$ , the associated Krylov sequences of vectors (and thereby subspaces)  $\bf v$ ,  $\bf Av$ ,  $\bf A^2v$ ,  $\bf A^3v$ , ..., represent successively larger Krylov subspaces.

#### Discussion of Jacobi's method for eigenvalues

Consider an example of an  $(n \times n)$  orthogonal transformation matrix

$$\mathbf{S} = \begin{pmatrix} 1 & 0 & \dots & 0 & 0 & \dots & 0 & 0 \\ 0 & 1 & \dots & 0 & 0 & \dots & 0 & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots & 0 & \dots \\ 0 & 0 & \dots & \cos \theta & 0 & \dots & 0 & \sin \theta \\ 0 & 0 & \dots & 0 & 1 & \dots & 0 & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots & 1 & \dots \\ 0 & 0 & \dots & -\sin \theta & 0 & \dots & 0 & \cos \theta \end{pmatrix}$$

with property  $\mathbf{S^T} = \mathbf{S^{-1}}$ . It performs a plane rotation around an angle  $\theta$  in the Euclidean n-dimensional space.

It means that its matrix elements that differ from zero are given by

$$s_{kk} = s_{ll} = \cos\theta, s_{kl} = -s_{lk} = -\sin\theta, s_{ii} = 1$$
  $i \neq k$   $i \neq l$ ,

A similarity transformation

$$\mathbf{B} = \mathbf{S}^T \mathbf{A} \mathbf{S}$$

results in

$$b_{ik} = a_{ik} \cos \theta - a_{il} \sin \theta, i \neq k, i \neq l$$

$$b_{il} = a_{il} \cos \theta + a_{ik} \sin \theta, i \neq k, i \neq l$$

$$b_{kk} = a_{kk} \cos^2 \theta - 2a_{kl} \cos \theta \sin \theta + a_{ll} \sin^2 \theta$$

$$b_{ll} = a_{ll} \cos^2 \theta + 2a_{kl} \cos \theta \sin \theta + a_{kk} \sin^2 \theta$$

$$b_{kl} = (a_{kk} - a_{ll}) \cos \theta \sin \theta + a_{kl} (\cos^2 \theta - \sin^2 \theta)$$

The angle  $\theta$  is arbitrary. The recipe is to choose  $\theta$  so that all non-diagonal matrix elements  $b_{kl}$  become zero.

The main idea is thus to reduce systematically the norm of the off-diagonal matrix elements of a matrix  ${\bf A}$ 

$$\mathrm{off}(\mathbf{A}) = \sqrt{\sum_{i=1}^{n} \sum_{j=1, j \neq i}^{n} a_{ij}^2}.$$

To demonstrate the algorithm, we consider the simple  $2 \times 2$  similarity transformation of the full matrix. The matrix is symmetric, we single out  $1 \le k < l \le n$  and use the abbreviations  $c = \cos \theta$  and  $s = \sin \theta$  to obtain

$$\left(\begin{array}{cc} b_{kk} & 0 \\ 0 & b_{ll} \end{array}\right) = \left(\begin{array}{cc} c & -s \\ s & c \end{array}\right) \left(\begin{array}{cc} a_{kk} & a_{kl} \\ a_{lk} & a_{ll} \end{array}\right) \left(\begin{array}{cc} c & s \\ -s & c \end{array}\right).$$

We require that the non-diagonal matrix elements  $b_{kl}=b_{lk}=0$ , implying that

$$a_{kl}(c^2 - s^2) + (a_{kk} - a_{ll})cs = b_{kl} = 0.$$

If  $a_{kl} = 0$  one sees immediately that  $\cos \theta = 1$  and  $\sin \theta = 0$ .

The Frobenius norm of an orthogonal transformation is always preserved. The Frobenius norm is defined as

$$norm(\mathbf{A})_F = \sqrt{\sum_{i=1}^n \sum_{j=1}^n |a_{ij}|^2}.$$

This means that for our  $2 \times 2$  case we have

$$2a_{kl}^2 + a_{kk}^2 + a_{ll}^2 = b_{kk}^2 + b_{ll}^2,$$

which leads to

off(**B**)<sup>2</sup> = norm(**B**)<sub>F</sub><sup>2</sup> - 
$$\sum_{i=1}^{n} b_{ii}^{2}$$
 = off(**A**)<sup>2</sup> -  $2a_{kl}^{2}$ ,

since

$$\operatorname{norm}(\mathbf{B})_F^2 - \sum_{i=1}^n b_{ii}^2 = \operatorname{norm}(\mathbf{A})_F^2 - \sum_{i=1}^n a_{ii}^2 + (a_{kk}^2 + a_{ll}^2 - b_{kk}^2 - b_{ll}^2).$$

This results means that the matrix A moves closer to diagonal form for each transformation.

Defining the quantities  $\tan \theta = t = s/c$  and

$$\cot 2\theta = \tau = \frac{a_{ll} - a_{kk}}{2a_{kl}},$$

we obtain the quadratic equation (using  $\cot 2\theta = 1/2(\cot \theta - \tan \theta)$ 

$$t^2 + 2\tau t - 1 = 0$$
.

resulting in

$$t = -\tau \pm \sqrt{1 + \tau^2},$$

and c and s are easily obtained via

$$c = \frac{1}{\sqrt{1+t^2}},$$

and s = tc. Convince yourself that we have  $|\theta| \le \pi/4$ . This has the effect of minimizing the difference between the matrices **B** and **A** since

$$\operatorname{norm}(\mathbf{B} - \mathbf{A})_F^2 = 4(1 - c) \sum_{i=1, i \neq k, l}^n (a_{ik}^2 + a_{il}^2) + \frac{2a_{kl}^2}{c^2}.$$

• Choose a tolerance  $\epsilon$ , making it a small number, typically  $10^{-8}$  or smaller.

• Setup a *while* test where one compares the norm of the newly computed off-diagonal matrix elements

off(**A**) = 
$$\sqrt{\sum_{i=1}^{n} \sum_{j=1, j \neq i}^{n} a_{ij}^2} > \epsilon$$
.

- Now choose the matrix elements  $a_{kl}$  so that we have those with largest value, that is  $|a_{kl}| = \max_{i \neq j} |a_{ij}|$ .
- Compute thereafter  $\tau = (a_{ll} a_{kk})/2a_{kl}$ ,  $\tan \theta$ ,  $\cos \theta$  and  $\sin \theta$ .
- Compute thereafter the similarity transformation for this set of values (k, l), obtaining the new matrix  $\mathbf{B} = \mathbf{S}(k, l, \theta)^T \mathbf{A} \mathbf{S}(k, l, \theta)$ .
- Compute the new norm of the off-diagonal matrix elements and continue till you have satisfied off( $\mathbf{B}$ )  $\leq \epsilon$

The convergence rate of the Jacobi method is however poor, one needs typically  $3n^2 - 5n^2$  rotations and each rotation requires 4n operations, resulting in a total of  $12n^3 - 20n^3$  operations in order to zero out non-diagonal matrix elements.

We specialize to a symmetric  $3 \times 3$  matrix **A**. We start the process as follows (assuming that  $a_{23} = a_{32}$  is the largest non-diagonal) with  $c = \cos \theta$  and  $s = \sin \theta$ 

$$\mathbf{B} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c & -s \\ 0 & s & c \end{pmatrix} \begin{pmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ a_{31} & a_{32} & a_{33} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & c & s \\ 0 & -s & c \end{pmatrix}.$$

We will choose the angle  $\theta$  in order to have  $a_{23} = a_{32} = 0$ . We get (symmetric matrix)

$$\mathbf{B} = \begin{pmatrix} a_{11} & a_{12}c - a_{13}s & a_{12}s + a_{13}c \\ a_{12}c - a_{13}s & a_{22}c^2 + a_{33}s^2 - 2a_{23}sc & (a_{22} - a_{33})sc + a_{23}(c^2 - s^2) \\ a_{12}s + a_{13}c & (a_{22} - a_{33})sc + a_{23}(c^2 - s^2) & a_{22}s^2 + a_{33}c^2 + 2a_{23}sc \end{pmatrix}.$$

Note that  $a_{11}$  is unchanged! As it should.

We have

$$\mathbf{B} = \begin{pmatrix} a_{11} & a_{12}c - a_{13}s & a_{12}s + a_{13}c \\ a_{12}c - a_{13}s & a_{22}c^2 + a_{33}s^2 - 2a_{23}sc & (a_{22} - a_{33})sc + a_{23}(c^2 - s^2) \\ a_{12}s + a_{13}c & (a_{22} - a_{33})sc + a_{23}(c^2 - s^2) & a_{22}s^2 + a_{33}c^2 + 2a_{23}sc \end{pmatrix}.$$

or

$$b_{11} = a_{11}$$

$$b_{12} = a_{12}\cos\theta - a_{13}\sin\theta, 1 \neq 2, 1 \neq 3$$

$$b_{13} = a_{13}\cos\theta + a_{12}\sin\theta, 1 \neq 2, 1 \neq 3$$

$$b_{22} = a_{22}\cos^2\theta - 2a_{23}\cos\theta\sin\theta + a_{33}\sin^2\theta$$

$$b_{33} = a_{33}\cos^2\theta + 2a_{23}\cos\theta\sin\theta + a_{22}\sin^2\theta$$

$$b_{23} = (a_{22} - a_{33})\cos\theta\sin\theta + a_{23}(\cos^2\theta - \sin^2\theta)$$

We will fix the angle  $\theta$  so that  $b_{23} = 0$ .

We get then a new matrix

$$\mathbf{B} = \left(\begin{array}{ccc} b_{11} & b_{12} & b_{13} \\ b_{12} & b_{22} & 0 \\ b_{13} & 0 & a_{33} \end{array}\right).$$

We repeat then assuming that  $b_{12}$  is the largest non-diagonal matrix element and get a new matrix

$$\mathbf{C} = \begin{pmatrix} c & -s & 0 \\ s & c & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} b_{11} & b_{12} & b_{13} \\ b_{12} & b_{22} & 0 \\ b_{13} & 0 & b_{33} \end{pmatrix} \begin{pmatrix} c & s & 0 \\ -s & c & 0 \\ 0 & 0 & 1 \end{pmatrix}.$$

We continue this process till all non-diagonal matrix elements are zero (ideally). You will notice that performing the above operations that the matrix element  $b_{23}$  which was previous zero becomes different from zero. This is one of the problems which slows down the jacobi procedure.

The more general expression for the new matrix elements are

$$b_{ii} = a_{ii}, i \neq k, i \neq l$$

$$b_{ik} = a_{ik} \cos \theta - a_{il} \sin \theta, i \neq k, i \neq l$$

$$b_{il} = a_{il} \cos \theta + a_{ik} \sin \theta, i \neq k, i \neq l$$

$$b_{kk} = a_{kk} \cos^2 \theta - 2a_{kl} \cos \theta \sin \theta + a_{ll} \sin^2 \theta$$

$$b_{ll} = a_{ll} \cos^2 \theta + 2a_{kl} \cos \theta \sin \theta + a_{kk} \sin^2 \theta$$

$$b_{kl} = (a_{kk} - a_{ll}) \cos \theta \sin \theta + a_{kl} (\cos^2 \theta - \sin^2 \theta)$$

This is what we will need to code.

#### Our system

We are first interested in the solution of the radial part of Schroedinger's equation for one electron. This equation reads

$$-\frac{\hbar^2}{2m}\left(\frac{1}{r^2}\frac{d}{dr}r^2\frac{d}{dr}-\frac{l(l+1)}{r^2}\right)R(r)+V(r)R(r)=ER(r).$$

In our case V(r) is the harmonic oscillator potential  $(1/2)kr^2$  with  $k=m\omega^2$  and E is the energy of the harmonic oscillator in three dimensions. The oscillator frequency is  $\omega$  and the energies are

$$E_{nl} = \hbar\omega \left(2n + l + \frac{3}{2}\right),\,$$

with  $n = 0, 1, 2, \dots$  and  $l = 0, 1, 2, \dots$ 

Since we have made a transformation to spherical coordinates it means that  $r \in [0, \infty)$ . The quantum number l is the orbital momentum of the electron. Then we substitute R(r) = (1/r)u(r) and obtain

$$-\frac{\hbar^2}{2m}\frac{d^2}{dr^2}u(r) + \left(V(r) + \frac{l(l+1)}{r^2}\frac{\hbar^2}{2m}\right)u(r) = Eu(r).$$

The boundary conditions are u(0) = 0 and  $u(\infty) = 0$ .

We introduce a dimensionless variable  $\rho = (1/\alpha)r$  where  $\alpha$  is a constant with dimension length and get

$$-\frac{\hbar^2}{2m\alpha^2}\frac{d^2}{d\rho^2}u(\rho) + \left(V(\rho) + \frac{l(l+1)}{\rho^2}\frac{\hbar^2}{2m\alpha^2}\right)u(\rho) = Eu(\rho).$$

In project 2 we choose l=0. Inserting  $V(\rho)=(1/2)k\alpha^2\rho^2$  we end up with

$$-\frac{\hbar^2}{2m\alpha^2}\frac{d^2}{d\rho^2}u(\rho) + \frac{k}{2}\alpha^2\rho^2u(\rho) = Eu(\rho).$$

We multiply thereafter with  $2m\alpha^2/\hbar^2$  on both sides and obtain

$$-\frac{d^2}{d\rho^2}u(\rho) + \frac{mk}{\hbar^2}\alpha^4\rho^2u(\rho) = \frac{2m\alpha^2}{\hbar^2}Eu(\rho).$$

We have thus

$$-\frac{d^2}{d\rho^2}u(\rho) + \frac{mk}{\hbar^2}\alpha^4\rho^2u(\rho) = \frac{2m\alpha^2}{\hbar^2}Eu(\rho).$$

The constant  $\alpha$  can now be fixed so that

$$\frac{mk}{\hbar^2}\alpha^4 = 1,$$

or

$$\alpha = \left(\frac{\hbar^2}{mk}\right)^{1/4}.$$

Defining

$$\lambda = \frac{2m\alpha^2}{\hbar^2}E,$$

we can rewrite Schroedinger's equation as

$$-\frac{d^2}{d\rho^2}u(\rho) + \rho^2 u(\rho) = \lambda u(\rho).$$

This is the first equation to solve numerically. In three dimensions the eigenvalues for l=0 are  $\lambda_0=3, \lambda_1=7, \lambda_2=11, \ldots$ 

We use the by now standard expression for the second derivative of a function  $\boldsymbol{u}$ 

$$u'' = \frac{u(\rho + h) - 2u(\rho) + u(\rho - h)}{h^2} + O(h^2), \tag{3}$$

where h is our step. Next we define minimum and maximum values for the variable  $\rho$ ,  $\rho_{\min} = 0$  and  $\rho_{\max}$ , respectively. You need to check your results for the energies against different values  $\rho_{\max}$ , since we cannot set  $\rho_{\max} = \infty$ .

With a given number of steps,  $n_{\text{step}}$ , we then define the step h as

$$h = \frac{\rho_{\text{max}} - \rho_{\text{min}}}{n_{\text{step}}}.$$

Define an arbitrary value of  $\rho$  as

$$\rho_i = \rho_{\min} + ih$$
 $i = 0, 1, 2, \dots, n_{\text{step}}$ 

we can rewrite the Schrödinger equation for  $\rho_i$  as

$$-\frac{u(\rho_i + h) - 2u(\rho_i) + u(\rho_i - h)}{h^2} + \rho_i^2 u(\rho_i) = \lambda u(\rho_i),$$

or in a more compact way

$$-\frac{u_{i+1} - 2u_i + u_{i-1}}{h^2} + \rho_i^2 u_i = -\frac{u_{i+1} - 2u_i + u_{i-1}}{h^2} + V_i u_i = \lambda u_i,$$

where  $V_i = \rho_i^2$  is the harmonic oscillator potential.

Define first the diagonal matrix element

$$d_i = \frac{2}{h^2} + V_i,$$

and the non-diagonal matrix element

$$e_i = -\frac{1}{h^2}.$$

In this case the non-diagonal matrix elements are given by a mere constant. All non-diagonal matrix elements are equal.

With these definitions the Schroedinger equation takes the following form

$$d_i u_i + e_{i-1} u_{i-1} + e_{i+1} u_{i+1} = \lambda u_i,$$

where  $u_i$  is unknown. We can write the latter equation as a matrix eigenvalue problem

$$\begin{pmatrix} d_1 & e_1 & 0 & 0 & \dots & 0 & 0 \\ e_1 & d_2 & e_2 & 0 & \dots & 0 & 0 \\ 0 & e_2 & d_3 & e_3 & 0 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & \dots & \dots & \dots & \dots & d_{n_{\text{step}}-2} & e_{n_{\text{step}}-1} \\ 0 & \dots & \dots & \dots & \dots & e_{n_{\text{step}}-1} & d_{n_{\text{step}}-1} \end{pmatrix} \begin{pmatrix} u_1 \\ u_2 \\ \dots \\ \dots \\ u_{n_{\text{step}}-1} \end{pmatrix} = \lambda \begin{pmatrix} u_1 \\ u_2 \\ \dots \\ \dots \\ u_{n_{\text{step}}-1} \end{pmatrix}$$

$$(4)$$

or if we wish to be more detailed, we can write the tridiagonal matrix as

$$\begin{pmatrix}
\frac{2}{h^2} + V_1 & -\frac{1}{h^2} & 0 & 0 & \dots & 0 & 0 \\
-\frac{1}{h^2} & \frac{2}{h^2} + V_2 & -\frac{1}{h^2} & 0 & \dots & 0 & 0 \\
0 & -\frac{1}{h^2} & \frac{2}{h^2} + V_3 & -\frac{1}{h^2} & 0 & \dots & 0 \\
\dots & \dots & \dots & \dots & \dots & \dots & \dots \\
0 & \dots & \dots & \dots & \dots & \dots & \dots \\
0 & \dots & \dots & \dots & \dots & \frac{2}{h^2} + V_{n_{\text{step}}-2} & -\frac{1}{h^2} \\
0 & \dots & \dots & \dots & \dots & \frac{2}{h^2} + V_{n_{\text{step}}-1}
\end{pmatrix}$$
(5)

Recall that the solutions are known via the boundary conditions at  $i = n_{\text{step}}$  and at the other end point, that is for  $\rho_0$ . The solution is zero in both cases.

#### Two electrons that interact

We are going to study two electrons in a harmonic oscillator well which also interact via a repulsive Coulomb interaction. Let us start with the single-electron equation written as

$$-\frac{\hbar^2}{2m}\frac{d^2}{dr^2}u(r) + \frac{1}{2}kr^2u(r) = E^{(1)}u(r),$$

where  $E^{(1)}$  stands for the energy with one electron only. For two electrons with no repulsive Coulomb interaction, we have the following Schroedinger equation

$$\left(-\frac{\hbar^2}{2m}\frac{d^2}{dr_1^2} - \frac{\hbar^2}{2m}\frac{d^2}{dr_2^2} + \frac{1}{2}kr_1^2 + \frac{1}{2}kr_2^2\right)u(r_1, r_2) = E^{(2)}u(r_1, r_2).$$

Note that we deal with a two-electron wave function  $u(r_1, r_2)$  and two-electron energy  $E^{(2)}$ .

With no interaction this can be written out as the product of two singleelectron wave functions, that is we have a solution on closed form.

We introduce the relative coordinate  $\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$  and the center-of-mass coordinate  $\mathbf{R} = 1/2(\mathbf{r}_1 + \mathbf{r}_2)$ . With these new coordinates, the radial Schroedinger equation reads

$$\left(-\frac{\hbar^2}{m}\frac{d^2}{dr^2} - \frac{\hbar^2}{4m}\frac{d^2}{dR^2} + \frac{1}{4}kr^2 + kR^2\right)u(r,R) = E^{(2)}u(r,R).$$

The equations for r and R can be separated via the ansatz for the wave function  $u(r,R) = \psi(r)\phi(R)$  and the energy is given by the sum of the relative energy  $E_r$  and the center-of-mass energy  $E_R$ , that is

$$E^{(2)} = E_r + E_R.$$

We add then the repulsive Coulomb interaction between two electrons, namely a term

$$V(r_1, r_2) = \frac{\beta e^2}{|\mathbf{r}_1 - \mathbf{r}_2|} = \frac{\beta e^2}{r},$$

with  $\beta e^2 = 1.44$  eVnm.

Adding this term, the r-dependent Schroedinger equation becomes

$$\left(-\frac{\hbar^2}{m}\frac{d^2}{dr^2} + \frac{1}{4}kr^2 + \frac{\beta e^2}{r}\right)\psi(r) = E_r\psi(r).$$

This equation is similar to the one we had previously in parts (a) and (b) and we introduce again a dimensionless variable  $\rho = r/\alpha$ . Repeating the same steps, we arrive at

$$-\frac{d^2}{d\rho^2}\psi(\rho) + \frac{mk}{4\hbar^2}\alpha^4\rho^2\psi(\rho) + \frac{m\alpha\beta e^2}{\rho\hbar^2}\psi(\rho) = \frac{m\alpha^2}{\hbar^2}E_r\psi(\rho).$$

We want to manipulate this equation further to make it as similar to that in (a) as possible. We define a 'frequency'

$$\omega_r^2 = \frac{1}{4} \frac{mk}{\hbar^2} \alpha^4,$$

and fix the constant  $\alpha$  by requiring

$$\frac{m\alpha\beta e^2}{\hbar^2} = 1$$

or

$$\alpha = \frac{\hbar^2}{m\beta e^2}.$$

Defining

$$\lambda = \frac{m\alpha^2}{\hbar^2} E,$$

we can rewrite Schroedinger's equation as

$$-\frac{d^2}{d\rho^2}\psi(\rho) + \omega_r^2 \rho^2 \psi(\rho) + \frac{1}{\rho}\psi(\rho) = \lambda \psi(\rho).$$

We treat  $\omega_r$  as a parameter which reflects the strength of the oscillator potential.

Here we will study the cases  $\omega_r = 0.01$ ,  $\omega_r = 0.5$ ,  $\omega_r = 1$ , and  $\omega_r = 5$  for the ground state only, that is the lowest-lying state.

With no repulsive Coulomb interaction you should get a result which corresponds to the relative energy of a non-interacting system. Make sure your results are stable as functions of  $\rho_{\text{max}}$  and the number of steps.

We are only interested in the ground state with l=0. We omit the center-of-mass energy.

For specific oscillator frequencies, the above equation has analytic answers, see the article by M. Taut, Phys. Rev. A 48, 3561 - 3566 (1993). The article can be retrieved from the following web address http://prola.aps.org/abstract/PRA/v48/i5/p3561\_1.

# Discussion of Householder's method for eigenvalues

The drawbacks with Jacobi's method are rather obvious, with perhaps the most negative feature being the fact that we cannot tell \* a priori\* how many transformations are needed. Can we do better? The answer to this is yes and is given by a clever algorithm outlined by Householder. It was ranked among the top ten algorithms in the previous century. We will discuss this algorithm in more detail below.

The first step consists in finding an orthogonal matrix S which is the product of (n-2) orthogonal matrices

$$\mathbf{S} = \mathbf{S}_1 \mathbf{S}_2 \dots \mathbf{S}_{n-2},$$

each of which successively transforms one row and one column of  $\mathbf{A}$  into the required tridiagonal form. Only n-2 transformations are required, since the last two elements are already in tridiagonal form.

In order to determine each  $S_i$  let us see what happens after the first multiplication, namely,

$$\mathbf{S}_{1}^{T}\mathbf{A}\mathbf{S}_{1} = \begin{pmatrix} a_{11} & e_{1} & 0 & 0 & \dots & 0 & 0 \\ e_{1} & a_{22}' & a_{23}' & \dots & \dots & a_{2n}' \\ 0 & a_{32}' & a_{33}' & \dots & \dots & \dots & a_{3n}' \\ 0 & \dots & \dots & \dots & \dots & \dots \\ 0 & a_{n2}' & a_{n3}' & \dots & \dots & \dots & a_{nn}' \end{pmatrix}$$

where the primed quantities represent a matrix  $\mathbf{A}'$  of dimension n-1 which will subsequentely be transformed by  $\mathbf{S}_2$ .

The factor  $e_1$  is a possibly non-vanishing element. The next transformation produced by  $\mathbf{S}_2$  has the same effect as  $\mathbf{S}s$  but now on the submatirx  $\mathbf{A}'$  only

$$(\mathbf{S}_1 \mathbf{S}_2)^T \mathbf{A} \mathbf{S}_1 \mathbf{S}_2 = \begin{pmatrix} a_{11} & e_1 & 0 & 0 & \dots & 0 & 0 \\ e_1 & a'_{22} & e_2 & 0 & \dots & \dots & 0 \\ 0 & e_2 & a''_{33} & \dots & \dots & \dots & a''_{3n} \\ 0 & \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & 0 & a''_{n3} & \dots & \dots & \dots & a''_{nn} \end{pmatrix}$$

Note that the effective size of the matrix on which we apply the transformation reduces for every new step. In the previous Jacobi method each similarity transformation is in principle performed on the full size of the original matrix.

After a series of such transformations, we end with a set of diagonal matrix elements

$$a_{11}, a'_{22}, a''_{33} \dots a_{nn}^{n-1},$$

and off-diagonal matrix elements

$$e_1, e_2, e_3, \ldots, e_{n-1}.$$

The resulting matrix reads

$$\mathbf{S}^{T}\mathbf{A}\mathbf{S} = \begin{pmatrix} a_{11} & e_{1} & 0 & 0 & \dots & 0 & 0 \\ e_{1} & a'_{22} & e_{2} & 0 & \dots & 0 & 0 \\ 0 & e_{2} & a''_{33} & e_{3} & 0 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & \dots & \dots & \dots & \dots & a_{n-2}^{(n-1)} & e_{n-1} \\ 0 & \dots & \dots & \dots & \dots & e_{n-1} & a_{nn}^{(n-1)} \end{pmatrix}.$$

It remains to find a recipe for determining the transformation  $S_n$ . We illustrate the method for  $S_1$  which we assume takes the form

$$\mathbf{S_1} = \left( \begin{array}{cc} 1 & \mathbf{0^T} \\ \mathbf{0} & \mathbf{P} \end{array} \right),$$

with  $\mathbf{0^T}$  being a zero row vector,  $\mathbf{0^T} = \{0, 0, \cdots\}$  of dimension (n-1). The matrix  $\mathbf{P}$  is symmetric with dimension  $((n-1) \times (n-1))$  satisfying  $\mathbf{P^2} = \mathbf{I}$  and  $\mathbf{P^T} = \mathbf{P}$ . A possible choice which fullfils the latter two requirements is

$$\mathbf{P} = \mathbf{I} - 2\mathbf{u}\mathbf{u}^T$$
.

where **I** is the (n-1) unity matrix and **u** is an n-1 column vector with norm  $\mathbf{u}^T\mathbf{u}$  (inner product).

Note that  $\mathbf{u}\mathbf{u}^T$  is an outer product giving a matrix of dimension  $((n-1)\times(n-1))$ . Each matrix element of **P** then reads

$$P_{ij} = \delta_{ij} - 2u_i u_j$$

where i and j range from 1 to n-1. Applying the transformation  $S_1$  results in

$$\mathbf{S}_1^T \mathbf{A} \mathbf{S}_1 = \left( \begin{array}{cc} a_{11} & (\mathbf{P} \mathbf{v})^T \\ \mathbf{P} \mathbf{v} & \mathbf{A}' \end{array} \right),$$

where  $\mathbf{v^T} = \{a_{21}, a_{31}, \cdots, a_{n1}\}$  and Ps must satisfy  $(\mathbf{Pv})^T = \{k, 0, 0, \cdots\}$ . Then

$$\mathbf{P}\mathbf{v} = \mathbf{v} - 2\mathbf{u}(\mathbf{u}^T \mathbf{v}) = k\mathbf{e},\tag{6}$$

with  $\mathbf{e^T} = \{1, 0, 0, \dots 0\}.$ 

Solving the latter equation gives us  $\mathbf{u}$  and thus the needed transformation  $\mathbf{P}$ . We do first however need to compute the scalar k by taking the scalar product of the last equation with its transpose and using the fact that  $\mathbf{P}^2 = \mathbf{I}$ . We get then

$$(\mathbf{P}\mathbf{v})^T \mathbf{P}\mathbf{v} = k^2 = \mathbf{v}^T \mathbf{v} = |v|^2 = \sum_{i=2}^n a_{i1}^2,$$

which determines the constant  $k = \pm v$ .

Now we can rewrite Eq. (6) as

$$\mathbf{v} - k\mathbf{e} = 2\mathbf{u}(\mathbf{u}^T\mathbf{v}),$$

and taking the scalar product of this equation with itself and obtain

$$2(\mathbf{u}^T \mathbf{v})^2 = (v^2 \pm a_{21}v),\tag{7}$$

which finally determines

$$\mathbf{u} = \frac{\mathbf{v} - k\mathbf{e}}{2(\mathbf{u}^T\mathbf{v})}.$$

In solving Eq. (7) great care has to be exercised so as to choose those values which make the right-hand largest in order to avoid loss of numerical precision. The above steps are then repeated for every transformations till we have a tridiagonal matrix suitable for obtaining the eigenvalues.

Our Householder transformation has given us a tridiagonal matrix. We discuss here how one can use Householder's iterative procedure to obtain the eigenvalues. Let us specialize to a  $4\times 4$  matrix. The tridiagonal matrix takes the form

$$\mathbf{A} = \begin{pmatrix} d_1 & e_1 & 0 & 0 \\ e_1 & d_2 & e_2 & 0 \\ 0 & e_2 & d_3 & e_3 \\ 0 & 0 & e_3 & d_4 \end{pmatrix}.$$

As a first observation, if any of the elements  $e_i$  are zero the matrix can be separated into smaller pieces before diagonalization. Specifically, if  $e_1 = 0$  then  $d_1$  is an eigenvalue.

Thus, let us introduce a transformation  $S_1$  which operates like

$$\mathbf{S_1} = \begin{pmatrix} \cos \theta & 0 & 0 & \sin \theta \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ \cos \theta & 0 & 0 & \cos \theta \end{pmatrix}$$

Then the similarity transformation

$$\mathbf{S_1^T A S_1} = \mathbf{A'} = \left( egin{array}{cccc} d'_1 & e'_1 & 0 & 0 \ e'_1 & d_2 & e_2 & 0 \ 0 & e_2 & d_3 & e'3 \ 0 & 0 & e'_3 & d'_4 \end{array} 
ight)$$

produces a matrix where the primed elements in  $\mathbf{A}'$  have been changed by the transformation whereas the unprimed elements are unchanged. If we now choose  $\theta$  to give the element  $a_{21}' = e' = 0$  then we have the first eigenvalue  $= a_{11}' = d_1'$ . (This is actually what you are doing in project 2!!)

This procedure can be continued on the remaining three-dimensional submatrix for the next eigenvalue. Thus after few transformations we have the wanted diagonal form.

What we see here is just a special case of the more general procedure developed by Francis in two articles in 1961 and 1962.

The algorithm is based on the so-called QR method (or just QR-algorithm). It follows from a theorem by Schur which states that any square matrix can be

written out in terms of an orthogonal matrix  $\mathbf{Q}$  and an upper triangular matrix  $\mathbf{U}$ . Historically R was used instead of U since the wording right triangular matrix was first used. The method is based on an iterative procedure similar to Jacobi's method, by a succession of planar rotations. For a tridiagonal matrix it is simple to carry out in principle, but complicated in detail! We will discuss this in more detail during week 38.

#### **Eigenvalues**

The eigenvalues of a tridiagonal matrix can be obtained using the characteristic polynomial

$$P(\lambda) = det(\lambda \mathbf{I} - \mathbf{A}) = \prod_{i=1}^{n} (\lambda_i - \lambda),$$

which rewritten in matrix form reads

$$P(\lambda) = \begin{pmatrix} d_1 - \lambda & e_1 & 0 & 0 & \dots & 0 & 0 \\ e_1 & d_2 - \lambda & e_2 & 0 & \dots & 0 & 0 \\ 0 & e_2 & d_3 - \lambda & e_3 & 0 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & \dots & \dots & \dots & \dots & d_{N_{\text{step}}-2} - \lambda & e_{N_{\text{step}}-1} \\ 0 & \dots & \dots & \dots & \dots & e_{N_{\text{step}}-1} & d_{N_{\text{step}}-1} - \lambda \end{pmatrix}$$

We can solve this equation in an iterative manner. We let  $P_k(\lambda)$  be the value of k subdeterminant of the above matrix of dimension  $n \times n$ . The polynomial  $P_k(\lambda)$  is clearly a polynomial of degree k. Starting with  $P_1(\lambda)$  we have  $P_1(\lambda) = d_1 - \lambda$ . The next polynomial reads  $P_2(\lambda) = (d_2 - \lambda)P_1(\lambda) - e_1^2$ . By expanding the determinant for  $P_k(\lambda)$  in terms of the minors of the nth column we arrive at the recursion relation

$$P_k(\lambda) = (d_k - \lambda)P_{k-1}(\lambda) - e_{k-1}^2 P_{k-2}(\lambda).$$

Together with the starting values  $P_1(\lambda)$  and  $P_2(\lambda)$  and good root searching methods we arrive at an efficient computational scheme for finding the roots of  $P_n(\lambda)$ . However, for large matrices this algorithm is rather inefficient and time-consuming.

#### Eigenvalues and Lanczos' method

Basic features with a real symmetric matrix (and normally huge  $n>10^6$  and sparse)  $\hat{A}$  of dimension  $n\times n$ :

- Lanczos' algorithm generates a sequence of real tridiagonal matrices  $T_k$  of dimension  $k \times k$  with  $k \le n$ , with the property that the extremal eigenvalues of  $T_k$  are progressively better estimates of  $\hat{A}$ ' extremal eigenvalues.\* The method converges to the extremal eigenvalues.
- The similarity transformation is

$$\hat{T} = \hat{Q}^T \hat{A} \hat{Q},$$

with the first vector  $\hat{Q}\hat{e}_1 = \hat{q}_1$ .

We are going to solve iteratively

$$\hat{T} = \hat{Q}^T \hat{A} \hat{Q},$$

with the first vector  $\hat{Q}\hat{e}_1 = \hat{q}_1$ . We can write out the matrix  $\hat{Q}$  in terms of its column vectors

$$\hat{Q} = [\hat{q}_1 \hat{q}_2 \dots \hat{q}_n] .$$

The matrix

$$\hat{T} = \hat{Q}^T \hat{A} \hat{Q},$$

can be written as

$$\hat{T} = \begin{pmatrix} \alpha_1 & \beta_1 & 0 & \dots & \dots & 0 \\ \beta_1 & \alpha_2 & \beta_2 & 0 & \dots & 0 \\ 0 & \beta_2 & \alpha_3 & \beta_3 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & 0 \\ 0 & \dots & \dots & 0 & \beta_{n-1} & \beta_{n-1} \\ 0 & \dots & \dots & 0 & \beta_{n-1} & \alpha_n \end{pmatrix}$$

Using the fact that

$$\hat{Q}\hat{Q}^T = \hat{I},$$

we can rewrite

$$\hat{T} = \hat{Q}^T \hat{A} \hat{Q},$$

as

$$\hat{Q}\hat{T} = \hat{A}\hat{Q}$$
.

If we equate columns

$$\hat{T} = \begin{pmatrix} \alpha_1 & \beta_1 & 0 & \dots & \dots & 0 \\ \beta_1 & \alpha_2 & \beta_2 & 0 & \dots & 0 \\ 0 & \beta_2 & \alpha_3 & \beta_3 & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & 0 \\ 0 & \dots & \dots & 0 & \beta_{n-1} & \beta_{n-1} \\ 0 & \dots & \dots & 0 & \beta_{n-1} & \alpha_n \end{pmatrix}$$

we obtain

$$\hat{A}\hat{q}_{k} = \beta_{k-1}\hat{q}_{k-1} + \alpha_{k}\hat{q}_{k} + \beta_{k}\hat{q}_{k+1}.$$

We have thus

$$\hat{A}\hat{q}_k = \beta_{k-1}\hat{q}_{k-1} + \alpha_k\hat{q}_k + \beta_k\hat{q}_{k+1},$$

with  $\beta_0 \hat{q}_0 = 0$  for k = 1 : n - 1. Remember that the vectors  $\hat{q}_k$  are orthornormal and this implies

$$\alpha_k = \hat{q}_k^T \hat{A} \hat{q}_k,$$

and these vectors are called Lanczos vectors.

We have thus

$$\hat{A}\hat{q}_k = \beta_{k-1}\hat{q}_{k-1} + \alpha_k\hat{q}_k + \beta_k\hat{q}_{k+1},$$

with  $\beta_0 \hat{q}_0 = 0$  for k = 1: n-1 and

$$\alpha_k = \hat{q}_k^T \hat{A} \hat{q}_k.$$

If

$$\hat{r}_k = (\hat{A} - \alpha_k \hat{I})\hat{q}_k - \beta_{k-1}\hat{q}_{k-1},$$

is non-zero, then

$$\hat{q}_{k+1} = \hat{r}_k / \beta_k,$$

with  $\beta_k = \pm ||\hat{r}_k||_2$ .

# Variational Monte Carlo methods

# Quantum Monte Carlo Motivation

We start with the variational principle. Given a hamiltonian H and a trial wave function  $\Psi_T$ , the variational principle states that the expectation value of  $\langle H \rangle$ , defined through

$$E[H] = \langle H \rangle = \frac{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) H(\mathbf{R}) \Psi_T(\mathbf{R})}{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) \Psi_T(\mathbf{R})},$$

is an upper bound to the ground state energy  $E_0$  of the hamiltonian H, that is

$$E_0 \leq \langle H \rangle$$
.

In general, the integrals involved in the calculation of various expectation values are multi-dimensional ones. Traditional integration methods such as the Gauss-Legendre will not be adequate for say the computation of the energy of a many-body system.

The trial wave function can be expanded in the eigenstates of the hamiltonian since they form a complete set, viz.,

$$\Psi_T(\mathbf{R}) = \sum_i a_i \Psi_i(\mathbf{R}),$$

and assuming the set of eigenfunctions to be normalized one obtains

$$\frac{\sum_{nm} a_m^* a_n \int d\mathbf{R} \Psi_m^*(\mathbf{R}) H(\mathbf{R}) \Psi_n(\mathbf{R})}{\sum_{nm} a_m^* a_n \int d\mathbf{R} \Psi_m^*(\mathbf{R}) \Psi_n(\mathbf{R})} = \frac{\sum_n a_n^2 E_n}{\sum_n a_n^2} \ge E_0,$$

where we used that  $H(\mathbf{R})\Psi_n(\mathbf{R}) = E_n\Psi_n(\mathbf{R})$ . In general, the integrals involved in the calculation of various expectation values are multi-dimensional ones. The variational principle yields the lowest state of a given symmetry.

In most cases, a wave function has only small values in large parts of configuration space, and a straightforward procedure which uses homogenously

distributed random points in configuration space will most likely lead to poor results. This may suggest that some kind of importance sampling combined with e.g., the Metropolis algorithm may be a more efficient way of obtaining the ground state energy. The hope is then that those regions of configurations space where the wave function assumes appreciable values are sampled more efficiently.

The tedious part in a VMC calculation is the search for the variational minimum. A good knowledge of the system is required in order to carry out reasonable VMC calculations. This is not always the case, and often VMC calculations serve rather as the starting point for so-called diffusion Monte Carlo calculations (DMC). DMC is a way of solving exactly the many-body Schroedinger equation by means of a stochastic procedure. A good guess on the binding energy and its wave function is however necessary. A carefully performed VMC calculation can aid in this context.

The basic recipe in a VMC calculation consists of the following elements:

- Construct first a trial wave function  $\psi_T(\mathbf{R}, \boldsymbol{\alpha})$ , for a many-body system consisting of N particles located at positions  $\mathbf{R} = (\mathbf{R}_1, \dots, \mathbf{R}_N)$ . The trial wave function depends on  $\alpha$  variational parameters  $\boldsymbol{\alpha} = (\alpha_1, \dots, \alpha_M)$ .
- ullet Then we evaluate the expectation value of the hamiltonian H

$$E[H] = \langle H \rangle = \frac{\int d\mathbf{R} \Psi_T^*(\mathbf{R}, \boldsymbol{\alpha}) H(\mathbf{R}) \Psi_T(\mathbf{R}, \boldsymbol{\alpha})}{\int d\mathbf{R} \Psi_T^*(\mathbf{R}, \boldsymbol{\alpha}) \Psi_T(\mathbf{R}, \boldsymbol{\alpha})}.$$

• Thereafter we vary  $\alpha$  according to some minimization algorithm and return to the first step.

With a trial wave function  $\psi_T(\mathbf{R})$  we can in turn construct the quantum mechanical probability distribution

$$P(\mathbf{R}) = \frac{\left|\psi_T(\mathbf{R})\right|^2}{\int \left|\psi_T(\mathbf{R})\right|^2 d\mathbf{R}}.$$

This is our new probability distribution function (PDF). The approximation to the expectation value of the Hamiltonian is now

$$E[H(\boldsymbol{\alpha})] = \frac{\int d\boldsymbol{R} \Psi_T^*(\boldsymbol{R}, \boldsymbol{\alpha}) H(\boldsymbol{R}) \Psi_T(\boldsymbol{R}, \boldsymbol{\alpha})}{\int d\boldsymbol{R} \Psi_T^*(\boldsymbol{R}, \boldsymbol{\alpha}) \Psi_T(\boldsymbol{R}, \boldsymbol{\alpha})}.$$

Define a new quantity

$$E_L(\boldsymbol{R}, \boldsymbol{\alpha}) = \frac{1}{\psi_T(\boldsymbol{R}, \boldsymbol{\alpha})} H \psi_T(\boldsymbol{R}, \boldsymbol{\alpha}),$$

called the local energy, which, together with our trial PDF yields

$$E[H(\boldsymbol{\alpha})] = \int P(\boldsymbol{R}) E_L(\boldsymbol{R}) d\boldsymbol{R} \approx \frac{1}{N} \sum_{i=1}^{N} P(\boldsymbol{R_i}, \boldsymbol{\alpha}) E_L(\boldsymbol{R_i}, \boldsymbol{\alpha})$$

with N being the number of Monte Carlo samples.

The Algorithm for performing a variational Monte Carlo calculations runs thus as this

- Initialisation: Fix the number of Monte Carlo steps. Choose an initial R and variational parameters  $\alpha$  and calculate  $|\psi_T^{\alpha}(R)|^2$ .
- Initialise the energy and the variance and start the Monte Carlo calculation.
  - Calculate a trial position  $\mathbf{R}_p = \mathbf{R} + r * step$  where r is a random variable  $r \in [0, 1]$ .
  - Metropolis algorithm to accept or reject this move  $w = P(\mathbf{R}_p)/P(\mathbf{R})$ .
  - If the step is accepted, then we set  $\mathbf{R} = \mathbf{R}_p$ .
  - Update averages
- Finish and compute final averages.

Observe that the jumping in space is governed by the variable *step*. This is Called brute-force sampling. Need importance sampling to get more relevant sampling, see lectures below.

**Quantum Monte Carlo: hydrogen atom.** The radial Schroedinger equation for the hydrogen atom can be written as

$$-\frac{\hbar^2}{2m}\frac{\partial^2 u(r)}{\partial r^2} - \left(\frac{ke^2}{r} - \frac{\hbar^2 l(l+1)}{2mr^2}\right)u(r) = Eu(r),$$

or with dimensionless variables

$$-\frac{1}{2}\frac{\partial^2 u(\rho)}{\partial \rho^2} - \frac{u(\rho)}{\rho} + \frac{l(l+1)}{2\rho^2}u(\rho) - \lambda u(\rho) = 0,$$

with the hamiltonian

$$H = -\frac{1}{2} \frac{\partial^2}{\partial \rho^2} - \frac{1}{\rho} + \frac{l(l+1)}{2\rho^2}.$$

Use variational parameter  $\alpha$  in the trial wave function

$$u_T^{\alpha}(\rho) = \alpha \rho e^{-\alpha \rho}.$$

Inserting this wave function into the expression for the local energy  ${\cal E}_L$  gives

$$E_L(\rho) = -\frac{1}{\rho} - \frac{\alpha}{2} \left( \alpha - \frac{2}{\rho} \right).$$

A simple variational Monte Carlo calculation results in

$\alpha$	$\langle H \rangle$	$\sigma^2$	$\sigma/\sqrt{N}$
7.00000E-01	-4.57759E-01	4.51201E-02	6.71715E-04
8.00000E-01	-4.81461E-01	3.05736E-02	5.52934E- $04$
9.00000E- $01$	-4.95899E-01	8.20497E-03	2.86443E-04
1.00000E-00	-5.00000E $-01$	0.00000E+00	0.00000E+00
1.10000E+00	-4.93738E $-01$	1.16989E-02	3.42036E-04
1.20000E+00	-4.75563E- $01$	8.85899E-02	9.41222E-04
1.30000E+00	-4.54341E- $01$	1.45171E-01	1.20487E-03

We note that at  $\alpha=1$  we obtain the exact result, and the variance is zero, as it should. The reason is that we then have the exact wave function, and the action of the hamiltionan on the wave function

$$H\psi = \text{constant} \times \psi$$
,

yields just a constant. The integral which defines various expectation values involving moments of the hamiltonian becomes then

$$\langle H^n \rangle = \frac{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) H^n(\mathbf{R}) \Psi_T(\mathbf{R})}{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) \Psi_T(\mathbf{R})} = \text{constant} \times \frac{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) \Psi_T(\mathbf{R})}{\int d\mathbf{R} \Psi_T^*(\mathbf{R}) \Psi_T(\mathbf{R})} = \text{constant}.$$

This gives an important information: the exact wave function leads to zero variance! Variation is then performed by minimizing both the energy and the variance.

For bosons in a harmonic oscillator-like trap we will use is a spherical (S) or an elliptical (E) harmonic trap in one, two and finally three dimensions, with the latter given by

$$V_{ext}(\mathbf{r}) = \begin{cases} \frac{1}{2} m \omega_{ho}^2 r^2 & (S) \\ \frac{1}{2} m [\omega_{ho}^2 (x^2 + y^2) + \omega_z^2 z^2] & (E) \end{cases}$$
 (8)

where (S) stands for symmetric and

$$\hat{H} = \sum_{i}^{N} \left( \frac{-\hbar^2}{2m} \nabla_i^2 + V_{ext}(\mathbf{r}_i) \right) + \sum_{i < j}^{N} V_{int}(\mathbf{r}_i, \mathbf{r}_j), \tag{9}$$

as the two-body Hamiltonian of the system.

We will represent the inter-boson interaction by a pairwise, repulsive potential

$$V_{int}(|\mathbf{r}_i - \mathbf{r}_j|) = \begin{cases} \infty & |\mathbf{r}_i - \mathbf{r}_j| \le a \\ 0 & |\mathbf{r}_i - \mathbf{r}_j| > a \end{cases}$$
(10)

where a is the so-called hard-core diameter of the bosons. Clearly,  $V_{int}(|\mathbf{r}_i - \mathbf{r}_j|)$  is zero if the bosons are separated by a distance  $|\mathbf{r}_i - \mathbf{r}_j|$  greater than a but infinite if they attempt to come within a distance  $|\mathbf{r}_i - \mathbf{r}_j| \leq a$ .

Our trial wave function for the ground state with N atoms is given by

$$\Psi_T(\mathbf{R}) = \Psi_T(\mathbf{r}_1, \mathbf{r}_2, \dots \mathbf{r}_N, \alpha, \beta) = \prod_i g(\alpha, \beta, \mathbf{r}_i) \prod_{i < j} f(a, |\mathbf{r}_i - \mathbf{r}_j|), \quad (11)$$

where  $\alpha$  and  $\beta$  are variational parameters. The single-particle wave function is proportional to the harmonic oscillator function for the ground state

$$g(\alpha, \beta, \mathbf{r}_i) = \exp\left[-\alpha(x_i^2 + y_i^2 + \beta z_i^2)\right]. \tag{12}$$

For spherical traps we have  $\beta = 1$  and for non-interacting bosons (a = 0) we have  $\alpha = 1/2a_{ho}^2$ . The correlation wave function is

$$f(a, |\mathbf{r}_i - \mathbf{r}_j|) = \begin{cases} 0 & |\mathbf{r}_i - \mathbf{r}_j| \le a \\ (1 - \frac{a}{|\mathbf{r}_i - \mathbf{r}_j|}) & |\mathbf{r}_i - \mathbf{r}_j| > a. \end{cases}$$
(13)

A simple Python code that solves the two-boson or two-fermion case in two-dimensions.

```
# Importing various packages
from math import exp, sqrt
from random import random, seed
import numpy as np
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D
from matplotlib import cm
from matplotlib.ticker import LinearLocator, FormatStrFormatter
import sys
#Trial wave function for quantum dots in two dims
def WaveFunction(r,alpha,beta):
   r1 = r[0,0]**2 + r[0,1]**2
   r2 = r[1,0]**2 + r[1,1]**2
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
   deno = r12/(1+beta*r12)
   return exp(-0.5*alpha*(r1+r2)+deno)
#Local energy for quantum dots in two dims, using analytical local energy
def LocalEnergy(r,alpha,beta):
   r1 = (r[0,0]**2 + r[0,1]**2)
   r2 = (r[1,0]**2 + r[1,1]**2)
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
   deno2 = deno*deno
   return 0.5*(1-alpha*alpha)*(r1 + r2) +2.0*alpha + 1.0/r12+deno2*(alpha*r12-deno2+2*beta*deno-
# The Monte Carlo sampling with the Metropolis algo
def MonteCarloSampling():
   NumberMCcycles= 100000
   StepSize = 1.0
    # positions
   PositionOld = np.zeros((NumberParticles, Dimension), np.double)
   PositionNew = np.zeros((NumberParticles, Dimension), np.double)
    # seed for rng generator
   seed()
    # start variational parameter
   alpha = 0.9
   for ia in range(MaxVariations):
    alpha += .025
        AlphaValues[ia] = alpha
        beta = 0.2
        for jb in range(MaxVariations):
            beta += .01
            BetaValues[jb] = beta
            energy = energy2 = 0.0
DeltaE = 0.0
            \#Initial\ position
            for i in range(NumberParticles):
                for j in range(Dimension):
                    PositionOld[i,j] = StepSize * (random() - .5)
            wfold = WaveFunction(PositionOld,alpha,beta)
            #Loop over MC MCcycles
            for MCcycle in range(NumberMCcycles):
                #Trial position
```

```
for i in range(NumberParticles):
                      for j in range(Dimension):
                          PositionNew[i,j] = PositionOld[i,j] + StepSize * (random() - .5)
                 wfnew = WaveFunction(PositionNew,alpha,beta)
                  #Metropolis test to see whether we accept the move
                 if random() < wfnew**2 / wfold**2:</pre>
                     PositionOld = PositionNew.copy()
                     wfold = wfnew
                     DeltaE = LocalEnergy(PositionOld,alpha,beta)
                 energy += DeltaE
energy2 += DeltaE**2
             \# We \ calculate \ mean, \ variance \ and \ error \ \dots
             energy /= NumberMCcycles
             energy2 /= NumberMCcycles
             variance = energy2 - energy**2
             error = sqrt(variance/NumberMCcycles)
             Energies[ia,jb] = energy
    return Energies, AlphaValues, BetaValues
#Here starts the main program with variable declarations
NumberParticles = 2
Dimension = 2
MaxVariations = 10
Energies = np.zeros((MaxVariations, MaxVariations))
AlphaValues = np.zeros(MaxVariations)
BetaValues = np.zeros(MaxVariations)
(Energies, AlphaValues, BetaValues) = MonteCarloSampling()
# Prepare for plots
fig = plt.figure()
ax = fig.gca(projection='3d')
# Plot the surface.
X, Y = np.meshgrid(AlphaValues, BetaValues)
surf = ax.plot_surface(X, Y, Energies,cmap=cm.coolwarm,linewidth=0, antialiased=False)
# Customize the z axis.
zmin = np.matrix(Energies).min()
zmax = np.matrix(Energies).max()
ax.set_zlim(zmin, zmax)
ax.set_xlabel(r'$\alpha$')
ax.set_ylabel(r'$\beta$')
ax.set_zlabel(r'$\langle E \rangle$')
ax.zaxis.set_major_locator(LinearLocator(10))
ax.zaxis.set_major_formatter(FormatStrFormatter('%.02f'))
# Add a color bar which maps values to colors. fig.colorbar(surf, shrink=0.5, aspect=5)
plt.show()
```

#### Quantum Monte Carlo: the helium atom

The helium atom consists of two electrons and a nucleus with charge Z=2. The contribution to the potential energy due to the attraction from the nucleus is

$$-\frac{2ke^2}{r_1} - \frac{2ke^2}{r_2},$$

and if we add the repulsion arising from the two interacting electrons, we obtain the potential energy

$$V(r_1, r_2) = -\frac{2ke^2}{r_1} - \frac{2ke^2}{r_2} + \frac{ke^2}{r_{12}},$$

with the electrons separated at a distance  $r_{12} = |\mathbf{r}_1 - \mathbf{r}_2|$ .

The hamiltonian becomes then

$$\hat{H} = -\frac{\hbar^2 \nabla_1^2}{2m} - \frac{\hbar^2 \nabla_2^2}{2m} - \frac{2ke^2}{r_1} - \frac{2ke^2}{r_2} + \frac{ke^2}{r_{12}},$$

and Schroedingers equation reads

$$\hat{H}\psi = E\psi.$$

All observables are evaluated with respect to the probability distribution

$$P(\mathbf{R}) = \frac{\left|\psi_T(\mathbf{R})\right|^2}{\int \left|\psi_T(\mathbf{R})\right|^2 d\mathbf{R}}.$$

generated by the trial wave function. The trial wave function must approximate an exact eigenstate in order that accurate results are to be obtained.

Choice of trial wave function for Helium: Assume  $r_1 \to 0$ .

$$E_L(\boldsymbol{R}) = \frac{1}{\psi_T(\boldsymbol{R})} H \psi_T(\boldsymbol{R}) = \frac{1}{\psi_T(\boldsymbol{R})} \left( -\frac{1}{2} \nabla_1^2 - \frac{Z}{r_1} \right) \psi_T(\boldsymbol{R}) + \text{finite terms}.$$

$$E_L(R) = \frac{1}{\mathbf{R}_T(r_1)} \left( -\frac{1}{2} \frac{d^2}{dr_1^2} - \frac{1}{r_1} \frac{d}{dr_1} - \frac{Z}{r_1} \right) \mathbf{R}_T(r_1) + \text{finite terms}$$

For small values of  $r_1$ , the terms which dominate are

$$\lim_{r_1 \to 0} E_L(R) = \frac{1}{\mathbf{R}_T(r_1)} \left( -\frac{1}{r_1} \frac{d}{dr_1} - \frac{Z}{r_1} \right) \mathbf{R}_T(r_1),$$

since the second derivative does not diverge due to the finiteness of  $\Psi$  at the origin.

This results in

$$\frac{1}{\mathbf{R}_T(r_1)} \frac{d\mathbf{R}_T(r_1)}{dr_1} = -Z,$$

and

$$\mathbf{R}_T(r_1) \propto e^{-Zr_1}$$
.

A similar condition applies to electron 2 as well. For orbital momenta l>0 we have

$$\frac{1}{\mathbf{R}_T(r)} \frac{d\mathbf{R}_T(r)}{dr} = -\frac{Z}{l+1}.$$

Similarly, studying the case  $r_{12} \to 0$  we can write a possible trial wave function as

$$\psi_T(\mathbf{R}) = e^{-\alpha(r_1 + r_2)} e^{\beta r_{12}}.$$

The last equation can be generalized to

$$\psi_T(\mathbf{R}) = \phi(\mathbf{r}_1)\phi(\mathbf{r}_2)\dots\phi(\mathbf{r}_N)\prod_{i< j}f(r_{ij}),$$

for a system with N electrons or particles.

During the development of our code we need to make several checks. It is also very instructive to compute a closed form expression for the local energy. Since our wave function is rather simple it is straightforward to find an analytic expressions. Consider first the case of the simple helium function

$$\Psi_T(\boldsymbol{r}_1, \boldsymbol{r}_2) = e^{-\alpha(r_1 + r_2)}$$

The local energy is for this case

$$E_{L1} = (\alpha - Z) \left( \frac{1}{r_1} + \frac{1}{r_2} \right) + \frac{1}{r_{12}} - \alpha^2$$

which gives an expectation value for the local energy given by

$$\langle E_{L1} \rangle = \alpha^2 - 2\alpha \left( Z - \frac{5}{16} \right)$$

With closed form formulae we can speed up the computation of the correlation. In our case we write it as

$$\Psi_C = \exp\left\{\sum_{i < j} \frac{ar_{ij}}{1 + \beta r_{ij}}\right\},\,$$

which means that the gradient needed for the so-called quantum force and local energy can be calculated analytically. This will speed up your code since the computation of the correlation part and the Slater determinant are the most time consuming parts in your code.

We will refer to this correlation function as  $\Psi_C$  or the *linear Pade-Jastrow*. We can test this by computing the local energy for our helium wave function

$$\psi_T(\mathbf{r}_1, \mathbf{r}_2) = \exp(-\alpha(r_1 + r_2)) \exp\left(\frac{r_{12}}{2(1 + \beta r_{12})}\right),$$

with  $\alpha$  and  $\beta$  as variational parameters.

The local energy is for this case

$$E_{L2} = E_{L1} + \frac{1}{2(1+\beta r_{12})^2} \left\{ \frac{\alpha(r_1 + r_2)}{r_{12}} \left(1 - \frac{r_1 r_2}{r_1 r_2}\right) - \frac{1}{2(1+\beta r_{12})^2} - \frac{2}{r_{12}} + \frac{2\beta}{1+\beta r_{12}} \right\}$$

It is very useful to test your code against these expressions. It means also that you don't need to compute a derivative numerically as discussed in the code example below.

For the computation of various derivatives with different types of wave functions, you will find it useful to use python with symbolic python, that is sympy, see online manual. Using sympy allows you autogenerate both Latex code as well c++, python or Fortran codes. Here you will find some simple examples. We choose the 2s hydrogen-orbital (not normalized) as an example

$$\phi_{2s}(\boldsymbol{r}) = (Zr-2)\exp{-(\frac{1}{2}Zr)},$$
 with  $r^2 = x^2 + y^2 + z^2$ . from sympy import symbols, diff, exp, sqrt x, y, z, Z = symbols('x y z Z') r = sqrt(x\*x + y\*y + z\*z) r phi = (Z\*r - 2)\*exp(-Z\*r/2) phi diff(phi, x)

This doesn't look very nice, but sympy provides several functions that allow for improving and simplifying the output.

We can improve our output by factorizing and substituting expressions

```
from sympy import symbols, diff, exp, sqrt, factor, Symbol, printing
x, y, z, Z = symbols('x y z Z')
r = sqrt(x*x + y*y + z*z)
phi = (Z*r - 2)*exp(-Z*r/2)
R = Symbol('r') #Creates a symbolic equivalent of r
#print latex and c++ code
print printing.latex(diff(phi, x).factor().subs(r, R))
print printing.ccode(diff(phi, x).factor().subs(r, R))
```

We can in turn look at second derivatives

```
from sympy import symbols, diff, exp, sqrt, factor, Symbol, printing
x, y, z, Z = symbols('x y z Z')
r = sqrt(x*x + y*y + z*z)
phi = (Z*r - 2)*exp(-Z*r/2)
R = Symbol('r') #Creates a symbolic equivalent of r
(diff(diff(phi, x), x) + diff(diff(phi, y), y) + diff(diff(phi, z), z)).factor().subs(r, R)
# Collect the Z values
(diff(diff(phi, x), x) + diff(diff(phi, y), y) + diff(diff(phi, z), z)).factor().collect(Z).subs(r, R)
# Factorize also the r**2 terms
(diff(diff(phi, x), x) + diff(diff(phi, y), y) + diff(diff(phi, z), z)).factor().collect(Z).subs(r, R)
print printing.ccode((diff(diff(phi, x), x) + diff(diff(phi, y), y) + diff(diff(phi, z), z)).factor().collect(Z).subs(r, R)
```

With some practice this allows one to be able to check one's own calculation and translate automatically into code lines.

# The Metropolis algorithm

The Metropolis algorithm , see the original article was invented by Metropolis et. al and is often simply called the Metropolis algorithm. It is a method to sample

a normalized probability distribution by a stochastic process. We define  $\mathbf{P}_i^{(n)}$  to be the probability for finding the system in the state i at step n. The algorithm is then

- Sample a possible new state j with some probability  $T_{i\to j}$ .
- Accept the new state j with probability  $A_{i\to j}$  and use it as the next sample. With probability  $1-A_{i\to j}$  the move is rejected and the original state i is used again as a sample.

We wish to derive the required properties of T and A such that  $\mathbf{P}_i^{(n\to\infty)} \to p_i$  so that starting from any distribution, the method converges to the correct distribution. Note that the description here is for a discrete probability distribution. Replacing probabilities  $p_i$  with expressions like  $p(x_i)dx_i$  will take all of these over to the corresponding continuum expressions.

The dynamical equation for  $\mathbf{P}_i^{(n)}$  can be written directly from the description above. The probability of being in the state i at step n is given by the probability of being in any state j at the previous step, and making an accepted transition to i added to the probability of being in the state i, making a transition to any state j and rejecting the move:

$$\mathbf{P}_{i}^{(n)} = \sum_{j} \left[ \mathbf{P}_{j}^{(n-1)} T_{j \to i} A_{j \to i} + \mathbf{P}_{i}^{(n-1)} T_{i \to j} \left( 1 - A_{i \to j} \right) \right].$$

Since the probability of making some transition must be 1,  $\sum_{j} T_{i \to j} = 1$ , and the above equation becomes

$$\mathbf{P}_{i}^{(n)} = \mathbf{P}_{i}^{(n-1)} + \sum_{j} \left[ \mathbf{P}_{j}^{(n-1)} T_{j \to i} A_{j \to i} - \mathbf{P}_{i}^{(n-1)} T_{i \to j} A_{i \to j} \right].$$

For large n we require that  $\mathbf{P}_i^{(n\to\infty)}=p_i$ , the desired probability distribution. Taking this limit, gives the balance requirement

$$\sum_{j} \left[ p_j T_{j \to i} A_{j \to i} - p_i T_{i \to j} A_{i \to j} \right] = 0.$$

The balance requirement is very weak. Typically the much stronger detailed balance requirement is enforced, that is rather than the sum being set to zero, we set each term separately to zero and use this to determine the acceptance probabilities. Rearranging, the result is

$$\frac{A_{j\to i}}{A_{i\to j}} = \frac{p_i T_{i\to j}}{p_i T_{j\to i}} \,.$$

The Metropolis choice is to maximize the A values, that is

$$A_{j \to i} = \min\left(1, \frac{p_i T_{i \to j}}{p_j T_{j \to i}}\right).$$

Other choices are possible, but they all correspond to multiplying  $A_{i\to j}$  and  $A_{j\to i}$  by the same constant smaller than unity.<sup>1</sup>

Having chosen the acceptance probabilities, we have guaranteed that if the  $\mathbf{P}_{i}^{(n)}$  has equilibrated, that is if it is equal to  $p_{i}$ , it will remain equilibrated. Next we need to find the circumstances for convergence to equilibrium.

The dynamical equation can be written as

$$\mathbf{P}_i^{(n)} = \sum_j M_{ij} \mathbf{P}_j^{(n-1)}$$

with the matrix M given by

$$M_{ij} = \delta_{ij} \left[ 1 - \sum_{k} T_{i \to k} A_{i \to k} \right] + T_{j \to i} A_{j \to i}.$$

Summing over i shows that  $\sum_i M_{ij} = 1$ , and since  $\sum_k T_{i \to k} = 1$ , and  $A_{i \to k} \leq 1$ , the elements of the matrix satisfy  $M_{ij} \geq 0$ . The matrix M is therefore a stochastic matrix.

The Metropolis method is simply the power method for computing the right eigenvector of M with the largest magnitude eigenvalue. By construction, the correct probability distribution is a right eigenvector with eigenvalue 1. Therefore, for the Metropolis method to converge to this result, we must show that M has only one eigenvalue with this magnitude, and all other eigenvalues are smaller.

#### Importance sampling

We need to replace the brute force Metropolis algorithm with a walk in coordinate space biased by the trial wave function. This approach is based on the Fokker-Planck equation and the Langevin equation for generating a trajectory in coordinate space. The link between the Fokker-Planck equation and the Langevin equations are explained, only partly, in the slides below. An excellent reference on topics like Brownian motion, Markov chains, the Fokker-Planck equation and the Langevin equation is the text by Van Kampen Here we will focus first on the implementation part first.

For a diffusion process characterized by a time-dependent probability density P(x,t) in one dimension the Fokker-Planck equation reads (for one particle /walker)

$$\frac{\partial P}{\partial t} = D \frac{\partial}{\partial x} \left( \frac{\partial}{\partial x} - F \right) P(x, t),$$

where F is a drift term and D is the diffusion coefficient.

The new positions in coordinate space are given as the solutions of the Langevin equation using Euler's method, namely, we go from the Langevin equation

$$\frac{\partial x(t)}{\partial t} = DF(x(t)) + \eta,$$

 $<sup>^{1}</sup>$ The penalty function method uses just such a factor to compensate for  $p_{i}$  that are evaluated stochastically and are therefore noisy.

with  $\eta$  a random variable, yielding a new position

$$y = x + DF(x)\Delta t + \xi \sqrt{\Delta t}$$

where  $\xi$  is gaussian random variable and  $\Delta t$  is a chosen time step. The quantity D is, in atomic units, equal to 1/2 and comes from the factor 1/2 in the kinetic energy operator. Note that  $\Delta t$  is to be viewed as a parameter. Values of  $\Delta t \in [0.001, 0.01]$  yield in general rather stable values of the ground state energy.

The process of isotropic diffusion characterized by a time-dependent probability density  $P(\mathbf{x},t)$  obeys (as an approximation) the so-called Fokker-Planck equation

$$\frac{\partial P}{\partial t} = \sum_{i} D \frac{\partial}{\partial \mathbf{x_i}} \left( \frac{\partial}{\partial \mathbf{x_i}} - \mathbf{F_i} \right) P(\mathbf{x}, t),$$

where  $\mathbf{F_i}$  is the  $i^{th}$  component of the drift term (drift velocity) caused by an external potential, and D is the diffusion coefficient. The convergence to a stationary probability density can be obtained by setting the left hand side to zero. The resulting equation will be satisfied if and only if all the terms of the sum are equal zero,

$$\frac{\partial^2 P}{\partial \mathbf{x_i}^2} = P \frac{\partial}{\partial \mathbf{x_i}} \mathbf{F_i} + \mathbf{F_i} \frac{\partial}{\partial \mathbf{x_i}} P.$$

The drift vector should be of the form  $\mathbf{F} = g(\mathbf{x}) \frac{\partial P}{\partial \mathbf{x}}$ . Then,

$$\frac{\partial^2 P}{\partial \mathbf{x_i}^2} = P \frac{\partial g}{\partial P} \left( \frac{\partial P}{\partial \mathbf{x}_i} \right)^2 + P g \frac{\partial^2 P}{\partial \mathbf{x}_i^2} + g \left( \frac{\partial P}{\partial \mathbf{x}_i} \right)^2.$$

The condition of stationary density means that the left hand side equals zero. In other words, the terms containing first and second derivatives have to cancel each other. It is possible only if  $g = \frac{1}{P}$ , which yields

$$\mathbf{F} = 2\frac{1}{\Psi_T} \nabla \Psi_T,$$

which is known as the so-called *quantum force*. This term is responsible for pushing the walker towards regions of configuration space where the trial wave function is large, increasing the efficiency of the simulation in contrast to the Metropolis algorithm where the walker has the same probability of moving in every direction.

The Fokker-Planck equation yields a (the solution to the equation) transition probability given by the Green's function

$$G(y, x, \Delta t) = \frac{1}{(4\pi D\Delta t)^{3N/2}} \exp\left(-(y - x - D\Delta t F(x))^2/4D\Delta t\right)$$

which in turn means that our brute force Metropolis algorithm

$$A(y,x) = \min(1, q(y,x)),$$

with  $q(y,x) = |\Psi_T(y)|^2/|\Psi_T(x)|^2$  is now replaced by the Metropolis-Hastings algorithm as well as Hasting's article,

$$q(y,x) = \frac{G(x,y,\Delta t)|\Psi_T(y)|^2}{G(y,x,\Delta t)|\Psi_T(x)|^2}$$

# Importance sampling, program elements

The general derivative formula of the Jastrow factor is (the subscript C stands for Correlation)

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{\partial g_{ik}}{\partial x_k} + \sum_{i=k+1}^{N} \frac{\partial g_{ki}}{\partial x_k}$$

However, with our written in way which can be reused later as

$$\Psi_C = \prod_{i < j} g(r_{ij}) = \exp\left\{ \sum_{i < j} f(r_{ij}) \right\},\,$$

the gradient needed for the quantum force and local energy is easy to compute. The function  $f(r_{ij})$  will depend on the system under study. In the equations below we will keep this general form.

In the Metropolis/Hasting algorithm, the acceptance ratio determines the probability for a particle to be accepted at a new position. The ratio of the trial wave functions evaluated at the new and current positions is given by (OB for the onebody part)

$$R \equiv \frac{\Psi_T^{new}}{\Psi_T^{old}} = \frac{\Psi_{OB}^{new}}{\Psi_{OB}^{old}} \frac{\Psi_C^{new}}{\Psi_C^{old}}$$

Here  $\Psi_{OB}$  is our onebody part (Slater determinant or product of boson single-particle states) while  $\Psi_C$  is our correlation function, or Jastrow factor. We need to optimize the  $\nabla \Psi_T/\Psi_T$  ratio and the second derivative as well, that is the  $\nabla^2 \Psi_T/\Psi_T$  ratio. The first is needed when we compute the so-called quantum force in importance sampling. The second is needed when we compute the kinetic energy term of the local energy.

$$\frac{\nabla \Psi}{\Psi} = \frac{\nabla (\Psi_{OB} \, \Psi_C)}{\Psi_{OB} \, \Psi_C} = \frac{\Psi_C \nabla \Psi_{OB} + \Psi_{OB} \nabla \Psi_C}{\Psi_{OB} \Psi_C} = \frac{\nabla \Psi_{OB}}{\Psi_{OB}} + \frac{\nabla \Psi_C}{\Psi_C}$$

The expectation value of the kinetic energy expressed in atomic units for electron i is

$$\begin{split} \langle \hat{K}_i \rangle &= -\frac{1}{2} \frac{\langle \Psi | \nabla_i^2 | \Psi \rangle}{\langle \Psi | \Psi \rangle}, \\ \hat{K}_i &= -\frac{1}{2} \frac{\nabla_i^2 \Psi}{\Psi}. \end{split}$$

The second derivative which enters the definition of the local energy is

$$\frac{\nabla^2 \Psi}{\Psi} = \frac{\nabla^2 \Psi_{OB}}{\Psi_{OB}} + \frac{\nabla^2 \Psi_C}{\Psi_C} + 2 \frac{\nabla \Psi_{OB}}{\Psi_{OB}} \cdot \frac{\nabla \Psi_C}{\Psi_C}$$

We discuss here how to calculate these quantities in an optimal way, We have defined the correlated function as

$$\Psi_C = \prod_{i < j} g(r_{ij}) = \prod_{i < j}^N g(r_{ij}) = \prod_{i=1}^N \prod_{j=i+1}^N g(r_{ij}),$$

with  $r_{ij} = |\mathbf{r}_i - \mathbf{r}_j| = \sqrt{(x_i - x_j)^2 + (y_i - y_j)^2 + (z_i - z_j)^2}$  in three dimensions or  $r_{ij} = |\mathbf{r}_i - \mathbf{r}_j| = \sqrt{(x_i - x_j)^2 + (y_i - y_j)^2}$  if we work with two-dimensional systems.

In our particular case we have

$$\Psi_C = \prod_{i < j} g(r_{ij}) = \exp\left\{ \sum_{i < j} f(r_{ij}) \right\}.$$

The total number of different relative distances  $r_{ij}$  is N(N-1)/2. In a matrix storage format, the relative distances form a strictly upper triangular matrix

$$\mathbf{r} \equiv \begin{pmatrix} 0 & r_{1,2} & r_{1,3} & \cdots & r_{1,N} \\ \vdots & 0 & r_{2,3} & \cdots & r_{2,N} \\ \vdots & \vdots & 0 & \ddots & \vdots \\ \vdots & \vdots & \vdots & \ddots & r_{N-1,N} \\ 0 & 0 & 0 & \cdots & 0 \end{pmatrix}.$$

This applies to  $\mathbf{g} = \mathbf{g}(r_{ij})$  as well.

In our algorithm we will move one particle at the time, say the kth-particle. This sampling will be seen to be particularly efficient when we are going to compute a Slater determinant.

We have that the ratio between Jastrow factors  $R_C$  is given by

$$R_C = \frac{\Psi_C^{\text{new}}}{\Psi_C^{\text{cur}}} = \prod_{i=1}^{k-1} \frac{g_{ik}^{\text{new}}}{g_{ik}^{\text{cur}}} \prod_{i=k+1}^{N} \frac{g_{ki}^{\text{new}}}{g_{ki}^{\text{cur}}}.$$

For the Pade-Jastrow form

$$R_C = \frac{\Psi_C^{\text{new}}}{\Psi_C^{\text{cur}}} = \frac{\exp U_{new}}{\exp U_{cur}} = \exp \Delta U,$$

where

$$\Delta U = \sum_{i=1}^{k-1} (f_{ik}^{\text{new}} - f_{ik}^{\text{cur}}) + \sum_{i=k+1}^{N} (f_{ki}^{\text{new}} - f_{ki}^{\text{cur}})$$

One needs to develop a special algorithm that runs only through the elements of the upper triangular matrix  $\mathbf{g}$  and have k as an index.

The expression to be derived in the following is of interest when computing the quantum force and the kinetic energy. It has the form

$$\frac{\nabla_i \Psi_C}{\Psi_C} = \frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_i},$$

for all dimensions and with i running over all particles.

For the first derivative only N-1 terms survive the ratio because the g-terms that are not differentiated cancel with their corresponding ones in the denominator. Then,

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{1}{g_{ik}} \frac{\partial g_{ik}}{\partial x_k} + \sum_{i=k+1}^N \frac{1}{g_{ki}} \frac{\partial g_{ki}}{\partial x_k}.$$

An equivalent equation is obtained for the exponential form after replacing  $g_{ij}$  by  $\exp(f_{ij})$ , yielding:

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{\partial g_{ik}}{\partial x_k} + \sum_{i=k+1}^{N} \frac{\partial g_{ki}}{\partial x_k},$$

with both expressions scaling as  $\mathcal{O}(N)$ .

Using the identity

$$\frac{\partial}{\partial x_i}g_{ij} = -\frac{\partial}{\partial x_j}g_{ij},$$

we get expressions where all the derivatives acting on the particle are represented by the second index of g:

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{1}{g_{ik}} \frac{\partial g_{ik}}{\partial x_k} - \sum_{i=k+1}^N \frac{1}{g_{ki}} \frac{\partial g_{ki}}{\partial x_i},$$

and for the exponential case:

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{\partial g_{ik}}{\partial x_k} - \sum_{i=k+1}^N \frac{\partial g_{ki}}{\partial x_i}.$$

For correlation forms depending only on the scalar distances  $r_{ij}$  we can use the chain rule. Noting that

$$\frac{\partial g_{ij}}{\partial x_j} = \frac{\partial g_{ij}}{\partial r_{ij}} \frac{\partial r_{ij}}{\partial x_j} = \frac{x_j - x_i}{r_{ij}} \frac{\partial g_{ij}}{\partial r_{ij}},$$

we arrive at

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{1}{g_{ik}} \frac{\mathbf{r_{ik}}}{r_{ik}} \frac{\partial g_{ik}}{\partial r_{ik}} - \sum_{i=k+1}^{N} \frac{1}{g_{ki}} \frac{\mathbf{r_{ki}}}{r_{ki}} \frac{\partial g_{ki}}{\partial r_{ki}}.$$

Note that for the Pade-Jastrow form we can set  $g_{ij} \equiv g(r_{ij}) = e^{f(r_{ij})} = e^{f_{ij}}$  and

$$\frac{\partial g_{ij}}{\partial r_{ij}} = g_{ij} \frac{\partial f_{ij}}{\partial r_{ij}}.$$

Therefore,

$$\frac{1}{\Psi_C} \frac{\partial \Psi_C}{\partial x_k} = \sum_{i=1}^{k-1} \frac{\mathbf{r_{ik}}}{r_{ik}} \frac{\partial f_{ik}}{\partial r_{ik}} - \sum_{i=k+1}^{N} \frac{\mathbf{r_{ki}}}{r_{ki}} \frac{\partial f_{ki}}{\partial r_{ki}},$$

where

$$\mathbf{r}_{ij} = |\mathbf{r}_j - \mathbf{r}_i| = (x_j - x_i)\mathbf{e}_1 + (y_j - y_i)\mathbf{e}_2 + (z_j - z_i)\mathbf{e}_3$$

is the relative distance.

The second derivative of the Jastrow factor divided by the Jastrow factor (the way it enters the kinetic energy) is

$$\left[\frac{\nabla^2 \Psi_C}{\Psi_C}\right]_x = 2 \sum_{k=1}^N \sum_{i=1}^{k-1} \frac{\partial^2 g_{ik}}{\partial x_k^2} + \sum_{k=1}^N \left(\sum_{i=1}^{k-1} \frac{\partial g_{ik}}{\partial x_k} - \sum_{i=k+1}^N \frac{\partial g_{ki}}{\partial x_i}\right)^2$$

But we have a simple form for the function, namely

$$\Psi_C = \prod_{i < j} \exp f(r_{ij}),$$

and it is easy to see that for particle k we have

$$\frac{\nabla_k^2 \Psi_C}{\Psi_C} = \sum_{ij \neq k} \frac{(\mathbf{r}_k - \mathbf{r}_i)(\mathbf{r}_k - \mathbf{r}_j)}{r_{ki}r_{kj}} f'(r_{ki}) f'(r_{kj}) + \sum_{j \neq k} \left( f''(r_{kj}) + \frac{2}{r_{kj}} f'(r_{kj}) \right)$$

# Importance sampling, Fokker-Planck and Langevin equations

A stochastic process is simply a function of two variables, one is the time, the other is a stochastic variable X, defined by specifying

- the set  $\{x\}$  of possible values for X;
- the probability distribution,  $w_X(x)$ , over this set, or briefly w(x)

The set of values  $\{x\}$  for X may be discrete, or continuous. If the set of values is continuous, then  $w_X(x)$  is a probability density so that  $w_X(x)dx$  is the probability that one finds the stochastic variable X to have values in the range [x, x + dx].

An arbitrary number of other stochastic variables may be derived from X. For example, any Y given by a mapping of X, is also a stochastic variable. The mapping may also be time-dependent, that is, the mapping depends on an additional variable t

$$Y_X(t) = f(X, t).$$

The quantity  $Y_X(t)$  is called a random function, or, since t often is time, a stochastic process. A stochastic process is a function of two variables, one is the time, the other is a stochastic variable X. Let x be one of the possible values of X then

$$y(t) = f(x, t),$$

is a function of t, called a sample function or realization of the process. In physics one considers the stochastic process to be an ensemble of such sample functions.

For many physical systems initial distributions of a stochastic variable y tend to equilibrium distributions:  $w(y,t) \to w_0(y)$  as  $t \to \infty$ . In equilibrium detailed balance constrains the transition rates

$$W(y \to y')w(y) = W(y' \to y)w_0(y),$$

where  $W(y' \to y)$  is the probability, per unit time, that the system changes from a state  $|y\rangle$ , characterized by the value y for the stochastic variable Y, to a state  $|y'\rangle$ .

Note that for a system in equilibrium the transition rate  $W(y' \to y)$  and the reverse  $W(y \to y')$  may be very different.

Consider, for instance, a simple system that has only two energy levels  $\epsilon_0 = 0$  and  $\epsilon_1 = \Delta E$ .

For a system governed by the Boltzmann distribution we find (the partition function has been taken out)

$$W(0 \to 1) \exp{-(\epsilon_0/kT)} = W(1 \to 0) \exp{-(\epsilon_1/kT)}$$

We get then

$$\frac{W(1 \to 0)}{W(0 \to 1)} = \exp{-(\Delta E/kT)},$$

which goes to zero when T tends to zero.

If we assume a discrete set of events, our initial probability distribution function can be given by

$$w_i(0) = \delta_{i,0},$$

and its time-development after a given time step  $\Delta t = \epsilon$  is

$$w_i(t) = \sum_j W(j \to i) w_j(t=0).$$

The continuous analog to  $w_i(0)$  is

$$w(\mathbf{x}) \to \delta(\mathbf{x})$$
.

where we now have generalized the one-dimensional position x to a generic-dimensional vector  $\mathbf{x}$ . The Kroenecker  $\delta$  function is replaced by the  $\delta$  distribution function  $\delta(\mathbf{x})$  at t=0.

The transition from a state j to a state i is now replaced by a transition to a state with position  $\mathbf{y}$  from a state with position  $\mathbf{x}$ . The discrete sum of

transition probabilities can then be replaced by an integral and we obtain the new distribution at a time  $t+\Delta t$  as

$$w(\mathbf{y}, t + \Delta t) = \int W(\mathbf{y}, t + \Delta t | \mathbf{x}, t) w(\mathbf{x}, t) d\mathbf{x},$$

and after m time steps we have

$$w(\mathbf{y}, t + m\Delta t) = \int W(\mathbf{y}, t + m\Delta t | \mathbf{x}, t) w(\mathbf{x}, t) d\mathbf{x}.$$

When equilibrium is reached we have

$$w(\mathbf{y}) = \int W(\mathbf{y}|\mathbf{x}, t)w(\mathbf{x})d\mathbf{x},$$

that is no time-dependence. Note our change of notation for W

We can solve the equation for  $w(\mathbf{y},t)$  by making a Fourier transform to momentum space. The PDF  $w(\mathbf{x},t)$  is related to its Fourier transform  $\tilde{w}(\mathbf{k},t)$  through

$$w(\mathbf{x},t) = \int_{-\infty}^{\infty} d\mathbf{k} \exp{(i\mathbf{k}\mathbf{x})} \tilde{w}(\mathbf{k},t),$$

and using the definition of the  $\delta$ -function

$$\delta(\mathbf{x}) = \frac{1}{2\pi} \int_{-\infty}^{\infty} d\mathbf{k} \exp{(i\mathbf{k}\mathbf{x})},$$

we see that

$$\tilde{w}(\mathbf{k},0) = 1/2\pi.$$

We can then use the Fourier-transformed diffusion equation

$$\frac{\partial \tilde{w}(\mathbf{k}, t)}{\partial t} = -D\mathbf{k}^2 \tilde{w}(\mathbf{k}, t),$$

with the obvious solution

$$\tilde{w}(\mathbf{k},t) = \tilde{w}(\mathbf{k},0) \exp\left[-(D\mathbf{k}^2 t)\right] = \frac{1}{2\pi} \exp\left[-(D\mathbf{k}^2 t)\right].$$

With the Fourier transform we obtain

$$w(\mathbf{x},t) = \int_{-\infty}^{\infty} d\mathbf{k} \exp\left[i\mathbf{k}\mathbf{x}\right] \frac{1}{2\pi} \exp\left[-(D\mathbf{k}^2 t)\right] = \frac{1}{\sqrt{4\pi Dt}} \exp\left[-(\mathbf{x}^2/4Dt)\right],$$

with the normalization condition

$$\int_{-\infty}^{\infty} w(\mathbf{x}, t) d\mathbf{x} = 1.$$

The solution represents the probability of finding our random walker at position  $\mathbf{x}$  at time t if the initial distribution was placed at  $\mathbf{x} = 0$  at t = 0.

There is another interesting feature worth observing. The discrete transition probability W itself is given by a binomial distribution. The results from the central limit theorem state that transition probability in the limit  $n \to \infty$  converges to the normal distribution. It is then possible to show that

$$W(il - jl, n\epsilon) \to W(\mathbf{y}, t + \Delta t | \mathbf{x}, t) = \frac{1}{\sqrt{4\pi D\Delta t}} \exp\left[-((\mathbf{y} - \mathbf{x})^2/4D\Delta t)\right],$$

and that it satisfies the normalization condition and is itself a solution to the diffusion equation.

Let us now assume that we have three PDFs for times  $t_0 < t' < t$ , that is  $w(\mathbf{x}_0, t_0), w(\mathbf{x}', t')$  and  $w(\mathbf{x}, t)$ . We have then

$$w(\mathbf{x},t) = \int_{-\infty}^{\infty} W(\mathbf{x}.t|\mathbf{x}'.t')w(\mathbf{x}',t')d\mathbf{x}',$$

and

$$w(\mathbf{x},t) = \int_{-\infty}^{\infty} W(\mathbf{x}.t|\mathbf{x}_0.t_0)w(\mathbf{x}_0,t_0)d\mathbf{x}_0,$$

and

$$w(\mathbf{x}',t') = \int_{-\infty}^{\infty} W(\mathbf{x}'.t'|\mathbf{x}_0,t_0)w(\mathbf{x}_0,t_0)d\mathbf{x}_0.$$

We can combine these equations and arrive at the famous Einstein-Smoluchenski-Kolmogorov-Chapman (ESKC) relation

$$W(\mathbf{x}t|\mathbf{x}_0t_0) = \int_{-\infty}^{\infty} W(\mathbf{x},t|\mathbf{x}',t')W(\mathbf{x}',t'|\mathbf{x}_0,t_0)d\mathbf{x}'.$$

We can replace the spatial dependence with a dependence upon say the velocity (or momentum), that is we have

$$W(\mathbf{v},t|\mathbf{v}_0,t_0) = \int_{-\infty}^{\infty} W(\mathbf{v},t|\mathbf{v}',t')W(\mathbf{v}',t'|\mathbf{v}_0,t_0)d\mathbf{x}'.$$

We will now derive the Fokker-Planck equation. We start from the ESKC equation  $\,$ 

$$W(\mathbf{x},t|\mathbf{x}_0,t_0) = \int_{-\infty}^{\infty} W(\mathbf{x},t|\mathbf{x}',t')W(\mathbf{x}',t'|\mathbf{x}_0,t_0)d\mathbf{x}'.$$

Define  $s = t' - t_0$ ,  $\tau = t - t'$  and  $t - t_0 = s + \tau$ . We have then

$$W(\mathbf{x}, s + \tau | \mathbf{x}_0) = \int_{-\infty}^{\infty} W(\mathbf{x}, \tau | \mathbf{x}') W(\mathbf{x}', s | \mathbf{x}_0) d\mathbf{x}'.$$

Assume now that  $\tau$  is very small so that we can make an expansion in terms of a small step xi, with  $\mathbf{x}' = \mathbf{x} - \xi$ , that is

$$W(\mathbf{x}, s | \mathbf{x}_0) + \frac{\partial W}{\partial s} \tau + O(\tau^2) = \int_{-\infty}^{\infty} W(\mathbf{x}, \tau | \mathbf{x} - \xi) W(\mathbf{x} - \xi, s | \mathbf{x}_0) d\mathbf{x}'.$$

We assume that  $W(\mathbf{x}, \tau | \mathbf{x} - \xi)$  takes non-negligible values only when  $\xi$  is small. This is just another way of stating the Master equation!!

We say thus that  $\mathbf{x}$  changes only by a small amount in the time interval  $\tau$ . This means that we can make a Taylor expansion in terms of  $\xi$ , that is we expand

$$W(\mathbf{x},\tau|\mathbf{x}-\xi)W(\mathbf{x}-\xi,s|\mathbf{x}_0) = \sum_{n=0}^{\infty} \frac{(-\xi)^n}{n!} \frac{\partial^n}{\partial x^n} \left[ W(\mathbf{x}+\xi,\tau|\mathbf{x})W(\mathbf{x},s|\mathbf{x}_0) \right].$$

We can then rewrite the ESKC equation as

$$\frac{\partial W}{\partial s}\tau = -W(\mathbf{x}, s|\mathbf{x}_0) + \sum_{n=0}^{\infty} \frac{(-\xi)^n}{n!} \frac{\partial^n}{\partial x^n} \left[ W(\mathbf{x}, s|\mathbf{x}_0) \int_{-\infty}^{\infty} \xi^n W(\mathbf{x} + \xi, \tau|\mathbf{x}) d\xi \right].$$

We have neglected higher powers of  $\tau$  and have used that for n = 0 we get simply  $W(\mathbf{x}, s|\mathbf{x}_0)$  due to normalization.

We say thus that  $\mathbf{x}$  changes only by a small amount in the time interval  $\tau$ . This means that we can make a Taylor expansion in terms of  $\xi$ , that is we expand

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We can then rewrite the ESKC equation as

$$\frac{\partial W(\mathbf{x},s|\mathbf{x}_0)}{\partial s}\tau = -W(\mathbf{x},s|\mathbf{x}_0) + \sum_{n=0}^{\infty} \frac{(-\xi)^n}{n!} \frac{\partial^n}{\partial x^n} \left[ W(\mathbf{x},s|\mathbf{x}_0) \int_{-\infty}^{\infty} \xi^n W(\mathbf{x}+\xi,\tau|\mathbf{x}) d\xi \right].$$

We have neglected higher powers of  $\tau$  and have used that for n=0 we get simply  $W(\mathbf{x}, s|\mathbf{x}_0)$  due to normalization.

We simplify the above by introducing the moments

$$M_n = \frac{1}{\tau} \int_{-\infty}^{\infty} \xi^n W(\mathbf{x} + \xi, \tau | \mathbf{x}) d\xi = \frac{\langle [\Delta x(\tau)]^n \rangle}{\tau},$$

resulting in

$$\frac{\partial W(\mathbf{x}, s | \mathbf{x}_0)}{\partial s} = \sum_{n=1}^{\infty} \frac{(-\xi)^n}{n!} \frac{\partial^n}{\partial x^n} \left[ W(\mathbf{x}, s | \mathbf{x}_0) M_n \right].$$

When  $\tau \to 0$  we assume that  $\langle [\Delta x(\tau)]^n \rangle \to 0$  more rapidly than  $\tau$  itself if n > 2. When  $\tau$  is much larger than the standard correlation time of system then  $M_n$  for n > 2 can normally be neglected. This means that fluctuations become negligible at large time scales.

If we neglect such terms we can rewrite the ESKC equation as

$$\frac{\partial W(\mathbf{x}, s | \mathbf{x}_0)}{\partial s} = -\frac{\partial M_1 W(\mathbf{x}, s | \mathbf{x}_0)}{\partial x} + \frac{1}{2} \frac{\partial^2 M_2 W(\mathbf{x}, s | \mathbf{x}_0)}{\partial x^2}.$$

In a more compact form we have

$$\frac{\partial W}{\partial s} = -\frac{\partial M_1 W}{\partial x} + \frac{1}{2} \frac{\partial^2 M_2 W}{\partial x^2},$$

which is the Fokker-Planck equation! It is trivial to replace position with velocity (momentum).

Consider a particle suspended in a liquid. On its path through the liquid it will continuously collide with the liquid molecules. Because on average the particle will collide more often on the front side than on the back side, it will experience a systematic force proportional with its velocity, and directed opposite to its velocity. Besides this systematic force the particle will experience a stochastic force  $\mathbf{F}(t)$ . The equations of motion are

- $\frac{d\mathbf{r}}{dt} = \mathbf{v}$  and
- $\frac{d\mathbf{v}}{dt} = -\xi\mathbf{v} + \mathbf{F}$ .

From hydrodynamics we know that the friction constant  $\xi$  is given by

$$\xi = 6\pi \eta a/m$$

where  $\eta$  is the viscosity of the solvent and a is the radius of the particle . Solving the second equation in the previous slide we get

$$\mathbf{v}(t) = \mathbf{v}_0 e^{-\xi t} + \int_0^t d\tau e^{-\xi(t-\tau)} \mathbf{F}(\tau).$$

If we want to get some useful information out of this, we have to average over all possible realizations of  $\mathbf{F}(t)$ , with the initial velocity as a condition. A useful quantity for example is

$$\langle \mathbf{v}(t) \cdot \mathbf{v}(t) \rangle_{\mathbf{v}_0} = v_0^{-\xi 2t} + 2 \int_0^t d\tau e^{-\xi (2t - \tau)} \mathbf{v}_0 \cdot \langle \mathbf{F}(\tau) \rangle_{\mathbf{v}_0}$$
$$+ \int_0^t d\tau' \int_0^t d\tau e^{-\xi (2t - \tau - \tau')} \langle \mathbf{F}(\tau) \cdot \mathbf{F}(\tau') \rangle_{\mathbf{v}_0}.$$

In order to continue we have to make some assumptions about the conditional averages of the stochastic forces. In view of the chaotic character of the stochastic forces the following assumptions seem to be appropriate

$$\langle \mathbf{F}(t) \rangle = 0$$
,

and

$$\langle \mathbf{F}(t) \cdot \mathbf{F}(t') \rangle_{\mathbf{v}_0} = C_{\mathbf{v}_0} \delta(t - t').$$

We omit the subscript  $\mathbf{v}_0$ , when the quantity of interest turns out to be independent of  $\mathbf{v}_0$ . Using the last three equations we get

$$\langle \mathbf{v}(t) \cdot \mathbf{v}(t) \rangle_{\mathbf{v}_0} = v_0^2 e^{-2\xi t} + \frac{C_{\mathbf{v}_0}}{2\xi} (1 - e^{-2\xi t}).$$

For large t this should be equal to 3kT/m, from which it follows that

$$\langle \mathbf{F}(t) \cdot \mathbf{F}(t') \rangle = 6 \frac{kT}{m} \xi \delta(t - t').$$

This result is called the fluctuation-dissipation theorem .

Integrating

$$\mathbf{v}(t) = \mathbf{v}_0 e^{-\xi t} + \int_0^t d\tau e^{-\xi(t-\tau)} \mathbf{F}(\tau),$$

we get

$$\mathbf{r}(t) = \mathbf{r}_0 + \mathbf{v}_0 \frac{1}{\xi} (1 - e^{-\xi t}) + \int_0^t d\tau \int_0^\tau \tau' e^{-\xi(\tau - \tau')} \mathbf{F}(\tau'),$$

from which we calculate the mean square displacement

$$\langle (\mathbf{r}(t) - \mathbf{r}_0)^2 \rangle_{\mathbf{v}_0} = \frac{v_0^2}{\xi} (1 - e^{-\xi t})^2 + \frac{3kT}{m\xi^2} (2\xi t - 3 + 4e^{-\xi t} - e^{-2\xi t}).$$

For very large t this becomes

$$\langle (\mathbf{r}(t) - \mathbf{r}_0)^2 \rangle = \frac{6kT}{m\xi}t$$

from which we get the Einstein relation

$$D = \frac{kT}{m\xi}$$

where we have used  $\langle (\mathbf{r}(t) - \mathbf{r}_0)^2 \rangle = 6Dt$ .

## Code example for two electrons in a quantum dots

```
# 2-electron VMC code for 2dim quantum dot with importance sampling
# Using gaussian rng for new positions and Metropolis- Hastings
# No energy minimization
from math import exp, sqrt
from random import random, seed, normalvariate import numpy as np
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D
from matplotlib import cm
from matplotlib.ticker import LinearLocator, FormatStrFormatter
import sys
from numba import jit,njit
#Read name of output file from command line
if len(sys.argv) == 2:
    outfilename = sys.argv[1]
else:
    print('\nError: Name of output file must be given as command line argument.\n')
outfile = open(outfilename,'w')
# Trial wave function for the 2-electron quantum dot in two dims
```

```
def WaveFunction(r,alpha,beta):
    r1 = r[0,0]**2 + r[0,1]**2

r2 = r[1,0]**2 + r[1,1]**2
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = r12/(1+beta*r12)
    return exp(-0.5*alpha*(r1+r2)+deno)
\# Local energy for the 2-electron quantum dot in two dims, using analytical local energy def LocalEnergy(r,alpha,beta):
    r1 = (r[0,0]**2 + r[0,1]**2)
    r2 = (r[1,0]**2 + r[1,1]**2)
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    deno2 = deno*deno
    return 0.5*(1-alpha*alpha)*(r1 + r2) +2.0*alpha + 1.0/r12+deno2*(alpha*r12-deno2+2*beta*deno-
# Setting up the quantum force for the two-electron quantum dot, recall that it is a vector
def QuantumForce(r,alpha,beta):
    qforce = np.zeros((NumberParticles,Dimension), np.double)
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    qforce [0,:] = -2*r[0,:]*alpha*(r[0,:]-r[1,:])*deno*deno/r12
    qforce[1,:] = -2*r[1,:]*alpha*(r[1,:]-r[0,:])*deno*deno/r12
    return qforce
# The Monte Carlo sampling with the Metropolis algo
# jit decorator tells Numba to compile this function.
# The argument types will be inferred by Numba when function is called.
@jit()
def MonteCarloSampling():
    NumberMCcycles= 100000
    # Parameters in the Fokker-Planck simulation of the quantum force
    D = 0.5
    TimeStep = 0.05
    # positions
    PositionOld = np.zeros((NumberParticles, Dimension), np.double)
    PositionNew = np.zeros((NumberParticles, Dimension), np.double)
    # Quantum force
    QuantumForceOld = np.zeros((NumberParticles, Dimension), np.double)
    QuantumForceNew = np.zeros((NumberParticles, Dimension), np.double)
    # seed for rng generator
    seed()
    # start variational parameter loops, two parameters here
    alpha = 0.9
    for ia in range(MaxVariations):
        alpha += .025
        AlphaValues[ia] = alpha
        beta = 0.2
        for jb in range(MaxVariations):
            beta += .01
            BetaValues[jb] = beta
            energy = energy2 = 0.0
            DeltaE = 0.0
            #Initial position
            for i in range(NumberParticles):
                for j in range(Dimension):
                     PositionOld[i,j] = normalvariate(0.0,1.0)*sqrt(TimeStep)
```

```
wfold = WaveFunction(PositionOld,alpha,beta)
            QuantumForceOld = QuantumForce(PositionOld,alpha, beta)
            #Loop over MC MCcycles
            for MCcycle in range(NumberMCcycles):
                #Trial position moving one particle at the time
                for i in range(NumberParticles):
                    for j in range (Dimension):
                        PositionNew[i,j] = PositionOld[i,j]+normalvariate(0.0,1.0)*sqrt(TimeStep)
                                            QuantumForceOld[i,j]*TimeStep*D
                    wfnew = WaveFunction(PositionNew,alpha,beta)
                    QuantumForceNew = QuantumForce(PositionNew,alpha, beta)
                    GreensFunction = 0.0
                    for j in range(Dimension):
                        GreensFunction += 0.5*(QuantumForceOld[i,j]+QuantumForceNew[i,j])*\
                                       (D*TimeStep*0.5*(QuantumForceOld[i,j]-QuantumForceNew[i,j])
                                       PositionNew[i,j]+PositionOld[i,j])
                    GreensFunction = exp(GreensFunction)
                    ProbabilityRatio = GreensFunction*wfnew**2/wfold**2
                    #Metropolis-Hastings test to see whether we accept the move
                    if random() <= ProbabilityRatio:</pre>
                       for j in range(Dimension):
                           PositionOld[i,j] = PositionNew[i,j]
                           QuantumForceOld[i,j] = QuantumForceNew[i,j]
                       wfold = wfnew
                DeltaE = LocalEnergy(PositionOld,alpha,beta)
                energy += DeltaE
                energy2 += DeltaE**2
            # We calculate mean, variance and error (no blocking applied)
            energy /= NumberMCcycles
            energy2 /= NumberMCcycles
            variance = energy2 - energy**2
            error = sqrt(variance/NumberMCcycles)
            Energies[ia,jb] = energy
outfile.write('%f %f %f %f %f \n' %(alpha,beta,energy,variance,error))
   return Energies, AlphaValues, BetaValues
#Here starts the main program with variable declarations
NumberParticles = 2
Dimension = 2
MaxVariations = 10
Energies = np.zeros((MaxVariations, MaxVariations))
AlphaValues = np.zeros(MaxVariations)
BetaValues = np.zeros(MaxVariations)
(Energies, AlphaValues, BetaValues) = MonteCarloSampling()
outfile.close()
# Prepare for plots
fig = plt.figure()
ax = fig.gca(projection='3d')
# Plot the surface.
X, Y = np.meshgrid(AlphaValues, BetaValues)
surf = ax.plot_surface(X, Y, Energies,cmap=cm.coolwarm,linewidth=0, antialiased=False)
# Customize the z axis
zmin = np.matrix(Energies).min()
zmax = np.matrix(Energies).max()
ax.set_zlim(zmin, zmax)
ax.set_xlabel(r'$\alpha$')
ax.set_ylabel(r'$\beta$')
ax.set_zlabel(r'$\langle E \rangle$')
```

```
ax.zaxis.set_major_locator(LinearLocator(10))
ax.zaxis.set_major_formatter(FormatStrFormatter('%.02f'))
# Add a color bar which maps values to colors.
fig.colorbar(surf, shrink=0.5, aspect=5)
plt.show()
```

Bringing the gradient optmization. The simple one-particle case in a harmonic oscillator trap

```
# Gradient descent stepping with analytical derivative
import numpy as np
from scipy.optimize import minimize
def DerivativeE(x):
    return x-1.0/(4*x*x*x);
def Energy(x):
  return x*x*0.5+1.0/(8*x*x);
x0 = 1.0
eta = 0.1
Niterations = 100
for iter in range(Niterations):
    gradients = DerivativeE(x0)
    x0 -= eta*gradients
print(x0)
# 2-electron VMC code for 2dim quantum dot with importance sampling
# Using gaussian rng for new positions and Metropolis- Hastings
from math import exp, sqrt
from random import random, seed, normalvariate
import numpy as np
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D
from matplotlib import cm
from matplotlib.ticker import LinearLocator, FormatStrFormatter
import sys
from numba import jit
# Trial wave function for the 2-electron quantum dot in two dims
def WaveFunction(r,alpha):
    r1 = r[0,0]**2 + r[0,1]**2
    r2 = r[1,0]**2 + r[1,1]**2
    return exp(-0.5*alpha*(r1+r2))
\hbox{\it\#Local energy for the 2-electron quantum dot in two dims, using analytical local energy $$ def LocalEnergy(r,alpha):}
    r1 = (r[0,0]**2 + r[0,1]**2)
    r2 = (r[1,0]**2 + r[1,1]**2)
    return 0.5*(1-alpha*alpha)*(r1 + r2) +2.0*alpha
# Derivate of wave function ansatz as function of variational parameters
def DerivativeWFansatz(r,alpha):
```

```
r1 = (r[0,0]**2 + r[0,1]**2)
       r2 = (r[1,0]**2 + r[1,1]**2)
       WfDer = -(r1+r2)
       return WfDer
# Setting up the quantum force for the two-electron quantum dot, recall that it is a vector
def QuantumForce(r,alpha):
        qforce = np.zeros((NumberParticles,Dimension), np.double)
        qforce[0,:] = -2*r[0,:]*alpha
       qforce[1,:] = -2*r[1,:]*alpha
       return qforce
# Computing the derivative of the energy and the energy
# jit decorator tells Numba to compile this function.
# The argument types will be inferred by Numba when function is called.
@jit
def EnergyMinimization(alpha):
       NumberMCcycles= 1000
        # Parameters in the Fokker-Planck simulation of the quantum force
       D = 0.5
       TimeStep = 0.05
        # positions
       PositionOld = np.zeros((NumberParticles, Dimension), np.double)
       PositionNew = np.zeros((NumberParticles, Dimension), np.double)
        # Quantum force
        QuantumForceOld = np.zeros((NumberParticles, Dimension), np.double)
        QuantumForceNew = np.zeros((NumberParticles, Dimension), np.double)
        # seed for rng generator
       seed()
       energy = 0.0
DeltaE = 0.0
       EnergyDer = 0.0
       DeltaPsi = 0.0
       DerivativePsiE = 0.0
        #Initial position
       for i in range(NumberParticles):
                for j in range(Dimension):
                        PositionOld[i,j] = normalvariate(0.0,1.0)*sqrt(TimeStep)
       wfold = WaveFunction(PositionOld,alpha)
       QuantumForceOld = QuantumForce(PositionOld,alpha)
        #Loop over MC MCcycles
        for MCcycle in range(NumberMCcycles):
                 #Trial position moving one particle at the time
                for i in range(NumberParticles):
                        for j in range(Dimension):
                                \label{eq:positionNew} \textbf{PositionNew[i,j] = PositionOld[i,j] + normal variate(0.0,1.0) * sqrt(TimeStep) + (0.0,1.0) * sqrt(TimeSte
                                                                                QuantumForceOld[i,j]*TimeStep*D
                        wfnew = WaveFunction(PositionNew,alpha)
                        QuantumForceNew = QuantumForce(PositionNew,alpha)
                        GreensFunction = 0.0
                        for j in range(Dimension):
                                GreensFunction += 0.5*(QuantumForceOld[i,j]+QuantumForceNew[i,j])*\
                                                                              (D*TimeStep*0.5*(QuantumForceOld[i,j]-QuantumForceNew[i,j])
                                                                             PositionNew[i,j]+PositionOld[i,j])
                        GreensFunction = exp(GreensFunction)
                        ProbabilityRatio = GreensFunction*wfnew**2/wfold**2
```

```
#Metropolis-Hastings test to see whether we accept the move
            if random() <= ProbabilityRatio:</pre>
                for j in range (Dimension):
                    PositionOld[i,j] = PositionNew[i,j]
                    QuantumForceOld[i,j] = QuantumForceNew[i,j]
                wfold = wfnew
        DeltaE = LocalEnergy(PositionOld,alpha)
        DeltaPsi = DerivativeWFansatz(PositionOld,alpha)
        energy += DeltaE
        DerivativePsiE += DeltaPsi*DeltaE
    # We calculate mean, variance and error (no blocking applied)
    energy /= NumberMCcycles
   DerivativePsiE /= NumberMCcycles
   DeltaPsi /= NumberMCcycles
   EnergyDer = 2*(DerivativePsiE-DeltaPsi*energy)
   return energy, EnergyDer
#Here starts the main program with variable declarations
NumberParticles = 2
Dimension = 2
# guess for variational parameters
x0 = 1.5
{\it \# Set up iteration using stochastic gradient method}
Energy = 0 ; EnergyDer = 0
Energy, EnergyDer = EnergyMinimization(x0)
print(Energy, EnergyDer)
eta = 0.01
Niterations = 100
for iter in range(Niterations):
   gradients = EnergyDer
   x0 -= eta*gradients
   Energy, EnergyDer = EnergyMinimization(x0)
print(x0)
```

# VMC for fermions: Efficient calculation of Slater determinants

The potentially most time-consuming part is the evaluation of the gradient and the Laplacian of an N-particle Slater determinant.

We have to differentiate the determinant with respect to all spatial coordinates of all particles. A brute force differentiation would involve  $N \cdot d$  evaluations of the entire determinant which would even worsen the already undesirable time scaling, making it  $Nd \cdot O(N^3) \sim O(d \cdot N^4)$ .

This poses serious hindrances to the overall efficiency of our code.

The efficiency can be improved however if we move only one electron at the time. The Slater determinant matrix  $\hat{D}$  is defined by the matrix elements

$$d_{ij} = \phi_i(x_i)$$

where  $\phi_j(\mathbf{r}_i)$  is a single particle wave function. The columns correspond to the position of a given particle while the rows stand for the various quantum numbers.

What we need to realize is that when differentiating a Slater determinant with respect to some given coordinate, only one row of the corresponding Slater matrix is changed.

Therefore, by recalculating the whole determinant we risk producing redundant information. The solution turns out to be an algorithm that requires to keep track of the *inverse* of the Slater matrix.

Let the current position in phase space be represented by the  $(N \cdot d)$ -element vector  $\mathbf{r}^{\text{old}}$  and the new suggested position by the vector  $\mathbf{r}^{\text{new}}$ .

The inverse of  $\hat{D}$  can be expressed in terms of its cofactors  $C_{ij}$  and its determinant (this our notation for a determinant)  $|\hat{D}|$ :

$$d_{ij}^{-1} = \frac{C_{ji}}{|\hat{D}|} \tag{14}$$

Notice that the interchanged indices indicate that the matrix of cofactors is to be transposed.

If  $\hat{D}$  is invertible, then we must obviously have  $\hat{D}^{-1}\hat{D} = \mathbf{1}$ , or explicitly in terms of the individual elements of  $\hat{D}$  and  $\hat{D}^{-1}$ :

$$\sum_{k=1}^{N} d_{ik} d_{kj}^{-1} = \delta_{ij} \tag{15}$$

Consider the ratio, which we shall call R, between  $|\hat{D}(\mathbf{r}^{\text{new}})|$  and  $|\hat{D}(\mathbf{r}^{\text{old}})|$ . By definition, each of these determinants can individually be expressed in terms of the i-th row of its cofactor matrix

$$R \equiv \frac{|\hat{D}(\mathbf{r}^{\text{new}})|}{|\hat{D}(\mathbf{r}^{\text{old}})|} = \frac{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{new}}) C_{ij}(\mathbf{r}^{\text{new}})}{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{old}}) C_{ij}(\mathbf{r}^{\text{old}})}$$
(16)

Suppose now that we move only one particle at a time, meaning that  $\mathbf{r}^{\text{new}}$  differs from  $\mathbf{r}^{\text{old}}$  by the position of only one, say the *i*-th, particle. This means that  $\hat{D}(\mathbf{r}^{\text{new}})$  and  $\hat{D}(\mathbf{r}^{\text{old}})$  differ only by the entries of the *i*-th row. Recall also that the *i*-th row of a cofactor matrix  $\hat{C}$  is independent of the entries of the *i*-th row of its corresponding matrix  $\hat{D}$ . In this particular case we therefore get that the *i*-th row of  $\hat{C}(\mathbf{r}^{\text{new}})$  and  $\hat{C}(\mathbf{r}^{\text{old}})$  must be equal. Explicitly, we have:

$$C_{ij}(\mathbf{r}^{\text{new}}) = C_{ij}(\mathbf{r}^{\text{old}}) \quad \forall \ j \in \{1, \dots, N\}$$
 (17)

Inserting this into the numerator of eq. (16) and using eq. (14) to substitute the cofactors with the elements of the inverse matrix, we get:

$$R = \frac{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{new}}) C_{ij}(\mathbf{r}^{\text{old}})}{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{old}}) C_{ij}(\mathbf{r}^{\text{old}})} = \frac{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{new}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}})}{\sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{old}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}})}$$
(18)

Now by eq. (15) the denominator of the rightmost expression must be unity, so that we finally arrive at:

$$R = \sum_{j=1}^{N} d_{ij}(\mathbf{r}^{\text{new}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}}) = \sum_{j=1}^{N} \phi_{j}(\mathbf{r}_{i}^{\text{new}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}})$$
(19)

What this means is that in order to get the ratio when only the *i*-th particle has been moved, we only need to calculate the dot product of the vector  $(\phi_1(\mathbf{r}_i^{\text{new}}), \ldots, \phi_N(\mathbf{r}_i^{\text{new}}))$  of single particle wave functions evaluated at this new position with the *i*-th column of the inverse matrix  $\hat{D}^{-1}$  evaluated at the original position. Such an operation has a time scaling of O(N). The only extra thing we need to do is to maintain the inverse matrix  $\hat{D}^{-1}(\mathbf{x}^{\text{old}})$ .

If the new position  $\mathbf{r}^{\text{new}}$  is accepted, then the inverse matrix can by suitably updated by an algorithm having a time scaling of  $O(N^2)$ . This algorithm goes as follows. First we update all but the *i*-th column of  $\hat{D}^{-1}$ . For each column  $j \neq i$ , we first calculate the quantity:

$$S_j = (\hat{D}(\mathbf{r}^{\text{new}}) \times \hat{D}^{-1}(\mathbf{r}^{\text{old}}))_{ij} = \sum_{l=1}^N d_{il}(\mathbf{r}^{\text{new}}) d_{lj}^{-1}(\mathbf{r}^{\text{old}})$$
(20)

The new elements of the j-th column of  $\hat{D}^{-1}$  are then given by:

$$d_{kj}^{-1}(\mathbf{r}^{\text{new}}) = d_{kj}^{-1}(\mathbf{r}^{\text{old}}) - \frac{S_j}{R} d_{ki}^{-1}(\mathbf{r}^{\text{old}}) \quad \begin{array}{c} \forall & k \in \{1, \dots, N\} \\ j \neq i \end{array}$$
 (21)

Finally the elements of the *i*-th column of  $\hat{D}^{-1}$  are updated simply as follows:

$$d_{ki}^{-1}(\mathbf{r}^{\text{new}}) = \frac{1}{R} d_{ki}^{-1}(\mathbf{r}^{\text{old}}) \quad \forall \quad k \in \{1, \dots, N\}$$
 (22)

We see from these formulas that the time scaling of an update of  $\hat{D}^{-1}$  after changing one row of  $\hat{D}$  is  $O(N^2)$ .

The scheme is also applicable for the calculation of the ratios involving derivatives. It turns out that differentiating the Slater determinant with respect to the coordinates of a single particle  $\mathbf{r}_i$  changes only the *i*-th row of the corresponding Slater matrix.

The gradient and the Laplacian. The gradient and the Laplacian can therefore be calculated as follows:

$$\frac{\vec{\nabla}_i |\hat{D}(\mathbf{r})|}{|\hat{D}(\mathbf{r})|} = \sum_{j=1}^N \vec{\nabla}_i d_{ij}(\mathbf{r}) d_{ji}^{-1}(\mathbf{r}) = \sum_{j=1}^N \vec{\nabla}_i \phi_j(\mathbf{r}_i) d_{ji}^{-1}(\mathbf{r})$$

and

$$\frac{\nabla_i^2 |\hat{D}(\mathbf{r})|}{|\hat{D}(\mathbf{r})|} = \sum_{j=1}^N \nabla_i^2 d_{ij}(\mathbf{r}) d_{ji}^{-1}(\mathbf{r}) = \sum_{j=1}^N \nabla_i^2 \phi_j(\mathbf{r}_i) d_{ji}^{-1}(\mathbf{r})$$

Thus, to calculate all the derivatives of the Slater determinant, we only need the derivatives of the single particle wave functions  $(\vec{\nabla}_i \phi_j(\mathbf{r}_i))$  and  $\nabla_i^2 \phi_j(\mathbf{r}_i)$  and the elements of the corresponding inverse Slater matrix  $(\hat{D}^{-1}(\mathbf{r}_i))$ . A calculation of a single derivative is by the above result an O(N) operation. Since there are  $d \cdot N$  derivatives, the time scaling of the total evaluation becomes  $O(d \cdot N^2)$ . With an  $O(N^2)$  updating algorithm for the inverse matrix, the total scaling is no worse, which is far better than the brute force approach yielding  $O(d \cdot N^4)$ .

**Important note**: In most cases you end with closed form expressions for the single-particle wave functions. It is then useful to calculate the various derivatives and make separate functions for them.

The Slater determinant takes the form

$$\Phi(\mathbf{r}_{1}, \mathbf{r}_{2}, , \mathbf{r}_{3}, \mathbf{r}_{4}, \alpha, \beta, \gamma, \delta) = \frac{1}{\sqrt{4!}} \begin{vmatrix} \psi_{100\uparrow}(\mathbf{r}_{1}) & \psi_{100\uparrow}(\mathbf{r}_{2}) & \psi_{100\uparrow}(\mathbf{r}_{3}) & \psi_{100\uparrow}(\mathbf{r}_{4}) \\ \psi_{100\downarrow}(\mathbf{r}_{1}) & \psi_{100\downarrow}(\mathbf{r}_{2}) & \psi_{100\downarrow}(\mathbf{r}_{3}) & \psi_{100\downarrow}(\mathbf{r}_{4}) \\ \psi_{200\uparrow}(\mathbf{r}_{1}) & \psi_{200\uparrow}(\mathbf{r}_{2}) & \psi_{200\uparrow}(\mathbf{r}_{3}) & \psi_{200\uparrow}(\mathbf{r}_{4}) \\ \psi_{200\downarrow}(\mathbf{r}_{1}) & \psi_{200\downarrow}(\mathbf{r}_{2}) & \psi_{200\downarrow}(\mathbf{r}_{3}) & \psi_{200\downarrow}(\mathbf{r}_{4}) \end{vmatrix}$$

The Slater determinant as written is zero since the spatial wave functions for the spin up and spin down states are equal. But we can rewrite it as the product of two Slater determinants, one for spin up and one for spin down.

We can rewrite it as

$$\Phi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4, \alpha, \beta, \gamma, \delta) = \det \uparrow (1, 2) \det \downarrow (3, 4) - \det \uparrow (1, 3) \det \downarrow (2, 4) \\
- \det \uparrow (1, 4) \det \downarrow (3, 2) + \det \uparrow (2, 3) \det \downarrow (1, 4) - \det \uparrow (2, 4) \det \downarrow (1, 3) \\
+ \det \uparrow (3, 4) \det \downarrow (1, 2),$$

where we have defined

$$\det \uparrow (1,2) = \frac{1}{\sqrt{2}} \left| \begin{array}{cc} \psi_{100\uparrow}(\mathbf{r}_1) & \psi_{100\uparrow}(\mathbf{r}_2) \\ \psi_{200\uparrow}(\mathbf{r}_1) & \psi_{200\uparrow}(\mathbf{r}_2) \end{array} \right|,$$

and

$$\det \downarrow (3,4) = \frac{1}{\sqrt{2}} \left| \begin{array}{cc} \psi_{100\downarrow}(\mathbf{r}_3) & \psi_{100\downarrow}(\mathbf{r}_4) \\ \psi_{200\downarrow}(\mathbf{r}_3) & \psi_{200\downarrow}(\mathbf{r}_4) \end{array} \right|.$$

We want to avoid to sum over spin variables, in particular when the interaction does not depend on spin.

It can be shown, see for example Moskowitz and Kalos, Int. J. Quantum Chem. **20** 1107 (1981), that for the variational energy we can approximate the Slater determinant as

$$\Phi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4, \alpha, \beta, \gamma, \delta) \propto \det \uparrow (1, 2) \det \downarrow (3, 4),$$

or more generally as

$$\Phi(\mathbf{r}_1, \mathbf{r}_2, \dots \mathbf{r}_N) \propto \det \uparrow \det \downarrow$$
,

where we have the Slater determinant as the product of a spin up part involving the number of electrons with spin up only (2 for beryllium and 5 for neon) and a spin down part involving the electrons with spin down.

This ansatz is not antisymmetric under the exchange of electrons with opposite spins but it can be shown (show this) that it gives the same expectation value for the energy as the full Slater determinant.

As long as the Hamiltonian is spin independent, the above is correct. It is rather straightforward to see this if you go back to the equations for the energy discussed earlier this semester.

We will thus factorize the full determinant  $|\hat{D}|$  into two smaller ones, where each can be identified with  $\uparrow$  and  $\downarrow$  respectively:

$$|\hat{D}| = |\hat{D}|_{\uparrow} \cdot |\hat{D}|_{\downarrow\downarrow}$$

The combined dimensionality of the two smaller determinants equals the dimensionality of the full determinant. Such a factorization is advantageous in that it makes it possible to perform the calculation of the ratio R and the updating of the inverse matrix separately for  $|\hat{D}|_{\uparrow}$  and  $|\hat{D}|_{\downarrow}$ :

$$\frac{|\hat{D}|^{\text{new}}}{|\hat{D}|^{\text{old}}} = \frac{|\hat{D}|^{\text{new}}_{\uparrow}}{|\hat{D}|^{\text{old}}_{\uparrow}} \cdot \frac{|\hat{D}|^{\text{new}}_{\downarrow}}{|\hat{D}|^{\text{old}}_{\downarrow}}$$

This reduces the calculation time by a constant factor. The maximal time reduction happens in a system of equal numbers of  $\uparrow$  and  $\downarrow$  particles, so that the two factorized determinants are half the size of the original one.

Consider the case of moving only one particle at a time which originally had the following time scaling for one transition:

$$O_R(N) + O_{\text{inverse}}(N^2)$$

For the factorized determinants one of the two determinants is obviously unaffected by the change so that it cancels from the ratio R.

Therefore, only one determinant of size N/2 is involved in each calculation of R and update of the inverse matrix. The scaling of each transition then becomes:

$$O_R(N/2) + O_{\text{inverse}}(N^2/4)$$

and the time scaling when the transitions for all N particles are put together:

$$O_R(N^2/2) + O_{\text{inverse}}(N^3/4)$$

which gives the same reduction as in the case of moving all particles at once.

Computing the ratios discussed above requires that we maintain the inverse of the Slater matrix evaluated at the current position. Each time a trial position is accepted, the row number i of the Slater matrix changes and updating its inverse has to be carried out. Getting the inverse of an  $N \times N$  matrix by Gaussian elimination has a complexity of order of  $\mathcal{O}(N^3)$  operations, a luxury that we cannot afford for each time a particle move is accepted. We will use the expression

$$d_{kj}^{-1}(\mathbf{x}^{\mathbf{new}}) = \begin{cases} d_{kj}^{-1}(\mathbf{x}^{\mathbf{old}}) - \frac{d_{ki}^{-1}(\mathbf{x}^{\mathbf{old}})}{R} \sum_{l=1}^{N} d_{il}(\mathbf{x}^{\mathbf{new}}) d_{lj}^{-1}(\mathbf{x}^{\mathbf{old}}) & \text{if } j \neq i \\ \frac{d_{ki}^{-1}(\mathbf{x}^{\mathbf{old}})}{R} \sum_{l=1}^{N} d_{il}(\mathbf{x}^{\mathbf{old}}) d_{lj}^{-1}(\mathbf{x}^{\mathbf{old}}) & \text{if } j = i \end{cases}$$

This equation scales as  $O(N^2)$ . The evaluation of the determinant of an  $N \times N$  matrix by standard Gaussian elimination requires  $\mathbf{O}(N^3)$  calculations. As there are Nd independent coordinates we need to evaluate Nd Slater determinants for the gradient (quantum force) and Nd for the Laplacian (kinetic energy). With the updating algorithm we need only to invert the Slater determinant matrix once. This can be done by standard LU decomposition methods.

Expectation value of the kinetic energy. The expectation value of the kinetic energy expressed in atomic units for electron i is

$$\langle \hat{K}_i \rangle = -\frac{1}{2} \frac{\langle \Psi | \nabla_i^2 | \Psi \rangle}{\langle \Psi | \Psi \rangle},$$

$$K_i = -\frac{1}{2} \frac{\nabla_i^2 \Psi}{\Psi}.$$
(23)

$$\begin{split} \frac{\nabla^2 \Psi}{\Psi} &= \frac{\nabla^2 (\Psi_D \, \Psi_C)}{\Psi_D \, \Psi_C} = \frac{\nabla \cdot \left[ \nabla (\Psi_D \, \Psi_C) \right]}{\Psi_D \, \Psi_C} = \frac{\nabla \cdot \left[ \Psi_C \nabla \Psi_D + \Psi_D \nabla \Psi_C \right]}{\Psi_D \, \Psi_C} \\ &= \frac{\nabla \Psi_C \cdot \nabla \Psi_D + \Psi_C \nabla^2 \Psi_D + \nabla \Psi_D \cdot \nabla \Psi_C + \Psi_D \nabla^2 \Psi_C}{\Psi_D \, \Psi_C} \end{split} \tag{24}$$

$$\frac{\nabla^2 \Psi}{\Psi} = \frac{\nabla^2 \Psi_D}{\Psi_D} + \frac{\nabla^2 \Psi_C}{\Psi_C} + 2 \frac{\nabla \Psi_D}{\Psi_D} \cdot \frac{\nabla \Psi_C}{\Psi_C}$$
 (25)

The second derivative of the Jastrow factor divided by the Jastrow factor (the way it enters the kinetic energy) is

$$\left[\frac{\nabla^2 \Psi_C}{\Psi_C}\right]_x = 2 \sum_{k=1}^N \sum_{i=1}^{k-1} \frac{\partial^2 g_{ik}}{\partial x_k^2} + \sum_{k=1}^N \left(\sum_{i=1}^{k-1} \frac{\partial g_{ik}}{\partial x_k} - \sum_{i=k+1}^N \frac{\partial g_{ki}}{\partial x_i}\right)^2$$

But we have a simple form for the function, namely

$$\Psi_C = \prod_{i < j} \exp f(r_{ij}) = \exp \left\{ \sum_{i < j} \frac{ar_{ij}}{1 + \beta r_{ij}} \right\},\,$$

and it is easy to see that for particle k we have

$$\frac{\nabla_k^2 \Psi_C}{\Psi_C} = \sum_{ij \neq k} \frac{(\mathbf{r}_k - \mathbf{r}_i)(\mathbf{r}_k - \mathbf{r}_j)}{r_{ki}r_{kj}} f'(r_{ki}) f'(r_{kj}) + \sum_{j \neq k} \left( f''(r_{kj}) + \frac{2}{r_{kj}} f'(r_{kj}) \right)$$

Using

$$f(r_{ij}) = \frac{ar_{ij}}{1 + \beta r_{ij}},$$

and  $g'(r_{kj}) = dg(r_{kj})/dr_{kj}$  and  $g''(r_{kj}) = d^2g(r_{kj})/dr_{kj}^2$  we find that for particle k we have

$$\frac{\nabla_k^2 \Psi_C}{\Psi_C} = \sum_{ij \neq k} \frac{(\mathbf{r}_k - \mathbf{r}_i)(\mathbf{r}_k - \mathbf{r}_j)}{r_{ki}r_{kj}} \frac{a}{(1 + \beta r_{ki})^2} \frac{a}{(1 + \beta r_{kj})^2} + \sum_{j \neq k} \left( \frac{2a}{r_{kj}(1 + \beta r_{kj})^2} - \frac{2a\beta}{(1 + \beta r_{kj})^3} \right)$$

The gradient and Laplacian can be calculated as follows:

$$\frac{\nabla_i |\hat{D}(\mathbf{r})|}{|\hat{D}(\mathbf{r})|} = \sum_{j=1}^N \vec{\nabla}_i d_{ij}(\mathbf{r}) d_{ji}^{-1}(\mathbf{r}) = \sum_{j=1}^N \vec{\nabla}_i \phi_j(\mathbf{r}_i) d_{ji}^{-1}(\mathbf{r})$$

and

$$\frac{\nabla_i^2 |\hat{D}(\mathbf{r})|}{|\hat{D}(\mathbf{r})|} = \sum_{i=1}^N \nabla_i^2 d_{ij}(\mathbf{r}) \, d_{ji}^{-1}(\mathbf{r}) = \sum_{i=1}^N \nabla_i^2 \phi_j(\mathbf{r}_i) \, d_{ji}^{-1}(\mathbf{r})$$

The gradient for the determinant is

$$\frac{\nabla_i |\hat{D}(\mathbf{r})|}{|\hat{D}(\mathbf{r})|} = \sum_{j=1}^N \nabla_i d_{ij}(\mathbf{r}) d_{ji}^{-1}(\mathbf{r}) = \sum_{j=1}^N \nabla_i \phi_j(\mathbf{r}_i) d_{ji}^{-1}(\mathbf{r}).$$

We have

$$\Psi_C = \prod_{i < j} g(r_{ij}) = \exp\left\{\sum_{i < j} \frac{ar_{ij}}{1 + \beta r_{ij}}\right\},\,$$

the gradient needed for the quantum force and local energy is easy to compute. We get for particle  $\boldsymbol{k}$ 

$$\frac{\nabla_k \Psi_C}{\Psi_C} = \sum_{j \neq k} \frac{\mathbf{r}_{kj}}{r_{kj}} \frac{a}{(1 + \beta r_{kj})^2},$$

which is rather easy to code. Remember to sum over all particles when you compute the local energy.

We need to compute the ratio between wave functions, in particular for the Slater determinants.

$$R = \sum_{i=1}^{N} d_{ij}(\mathbf{r}^{\text{new}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}}) = \sum_{i=1}^{N} \phi_{j}(\mathbf{r}_{i}^{\text{new}}) d_{ji}^{-1}(\mathbf{r}^{\text{old}})$$

What this means is that in order to get the ratio when only the *i*-th particle has been moved, we only need to calculate the dot product of the vector  $(\phi_1(\mathbf{r}_i^{\text{new}}), \ldots, \phi_N(\mathbf{r}_i^{\text{new}}))$  of single particle wave functions evaluated at this new position with the *i*-th column of the inverse matrix  $\hat{D}^{-1}$  evaluated at the original position. Such an operation has a time scaling of O(N). The only extra thing we need to do is to maintain the inverse matrix  $\hat{D}^{-1}(\mathbf{x}^{\text{old}})$ .

## Gradient Methods

# Top-down start

- We will start with a top-down view, with a simple harmonic oscillator problem in one dimension as case.
- Thereafter we continue with implementing the simplest possible steepest descent approach to our two-electron problem with an electrostatic (Coulomb) interaction. Our code includes also importance sampling. The simple Python code here illustrates the basic elements which need to be included in our own code.
- Then we move on to the mathematical description of various gradient methods.

#### Motivation

Our aim with this part is to be able to

- find an optimal value for the variational parameters using only some few Monte Carlo cycles
- use these optimal values for the variational parameters to perform a large-scale Monte Carlo calculation

## Simple example and demonstration

Let us illustrate what is needed in our calculations using a simple example, the harmonic oscillator in one dimension. For the harmonic oscillator in onedimension we have a trial wave function and probability

$$\psi_T(x;\alpha) = \exp{-(\frac{1}{2}\alpha^2x^2)},$$

which results in a local energy

$$\frac{1}{2}\left(\alpha^2+x^2(1-\alpha^4)\right).$$

We can compare our numerically calculated energies with the exact energy as function of  $\alpha$ 

$$\overline{E}[\alpha] = \frac{1}{4} \left( \alpha^2 + \frac{1}{\alpha^2} \right).$$

# Simple example and demonstration

The derivative of the energy with respect to  $\alpha$  gives

$$\frac{d\langle E_L[\alpha]\rangle}{d\alpha} = \frac{1}{2}\alpha - \frac{1}{2\alpha^3}$$

and a second derivative which is always positive (meaning that we find a minimum)

$$\frac{d^2 \langle E_L[\alpha] \rangle}{d\alpha^2} = \frac{1}{2} + \frac{3}{2\alpha^4}$$

The condition

$$\frac{d\langle E_L[\alpha]\rangle}{d\alpha} = 0,$$

gives the optimal  $\alpha = 1$ , as expected.

\*

Exercise 1: Find the local energy for the harmonic oscillator aragraph!paragraph>paragraph>-0.5em

- a) Derive the local energy for the harmonic oscillator in one dimension and find its expectation value.
  - aragraph!paragraph>paragraph>-0.5em
  - b) Show also that the optimal value of optimal  $\alpha=1$  aragraph!paragraph>paragraph>-0.5em
- c) Repeat the above steps in two dimensions for N bosons or electrons. What is the optimal value of  $\alpha$ ?

# Variance in the simple model

We can also minimize the variance. In our simple model the variance is

$$\sigma^2[\alpha] = \frac{1}{4} \left( 1 + (1 - \alpha^4)^2 \frac{3}{4\alpha^4} \right) - \overline{E}^2.$$

which yields a second derivative which is always positive.

# Computing the derivatives

In general we end up computing the expectation value of the energy in terms of some parameters  $\alpha_0, \alpha_1, \ldots, \alpha_n$  and we search for a minimum in this multivariable parameter space. This leads to an energy minimization problem where we need the derivative of the energy as a function of the variational parameters.

In the above example this was easy and we were able to find the expression for the derivative by simple derivations. However, in our actual calculations the energy is represented by a multi-dimensional integral with several variational parameters. How can we can then obtain the derivatives of the energy with respect to the variational parameters without having to resort to expensive numerical derivations?

# Expressions for finding the derivatives of the local energy

To find the derivatives of the local energy expectation value as function of the variational parameters, we can use the chain rule and the hermiticity of the Hamiltonian.

Let us define

$$\bar{E}_{\alpha} = \frac{d\langle E_L[\alpha] \rangle}{d\alpha}.$$

as the derivative of the energy with respect to the variational parameter  $\alpha$  (we limit ourselves to one parameter only). In the above example this was easy and we obtain a simple expression for the derivative. We define also the derivative of the trial function (skipping the subindex T) as

$$\bar{\psi}_{\alpha} = \frac{d\psi[\alpha]\rangle}{d\alpha}.$$

# Derivatives of the local energy

The elements of the gradient of the local energy are then (using the chain rule and the hermiticity of the Hamiltonian)

$$\bar{E}_{\alpha} = 2 \left( \langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} E_{L}[\alpha] \rangle - \langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} \rangle \langle E_{L}[\alpha] \rangle \right).$$

From a computational point of view it means that you need to compute the expectation values of

 $\langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} E_L[\alpha] \rangle$ ,

and

$$\langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} \rangle \langle E_L[\alpha] \rangle$$

\*

Exercise 2: General expression for the derivative of the energy aragraph!paragraph>paragraph>-0.5em

a) Show that

$$\bar{E}_{\alpha} = 2 \left( \langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} E_{L}[\alpha] \rangle - \langle \frac{\bar{\psi}_{\alpha}}{\psi[\alpha]} \rangle \langle E_{L}[\alpha] \rangle \right).$$

aragraph!paragraph>paragraph>-0.5em

b) Find the corresponding expression for the variance.

# Python program for 2-electrons in 2 dimensions

```
# 2-electron VMC code for 2dim quantum dot with importance sampling
# Using gaussian rng for new positions and Metropolis- Hastings
# Added energy minimization with gradient descent using fixed step size
# To do: replace with optimization codes from scipy and/or use stochastic gradient descent
from math import exp, sqrt
from random import random, seed, normalvariate
import numpy as np
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D
from matplotlib import cm
from matplotlib.ticker import LinearLocator, FormatStrFormatter
import sys
# Trial wave function for the 2-electron quantum dot in two dims
def WaveFunction(r,alpha,beta):
    r1 = r[0,0]**2 + r[0,1]**2
    r2 = r[1,0]**2 + r[1,1]**2
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = r12/(1+beta*r12)
    return exp(-0.5*alpha*(r1+r2)+deno)
# Local energy for the 2-electron quantum dot in two dims, using analytical local energy def LocalEnergy(r,alpha,beta):
    r1 = (r[0,0]**2 + r[0,1]**2)
    r2 = (r[1,0]**2 + r[1,1]**2)
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    deno2 = deno*deno
    return 0.5*(1-alpha*alpha)*(r1 + r2) +2.0*alpha + 1.0/r12+deno2*(alpha*r12-deno2+2*beta*deno-
# Derivate of wave function ansatz as function of variational parameters
def DerivativeWFansatz(r,alpha,beta):
    WfDer = np.zeros((2), np.double)
    r1 = (r[0,0]**2 + r[0,1]**2)
    r2 = (r[1,0]**2 + r[1,1]**2)
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    deno2 = deno*deno
    WfDer[0] = -0.5*(r1+r2)
    WfDer[1] = -r12*r12*deno2
    return WfDer
# Setting up the quantum force for the two-electron quantum dot, recall that it is a vector
def QuantumForce(r,alpha,beta):
    qforce = np.zeros((NumberParticles,Dimension), np.double)
    r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    qforce[0,:] = -2*r[0,:]*alpha*(r[0,:]-r[1,:])*deno*deno/r12
    qforce[1,:] = -2*r[1,:]*alpha*(r[1,:]-r[0,:])*deno*deno/r12
    return qforce
# Computing the derivative of the energy and the energy
def EnergyMinimization(alpha, beta):
```

```
NumberMCcycles= 10000
# Parameters in the Fokker-Planck simulation of the quantum force
D = 0.5
TimeStep = 0.05
# positions
PositionOld = np.zeros((NumberParticles, Dimension), np.double)
PositionNew = np.zeros((NumberParticles, Dimension), np.double)
# Quantum force
QuantumForceOld = np.zeros((NumberParticles, Dimension), np.double)
QuantumForceNew = np.zeros((NumberParticles, Dimension), np.double)
# seed for rng generator
seed()
energy = 0.0
DeltaE = 0.0
EnergyDer = np.zeros((2), np.double)
DeltaPsi = np.zeros((2), np.double)
DerivativePsiE = np.zeros((2), np.double)
#Initial position
for i in range(NumberParticles):
        for j in range(Dimension):
                PositionOld[i,j] = normalvariate(0.0,1.0)*sqrt(TimeStep)
wfold = WaveFunction(PositionOld, alpha, beta)
QuantumForceOld = QuantumForce(PositionOld,alpha, beta)
#Loop over MC MCcycles
for MCcycle in range(NumberMCcycles):
        #Trial position moving one particle at the time
        for i in range(NumberParticles):
                for j in range(Dimension):
                        PositionNew[i,j] = PositionOld[i,j] + normal variate(0.0,1.0) * sqrt(TimeStep) + Variation (0.0,1.0) * sqrt(TimeStep) + Variation(TimeStep) + Variation(
                                                                     {\tt QuantumForceOld[i,j]*TimeStep*D}
                wfnew = WaveFunction(PositionNew, alpha, beta)
                QuantumForceNew = QuantumForce(PositionNew,alpha, beta)
                GreensFunction = 0.0
                for j in range(Dimension):
                        GreensFunction += 0.5*(QuantumForceOld[i,j]+QuantumForceNew[i,j])*\
                                                                   (D*TimeStep*0.5*(QuantumForceOld[i,j]-QuantumForceNew[i,j])\\
                                                                   PositionNew[i,j]+PositionOld[i,j])
                GreensFunction = exp(GreensFunction)
                ProbabilityRatio = GreensFunction*wfnew**2/wfold**2
                for j in range (Dimension):
                                PositionOld[i,j] = PositionNew[i,j]
                                QuantumForceOld[i,j] = QuantumForceNew[i,j]
                        wfold = wfnew
       DeltaE = LocalEnergy(PositionOld,alpha,beta)
        DerPsi = DerivativeWFansatz(PositionOld,alpha,beta)
       DeltaPsi += DerPsi
        energy += DeltaE
        DerivativePsiE += DerPsi*DeltaE
# We calculate mean values
energy /= NumberMCcycles
DerivativePsiE /= NumberMCcycles
DeltaPsi /= NumberMCcycles
EnergyDer = 2*(DerivativePsiE-DeltaPsi*energy)
return energy, EnergyDer
```

```
#Here starts the main program with variable declarations
NumberParticles = 2
Dimension = 2
# guess for variational parameters
alpha = 0.9
beta = 0.2
# Set up iteration using gradient descent method
Energy = 0
EDerivative = np.zeros((2), np.double)
eta = 0.01
Niterations = 50
for iter in range(Niterations):
    Energy, EDerivative = EnergyMinimization(alpha, beta)
    alphagradient = EDerivative[0]
    betagradient = EDerivative[1]
    alpha -= eta*alphagradient
    beta -= eta*betagradient
print(alpha, beta)
print(Energy, EDerivative[0], EDerivative[1])
```

# Using Broyden's algorithm in scipy

The following function uses the above described BFGS algorithm. Here we have defined a function which calculates the energy and a function which computes the first derivative.

```
\# 2-electron VMC code for 2dim quantum dot with importance sampling
# Using gaussian rng for new positions and Metropolis- Hastings
# Added energy minimization using the BFGS algorithm, see p. 136 of https://www.springer.com/it/b
from math import exp, sqrt
from random import random, seed, normalvariate
import numpy as np
import matplotlib.pyplot as plt
from mpl_toolkits.mplot3d import Axes3D
from matplotlib import cm
from matplotlib.ticker import LinearLocator, FormatStrFormatter
from scipy.optimize import minimize
import sys
# Trial wave function for the 2-electron quantum dot in two dims
def WaveFunction(r,alpha,beta):
   r1 = r[0,0]**2 + r[0,1]**2
r2 = r[1,0]**2 + r[1,1]**2
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
   deno = r12/(1+beta*r12)
   return exp(-0.5*alpha*(r1+r2)+deno)
# Local energy for the 2-electron quantum dot in two dims, using analytical local energy
def LocalEnergy(r,alpha,beta):
```

```
r1 = (r[0,0]**2 + r[0,1]**2)
   r2 = (r[1,0]**2 + r[1,1]**2)
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
   deno = 1.0/(1+beta*r12)
   deno2 = deno*deno
   return 0.5*(1-alpha*alpha)*(r1 + r2) +2.0*alpha + 1.0/r12+deno2*(alpha*r12-deno2+2*beta*deno-
# Derivate of wave function ansatz as function of variational parameters
def DerivativeWFansatz(r,alpha,beta):
    WfDer = np.zeros((2), np.double)
   r1 = (r[0,0]**2 + r[0,1]**2)
   r2 = (r[1,0]**2 + r[1,1]**2)
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
   deno = 1.0/(1+beta*r12)
   deno2 = deno*deno
   WfDer[0] = -0.5*(r1+r2)
   WfDer[1] = -r12*r12*deno2
   return WfDer
\# Setting up the quantum force for the two-electron quantum dot, recall that it is a vector
def QuantumForce(r,alpha,beta):
    qforce = np.zeros((NumberParticles,Dimension), np.double)
   r12 = sqrt((r[0,0]-r[1,0])**2 + (r[0,1]-r[1,1])**2)
    deno = 1.0/(1+beta*r12)
    qforce[0,:] = -2*r[0,:]*alpha*(r[0,:]-r[1,:])*deno*deno/r12
    qforce[1,:] = -2*r[1,:]*alpha*(r[1,:]-r[0,:])*deno*deno/r12
   return qforce
# Computing the derivative of the energy and the energy
def EnergyDerivative(x0):
    # Parameters in the Fokker-Planck simulation of the quantum force
   D = 0.5
   TimeStep = 0.05
   NumberMCcycles= 10000
    # positions
   PositionOld = np.zeros((NumberParticles, Dimension), np.double)
   PositionNew = np.zeros((NumberParticles, Dimension), np.double)
    # Quantum force
    QuantumForceOld = np.zeros((NumberParticles, Dimension), np.double)
   QuantumForceNew = np.zeros((NumberParticles, Dimension), np.double)
   energy = 0.0
DeltaE = 0.0
   alpha = x0[0]
   beta = x0[1]
    EnergyDer = 0.0
   DeltaPsi = 0.0
   DerivativePsiE = 0.0
    #Initial position
   for i in range(NumberParticles):
       for j in range(Dimension):
            PositionOld[i,j] = normalvariate(0.0,1.0)*sqrt(TimeStep)
   wfold = WaveFunction(PositionOld,alpha,beta)
   QuantumForceOld = QuantumForce(PositionOld,alpha, beta)
    #Loop over MC MCcycles
```

```
for MCcycle in range(NumberMCcycles):
                #Trial position moving one particle at the time
                for i in range(NumberParticles):
                        for j in range(Dimension):
                                PositionNew[i,j] = PositionOld[i,j] + normal variate(0.0,1.0) * sqrt(TimeStep) + Variation (0.0,1.0) * sqrt(TimeStep) + Variation(TimeStep) + Variation(
                                                                                QuantumForceOld[i,j]*TimeStep*D
                         wfnew = WaveFunction(PositionNew,alpha,beta)
                         QuantumForceNew = QuantumForce(PositionNew,alpha, beta)
                        GreensFunction = 0.0
                         for j in range(Dimension):
                                GreensFunction += 0.5*(QuantumForceOld[i,j]+QuantumForceNew[i,j])*\
                                                                              (D*TimeStep*0.5*(QuantumForceOld[i,j]-QuantumForceNew[i,j])\\
                                                                              PositionNew[i,j]+PositionOld[i,j])
                        GreensFunction = exp(GreensFunction)
                        ProbabilityRatio = GreensFunction*wfnew**2/wfold**2
                         #Metropolis-Hastings test to see whether we accept the move
                         if random() <= ProbabilityRatio:</pre>
                                for j in range(Dimension):
                                         PositionOld[i,j] = PositionNew[i,j]
                                         QuantumForceOld[i,j] = QuantumForceNew[i,j]
                                wfold = wfnew
                DeltaE = LocalEnergy(PositionOld,alpha,beta)
                DerPsi = DerivativeWFansatz(PositionOld,alpha,beta)
                DeltaPsi += DerPsi
                energy += DeltaE
                DerivativePsiE += DerPsi*DeltaE
        # We calculate mean values
        energy /= NumberMCcycles
        DerivativePsiE /= NumberMCcycles
       DeltaPsi /= NumberMCcycles
EnergyDer = 2*(DerivativePsiE-DeltaPsi*energy)
        return EnergyDer
# Computing the expectation value of the local energy
def Energy(x0):
        # Parameters in the Fokker-Planck simulation of the quantum force
       D = 0.5
       TimeStep = 0.05
        # positions
       PositionOld = np.zeros((NumberParticles, Dimension), np.double)
       PositionNew = np.zeros((NumberParticles, Dimension), np.double)
        # Quantum force
        QuantumForceOld = np.zeros((NumberParticles, Dimension), np.double)
        QuantumForceNew = np.zeros((NumberParticles, Dimension), np.double)
        energy = 0.0
       DeltaE = 0.0
       alpha = x0[0]
beta = x0[1]
       NumberMCcycles= 10000
        #Initial position
       for i in range(NumberParticles):
                for j in range(Dimension):
                        PositionOld[i,j] = normalvariate(0.0,1.0)*sqrt(TimeStep)
       wfold = WaveFunction(PositionOld,alpha,beta)
        QuantumForceOld = QuantumForce(PositionOld,alpha, beta)
        #Loop over MC MCcycles
```

```
for MCcycle in range(NumberMCcycles):
                     #Trial position moving one particle at the time
                   for i in range(NumberParticles):
                              for j in range(Dimension):
                                        PositionNew[i,j] = PositionOld[i,j] + normal variate(0.0,1.0) * sqrt(TimeStep) + Volume (0.0,1.0) * 
                                                                                                   QuantumForceOld[i,j]*TimeStep*D
                              wfnew = WaveFunction(PositionNew,alpha,beta)
                              QuantumForceNew = QuantumForce(PositionNew,alpha, beta)
                              GreensFunction = 0.0
                              for j in range(Dimension):
                                         GreensFunction += 0.5*(QuantumForceOld[i,j]+QuantumForceNew[i,j])*\
                                                                                                 (D*TimeStep*0.5*(QuantumForceOld[i,j]-QuantumForceNew[i,j])
                                                                                                PositionNew[i,j]+PositionOld[i,j])
                              GreensFunction = exp(GreensFunction)
                              ProbabilityRatio = GreensFunction*wfnew**2/wfold**2
                              \#Metropolis	ext{-}Hastings test to see whether we accept the move
                              if random() <= ProbabilityRatio:</pre>
                                        for j in range(Dimension):
                                                  PositionOld[i,j] = PositionNew[i,j]
                                                  QuantumForceOld[i,j] = QuantumForceNew[i,j]
                                        wfold = wfnew
                   DeltaE = LocalEnergy(PositionOld,alpha,beta)
                    energy += DeltaE
          # We calculate mean values
          energy /= NumberMCcycles
         return energy
#Here starts the main program with variable declarations
NumberParticles = 2
Dimension = 2
# seed for rng generator
seed()
# quess for variational parameters
x0 = np.array([0.9,0.2])
# Using Broydens method
res = minimize(Energy, x0, method='BFGS', jac=EnergyDerivative, options={'gtol': 1e-4, 'disp': True
print(res.x)
```

Note that the **minimize** function returns the finale values for the variable  $\alpha = x0[0]$  and  $\beta = x0[1]$  in the array x.

#### Brief reminder on Newton-Raphson's method

Let us quickly remind ourselves how we derive the above method.

Perhaps the most celebrated of all one-dimensional root-finding routines is Newton's method, also called the Newton-Raphson method. This method requires the evaluation of both the function f and its derivative f' at arbitrary points. If you can only calculate the derivative numerically and/or your function is not of the smooth type, we normally discourage the use of this method.

## The equations

The Newton-Raphson formula consists geometrically of extending the tangent line at a current point until it crosses zero, then setting the next guess to the abscissa of that zero-crossing. The mathematics behind this method is rather simple. Employing a Taylor expansion for x sufficiently close to the solution s, we have

$$f(s) = 0 = f(x) + (s - x)f'(x) + \frac{(s - x)^2}{2}f''(x) + \dots$$

For small enough values of the function and for well-behaved functions, the terms beyond linear are unimportant, hence we obtain

$$f(x) + (s - x)f'(x) \approx 0,$$

yielding

$$s \approx x - \frac{f(x)}{f'(x)}$$
.

Having in mind an iterative procedure, it is natural to start iterating with

$$x_{n+1} = x_n - \frac{f(x_n)}{f'(x_n)}.$$

# Simple geometric interpretation

The above is Newton-Raphson's method. It has a simple geometric interpretation, namely  $x_{n+1}$  is the point where the tangent from  $(x_n, f(x_n))$  crosses the x-axis. Close to the solution, Newton-Raphson converges fast to the desired result. However, if we are far from a root, where the higher-order terms in the series are important, the Newton-Raphson formula can give grossly inaccurate results. For instance, the initial guess for the root might be so far from the true root as to let the search interval include a local maximum or minimum of the function. If an iteration places a trial guess near such a local extremum, so that the first derivative nearly vanishes, then Newton-Raphson may fail totally

#### Extending to more than one variable

Newton's method can be generalized to systems of several non-linear equations and variables. Consider the case with two equations

$$f_1(x_1, x_2) = 0 
 f_2(x_1, x_2) = 0,$$

which we Taylor expand to obtain

$$0 = f_1(x_1 + h_1, x_2 + h_2) = f_1(x_1, x_2) + h_1 \partial f_1 / \partial x_1 + h_2 \partial f_1 / \partial x_2 + \dots$$
  

$$0 = f_2(x_1 + h_1, x_2 + h_2) = f_2(x_1, x_2) + h_1 \partial f_2 / \partial x_1 + h_2 \partial f_2 / \partial x_2 + \dots$$

Defining the Jacobian matrix  $\hat{J}$  we have

$$\hat{J} = \left( \begin{array}{ccc} \partial f_1/\partial x_1 & \partial f_1/\partial x_2 \\ \partial f_2/\partial x_1 & \partial f_2/\partial x_2 \end{array} \right),$$

we can rephrase Newton's method as

$$\left(\begin{array}{c} x_1^{n+1} \\ x_2^{n+1} \end{array}\right) = \left(\begin{array}{c} x_1^n \\ x_2^n \end{array}\right) + \left(\begin{array}{c} h_1^n \\ h_2^n \end{array}\right),$$

where we have defined

$$\begin{pmatrix} h_1^n \\ h_2^n \end{pmatrix} = -\mathbf{\hat{J}}^{-1} \begin{pmatrix} f_1(x_1^n, x_2^n) \\ f_2(x_1^n, x_2^n) \end{pmatrix}.$$

We need thus to compute the inverse of the Jacobian matrix and it is to understand that difficulties may arise in case  $\hat{J}$  is nearly singular.

It is rather straightforward to extend the above scheme to systems of more than two non-linear equations. In our case, the Jacobian matrix is given by the Hessian that represents the second derivative of cost function.

## Steepest descent

The basic idea of gradient descent is that a function  $F(\mathbf{x})$ ,  $\mathbf{x} \equiv (x_1, \dots, x_n)$ , decreases fastest if one goes from  $\mathbf{x}$  in the direction of the negative gradient  $-\nabla F(\mathbf{x})$ .

It can be shown that if

$$\mathbf{x}_{k+1} = \mathbf{x}_k - \gamma_k \nabla F(\mathbf{x}_k),$$

with  $\gamma_k > 0$ .

For  $\gamma_k$  small enough, then  $F(\mathbf{x}_{k+1}) \leq F(\mathbf{x}_k)$ . This means that for a sufficiently small  $\gamma_k$  we are always moving towards smaller function values, i.e a minimum.

#### More on Steepest descent

The previous observation is the basis of the method of steepest descent, which is also referred to as just gradient descent (GD). One starts with an initial guess  $\mathbf{x}_0$  for a minimum of F and computes new approximations according to

$$\mathbf{x}_{k+1} = \mathbf{x}_k - \gamma_k \nabla F(\mathbf{x}_k), \quad k \ge 0.$$

The parameter  $\gamma_k$  is often referred to as the step length or the learning rate within the context of Machine Learning.

## The ideal

Ideally the sequence  $\{\mathbf{x}_k\}_{k=0}$  converges to a global minimum of the function F. In general we do not know if we are in a global or local minimum. In the special case when F is a convex function, all local minima are also global minima, so in this case gradient descent can converge to the global solution. The advantage of this scheme is that it is conceptually simple and straightforward to implement. However the method in this form has some severe limitations:

In machine learing we are often faced with non-convex high dimensional cost functions with many local minima. Since GD is deterministic we will get stuck in a local minimum, if the method converges, unless we have a very good intial guess. This also implies that the scheme is sensitive to the chosen initial condition.

Note that the gradient is a function of  $\mathbf{x} = (x_1, \dots, x_n)$  which makes it expensive to compute numerically.

## The sensitiveness of the gradient descent

The gradient descent method is sensitive to the choice of learning rate  $\gamma_k$ . This is due to the fact that we are only guaranteed that  $F(\mathbf{x}_{k+1}) \leq F(\mathbf{x}_k)$  for sufficiently small  $\gamma_k$ . The problem is to determine an optimal learning rate. If the learning rate is chosen too small the method will take a long time to converge and if it is too large we can experience erratic behavior.

Many of these shortcomings can be alleviated by introducing randomness. One such method is that of Stochastic Gradient Descent (SGD), see below.

#### Convex functions

Ideally we want our cost/loss function to be convex(concave).

First we give the definition of a convex set: A set C in  $\mathbb{R}^n$  is said to be convex if, for all x and y in C and all  $t \in (0,1)$ , the point (1-t)x+ty also belongs to C. Geometrically this means that every point on the line segment connecting x and y is in C as discussed below.

The convex subsets of  $\mathbb{R}$  are the intervals of  $\mathbb{R}$ . Examples of convex sets of  $\mathbb{R}^2$  are the regular polygons (triangles, rectangles, pentagons, etc...).

## Convex function

Convex function: Let  $X \subset \mathbb{R}^n$  be a convex set. Assume that the function  $f: X \to \mathbb{R}$  is continuous, then f is said to be convex if

$$f(tx_1 + (1-t)x_2) \le tf(x_1) + (1-t)f(x_2)$$

for all  $x_1, x_2 \in X$  and for all  $t \in [0, 1]$ . If  $\leq$  is replaced with a strict inequality in the definition, we demand  $x_1 \neq x_2$  and  $t \in (0, 1)$  then f is said to be strictly convex. For a single variable function, convexity means that if you draw a straight line connecting  $f(x_1)$  and  $f(x_2)$ , the value of the function on the interval  $[x_1, x_2]$  is always below the line as illustrated below.

#### Conditions on convex functions

In the following we state first and second-order conditions which ensures convexity of a function f. We write  $D_f$  to denote the domain of f, i.e the subset of  $R^n$  where f is defined. For more details and proofs we refer to: S. Boyd and L. Vandenberghe. Convex Optimization. Cambridge University Press.

**First order condition.** Suppose f is differentiable (i.e  $\nabla f(x)$  is well defined for all x in the domain of f). Then f is convex if and only if  $D_f$  is a convex set and

$$f(y) \ge f(x) + \nabla f(x)^T (y - x)$$

holds for all  $x, y \in D_f$ . This condition means that for a convex function the first order Taylor expansion (right hand side above) at any point a global under estimator of the function. To convince yourself you can make a drawing of  $f(x) = x^2 + 1$  and draw the tangent line to f(x) and note that it is always below the graph.

**Second order condition.** Assume that f is twice differentiable, i.e the Hessian matrix exists at each point in  $D_f$ . Then f is convex if and only if  $D_f$  is a convex set and its Hessian is positive semi-definite for all  $x \in D_f$ . For a single-variable function this reduces to  $f''(x) \geq 0$ . Geometrically this means that f has nonnegative curvature everywhere.

This condition is particularly useful since it gives us an procedure for determining if the function under consideration is convex, apart from using the definition.

#### More on convex functions

The next result is of great importance to us and the reason why we are going on about convex functions. In machine learning we frequently have to minimize a loss/cost function in order to find the best parameters for the model we are considering.

Ideally we want the global minimum (for high-dimensional models it is hard to know if we have local or global minimum). However, if the cost/loss function is convex the following result provides invaluable information:

Any minimum is global for convex functions. Consider the problem of finding  $x \in \mathbb{R}^n$  such that f(x) is minimal, where f is convex and differentiable. Then, any point  $x^*$  that satisfies  $\nabla f(x^*) = 0$  is a global minimum.

This result means that if we know that the cost/loss function is convex and we are able to find a minimum, we are guaranteed that it is a global minimum.

# Some simple problems

- 1. Show that  $f(x) = x^2$  is convex for  $x \in \mathbb{R}$  using the definition of convexity. Hint: If you re-write the definition, f is convex if the following holds for all  $x, y \in D_f$  and any  $\lambda \in [0, 1]$   $\lambda f(x) + (1 \lambda)f(y) f(\lambda x + (1 \lambda)y) \ge 0$ .
- 2. Using the second order condition show that the following functions are convex on the specified domain.
  - $f(x) = e^x$  is convex for  $x \in \mathbb{R}$ .
  - $g(x) = -\ln(x)$  is convex for  $x \in (0, \infty)$ .
- 3. Let  $f(x) = x^2$  and  $g(x) = e^x$ . Show that f(g(x)) and g(f(x)) is convex for  $x \in \mathbb{R}$ . Also show that if f(x) is any convex function than  $h(x) = e^{f(x)}$  is convex.
- 4. A norm is any function that satisfy the following properties
  - $f(\alpha x) = |\alpha| f(x)$  for all  $\alpha \in \mathbb{R}$ .
  - $f(x+y) \le f(x) + f(y)$
  - $f(x) \leq 0$  for all  $x \in \mathbb{R}^n$  with equality if and only if x = 0

Using the definition of convexity, try to show that a function satisfying the properties above is convex (the third condition is not needed to show this).

## Standard steepest descent

Before we proceed, we would like to discuss the approach called the **standard Steepest descent**, which again leads to us having to be able to compute a matrix. It belongs to the class of Conjugate Gradient methods (CG).

The success of the CG method for finding solutions of non-linear problems is based on the theory of conjugate gradients for linear systems of equations. It belongs to the class of iterative methods for solving problems from linear algebra of the type

$$\hat{A}\hat{x} = \hat{b}$$
.

In the iterative process we end up with a problem like

$$\hat{r} = \hat{b} - \hat{A}\hat{x},$$

where  $\hat{r}$  is the so-called residual or error in the iterative process.

When we have found the exact solution,  $\hat{r} = 0$ .

## Gradient method

The residual is zero when we reach the minimum of the quadratic equation

$$P(\hat{x}) = \frac{1}{2}\hat{x}^T \hat{A}\hat{x} - \hat{x}^T \hat{b},$$

with the constraint that the matrix  $\hat{A}$  is positive definite and symmetric. This defines also the Hessian and we want it to be positive definite.

# Steepest descent method

We denote the initial guess for  $\hat{x}$  as  $\hat{x}_0$ . We can assume without loss of generality that

$$\hat{x}_0 = 0$$
,

or consider the system

$$\hat{A}\hat{z} = \hat{b} - \hat{A}\hat{x}_0,$$

instead.

# Steepest descent method

One can show that the solution  $\hat{x}$  is also the unique minimizer of the quadratic form

$$f(\hat{x}) = \frac{1}{2}\hat{x}^T \hat{A}\hat{x} - \hat{x}^T \hat{x}, \quad \hat{x} \in \mathbf{R}^n.$$

This suggests taking the first basis vector  $\hat{r}_1$  (see below for definition) to be the gradient of f at  $\hat{x} = \hat{x}_0$ , which equals

$$\hat{A}\hat{x}_0 - \hat{b},$$

and  $\hat{x}_0 = 0$  it is equal  $-\hat{b}$ .

## Final expressions

We can compute the residual iteratively as

$$\hat{r}_{k+1} = \hat{b} - \hat{A}\hat{x}_{k+1}$$

which equals

$$\hat{b} - \hat{A}(\hat{x}_k + \alpha_k \hat{r}_k),$$

or

$$(\hat{b} - \hat{A}\hat{x}_k) - \alpha_k \hat{A}\hat{r}_k,$$

which gives

$$\alpha_k = \frac{\hat{r}_k^T \hat{r}_k}{\hat{r}_k^T \hat{A} \hat{r}_k}$$

leading to the iterative scheme

$$\hat{x}_{k+1} = \hat{x}_k - \alpha_k \hat{r}_k,$$

# Broyden-Fletcher-Goldfarb-Shanno algorithm

The optimization problem is to minimize  $f(\mathbf{x})$  where  $\mathbf{x}$  is a vector in  $\mathbb{R}^n$ , and f is a differentiable scalar function. There are no constraints on the values that  $\mathbf{x}$  can take.

The algorithm begins at an initial estimate for the optimal value  $\mathbf{x}_0$  and proceeds iteratively to get a better estimate at each stage.

The search direction  $p_k$  at stage k is given by the solution of the analogue of the Newton equation

$$B_k \mathbf{p}_k = -\nabla f(\mathbf{x}_k),$$

where  $B_k$  is an approximation to the Hessian matrix, which is updated iteratively at each stage, and  $\nabla f(\mathbf{x}_k)$  is the gradient of the function evaluated at  $x_k$ . A line search in the direction  $p_k$  is then used to find the next point  $x_{k+1}$  by minimising

$$f(\mathbf{x}_k + \alpha \mathbf{p}_k),$$

over the scalar  $\alpha > 0$ .

## Using gradient descent methods, limitations

- Gradient descent (GD) finds local minima of our function. Since the GD algorithm is deterministic, if it converges, it will converge to a local minimum of our cost/loss/risk function. Because in ML we are often dealing with extremely rugged landscapes with many local minima, this can lead to poor performance.
- GD is sensitive to initial conditions. One consequence of the local nature of GD is that initial conditions matter. Depending on where one starts, one will end up at a different local minima. Therefore, it is very important to think about how one initializes the training process. This is true for GD as well as more complicated variants of GD.
- Gradients are computationally expensive to calculate for large datasets. In many cases in statistics and ML, the cost/loss/risk function is a sum of terms, with one term for each data point. For example, in linear regression,  $E \propto \sum_{i=1}^{n} (y_i \mathbf{w}^T \cdot \mathbf{x}_i)^2$ ; for logistic regression, the square error is replaced by the cross entropy. To calculate the gradient we have to sum over all n data points. Doing this at every GD step becomes extremely computationally expensive. An ingenious solution to this, is to calculate the gradients using small subsets of the data called "mini batches". This has the added benefit of introducing stochasticity into our algorithm.
- GD is very sensitive to choices of learning rates. GD is extremely sensitive to the choice of learning rates. If the learning rate is very small, the training process take an extremely long time. For larger learning rates, GD can diverge and give poor results. Furthermore, depending on what the local landscape looks like, we have to modify the learning rates to

ensure convergence. Ideally, we would *adaptively* choose the learning rates to match the landscape.

- GD treats all directions in parameter space uniformly. Another major drawback of GD is that unlike Newton's method, the learning rate for GD is the same in all directions in parameter space. For this reason, the maximum learning rate is set by the behavior of the steepest direction and this can significantly slow down training. Ideally, we would like to take large steps in flat directions and small steps in steep directions. Since we are exploring rugged landscapes where curvatures change, this requires us to keep track of not only the gradient but second derivatives. The ideal scenario would be to calculate the Hessian but this proves to be too computationally expensive.
- GD can take exponential time to escape saddle points, even with random initialization. As we mentioned, GD is extremely sensitive to initial condition since it determines the particular local minimum GD would eventually reach. However, even with a good initialization scheme, through the introduction of randomness, GD can still take exponential time to escape saddle points.

## Improving gradient descent with momentum

We discuss here some simple examples where we introduce what is called 'memory'about previous steps, or what is normally called momentum gradient descent. The mathematics is explained below in connection with Stochastic gradient descent.

```
from numpy import asarray
from numpy import arange
from numpy.random import rand
from numpy.random import seed
from matplotlib import pyplot
# objective function
def objective(x):
        return x**2.0
# derivative of objective function
def derivative(x):
        return x * 2.0
# gradient descent algorithm
def gradient descent(objective, derivative, bounds, n iter, step size):
        # track all solutions
        solutions, scores = list(), list()
        # generate an initial point
solution = bounds[:, 0] + rand(len(bounds)) * (bounds[:, 1] - bounds[:, 0])
        # run the gradient descent
        for i in range(n_iter):
                 # calculate gradient
                 gradient = derivative(solution)
```

```
# take a step
                 solution = solution - step_size * gradient
                 # evaluate candidate point
                 solution_eval = objective(solution)
                 # store solution
                 solutions.append(solution)
                 scores.append(solution_eval)
                 # report progress
print('>%d f(%s) = %.5f' % (i, solution, solution_eval))
        return [solutions, scores]
# seed the pseudo random number generator
seed(4)
# define range for input
bounds = asarray([[-1.0, 1.0]])
# define the total iterations
n_{iter} = 30
# define the step size
step_size = 0.1
# perform the gradient descent search
solutions, scores = gradient_descent(objective, derivative, bounds, n_iter, step_size)
 \textit{\# sample input range uniform} \\ \overline{l}y \textit{ at 0.1 increments} 
inputs = arange(bounds[0,0], bounds[0,1]+0.1, 0.1)
# compute targets
results = objective(inputs)
# create a line plot of input vs result
pyplot.plot(inputs, results)
# plot the solutions found
pyplot.plot(solutions, scores, '.-', color='red')
# show the plot
pyplot.show()
```

## Same code but now with momentum gradient descent

```
from numpy import asarray
from numpy import arange
from numpy.random import rand
from numpy.random import seed
from matplotlib import pyplot
# objective function
def objective(x):
        return x**2.0
# derivative of objective function
def derivative(x):
        return x * 2.0
# gradient descent algorithm
def gradient_descent(objective, derivative, bounds, n_iter, step_size, momentum):
        # track all solutions
        solutions, scores = list(), list()
        # generate an initial point
solution = bounds[:, 0] + rand(len(bounds)) * (bounds[:, 1] - bounds[:, 0])
        # keep track of the change
        change = 0.0
        # run the gradient descent
        for i in range(n_iter):
```

```
# calculate gradient
                gradient = derivative(solution)
                 # calculate update
                new_change = step_size * gradient + momentum * change
                # take a step
                solution = solution - new_change
                # save the change
                change = new_change
                # evaluate candidate point
                solution_eval = objective(solution)
                # store solution
                solutions.append(solution)
                scores.append(solution_eval)
                # report progress
print('>%d f(%s) = %.5f' % (i, solution, solution_eval))
        return [solutions, scores]
# seed the pseudo random number generator
seed(4)
# define range for input
bounds = asarray([[-1.0, 1.0]])
# define the total iterations
n_{iter} = 30
# define the step size
step_size = 0.1
# define momentum
momentum = 0.3
# perform the gradient descent search with momentum
solutions, scores = gradient_descent(objective, derivative, bounds, n_iter, step_size, momentum)
# sample input range uniformly at 0.1 increments
inputs = arange(bounds[0,0], bounds[0,1]+0.1, 0.1)
# compute targets
results = objective(inputs)
# create a line plot of input vs result
pyplot.plot(inputs, results)
# plot the solutions found
pyplot.plot(solutions, scores, '.-', color='red')
# show the plot
pyplot.show()
```

#### Overview video on Stochastic Gradient Descent

What is Stochastic Gradient Descent

#### Batches and mini-batches

In gradient descent we compute the cost function and its gradient for all data points we have.

In large-scale applications such as the ILSVRC challenge, the training data can have on order of millions of examples. Hence, it seems wasteful to compute the full cost function over the entire training set in order to perform only a single parameter update. A very common approach to addressing this challenge is to compute the gradient over batches of the training data. For example, a typical batch could contain some thousand examples from an entire training set of several millions. This batch is then used to perform a parameter update.

# Stochastic Gradient Descent (SGD)

In stochastic gradient descent, the extreme case is the case where we have only one batch, that is we include the whole data set.

This process is called Stochastic Gradient Descent (SGD) (or also sometimes on-line gradient descent). This is relatively less common to see because in practice due to vectorized code optimizations it can be computationally much more efficient to evaluate the gradient for 100 examples, than the gradient for one example 100 times. Even though SGD technically refers to using a single example at a time to evaluate the gradient, you will hear people use the term SGD even when referring to mini-batch gradient descent (i.e. mentions of MGD for "Minibatch Gradient Descent", or BGD for "Batch gradient descent" are rare to see), where it is usually assumed that mini-batches are used. The size of the mini-batch is a hyperparameter but it is not very common to cross-validate or bootstrap it. It is usually based on memory constraints (if any), or set to some value, e.g. 32, 64 or 128. We use powers of 2 in practice because many vectorized operation implementations work faster when their inputs are sized in powers of 2.

In our notes with SGD we mean stochastic gradient descent with mini-batches.

## Stochastic Gradient Descent

Stochastic gradient descent (SGD) and variants thereof address some of the shortcomings of the Gradient descent method discussed above.

The underlying idea of SGD comes from the observation that the cost function, which we want to minimize, can almost always be written as a sum over n data points  $\{\mathbf{x}_i\}_{i=1}^n$ ,

$$C(\beta) = \sum_{i=1}^{n} c_i(\mathbf{x}_i, \beta).$$

## Computation of gradients

This in turn means that the gradient can be computed as a sum over i-gradients

$$\nabla_{\beta} C(\beta) = \sum_{i}^{n} \nabla_{\beta} c_{i}(\mathbf{x}_{i}, \beta).$$

Stochasticity/randomness is introduced by only taking the gradient on a subset of the data called minibatches. If there are n data points and the size of each minibatch is M, there will be n/M minibatches. We denote these minibatches by  $B_k$  where  $k = 1, \dots, n/M$ .

## SGD example

As an example, suppose we have 10 data points  $(\mathbf{x}_1, \dots, \mathbf{x}_{10})$  and we choose to have M = 5 minibathces, then each minibatch contains two data points. In

particular we have  $B_1 = (\mathbf{x}_1, \mathbf{x}_2), \dots, B_5 = (\mathbf{x}_9, \mathbf{x}_{10})$ . Note that if you choose M = 1 you have only a single batch with all data points and on the other extreme, you may choose M = n resulting in a minibatch for each datapoint, i.e  $B_k = \mathbf{x}_k$ .

The idea is now to approximate the gradient by replacing the sum over all data points with a sum over the data points in one the minibatches picked at random in each gradient descent step

$$\nabla_{\beta} C(\beta) = \sum_{i=1}^{n} \nabla_{\beta} c_i(\mathbf{x}_i, \beta) \to \sum_{i \in B_k}^{n} \nabla_{\beta} c_i(\mathbf{x}_i, \beta).$$

## The gradient step

Thus a gradient descent step now looks like

$$\beta_{j+1} = \beta_j - \gamma_j \sum_{i \in B_k}^n \nabla_{\beta} c_i(\mathbf{x}_i, \beta)$$

where k is picked at random with equal probability from [1, n/M]. An iteration over the number of minibathces (n/M) is commonly referred to as an epoch. Thus it is typical to choose a number of epochs and for each epoch iterate over the number of minibatches, as exemplified in the code below.

# Simple example code

```
import numpy as np

n = 100 #100 datapoints
M = 5  #size of each minibatch
m = int(n/M) #number of minibatches
n_epochs = 10 #number of epochs

j = 0
for epoch in range(1,n_epochs+1):
    for i in range(m):
        k = np.random.randint(m) #Pick the k-th minibatch at random
        #Compute the gradient using the data in minibatch Bk
        #Compute new suggestion for
        j += 1
```

Taking the gradient only on a subset of the data has two important benefits. First, it introduces randomness which decreases the chance that our opmization scheme gets stuck in a local minima. Second, if the size of the minibatches are small relative to the number of datapoints (M < n), the computation of the gradient is much cheaper since we sum over the datapoints in the k-th minibatch and not all n datapoints.

## When do we stop?

A natural question is when do we stop the search for a new minimum? One possibility is to compute the full gradient after a given number of epochs and check if the norm of the gradient is smaller than some threshold and stop if true. However, the condition that the gradient is zero is valid also for local minima, so this would only tell us that we are close to a local/global minimum. However, we could also evaluate the cost function at this point, store the result and continue the search. If the test kicks in at a later stage we can compare the values of the cost function and keep the  $\beta$  that gave the lowest value.

#### Slightly different approach

Another approach is to let the step length  $\gamma_j$  depend on the number of epochs in such a way that it becomes very small after a reasonable time such that we do not move at all. Such approaches are also called scaling. There are many such ways to scale the learning rate and discussions here. See also https://

towardsdatascience.com/learning-rate-schedules-and-adaptive-learning-rate-methods-for-deep-for a discussion of different scaling functions for the learning rate.

## Time decay rate

As an example, let  $e = 0, 1, 2, 3, \cdots$  denote the current epoch and let  $t_0, t_1 > 0$  be two fixed numbers. Furthermore, let  $t = e \cdot m + i$  where m is the number of minibatches and  $i = 0, \cdots, m-1$ . Then the function

$$\gamma_j(t;t_0,t_1) = \frac{t_0}{t+t_1}$$

goes to zero as the number of epochs gets large. I.e. we start with a step length  $\gamma_j(0;t_0,t_1)=t_0/t_1$  which decays in *time t*.

In this way we can fix the number of epochs, compute  $\beta$  and evaluate the cost function at the end. Repeating the computation will give a different result since the scheme is random by design. Then we pick the final  $\beta$  that gives the lowest value of the cost function.

```
import numpy as np

def step_length(t,t0,t1):
    return t0/(t+t1)

n = 100 #100 datapoints
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
n_epochs = 500 #number of epochs
t0 = 1.0
t1 = 10

gamma_j = t0/t1
j = 0
for epoch in range(1,n_epochs+1):
```

```
for i in range(m):
    k = np.random.randint(m) #Pick the k-th minibatch at random
    #Compute the gradient using the data in minibatch Bk
    #Compute new suggestion for beta
    t = epoch*m+i
    gamma_j = step_length(t,t0,t1)
    j += 1

print("gamma_j after %d epochs: %g" % (n_epochs,gamma_j))
```

#### Code with a Number of Minibatches which varies

In the code here we vary the number of mini-batches.

```
# Importing various packages
from math import exp, sqrt
from random import random, seed
import numpy as np
import matplotlib.pyplot as plt
n = 100
x = 2*np.random.rand(n,1)
y = 4+3*x+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]
XT_X = X.T @ X
theta_linreg = np.linalg.inv(X.T @ X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Hessian matrix
H = (2.0/n) * XT_X
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
theta = np.random.randn(2,1)
eta = 1.0/np.max(EigValues)
Niterations = 1000
for iter in range(Niterations):
     gradients = 2.0/n*X.T @ ((X @ theta)-y)
theta -= eta*gradients
print("theta from own gd")
print(theta)
xnew = np.array([[0],[2]])
Xnew = np.c_[np.ones((2,1)), xnew]
ypredict = Xnew.dot(theta)
ypredict2 = Xnew.dot(theta_linreg)
n_{epochs} = 50
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
t0, t1 = 5, 50
def learning_schedule(t):
     return t\overline{0}/(t+t1)
theta = np.random.randn(2,1)
for epoch in range(n_epochs):
```

```
# Can you figure out a better way of setting up the contributions to each batch?
    for i in range(m):
        random_index = M*np.random.randint(m)
        xi = X[random_index:random_index+M]
        yi = y[random_index:random_index+M]
        gradients = (2.0/M)* xi.T @ ((xi @ theta)-yi)
        eta = learning_schedule(epoch*m+i)
        theta = theta - eta*gradients
print("theta from own sdg")
print(theta)

plt.plot(xnew, ypredict, "r-")
plt.plot(xnew, ypredict2, "b-")
plt.plot(x, y ,'ro')
plt.axis([0,2.0,0, 15.0])
plt.xlabel(r'$x$')
plt.ylabel(r'$y$')
plt.title(r'Random numbers ')
plt.show()
```

## Replace or not

In the above code, we have use replacement in setting up the mini-batches. The discussion here may be useful.

#### Momentum based GD

The stochastic gradient descent (SGD) is almost always used with a *momentum* or inertia term that serves as a memory of the direction we are moving in parameter space. This is typically implemented as follows

$$\mathbf{v}_{t} = \gamma \mathbf{v}_{t-1} + \eta_{t} \nabla_{\theta} E(\boldsymbol{\theta}_{t})$$
  
$$\boldsymbol{\theta}_{t+1} = \boldsymbol{\theta}_{t} - \mathbf{v}_{t}, \tag{26}$$

where we have introduced a momentum parameter  $\gamma$ , with  $0 \le \gamma \le 1$ , and for brevity we dropped the explicit notation to indicate the gradient is to be taken over a different mini-batch at each step. We call this algorithm gradient descent with momentum (GDM). From these equations, it is clear that  $\mathbf{v}_t$  is a running average of recently encountered gradients and  $(1-\gamma)^{-1}$  sets the characteristic time scale for the memory used in the averaging procedure. Consistent with this, when  $\gamma = 0$ , this just reduces down to ordinary SGD as discussed earlier. An equivalent way of writing the updates is

$$\Delta \boldsymbol{\theta}_{t+1} = \gamma \Delta \boldsymbol{\theta}_t - \eta_t \nabla_{\boldsymbol{\theta}} E(\boldsymbol{\theta}_t),$$

where we have defined  $\Delta \theta_t = \theta_t - \theta_{t-1}$ .

## More on momentum based approaches

Let us try to get more intuition from these equations. It is helpful to consider a simple physical analogy with a particle of mass m moving in a viscous medium with drag coefficient  $\mu$  and potential  $E(\mathbf{w})$ . If we denote the particle's position by  $\mathbf{w}$ , then its motion is described by

$$m\frac{d^2\mathbf{w}}{dt^2} + \mu \frac{d\mathbf{w}}{dt} = -\nabla_w E(\mathbf{w}).$$

We can discretize this equation in the usual way to get

$$m\frac{\mathbf{w}_{t+\Delta t} - 2\mathbf{w}_t + \mathbf{w}_{t-\Delta t}}{(\Delta t)^2} + \mu \frac{\mathbf{w}_{t+\Delta t} - \mathbf{w}_t}{\Delta t} = -\nabla_w E(\mathbf{w}).$$

Rearranging this equation, we can rewrite this as

$$\Delta \mathbf{w}_{t+\Delta t} = -\frac{(\Delta t)^2}{m + \mu \Delta t} \nabla_w E(\mathbf{w}) + \frac{m}{m + \mu \Delta t} \Delta \mathbf{w}_t.$$

#### Momentum parameter

Notice that this equation is identical to previous one if we identify the position of the particle,  $\mathbf{w}$ , with the parameters  $\boldsymbol{\theta}$ . This allows us to identify the momentum parameter and learning rate with the mass of the particle and the viscous drag as:

$$\gamma = \frac{m}{m + \mu \Delta t}, \qquad \eta = \frac{(\Delta t)^2}{m + \mu \Delta t}.$$

Thus, as the name suggests, the momentum parameter is proportional to the mass of the particle and effectively provides inertia. Furthermore, in the large viscosity/small learning rate limit, our memory time scales as  $(1-\gamma)^{-1} \approx m/(\mu \Delta t)$ .

Why is momentum useful? SGD momentum helps the gradient descent algorithm gain speed in directions with persistent but small gradients even in the presence of stochasticity, while suppressing oscillations in high-curvature directions. This becomes especially important in situations where the landscape is shallow and flat in some directions and narrow and steep in others. It has been argued that first-order methods (with appropriate initial conditions) can perform comparable to more expensive second order methods, especially in the context of complex deep learning models.

These beneficial properties of momentum can sometimes become even more pronounced by using a slight modification of the classical momentum algorithm called Nesterov Accelerated Gradient (NAG).

In the NAG algorithm, rather than calculating the gradient at the current parameters,  $\nabla_{\theta} E(\boldsymbol{\theta}_t)$ , one calculates the gradient at the expected value of the parameters given our current momentum,  $\nabla_{\theta} E(\boldsymbol{\theta}_t + \gamma \mathbf{v}_{t-1})$ . This yields the NAG update rule

$$\mathbf{v}_{t} = \gamma \mathbf{v}_{t-1} + \eta_{t} \nabla_{\theta} E(\boldsymbol{\theta}_{t} + \gamma \mathbf{v}_{t-1})$$
  
$$\boldsymbol{\theta}_{t+1} = \boldsymbol{\theta}_{t} - \mathbf{v}_{t}.$$
 (27)

One of the major advantages of NAG is that it allows for the use of a larger learning rate than GDM for the same choice of  $\gamma$ .

## Second moment of the gradient

In stochastic gradient descent, with and without momentum, we still have to specify a schedule for tuning the learning rates  $\eta_t$  as a function of time. As discussed in the context of Newton's method, this presents a number of dilemmas. The learning rate is limited by the steepest direction which can change depending on the current position in the landscape. To circumvent this problem, ideally our algorithm would keep track of curvature and take large steps in shallow, flat directions and small steps in steep, narrow directions. Second-order methods accomplish this by calculating or approximating the Hessian and normalizing the learning rate by the curvature. However, this is very computationally expensive for extremely large models. Ideally, we would like to be able to adaptively change the step size to match the landscape without paying the steep computational price of calculating or approximating Hessians.

Recently, a number of methods have been introduced that accomplish this by tracking not only the gradient, but also the second moment of the gradient. These methods include AdaGrad, AdaDelta, Root Mean Squared Propagation (RMS-Prop), and ADAM.

#### RMS prop

In RMS prop, in addition to keeping a running average of the first moment of the gradient, we also keep track of the second moment denoted by  $\mathbf{s}_t = \mathbb{E}[\mathbf{g}_t^2]$ . The update rule for RMS prop is given by

$$\mathbf{g}_{t} = \nabla_{\theta} E(\boldsymbol{\theta})$$

$$\mathbf{s}_{t} = \beta \mathbf{s}_{t-1} + (1 - \beta) \mathbf{g}_{t}^{2}$$

$$\boldsymbol{\theta}_{t+1} = \boldsymbol{\theta}_{t} - \eta_{t} \frac{\mathbf{g}_{t}}{\sqrt{\mathbf{s}_{t} + \epsilon}},$$

$$(28)$$

where  $\beta$  controls the averaging time of the second moment and is typically taken to be about  $\beta=0.9,\ \eta_t$  is a learning rate typically chosen to be  $10^{-3}$ , and  $\epsilon\sim 10^{-8}$  is a small regularization constant to prevent divergences. Multiplication and division by vectors is understood as an element-wise operation. It is clear from this formula that the learning rate is reduced in directions where the norm of the gradient is consistently large. This greatly speeds up the convergence by allowing us to use a larger learning rate for flat directions.

## ADAM optimizer

A related algorithm is the ADAM optimizer. In ADAM, we keep a running average of both the first and second moment of the gradient and use this information to adaptively change the learning rate for different parameters. The method isefficient when working with large problems involving lots data and/or parameters. It is a combination of the gradient descent with momentum algorithm and the RMSprop algorithm discussed above.

In addition to keeping a running average of the first and second moments of the gradient (i.e.  $\mathbf{m}_t = \mathbb{E}[\mathbf{g}_t]$  and  $\mathbf{s}_t = \mathbb{E}[\mathbf{g}_t^2]$ , respectively), ADAM performs an additional bias correction to account for the fact that we are estimating the first two moments of the gradient using a running average (denoted by the hats in the update rule below). The update rule for ADAM is given by (where multiplication and division are once again understood to be element-wise operations below)

$$\mathbf{g}_{t} = \nabla_{\theta} E(\boldsymbol{\theta})$$

$$\mathbf{m}_{t} = \beta_{1} \mathbf{m}_{t-1} + (1 - \beta_{1}) \mathbf{g}_{t}$$

$$\mathbf{s}_{t} = \beta_{2} \mathbf{s}_{t-1} + (1 - \beta_{2}) \mathbf{g}_{t}^{2}$$

$$\mathbf{m}_{t} = \frac{\mathbf{m}_{t}}{1 - \beta_{1}^{t}}$$

$$\mathbf{s}_{t} = \frac{\mathbf{s}_{t}}{1 - \beta_{2}^{t}}$$

$$\boldsymbol{\theta}_{t+1} = \boldsymbol{\theta}_{t} - \eta_{t} \frac{\mathbf{m}_{t}}{\sqrt{\mathbf{s}_{t}} + \epsilon},$$

$$(30)$$

where  $\beta_1$  and  $\beta_2$  set the memory lifetime of the first and second moment and are typically taken to be 0.9 and 0.99 respectively, and  $\eta$  and  $\epsilon$  are identical to RMSprop.

Like in RMSprop, the effective step size of a parameter depends on the magnitude of its gradient squared. To understand this better, let us rewrite this expression in terms of the variance  $\sigma_t^2 = \mathbf{s}_t - (\mathbf{m}_t)^2$ . Consider a single parameter  $\theta_t$ . The update rule for this parameter is given by

$$\Delta\theta_{t+1} = -\eta_t \frac{\boldsymbol{m}_t}{\sqrt{\sigma_t^2 + m_t^2 + \epsilon}}.$$

## Algorithms and codes for Adagrad, RMSprop and Adam

The algorithms we have implemented are well described in the text by Goodfellow, Bengio and Courville, chapter 8.

The codes which implement these algorithms are discussed after our presentation of automatic differentiation.

## Practical tips

- Randomize the data when making mini-batches. It is always important to randomly shuffle the data when forming mini-batches. Otherwise, the gradient descent method can fit spurious correlations resulting from the order in which data is presented.
- Transform your inputs. Learning becomes difficult when our landscape has a mixture of steep and flat directions. One simple trick for minimizing these situations is to standardize the data by subtracting the mean and normalizing the variance of input variables. Whenever possible, also decorrelate the inputs. To understand why this is helpful, consider the case of linear regression. It is easy to show that for the squared error cost function, the Hessian of the cost function is just the correlation matrix between the inputs. Thus, by standardizing the inputs, we are ensuring that the landscape looks homogeneous in all directions in parameter space. Since most deep networks can be viewed as linear transformations followed by a non-linearity at each layer, we expect this intuition to hold beyond the linear case.
- Monitor the out-of-sample performance. Always monitor the performance of your model on a validation set (a small portion of the training data that is held out of the training process to serve as a proxy for the test set. If the validation error starts increasing, then the model is beginning to overfit. Terminate the learning process. This *early stopping* significantly improves performance in many settings.
- Adaptive optimization methods don't always have good generalization. Recent studies have shown that adaptive methods such as ADAM, RMSPorp, and AdaGrad tend to have poor generalization compared to SGD or SGD with momentum, particularly in the high-dimensional limit (i.e. the number of parameters exceeds the number of data points). Although it is not clear at this stage why these methods perform so well in training deep neural networks, simpler procedures like properly-tuned SGD may work as well or better in these applications.

Geron's text, see chapter 11, has several interesting discussions.

#### Automatic differentiation

Automatic differentiation (AD), also called algorithmic differentiation or computational differentiation, is a set of techniques to numerically evaluate the derivative of a function specified by a computer program. AD exploits the fact that every computer program, no matter how complicated, executes a sequence of elementary arithmetic operations (addition, subtraction, multiplication, division, etc.) and elementary functions (exp, log, sin, cos, etc.). By applying the chain rule repeatedly to these operations, derivatives of arbitrary order can be

computed automatically, accurately to working precision, and using at most a small constant factor more arithmetic operations than the original program.

Automatic differentiation is neither:

- Symbolic differentiation, nor
- Numerical differentiation (the method of finite differences).

Symbolic differentiation can lead to inefficient code and faces the difficulty of converting a computer program into a single expression, while numerical differentiation can introduce round-off errors in the discretization process and cancellation

Python has tools for so-called **automatic differentiation**. Consider the following example

$$f(x) = \sin\left(2\pi x + x^2\right)$$

which has the following derivative

$$f'(x) = \cos(2\pi x + x^2)(2\pi + 2x)$$

Using autograd we have

```
import autograd.numpy as np
# To do elementwise differentiation:
from autograd import elementwise_grad as egrad
# To plot:
import matplotlib.pyplot as plt
def f(x):
    return np.sin(2*np.pi*x + x**2)
def f_grad_analytic(x):
    return np.cos(2*np.pi*x + x**2)*(2*np.pi + 2*x)
# Do the comparison:
x = np.linspace(0,1,1000)
f_grad = egrad(f)
computed = f_grad(x)
analytic = f_grad_analytic(x)
plt.title('Derivative computed from Autograd compared with the analytical derivative')
plt.plot(x,computed,label='autograd')
plt.plot(x,analytic,label='analytic')
plt.xlabel('x')
plt.ylabel('y')
plt.legend()
plt.show()
print("The max absolute difference is: %g"%(np.max(np.abs(computed - analytic))))
```

## Using autograd

Here we experiment with what kind of functions Autograd is capable of finding the gradient of. The following Python functions are just meant to illustrate what Autograd can do, but please feel free to experiment with other, possibly more complicated, functions as well.

```
import autograd.numpy as np
from autograd import grad

def f1(x):
    return x**3 + 1

f1_grad = grad(f1)

# Remember to send in float as argument to the computed gradient from Autograd!
a = 1.0

# See the evaluated gradient at a using autograd:
print("The gradient of f1 evaluated at a = %g using autograd is: %g"%(a,f1_grad(a)))

# Compare with the analytical derivative, that is f1'(x) = 3*x**2
grad_analytical = 3*a**2
print("The gradient of f1 evaluated at a = %g by finding the analytic expression is: %g"%(a,grad_i)
```

## Autograd with more complicated functions

To differentiate with respect to two (or more) arguments of a Python function, Autograd need to know at which variable the function if being differentiated with respect to.

```
import autograd.numpy as np
from autograd import grad
def f2(x1,x2):
    return 3*x1**3 + x2*(x1 - 5) + 1
# By sending the argument 0, Autograd will compute the derivative w.r.t the first variable, in th
f2_grad_x1 = grad(f2,0)
# ... and differentiate w.r.t x2 by sending 1 as an additional arugment to grad
f2_grad_x2 = grad(f2,1)
x1 = 1.0
x2 = 3.0
print("Evaluating at x1 = %g, x2 = %g"%(x1,x2))
print("-"*30)
# Compare with the analytical derivatives:
# Derivative of f2 w.r.t x1 is: 9*x1**2 + x2:
f2_grad_x1_analytical = 9*x1**2 + x2
# Derivative of f2 w.r.t x2 is: x1 - 5:
f2_grad_x2_analytical = x1 - 5
# See the evaluated derivations:
```

```
print("The derivative of f2 w.r.t x1: %g"%( f2_grad_x1(x1,x2) ))
print("The analytical derivative of f2 w.r.t x1: %g"%( f2_grad_x1(x1,x2) ))
print()
print("The derivative of f2 w.r.t x2: %g"%( f2_grad_x2(x1,x2) ))
print("The analytical derivative of f2 w.r.t x2: %g"%( f2_grad_x2(x1,x2) ))
```

Note that the grad function will not produce the true gradient of the function. The true gradient of a function with two or more variables will produce a vector, where each element is the function differentiated w.r.t a variable.

# More complicated functions using the elements of their arguments directly

```
import autograd.numpy as np
from autograd import grad
def f3(x): # Assumes x is an array of length 5 or higher
    return 2*x[0] + 3*x[1] + 5*x[2] + 7*x[3] + 11*x[4]**2

f3_grad = grad(f3)

x = np.linspace(0,4,5)

# Print the computed gradient:
print("The computed gradient of f3 is: ", f3_grad(x))

# The analytical gradient is: (2, 3, 5, 7, 22*x[4])
f3_grad_analytical = np.array([2, 3, 5, 7, 22*x[4]])

# Print the analytical gradient:
print("The analytical gradient of f3 is: ", f3_grad_analytical)
```

Note that in this case, when sending an array as input argument, the output from Autograd is another array. This is the true gradient of the function, as opposed to the function in the previous example. By using arrays to represent the variables, the output from Autograd might be easier to work with, as the output is closer to what one could expect form a gradient-evaluting function.

#### Functions using mathematical functions from Numpy

```
import autograd.numpy as np
from autograd import grad
def f4(x):
    return np.sqrt(1+x**2) + np.exp(x) + np.sin(2*np.pi*x)

f4_grad = grad(f4)

x = 2.7

# Print the computed derivative:
print("The computed derivative of f4 at x = %g is: %g"%(x,f4_grad(x)))

# The analytical derivative is: x/sqrt(1 + x**2) + exp(x) + cos(2*pi*x)*2*pi
```

```
f4_grad_analytical = x/np.sqrt(1 + x**2) + np.exp(x) + np.cos(2*np.pi*x)*2*np.pi
# Print the analytical gradient:
print("The analytical gradient of f4 at x = %g is: %g"%(x,f4_grad_analytical))
```

#### More autograd

```
import autograd.numpy as np
from autograd import grad
def f5(x):
    if x >= 0:
        return x**2
    else:
        return -3*x + 1

f5_grad = grad(f5)

x = 2.7

# Print the computed derivative:
print("The computed derivative of f5 at x = %g is: %g"%(x,f5_grad(x)))
```

#### And with loops

```
import autograd.numpy as np
from autograd import grad
def f6_for(x):
    val = 0
    for i in range(10):
         val = val + x**i
    return val
def f6_while(x):
    val = 0
    i = 0
    while i < 10:
         val = val + x**i
         i = i + 1
    return val
f6_for_grad = grad(f6_for)
f6_while_grad = grad(f6_while)
x = 0.5
\# Print the computed derivaties of f6_for and f6_while
print("The computed derivative of f6_for at x = %g is: %g"%(x,f6_for_grad(x)))
print("The computed derivative of f6_while at x = %g is: %g"%(x,f6_while_grad(x)))
import autograd.numpy as np
from autograd import grad
# Both of the functions are implementation of the sum: sum(x**i) for i = 0, \ldots, 9
# The analytical derivative is: sum(i*x**(i-1))
f6_grad_analytical = 0
for i in range(10):
    f6_grad_analytical += i*x**(i-1)
```

```
print("The analytical derivative of f6 at x = %g is: %g"%(x,f6_grad_analytical))
```

## Using recursion

```
import autograd.numpy as np
from autograd import grad
def f7(n): # Assume that n is an integer
    if n == 1 or n == 0:
       return 1
    else:
        return n*f7(n-1)
f7_grad = grad(f7)
n = 2.0
print("The computed derivative of f7 at n = %d is: %g"%(n,f7_grad(n)))
# The function f7 is an implementation of the factorial of n.
# By using the product rule, one can find that the derivative is:
f7_grad_analytical = 0
for i in range(int(n)-1):
    tmp = 1
    for k in range(int(n)-1):
       if k != \check{i}:
           tmp *= (n - k)
    f7_grad_analytical += tmp
print("The analytical derivative of f7 at n = %d is: %g"%(n,f7_grad_analytical))
```

Note that if n is equal to zero or one, Autograd will give an error message. This message appears when the output is independent on input.

## Unsupported functions

Autograd supports many features. However, there are some functions that is not supported (yet) by Autograd.

Assigning a value to the variable being differentiated with respect to

```
import autograd.numpy as np
from autograd import grad
def f8(x): # Assume x is an array
    x[2] = 3
    return x*2

f8_grad = grad(f8)

x = 8.4

print("The derivative of f8 is:",f8_grad(x))
```

Here, Autograd tells us that an 'ArrayBox' does not support item assignment. The item assignment is done when the program tries to assign x[2] to the value 3. However, Autograd has implemented the computation of the derivative such that this assignment is not possible.

## The syntax a.dot(b) when finding the dot product

```
import autograd.numpy as np
from autograd import grad
def f9(a): # Assume a is an array with 2 elements
   b = np.array([1.0,2.0])
   return a.dot(b)

f9_grad = grad(f9)

x = np.array([1.0,0.0])

print("The derivative of f9 is:",f9_grad(x))
```

Here we are told that the 'dot' function does not belong to Autograd's version of a Numpy array. To overcome this, an alternative syntax which also computed the dot product can be used:

```
import autograd.numpy as np
from autograd import grad
def f9_alternative(x): # Assume a is an array with 2 elements
    b = np.array([1.0,2.0])
    return np.dot(x,b) # The same as x_1*b_1 + x_2*b_2

f9_alternative_grad = grad(f9_alternative)

x = np.array([3.0,0.0])

print("The gradient of f9 is:",f9_alternative_grad(x))

# The analytical gradient of the dot product of vectors x and b with two elements (x_1,x_2) and (
# w.r.t x is (b_1, b_2).
```

#### Recommended to avoid

The documentation recommends to avoid inplace operations such as

```
a += b
a -= b
a*= b
a /=b
```

## Using Autograd with OLS

We conclude the part on optmization by showing how we can make codes for linear regression and logistic regression using **autograd**. The first example shows results with ordinary leats squares.

```
# Using Autograd to calculate gradients for OLS
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
def CostOLS(beta):
    return (1.0/n)*np.sum((y-X @ beta)**2)
x = 2*np.random.rand(n,1)
y = 4+3*x+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]
XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Hessian matrix
H = (2.0/n) * XT_X
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
theta = np.random.randn(2,1)
eta = 1.0/np.max(EigValues)
Niterations = 1000
# define the gradient
training_gradient = grad(CostOLS)
for iter in range(Niterations):
    gradients = training_gradient(theta)
    theta -= eta*gradients
print("theta from own gd")
print(theta)
xnew = np.array([[0],[2]])
Xnew = np.c_[np.ones((2,1)), xnew]
ypredict = Xnew.dot(theta)
ypredict2 = Xnew.dot(theta_linreg)
plt.plot(xnew, ypredict, "r-")
plt.plot(xnew, ypredict2, "b-")
plt.plot(x, y, 'ro')
plt.axis([0,2.0,0, 15.0])
plt.xlabel(r'$x$')
plt.ylabel(r'$y$')
plt.title(r'Random numbers ')
plt.show()
```

## Same code but now with momentum gradient descent

```
# Using Autograd to calculate gradients for OLS
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
```

```
def CostOLS(beta):
    return (1.0/n)*np.sum((y-X @ beta)**2)
n = 100
x = 2*np.random.rand(n,1)
y = 4+3*x#+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]
XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Hessian matrix
H = (2.0/n) * XT_X
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
theta = np.random.randn(2,1)
eta = 1.0/np.max(EigValues)
Niterations = 30
# define the gradient
training_gradient = grad(CostOLS)
for iter in range(Niterations):
    gradients = training_gradient(theta)
    theta -= eta*gradients
    print(iter,gradients[0],gradients[1])
print("theta from own gd")
print(theta)
# Now improve with momentum gradient descent
change = 0.0
delta_momentum = 0.3
for iter in range(Niterations):
    # calculate gradient
    gradients = training_gradient(theta)
    # calculate update
    new_change = eta*gradients+delta_momentum*change
    # take a step
    theta -= new_change
    # save the change
    change = new_change
    print(iter,gradients[0],gradients[1])
print("theta from own gd wth momentum")
print(theta)
```

# But noen of these can compete with Newton's method

```
# Using Newton's method
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
def CostOLS(beta):
```

```
return (1.0/n)*np.sum((y-X @ beta)**2)
n = 100
x = 2*np.random.rand(n,1)
y = 4+3*x+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]
XT_X = X.T @ X
beta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(beta_linreg)
# Hessian matrix
H = (2.0/n) * XT_X
# Note that here the Hessian does not depend on the parameters beta
invH = np.linalg.pinv(H)
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
beta = np.random.randn(2,1)
Niterations = 5
# define the gradient
training_gradient = grad(CostOLS)
for iter in range(Niterations):
    gradients = training_gradient(beta)
    beta -= invH @ gradients
    print(iter,gradients[0],gradients[1])
print("beta from own Newton code")
print(beta)
```

## Including Stochastic Gradient Descent with Autograd

In this code we include the stochastic gradient descent approach discussed above. Note here that we specify which argument we are taking the derivative with respect to when using **autograd**.

```
# Using Autograd to calculate gradients using SGD
# OLS example
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
# Note change from previous example
def CostOLS(y,X,theta):
    return np.sum((y-X @ theta)**2)
n = 100
x = 2*np.random.rand(n,1)
y = 4+3*x+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]

XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
```

```
# Hessian matrix
H = (2.0/n) * XT_X
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
theta = np.random.randn(2,1)
eta = 1.0/np.max(EigValues)
Niterations = 1000
# Note that we request the derivative wrt third argument (theta, 2 here)
training_gradient = grad(CostOLS, 2)
for iter in range(Niterations):
    gradients = (1.0/n)*training_gradient(y, X, theta)
    theta -= eta*gradients
print("theta from own gd")
print(theta)
xnew = np.array([[0],[2]])
Xnew = np.c_[np.ones((2,1)), xnew]
ypredict = Xnew.dot(theta)
ypredict2 = Xnew.dot(theta_linreg)
plt.plot(xnew, ypredict, "r-")
plt.plot(xnew, ypredict2, "b-")
plt.plot(x, y, 'ro')
plt.axis([0,2.0,0, 15.0])
plt.xlabel(r'$x$')
plt.ylabel(r'$y$')
plt.title(r'Random numbers ')
plt.show()
n_{epochs} = 50
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
t0, t1 = 5, 50
def learning_schedule(t):
    return t0/(t+t1)
theta = np.random.randn(2,1)
for epoch in range(n_epochs):
# Can you figure out a better way of setting up the contributions to each batch?
    for i in range(m):
        random_index = M*np.random.randint(m)
        xi = X[random_index:random_index+M]
        yi = y[random_index:random_index+M]
         gradients = (1.0/M)*training_gradient(yi, xi, theta)
        eta = learning_schedule(epoch*m+i)
        \verb| theta = theta - eta*gradients| \\
print("theta from own sdg")
print(theta)
```

#### Same code but now with momentum gradient descent

```
\mbox{\# Using Autograd to calculate gradients using SGD} \mbox{\# OLS example}
```

```
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
# Note change from previous example
def CostOLS(y,X,theta):
    return np.sum((y-X @ theta)**2)
n = 100
x = 2*np.random.rand(n,1)
y = 4+3*x+np.random.randn(n,1)
X = np.c_[np.ones((n,1)), x]
XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Hessian matrix
H = (2.0/n) * XT_X
EigValues, EigVectors = np.linalg.eig(H)
print(f"Eigenvalues of Hessian Matrix:{EigValues}")
theta = np.random.randn(2,1)
eta = 1.0/np.max(EigValues)
Niterations = 100
# Note that we request the derivative wrt third argument (theta, 2 here)
training_gradient = grad(CostOLS,2)
for iter in range(Niterations):
    gradients = (1.0/n)*training_gradient(y, X, theta)
    theta -= eta*gradients
print("theta from own gd")
print(theta)
n_{epochs} = 50
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
t0, t1 = 5, 50
def learning_schedule(t):
   return t0/(t+t1)
theta = np.random.randn(2,1)
change = 0.0
delta_momentum = 0.3
for epoch in range(n_epochs):
    for i in range(m):
        random_index = M*np.random.randint(m)
        xi = X[random_index:random_index+M]
        yi = y[random_index:random_index+M]
        gradients = (1.0/M)*training_gradient(yi, xi, theta)
        eta = learning_schedule(epoch*m+i)
        # calculate update
        new_change = eta*gradients+delta_momentum*change
        # take a step
        theta -= new_change
```

```
# save the change
      change = new_change
print("theta from own sdg with momentum")
print(theta)
```

# Similar (second order function now) problem but now with AdaGrad

```
# Using Autograd to calculate gradients using AdaGrad and Stochastic Gradient descent
# OLS example
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
# Note change from previous example
def CostOLS(y,X,theta):
    return np.sum((y-X @ theta)**2)
n = 10000
x = np.random.rand(n,1)
y = 2.0+3*x +4*x*x# +np.random.randn(n, 1)
X = np.c_[np.ones((n,1)), x, x*x]
XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Note that we request the derivative wrt third argument (theta, 2 here)
training_gradient = grad(CostOLS,2)
# Define parameters for Stochastic Gradient Descent
n_epochs = 50
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
# Guess for unknown parameters theta
theta = np.random.randn(3,1)
# Value for learning rate
eta = 0.01
# Including AdaGrad parameter to avoid possible division by zero
delta = 1e-8
for epoch in range(n_epochs):
    # The outer product is calculated from scratch for each epoch
    Giter = np.zeros(shape=(3,3))
    for i in range(m):
        random_index = M*np.random.randint(m)
        xi = X[random_index:random_index+M]
        yi = y[random_index:random_index+M]
gradients = (1.0/M)*training_gradient(yi, xi, theta)
        # Calculate the outer product of the gradients
        Giter +=gradients @ gradients.T
        # Simpler algorithm with only diagonal elements
        Ginverse = np.c_[eta/(delta+np.sqrt(np.diagonal(Giter)))]
        # compute update
        update = np.multiply(Ginverse,gradients)
```

```
theta -= update
print("theta from own AdaGrad")
print(theta)
```

Running this code we note an almost perfect agreement with the results from matrix inversion.

# RMSprop for adaptive learning rate with Stochastic Gradient Descent

```
# Using Autograd to calculate gradients using RMSprop and Stochastic Gradient descent
# OLS example
from random import random, seed
import numpy as np
import autograd.numpy as np
import matplotlib.pyplot as plt
from autograd import grad
# Note change from previous example
def CostOLS(y,X,theta):
    return np.sum((y-X @ theta)**2)
x = np.random.rand(n,1)
y = 2.0+3*x +4*x*x# +np.random.randn(n, 1)
X = np.c_[np.ones((n,1)), x, x*x]
XT_X = X.T @ X
theta_linreg = np.linalg.pinv(XT_X) @ (X.T @ y)
print("Own inversion")
print(theta_linreg)
# Note that we request the derivative wrt third argument (theta, 2 here)
training_gradient = grad(CostOLS,2)
# Define parameters for Stochastic Gradient Descent
n_{epochs} = 50
M = 5 #size of each minibatch
m = int(n/M) #number of minibatches
# Guess for unknown parameters theta
theta = np.random.randn(3,1)
# Value for learning rate
eta = 0.01
# Value for parameter rho
rho = 0.99
# Including AdaGrad parameter to avoid possible division by zero
delta = 1e-8
for epoch in range(n_epochs):
    Giter = np.zeros(shape=(3,3))
    for i in range(m):
        random_index = M*np.random.randint(m)
        xi = X[random_index:random_index+M]
        yi = y[random_index:random_index+M]
        gradients = (1.0/M)*training_gradient(yi, xi, theta)
        # Previous value for the outer product of gradients
        Previous = Giter
        # Accumulated gradient
```

```
Giter +=gradients @ gradients.T
    # Scaling with rho the new and the previous results
Gnew = (rho*Previous+(1-rho)*Giter)
    # Taking the diagonal only and inverting
Ginverse = np.c_[eta/(delta+np.sqrt(np.diagonal(Gnew)))]
    # Hadamard product
    update = np.multiply(Ginverse,gradients)
    theta -= update
print("theta from own RMSprop")
print(theta)
```

## And Logistic Regression

```
import autograd.numpy as np
from autograd import grad
def sigmoid(x):
    return 0.5 * (np.tanh(x / 2.) + 1)
def logistic_predictions(weights, inputs):
     # Outputs probability of a label being true according to logistic model.
    return sigmoid(np.dot(inputs, weights))
def training_loss(weights):
    # Training loss is the negative log-likelihood of the training labels.
    preds = logistic_predictions(weights, inputs)
    label_probabilities = preds * targets + (1 - preds) * (1 - targets)
    return -np.sum(np.log(label_probabilities))
# Build a toy dataset.
[0.52, 0.06, -1.30],
[0.74, -2.49, 1.39]])
targets = np.array([True, True, False, True])
{\it \# Define \ a \ function \ that \ returns \ gradients \ of \ training \ loss \ using \ Autograd}.
training_gradient_fun = grad(training_loss)
# Optimize weights using gradient descent.
weights = np.array([0.0, 0.0, 0.0])
print("Initial loss:", training_loss(weights))
for i in range(100):
    weights -= training_gradient_fun(weights) * 0.01
print("Trained loss:", training_loss(weights))
```

## Introducing JAX

Presently, instead of using **autograd**, we recommend using JAX

**JAX** is Autograd and XLA (Accelerated Linear Algebra)), brought together for high-performance numerical computing and machine learning research. It provides composable transformations of Python+NumPy programs: differentiate, vectorize, parallelize, Just-In-Time compile to GPU/TPU, and more.

Here's a simple example on how you can use  $\mathbf{JAX}$  to compute the derivate of the logistic function.

```
import jax.numpy as jnp
from jax import grad, jit, vmap

def sum_logistic(x):
   return jnp.sum(1.0 / (1.0 + jnp.exp(-x)))

x_small = jnp.arange(3.)
derivative_fn = grad(sum_logistic)
print(derivative_fn(x_small))
```