

Why $N_{\text{eff}} \neq 3.046$

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Outline

- Motivations
- Theory;
 - Active Transport
 - Finite Temperature QED
- Our work:
 - Solve ODE
 - Add next order

Motivations

- For years (2005), SM has predicted $N_{\text{eff}} = 3.046$ [1]
- Recently, [2] gave a different answer $N_{\text{eff}} = 3.052$ using a “non-perturbative” method.
- Who is right? Can we improve upon correct result? What is N_{eff} ?

Effective number of neutrinos

- Simplistic view: How is the total energy density is divided up?
- Set a time: after electron annihilation epoch; before neutrinos become non-relativistic

$$\rho_R = \left[1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{\text{eff}} \right] \rho_\gamma$$

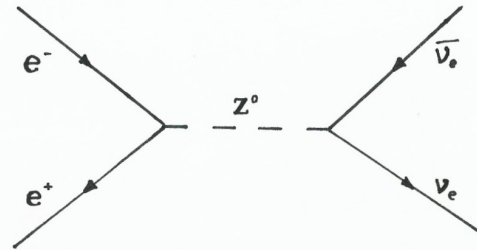
- Where have these terms come from?

Deviation from $N_{\text{eff}} = 3$

$$\rho_R = \left[1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{\text{eff}} \right] \rho_\gamma$$

- Exchange energy with photons?

- Not coupled directly.
- Indirect – four diagrams at tree level



<http://www.mushtukov.com/>

- T=MeV Neutrinos decouple
- T=0.5 MeV e+- decouple.

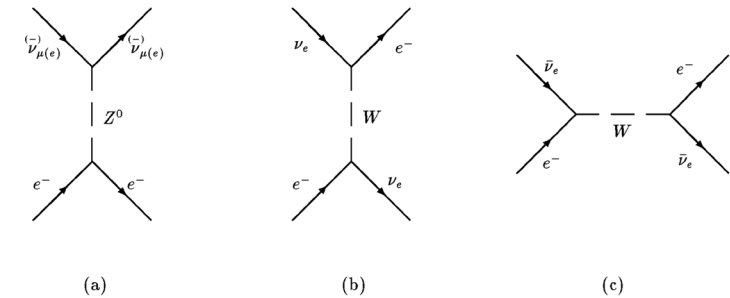


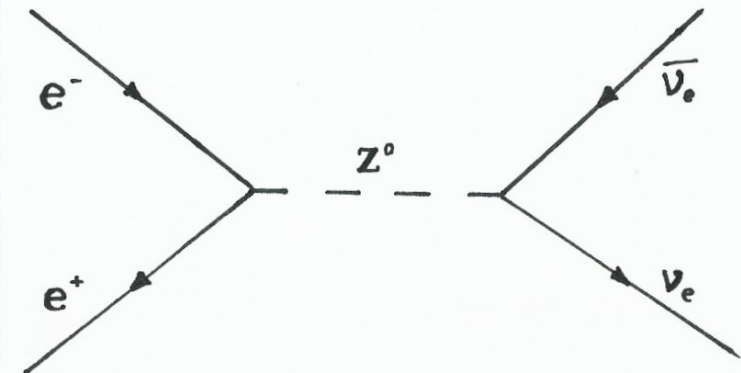
Figure 1: Feynman diagrams for the processes of neutral current (NC) ν_e -scattering (a), and charged current (CC) ν_e -scattering via the exchange of a W-Boson (b,c).

<https://cds.cern.ch/record/248487/>

Active transport

Neutrino decouple $\sim \text{MeV}$
 e^\pm annihilation $\sim 0.5\text{MeV}$

- Do all electrons annihilate instantaneously?
 - Fermi-Dirac distributions
- Are all neutrinos decoupled when electrons annihilate?
- Some energy transferred to neutrino sector
- Use continuity equation



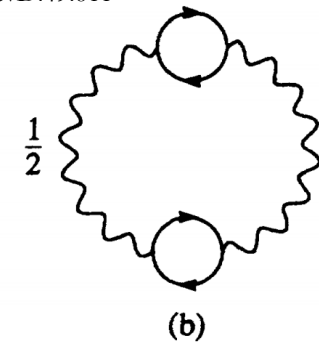
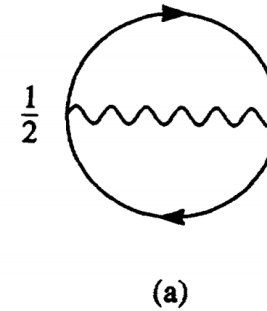
<http://www.mushtukov.com/>

Finite-temperature QED

$$\rho_R = \left[1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{\text{eff}} \right] \rho_\gamma$$

- Electrons and photons gain self energy through QED.
- Thermal masses instructive mentally
- $m_e^2 \rightarrow m_e^2 + \delta m_e^2(T)$, $m_\gamma^2 \rightarrow \delta m_\gamma^2(T)$.
- Not fully relativistic – lowers ρ_γ but ρ_ν same $\therefore N_{\text{eff}}$ increases

journals.aps.org/prd/pdf/10.1103/PhysRevD.49.611



New Variables

$$\begin{aligned}x &= m_e R(T) \\ y &= p R(T) \\ z &= T_\gamma R(T)\end{aligned}$$

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- Go to co-moving variables

$$\begin{aligned}x &= m_e R(T) \\ y &= p R(T) \\ z &= T_\gamma R(T)\end{aligned}$$

- $R(T)$ is some normalized scale factor $\propto a$

- $R(T_d) = \frac{1}{T_\nu}$

- Relativistic temperatures evolve $\propto \frac{1}{a}$

- $p \propto \frac{1}{a}$

- Need P and ρ . Related.

$$\rho = -P + T \frac{\partial P}{\partial T}$$

Continuity Equation

$$\begin{aligned}x &= m_e/T_v \\ y &= p/T_v \\ z &= T_\gamma/T_v\end{aligned}$$

- Normally $\frac{d}{dt}\rho = -3H(\rho + P)$.

$$\bar{P} = P \left(\frac{x}{m_e} \right)^4$$

- Need to track H . Tough.

$$\bar{\rho} = \rho \left(\frac{x}{m_e} \right)^4$$

- Co-moving continuity equation $\frac{d}{dx}\bar{\rho} = \frac{1}{x}(\bar{\rho} - 3\bar{P})$
- When does this approach work?



$$\frac{d}{dx}\bar{\rho} = \frac{1}{x}(\bar{\rho} - 3\bar{P})$$

ODE

$$\begin{aligned} x &= m_e/T_\nu \\ y &= p/T_\nu \\ z &= T_\gamma/T_\nu \end{aligned}$$

$$\bullet \quad \frac{dz}{dx} = \frac{\frac{1}{x}(\bar{\rho}^0 - 3\bar{P}^0) - \frac{\partial \bar{\rho}^0}{\partial x} + G_1(x,z)}{\frac{\partial \bar{\rho}^0}{\partial z} + G_2(x,z)}$$

$$\text{eg } P^0 = \frac{T}{\pi^2} \int_0^\infty dp \, p^2 \ln \left[\frac{(1 + e^{-E_e/T})^2}{(1 - e^{-E_\gamma/T})} \right]$$

- G's encode QED corrections. Interplay.

- Want $\frac{dz}{dx}$, tells us how z (T_γ/T_ν) evolves. $\Delta N_{\text{eff}} = -12 \frac{\delta z}{z_0}$, $\delta z = z_{\text{fin}} - z_0$, $z_0 = \left(\frac{11}{4}\right)^{1/3}$.

$$\frac{dz}{dx} = \frac{\frac{1}{x} (\bar{\rho}^0 - 3\bar{P}^0) - \frac{\partial \bar{\rho}^0}{\partial x} + G_1}{\frac{\partial \bar{\rho}^0}{\partial z} + G_2}$$

Higher order

$$\begin{aligned} x &= m_e/T_\nu \\ y &= p/T_\nu \\ z &= T_\gamma/T_\nu \end{aligned}$$

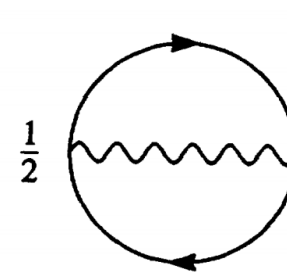
- $G_1^{(i)} = \frac{1}{x} (\bar{\rho}^{(i)} - 3\bar{P}^{(i)}) - \frac{\partial \bar{\rho}^{(i)}}{\partial x}$

- $G_2^{(i)} = \frac{\partial \bar{\rho}^{(i)}}{\partial z}$

- (i) indicates $\mathcal{O}(e^i)$

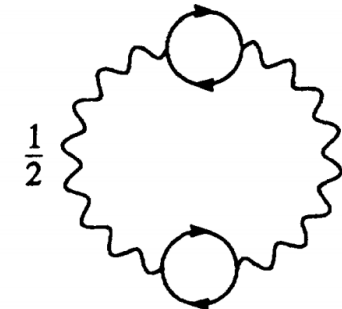
- e.g $P^{(2)} = - \int_0^\infty \frac{dp}{2\pi^2} \left[\frac{p^2}{E_e} \frac{\delta m_e^2}{1 + e^{E_e/T}} + \frac{p}{2} \frac{\delta m_\gamma^2}{e^{p/T} - 1} \right]$

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(a)

$\mathcal{O}(e^2)$

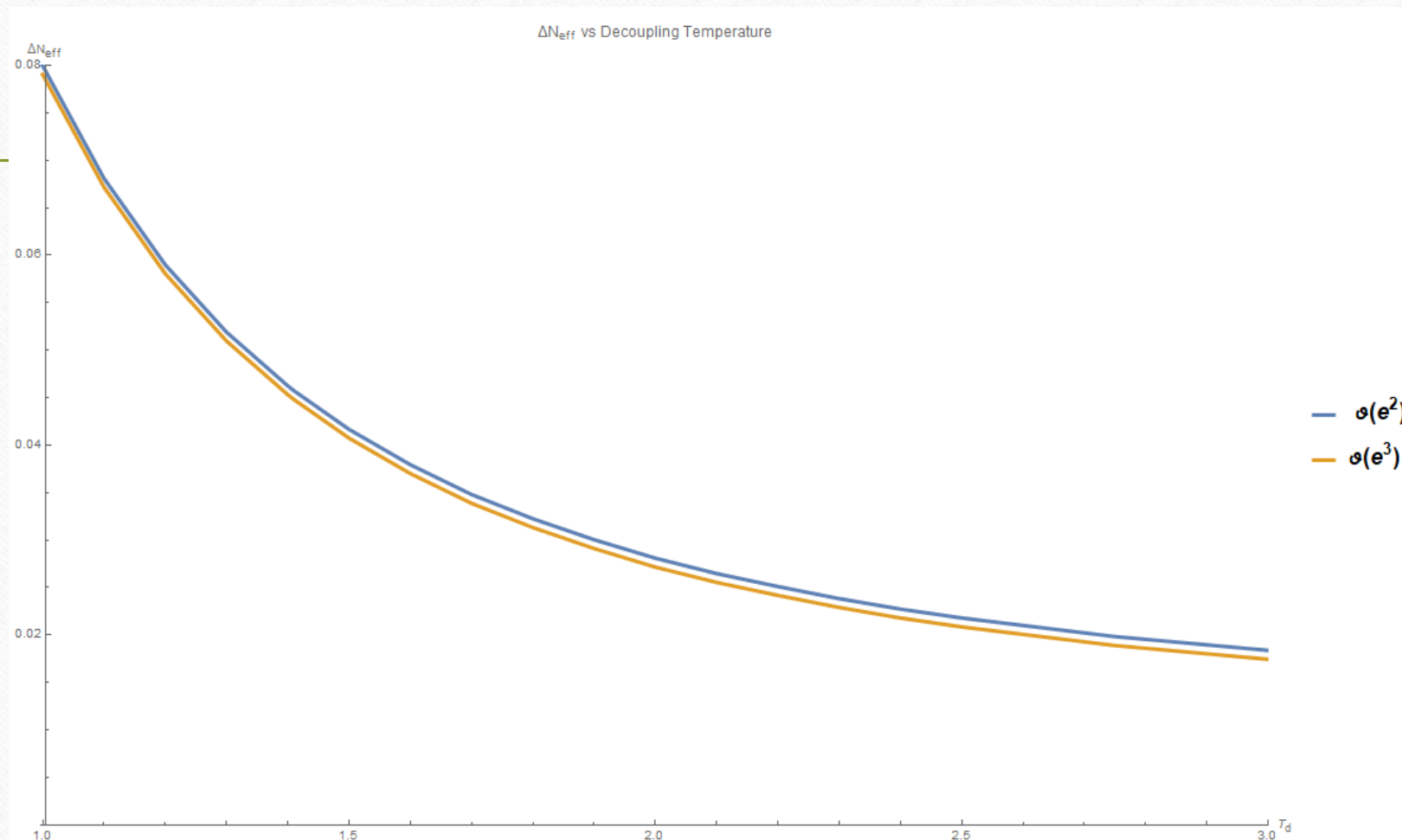


(b)

$\mathcal{O}(e^3)$

Solutions

- Wrote C++ code to solve
- Sensitive to T_D .
- $T_D = 1.41 \text{ MeV}$ [3],
- $N_{\text{eff}}^{(2)} = 3.046$;
- $N_{\text{eff}}^{(3)} = 3.045$
- Similar to inc. ν osc.

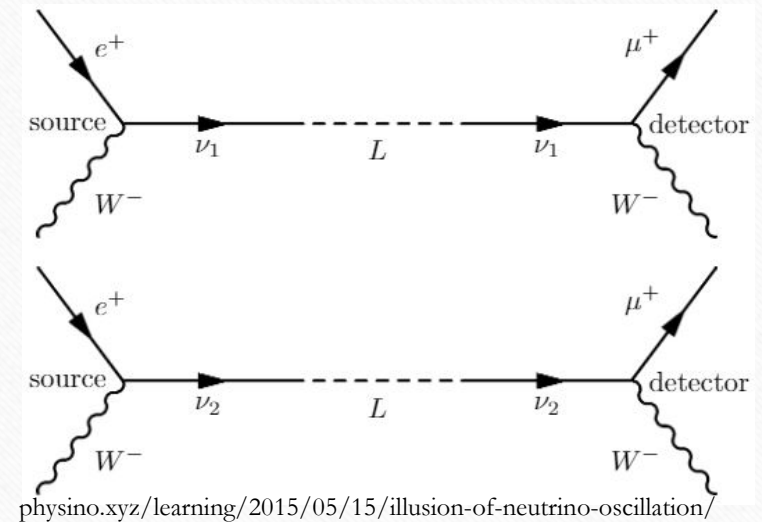


$$\frac{dz}{dx} = \frac{\frac{1}{x} (\bar{\rho}^0 - 3\bar{P}^0) - \frac{\partial \bar{\rho}^0}{\partial x} + G_1}{\frac{\partial \bar{\rho}^0}{\partial z} + G_2}$$

Future

$$\begin{aligned} x &= m_e/T_\nu \\ y &= p/T_\nu \\ z &= T_\gamma/T_\nu \end{aligned}$$

- Even higher order pointless?
- Include non-instantaneous neutrino decoupling
- Neutrino oscillations



$$\frac{dz}{dx} = \frac{\frac{1}{x} (\bar{\rho}^0 - 3\bar{P}^0) - \frac{\partial \bar{\rho}^0}{\partial x} + G_1}{\frac{\partial \bar{\rho}^0}{\partial z} + G_2}$$

Conclusion

$$\begin{aligned} x &= m_e/T_\nu \\ y &= p/T_\nu \\ z &= T_\gamma/T_\nu \end{aligned}$$

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- What is ρ_ν/ρ_γ ?
 - Find P and ρ .
 - Annihilations. Thermal effects
 - Use continuity equation, solve for z . Get N_{eff}