



Rethinking the origin of neutrino masses: the role of gravity

Lena Funcke

in collaboration with Gia Dvali *et al.*
(1602.03191, 1811.01991, 1905.01264)

CoEPP Webinar, 25 September 2019

Question: Origin of Small Neutrino Masses?

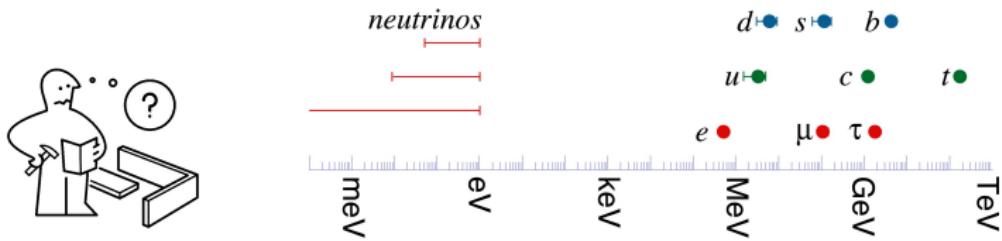


Image credits: IKEA and Murayama (2018).

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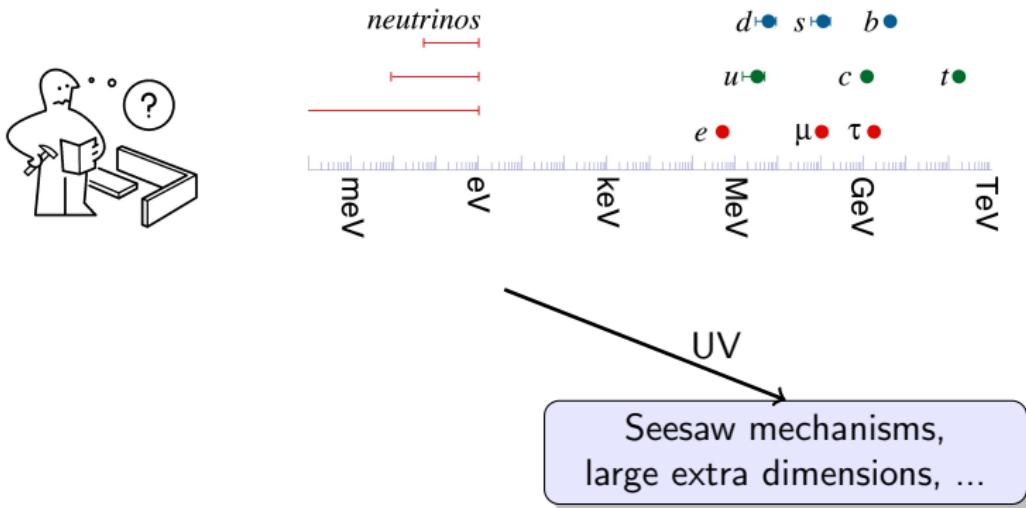


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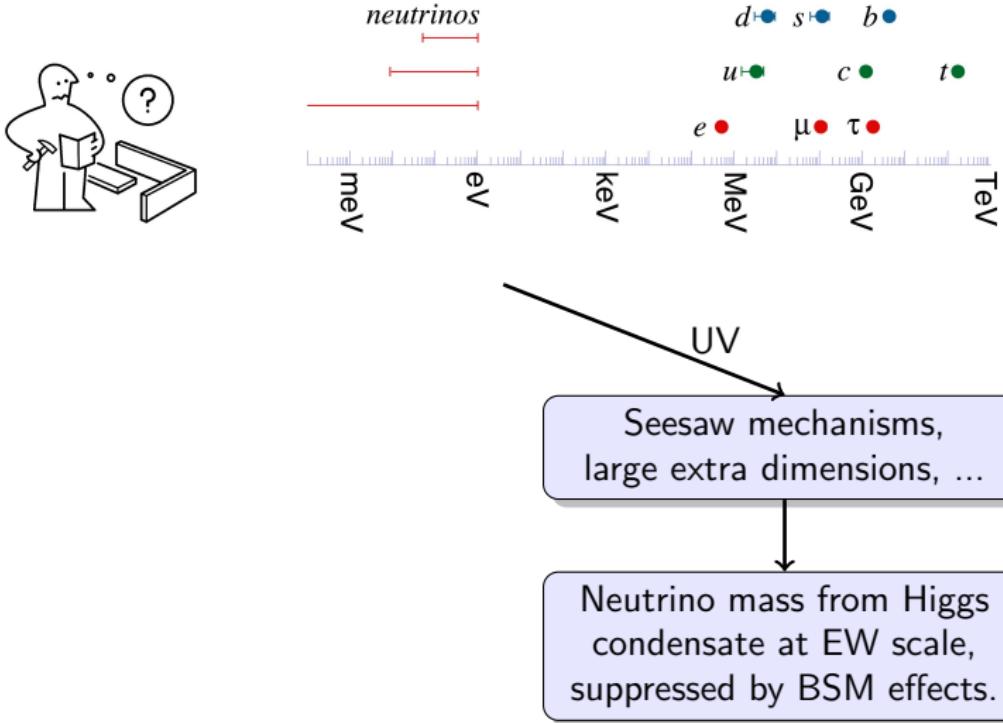


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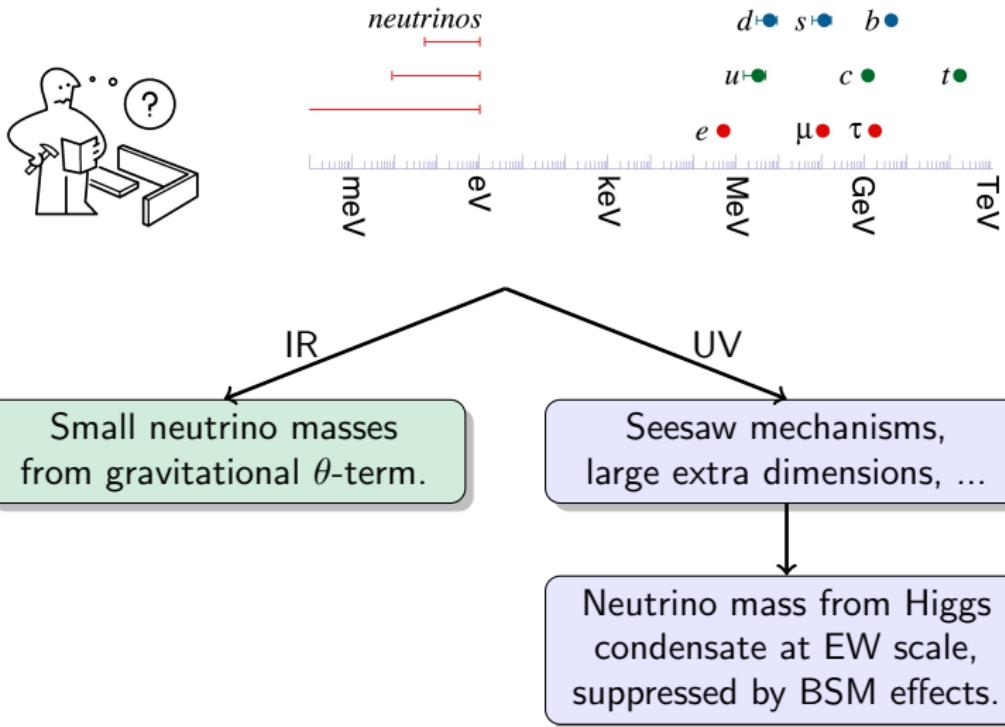


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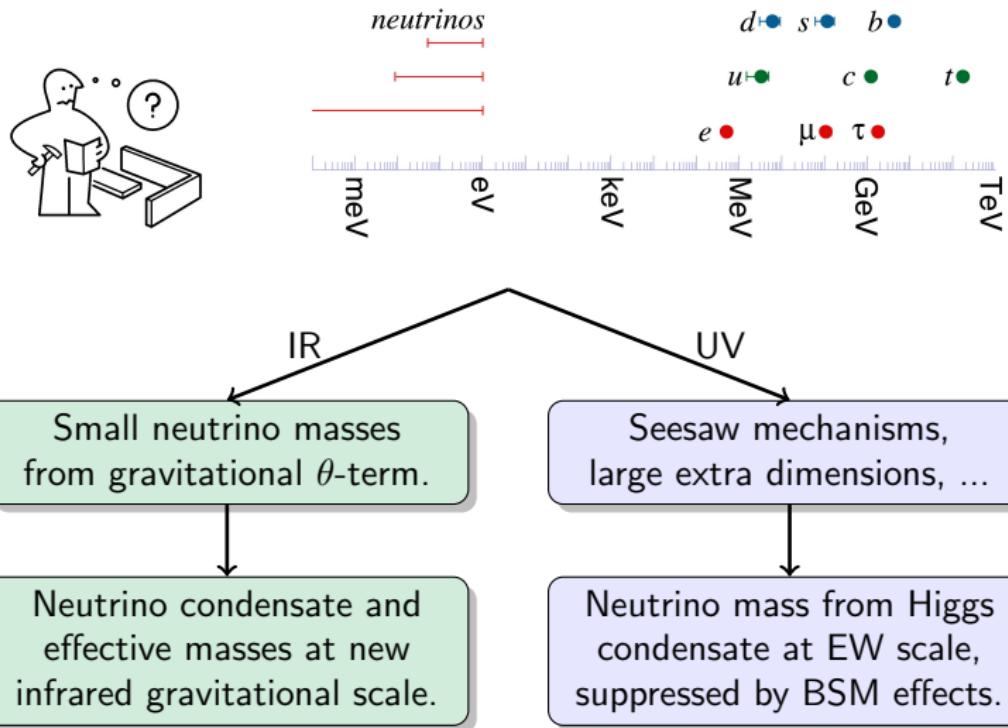


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Analogy: Non-Perturbative QCD Vacuum



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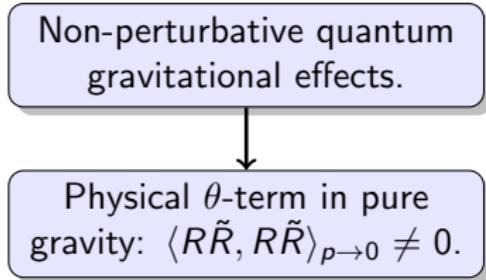
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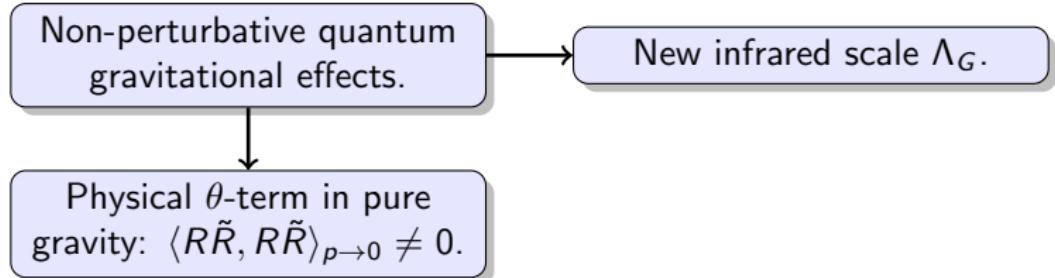
The Model: Neutrino Condensation

Non-perturbative quantum
gravitational effects.

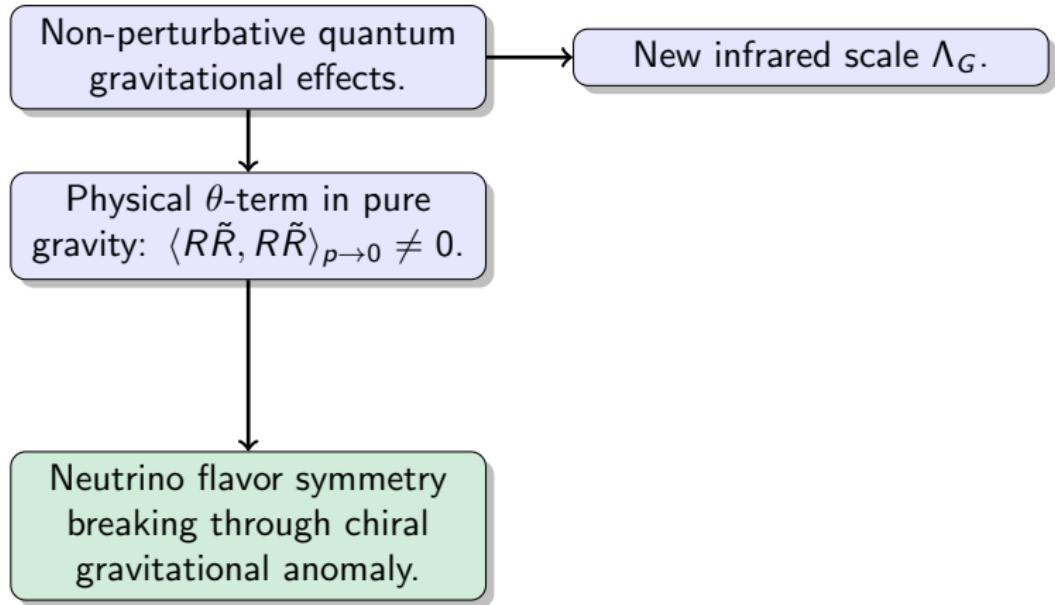
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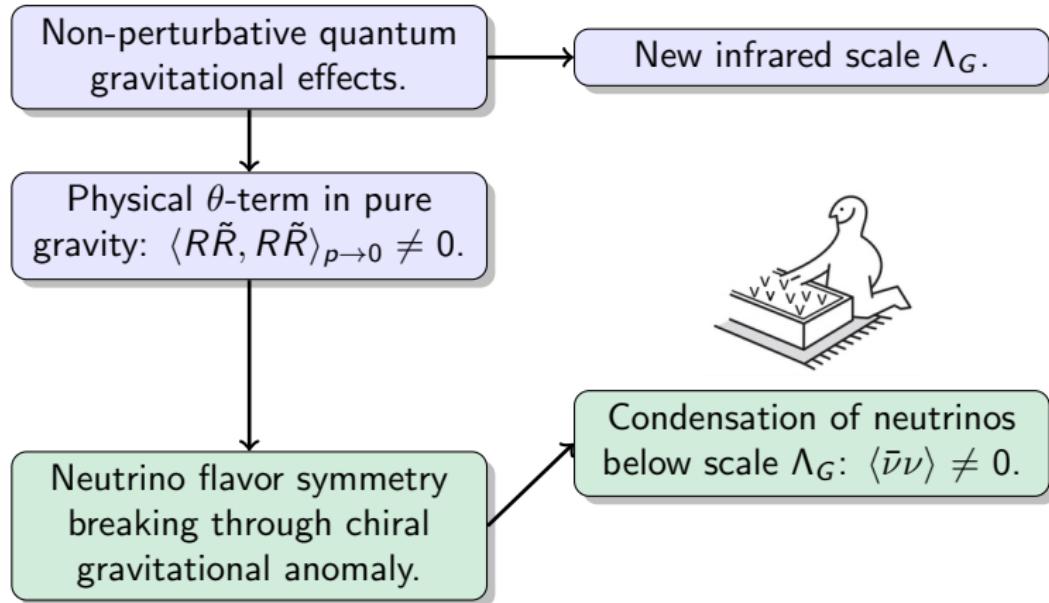
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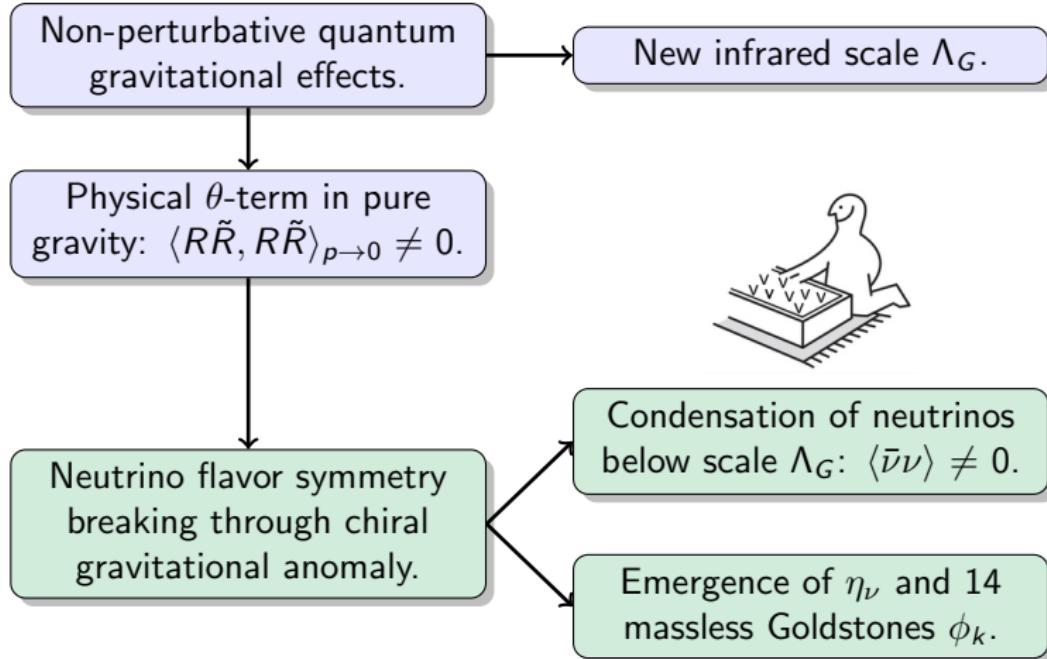
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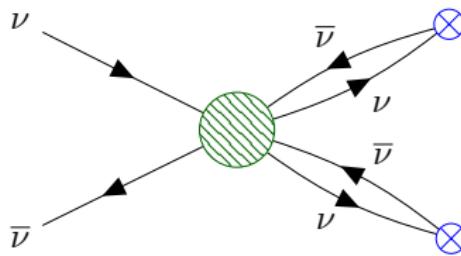


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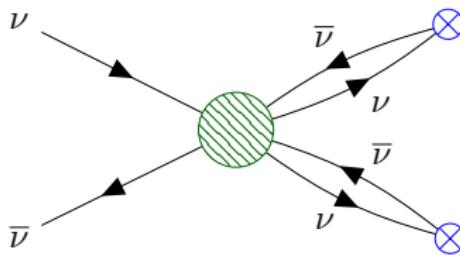
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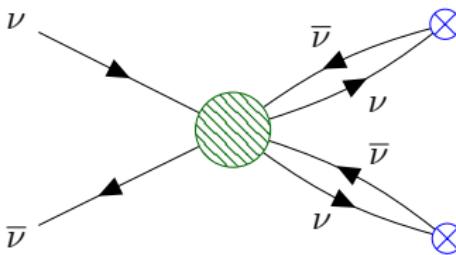
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- ▶ Coupling analogous to 't Hooft vertex in QCD [8].



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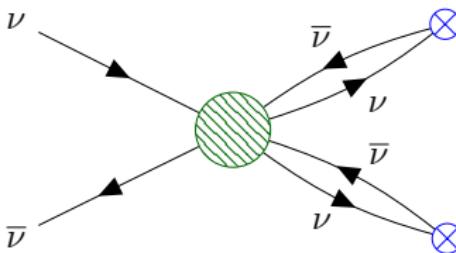


- ▶ Effective potential allows for neutrino mass hierarchy:
 $V(\hat{X}) = \sum_n \frac{1}{n} c_n \text{Tr}[(\hat{X}^+ \hat{X})^n]$ with $\hat{X} \equiv \langle \bar{\nu}_{\alpha_L} \nu_{\alpha_R} \rangle$
 $\Rightarrow \hat{X} = \text{diag}(x_1, x_2, x_3)$ determined by $\partial V / \partial x_i = 0$.

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- ▶ Mechanism works for Dirac and Majorana masses.

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Assumption: condensate $|\langle \bar{\nu} \nu \rangle| = \text{scale } \Lambda_G^3 = \text{temperature } T_{\chi\text{SB}}^3$.

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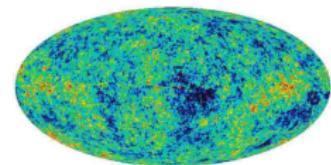
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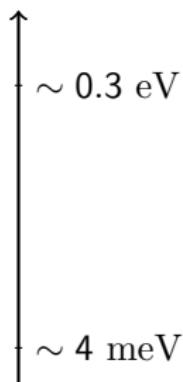
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Image credits: NASA / WMAP Science Team [<http://map.gsfc.nasa.gov/>]

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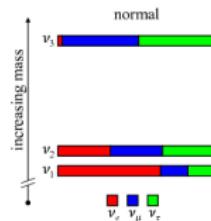
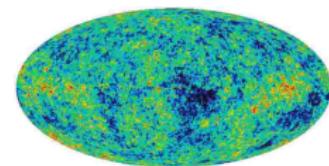
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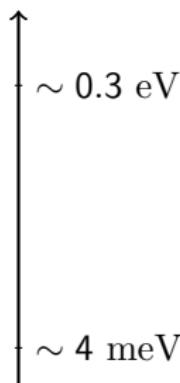
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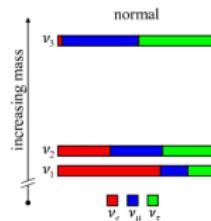
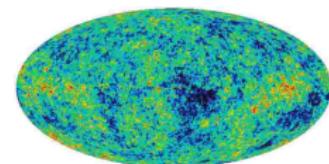
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→ Neutrino vacuum condensate $\langle \bar{\nu} \nu \rangle$ on dark energy scale.

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 \Rightarrow Masses $m_{\nu_e} \lesssim 1.1 \text{ eV}$ [12] still allowed, measurable at



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Image credit: KATRIN [<http://www.ikp.kit.edu/>].

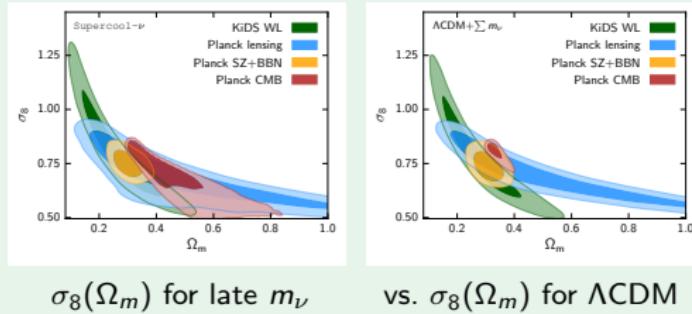
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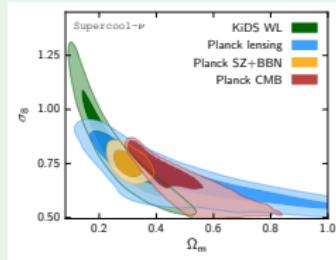
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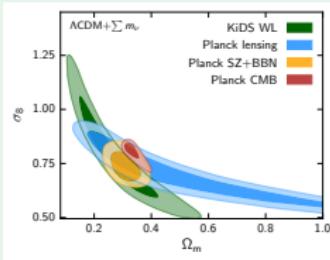
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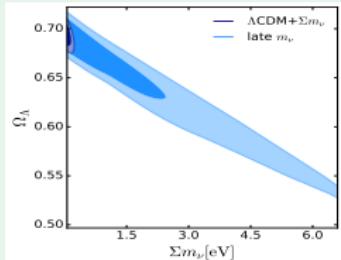
Impact on other cosmic parameters. Dark energy decay?



$\sigma_8(\Omega_m)$ for late m_{ν}



vs. $\sigma_8(\Omega_m)$ for Λ CDM



$\Omega_\Lambda(m_{\nu})$ for both models

[11] Aghanim *et al.* (Planck) (2018). [12] Aker *et al.* (2019).

Image credit: KATRIN [<http://www.ikp.kit.edu/>]. Plots: Lorenz, LF, Calabrese, Hannestad (2018).

Phenomenological Implications

Asymmetries and sterile neutrinos:

- Asymmetric local neutrino clustering [14]?



[14] with Mirzagoli, work in progress.

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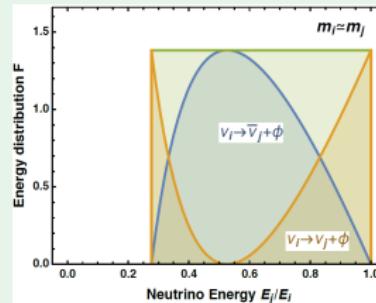
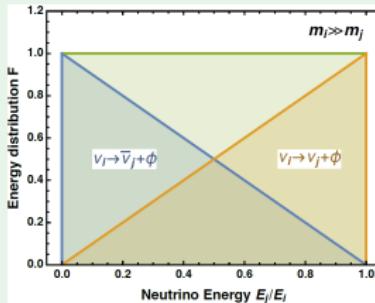
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Gravity measurements:

- ▶ Different polarization intensities of gravitational waves [16].



[16] Jackiw, Pi (2003).

Image credits: The SXS Project [<https://www.ligo.caltech.edu/>]

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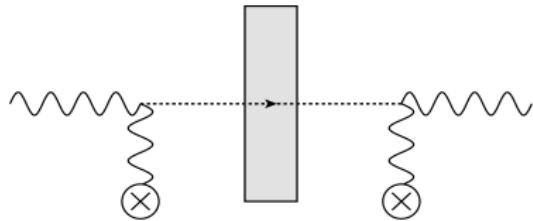


[16] Jackiw, Pi (2003). [17] Dvali, LF (2016b), "Domestic Axion" solution to strong CP problem.
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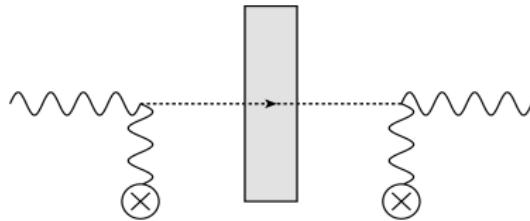
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Do you have any questions?