

# Plasma Physics Summary

Mathematical Notes

October 28, 2025

## Contents

<b>1</b>	<b>Introduction to Plasma Physics</b>	<b>2</b>
1.1	What is Plasma? . . . . .	2
1.2	Plasma Parameters . . . . .	2
1.3	Plasma Conditions . . . . .	2
<b>2</b>	<b>Magnetohydrodynamics (MHD)</b>	<b>2</b>
2.1	MHD Equations . . . . .	2
2.2	MHD Approximations . . . . .	3
2.3	Magnetic Reynolds Number . . . . .	3
2.4	Frozen-in Theorem . . . . .	3
<b>3</b>	<b>Plasma Waves and Instabilities</b>	<b>3</b>
3.1	Electromagnetic Waves in Plasma . . . . .	3
3.2	Electrostatic Waves . . . . .	3
3.3	Magnetohydrodynamic Waves . . . . .	3
3.4	Plasma Instabilities . . . . .	4
3.4.1	Rayleigh-Taylor Instability . . . . .	4
3.4.2	Kelvin-Helmholtz Instability . . . . .	4
3.4.3	Two-Stream Instability . . . . .	4
<b>4</b>	<b>Fusion Physics</b>	<b>4</b>
4.1	Thermonuclear Fusion . . . . .	4
4.2	Fusion Reactions . . . . .	4
4.3	Fusion Conditions . . . . .	5
4.4	Tokamak Physics . . . . .	5
4.4.1	Magnetic Confinement . . . . .	5
4.4.2	Plasma Equilibrium . . . . .	5
4.5	Fusion Power Density . . . . .	5
<b>5</b>	<b>Space and Astrophysical Plasmas</b>	<b>5</b>
5.1	Solar Wind . . . . .	5
5.1.1	Solar Wind Properties . . . . .	5
5.2	Magnetosphere . . . . .	6
5.2.1	Earth's Magnetosphere . . . . .	6
5.3	Stellar Atmospheres . . . . .	6
5.4	Interstellar Medium . . . . .	6

<b>6</b>	<b>Computational Plasma Physics</b>	<b>6</b>
6.1	Particle-in-Cell (PIC) Method . . . . .	6
6.1.1	PIC Algorithm . . . . .	6
6.2	Magnetohydrodynamic Simulations . . . . .	7
6.3	Kinetic Simulations . . . . .	7
<b>7</b>	<b>Plasma Diagnostics</b>	<b>7</b>
7.1	Electromagnetic Diagnostics . . . . .	7
7.2	Particle Diagnostics . . . . .	7
7.3	Spectroscopic Diagnostics . . . . .	7
<b>8</b>	<b>Applications</b>	<b>7</b>
8.1	Fusion Energy . . . . .	7
8.2	Space Physics . . . . .	8
8.3	Industrial Applications . . . . .	8
<b>9</b>	<b>Important Constants and Parameters</b>	<b>8</b>
9.1	Fundamental Constants . . . . .	8
9.2	Plasma Parameters . . . . .	8
9.3	Fusion Parameters . . . . .	8

# 1 Introduction to Plasma Physics

## 1.1 What is Plasma?

**Definition 1.1.** A **plasma** is a state of matter consisting of a collection of charged particles (ions and electrons) that is electrically neutral on average and exhibits collective behavior.

## 1.2 Plasma Parameters

- **Plasma frequency:**  $\omega_p = \sqrt{\frac{n_e e^2}{m_e \epsilon_0}}$
- **Debye length:**  $\lambda_D = \sqrt{\frac{\epsilon_0 k_B T_e}{n_e e^2}}$
- **Plasma parameter:**  $g = \frac{1}{n_e \lambda_D^3}$
- **Collision frequency:**  $\nu_{ei} = \frac{n_i e^4 \ln \Lambda}{4\pi \epsilon_0^2 m_e^{1/2} (k_B T_e)^{3/2}}$

## 1.3 Plasma Conditions

For a gas to be considered a plasma:

1.  $\lambda_D \ll L$  (Debye length much smaller than system size)
2.  $N_D = n_e \lambda_D^3 \gg 1$  (many particles in Debye sphere)
3.  $\omega_p \tau \gg 1$  (plasma frequency much larger than collision time)

# 2 Magnetohydrodynamics (MHD)

## 2.1 MHD Equations

The MHD equations describe the macroscopic behavior of plasmas:

**Theorem 2.1** (MHD Continuity Equation).

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0$$

**Theorem 2.2** (MHD Momentum Equation).

$$\rho \frac{D\vec{v}}{Dt} = -\nabla p + \vec{J} \times \vec{B} + \rho \vec{g}$$

**Theorem 2.3** (MHD Induction Equation).

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times (\vec{v} \times \vec{B}) + \eta \nabla^2 \vec{B}$$

where  $\eta = \frac{1}{\mu_0 \sigma}$  is the magnetic diffusivity.

**Theorem 2.4** (MHD Energy Equation).

$$\frac{D}{Dt} \left( \frac{p}{\rho^\gamma} \right) = 0$$

where  $\gamma$  is the adiabatic index.

## 2.2 MHD Approximations

- **Ideal MHD:**  $\sigma \rightarrow \infty$  (perfect conductor)
- **Resistive MHD:** Finite conductivity included
- **Hall MHD:** Electron inertia effects included

## 2.3 Magnetic Reynolds Number

**Definition 2.1.** The magnetic Reynolds number is:

$$R_m = \frac{vL}{\eta} = \mu_0 \sigma vL$$

where  $v$  is characteristic velocity and  $L$  is characteristic length.

## 2.4 Frozen-in Theorem

**Theorem 2.5** (Alfvén's Frozen-in Theorem). In ideal MHD, magnetic field lines are "frozen" into the plasma flow:

$$\frac{d\Phi}{dt} = 0$$

where  $\Phi$  is magnetic flux through a surface moving with the plasma.

# 3 Plasma Waves and Instabilities

## 3.1 Electromagnetic Waves in Plasma

**Theorem 3.1** (Dispersion Relation for Electromagnetic Waves).

$$\omega^2 = \omega_p^2 + c^2 k^2$$

## 3.2 Electrostatic Waves

- **Electron plasma waves:**  $\omega^2 = \omega_p^2 + \frac{3k_B T_e}{m_e} k^2$
- **Ion acoustic waves:**  $\omega^2 = \frac{k_B T_e}{m_i} k^2$
- **Upper hybrid waves:**  $\omega^2 = \omega_p^2 + \omega_c^2$
- **Lower hybrid waves:**  $\omega^2 = \frac{\omega_{ci} \omega_{ce}}{1 + \frac{\omega_p^2}{\omega_c^2}}$

## 3.3 Magnetohydrodynamic Waves

- **Alfvén waves:**  $v_A = \frac{B}{\sqrt{\mu_0 \rho}}$
- **Magnetosonic waves:**  $v_{ms} = \sqrt{v_A^2 + c_s^2}$
- **Slow magnetosonic waves:**  $v_s = \frac{v_A c_s}{\sqrt{v_A^2 + c_s^2}}$

### 3.4 Plasma Instabilities

**Definition 3.1.** A **plasma instability** occurs when small perturbations grow exponentially in time.

#### 3.4.1 Rayleigh-Taylor Instability

**Theorem 3.2** (Rayleigh-Taylor Growth Rate).

$$\gamma = \sqrt{gk \frac{\rho_2 - \rho_1}{\rho_2 + \rho_1}}$$

where  $\rho_2 > \rho_1$  and  $g$  is the acceleration.

#### 3.4.2 Kelvin-Helmholtz Instability

**Theorem 3.3** (Kelvin-Helmholtz Growth Rate).

$$\gamma = \frac{k\Delta v}{2} \sqrt{\frac{\rho_1 \rho_2}{(\rho_1 + \rho_2)^2}}$$

where  $\Delta v$  is the velocity shear.

#### 3.4.3 Two-Stream Instability

**Theorem 3.4** (Two-Stream Instability Condition).

$$\omega^2 = \omega_p^2 \left( 1 + \frac{v_0^2}{v_{th}^2} \right)$$

where  $v_0$  is the relative velocity between streams and  $v_{th}$  is thermal velocity.

## 4 Fusion Physics

### 4.1 Thermonuclear Fusion

**Definition 4.1.** **Thermonuclear fusion** is the process of combining light atomic nuclei to form heavier nuclei, releasing energy.

### 4.2 Fusion Reactions

- **Deuterium-Tritium:**  $D + T \rightarrow {}^4\text{He} + n + 17.6 \text{ MeV}$
- **Deuterium-Deuterium:**  $D + D \rightarrow {}^3\text{He} + n + 3.3 \text{ MeV}$
- **Deuterium-Deuterium:**  $D + D \rightarrow T + p + 4.0 \text{ MeV}$
- **Proton-Proton:**  $p + p \rightarrow D + e^+ + \nu_e + 0.42 \text{ MeV}$

### 4.3 Fusion Conditions

**Theorem 4.1** (Lawson Criterion). For energy breakeven:

$$n\tau \geq \frac{12k_B T}{\langle \sigma v \rangle E_f}$$

where  $n$  is density,  $\tau$  is confinement time,  $T$  is temperature,  $\langle \sigma v \rangle$  is reaction rate, and  $E_f$  is fusion energy per reaction.

### 4.4 Tokamak Physics

**Definition 4.2.** A **tokamak** is a toroidal magnetic confinement device for controlled fusion.

#### 4.4.1 Magnetic Confinement

- **Toroidal field:**  $B_\phi = \frac{B_0 R_0}{R}$
- **Poloidal field:** Generated by plasma current
- **Safety factor:**  $q = \frac{r B_\phi}{R B_\theta}$

#### 4.4.2 Plasma Equilibrium

**Theorem 4.2** (Grad-Shafranov Equation).

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - F \frac{dF}{d\psi}$$

where  $\psi$  is the poloidal flux function and  $F = RB_\phi$ .

### 4.5 Fusion Power Density

**Theorem 4.3** (Fusion Power Density).

$$P_f = \frac{1}{4} n^2 \langle \sigma v \rangle E_f$$

## 5 Space and Astrophysical Plasmas

### 5.1 Solar Wind

**Definition 5.1.** The **solar wind** is a stream of charged particles ejected from the Sun's corona.

#### 5.1.1 Solar Wind Properties

- **Slow solar wind:**  $v \approx 400$  km/s,  $n \approx 10^7$  cm<sup>-3</sup>
- **Fast solar wind:**  $v \approx 800$  km/s,  $n \approx 3 \times 10^6$  cm<sup>-3</sup>
- **Temperature:**  $T \approx 10^5 - 10^6$  K

## 5.2 Magnetosphere

**Definition 5.2.** The **magnetosphere** is the region around a planet dominated by its magnetic field.

### 5.2.1 Earth's Magnetosphere

- **Bow shock:** Standoff distance  $\approx 10 - 15R_E$
- **Magnetopause:** Boundary between solar wind and magnetosphere
- **Magnetotail:** Extended region downstream from Earth

## 5.3 Stellar Atmospheres

- **Photosphere:** Visible surface,  $T \approx 5800$  K
- **Chromosphere:** Transition region,  $T \approx 10^4$  K
- **Corona:** Outer atmosphere,  $T \approx 10^6$  K

## 5.4 Interstellar Medium

- **Cold neutral medium:**  $T \approx 100$  K,  $n \approx 30$  cm $^{-3}$
- **Warm neutral medium:**  $T \approx 8000$  K,  $n \approx 0.3$  cm $^{-3}$
- **Warm ionized medium:**  $T \approx 8000$  K,  $n \approx 0.1$  cm $^{-3}$
- **Hot ionized medium:**  $T \approx 10^6$  K,  $n \approx 0.003$  cm $^{-3}$

# 6 Computational Plasma Physics

## 6.1 Particle-in-Cell (PIC) Method

**Definition 6.1.** The **Particle-in-Cell method** simulates plasmas by following individual particles in self-consistent electromagnetic fields.

### 6.1.1 PIC Algorithm

1. Initialize particles and fields
2. Push particles using Lorentz force
3. Deposit charge and current densities
4. Solve Maxwell's equations for fields
5. Repeat for next time step

## 6.2 Magnetohydrodynamic Simulations

- **Finite difference methods:** Direct discretization of MHD equations
- **Finite volume methods:** Conservative formulation
- **Spectral methods:** Fourier decomposition
- **Adaptive mesh refinement:** Variable resolution

## 6.3 Kinetic Simulations

- **Vlasov equation:**  $\frac{\partial f}{\partial t} + \vec{v} \cdot \nabla f + \frac{q}{m}(\vec{E} + \vec{v} \times \vec{B}) \cdot \nabla_v f = 0$
- **Gyrokinetic theory:** Averaged over gyromotion
- **Drift kinetic theory:** Includes drift motion

# 7 Plasma Diagnostics

## 7.1 Electromagnetic Diagnostics

- **Magnetic probes:** Measure magnetic field fluctuations
- **Electric probes:** Measure electric field and plasma potential
- **Interferometry:** Measure electron density
- **Thomson scattering:** Measure electron temperature and density

## 7.2 Particle Diagnostics

- **Langmuir probes:** Current-voltage characteristics
- **Retarding field analyzers:** Energy distribution functions
- **Neutral particle analyzers:** Ion temperature and density
- **Fast ion diagnostics:** Energetic particle measurements

## 7.3 Spectroscopic Diagnostics

- **Optical emission spectroscopy:** Line intensities and widths
- **X-ray spectroscopy:** Bremsstrahlung and line emission
- **Neutron diagnostics:** Fusion product measurements

# 8 Applications

## 8.1 Fusion Energy

- **Magnetic confinement fusion:** Tokamaks, stellarators
- **Inertial confinement fusion:** Laser and particle beam drivers
- **Magnetized target fusion:** Hybrid approaches



## 8.2 Space Physics

- **Solar-terrestrial interactions:** Space weather
- **Planetary magnetospheres:** Comparative planetology
- **Astrophysical plasmas:** Stellar and galactic physics

## 8.3 Industrial Applications

- **Plasma processing:** Semiconductor manufacturing
- **Plasma propulsion:** Electric spacecraft propulsion
- **Plasma medicine:** Medical applications
- **Materials processing:** Surface modification

# 9 Important Constants and Parameters

## 9.1 Fundamental Constants

- Electron charge:  $e = 1.602 \times 10^{-19}$  C
- Electron mass:  $m_e = 9.109 \times 10^{-31}$  kg
- Proton mass:  $m_p = 1.673 \times 10^{-27}$  kg
- Permittivity of free space:  $\epsilon_0 = 8.854 \times 10^{-12}$  F/m
- Permeability of free space:  $\mu_0 = 4\pi \times 10^{-7}$  H/m
- Boltzmann constant:  $k_B = 1.381 \times 10^{-23}$  J/K
- Speed of light:  $c = 2.998 \times 10^8$  m/s

## 9.2 Plasma Parameters

- Classical electron radius:  $r_e = \frac{e^2}{4\pi\epsilon_0 m_e c^2} = 2.818 \times 10^{-15}$  m
- Bohr radius:  $a_0 = \frac{4\pi\epsilon_0 \hbar^2}{m_e e^2} = 5.292 \times 10^{-11}$  m
- Fine structure constant:  $\alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c} = 7.297 \times 10^{-3}$
- Electron cyclotron frequency:  $\omega_{ce} = \frac{eB}{m_e}$
- Ion cyclotron frequency:  $\omega_{ci} = \frac{eB}{m_i}$

## 9.3 Fusion Parameters

- DT fusion cross-section peak:  $\sigma_{max} \approx 5 \times 10^{-28}$  m<sup>2</sup> at  $T \approx 100$  keV
- Lawson criterion for DT:  $n\tau \geq 10^{20}$  s/m<sup>3</sup> at  $T \approx 10$  keV
- ITER parameters:  $R = 6.2$  m,  $a = 2.0$  m,  $B = 5.3$  T,  $I_p = 15$  MA