## Scripts for XR postprocessing with USPEX

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#### 1 split\_CIFs.py

It splits the structures with lowest enthalpy of a variable composition USPEX run into multiple CIF files. The script takes the following arguments from command line:

- 1. the output file extended\_convex\_hull from USPEX
- 2. the output file extended\_convex\_hull\_POSCARS from USPEX
- 3. the value of the external pressure used for the USPEX run

The script performs the following operations:

- 1. for each structure, it reads the parameters *enthalpy* and *fitness* from extended\_convex\_hull and the geometry from extended\_convex\_hull\_POSCARS
- 2. it selects, for each reduced formula, the 5 structures with lowest enthalpy
- 3. it outputs the selected structures as CIF files in a new folder results

The name of each CIF file has the format i\_ID\_fitness\_enthalpy\_iupacformula\_pressure\_symmetry.cif, where:

i is a natural number which orders the output with increasing fitness

**ID** is the structure ID from the USPEX run

fitness is the fitness of the structure

**enthalpy** is the *enthalpy* of the structure

**iupacformula** is the *IUPAC formula* of the structure

pressure is the pressure used for the USPEX run

symmetry is the space group number, determined with tolerance 0.2

Example: python split\_CIFs.py extended\_convex\_hull extended\_convex\_hull\_POSCARS 50GPa

#### 2 sublattice\_split\_CIFs.py

It reads the results of a variable composition USPEX run, it removes all hydrogen atoms, and then it splits the structures with lowest enthalpy into multiple CIF files. The script takes the following arguments from command line:

- 1. the output file extended\_convex\_hull from USPEX
- 2. the output file extended\_convex\_hull\_POSCARS from USPEX
- 3. the value of the external pressure used for the USPEX run

The script performs the following operations:

- 1. for each structure, it reads the parameters *enthalpy* and *fitness* from extended\_convex\_hull and the geometry from extended\_convex\_hull\_POSCARS
- 2. it deletes all hydrogen atoms
- 3. it selects, for each reduced formula, the 5 structures with lowest enthalpy
- 4. it outputs the selected structures as CIF files in a new folder results

The name of each CIF file has the format i\_ID\_fitness\_enthalpy\_iupacformula\_pressure\_symmetry.cif, where:

i is a natural number which orders the output with increasing fitness

**ID** is the structure ID from the USPEX run

fitness is the fitness of the structure

**enthalpy** is the *enthalpy* of the structure

iupacformula is the IUPAC formula of the structure, with hydrogens

pressure is the pressure used for the USPEX run

**symmetry** is the space group number, determined with a tolerance of 0.2 and without hydrogens

Example: python sublattice\_split\_CIFs.py extended\_convex\_hull extended\_convex\_hull\_POSCARS 50GPa

# 3 fixcomp\_split\_CIFs.py

It splits all the structures of a fixed composition USPEX run into multiple CIF files. The script takes the following arguments from command line:

- 1. the output file Individuals from USPEX
- 2. the output file  ${\tt gatheredPOSCARS}$  from USPEX

3. the value of the external pressure used for the USPEX run

The script performs the following operations:

- 1. for each structure, it reads the parameter *enthalpy* from Individuals and the geometry from gatheredPOSCARS
- 2. it computes real\_fitness = enthalpy / total\_number\_of\_atoms
- 3. it outputs the structures as CIF files in a new folder results

The name of each CIF file has the format i\_ID\_fitness\_enthalpy\_iupacformula\_pressure\_symmetry.cif, where:

i is a natural number which orders the output with increasing real\_fitness

**ID** is the structure ID from the USPEX run

fitness is the real\_fitness of the structure

**enthalpy** is the *enthalpy* of the structure

iupacformula is the IUPAC formula of the structure

pressure is the pressure used for the USPEX run

symmetry is the space group number, determined with a tolerance of 0.2

Example: python fixcomp\_split\_CIFs.py Individuals gatheredPOSCARS 50GPa

### 4 xr\_screening.py

It performs a screening of USPEX results, looking for the structures that best match an experimental X-ray spectrum. For the theory behind this script, see appendix A. In a few words, given an input experimental spectrum, the theoretical spectrum is calculated for each structure in the USPEX run and a value F is computed. The smaller is F, the better is the agreement between theoretical and experimental spectra. The script contains the following input parameters:

- 1. the importance coefficients  $f(h_i)$  used in the computation of F
- 2. the experimental pressure
- 3. the pressure of the USPEX run
- 4. the start and end angles for the computation of theoretical spectra
- 5. the experimental wavelength

- 6. the value of  $\sigma$  for the gaussian smearing of peaks, used for generating output pictures
- 7. the name of the *spectrum file*, containing angles and intensities of the experimental spectrum
- 8. the name of the file extended\_convex\_hull from USPEX
- 9. the name of the file extended\_convex\_hull\_POSCARS from USPEX
- 10. the parameter *match\_tol*, that is the tolerance for matching experimental peaks with theoretical peaks, in degrees

The script performs the following operations:

- 1. for each structure, it reads the parameter *fitness* from extended\_convex\_hull and the geometry from extended\_convex\_hull\_POSCARS
- 2. it computes the theoretical X-ray spectrum
- 3. it computes the agreement F between the spectra
- 4. it outputs a CIF file with the symmetrized structure (tolerance 0.2)
- 5. it outputs a PNG graph with the theoretical and experimental spectra superimposed for comparison

The name of each CIF and PNG file have the format  $F_ID_fitness_iupacformula_pressure_symmetry$ , where:

 $\mathbf{F}$  is the agreement F between theoretical and experimental spectra

**ID** is the structure ID from the USPEX run

fitness is the fitness of the structure from the USPEX run

**iupacformula** is the *IUPAC formula* of the structure

pressure is the pressure used for the USPEX run

**symmetry** is the space group number, determined with a tolerance of 0.2

Example: python xr\_screening.py

### 5 exclusion.py

It allows to quickly filter a multitude of CIF files, by removing those which have significant peaks in a user-defined exclusion region of the X-ray spectrum. The script takes the following arguments from command line:

1. the wavelength of the incident radiation in Å

- 2. the peak cut-off, in % of the maximum intensity
- 3. a number of intervals in degrees, expressed as two angles separated by a hyphen (-), defining the exclusion region

The script works in a folder with many CIF files, and performs the following:

- 1. it opens, one by one, all CIF files and it predicts the XRD pattern of the structure according to the given wavelength
- 2. if the predicted pattern contains any peak in the exclusion regions that is bigger than the given cut-off, it deletes the CIF file

Example: python exclusion.py 0.6199 25 25-28 31-32

#### 6 find\_peak.py

It allows to quickly filter a multitude of CIF files, by removing those which do not have significant peaks in all user-defined search intervals of the X-ray spectrum. The script takes the following arguments from command line:

- 1. the wavelength of the incident radiation in Å
- 2. the peak cut-off, in % of the maximum intensity
- 3. a number of search intervals in degrees, expressed as two angles separated by a hyphen (-)

The script works in a folder with many CIF files, and performs the following:

- 1. it opens, one by one, all CIF files and it predicts the XRD pattern of the structure according to the given wavelength
- 2. if there is at least one search region which does not contain any peak bigger than the given cut-off, it deletes the CIF file

Example: python find\_peak.py 0.6199 15 25-28 31-32

### 7 change\_pressure.py

It translates the structures contained in many CIF files to a different pressure, by deforming the lattice parameters using a second order Taylor expansion of the Birch-Murnaghan equation. More details about the underlying theory can be found in appendix A. The script takes the following arguments from command line:

- 1. the initial pressure
- 2. the final pressure

The script works in a folder with many CIF files, and performs the following:

- 1. it reads the structures contained in all CIF files
- 2. it deforms the lattice parameters according to the new pressure
- 3. it creates new CIF files containing the deformed structures

The name of the new CIF files have the format OLDNAME\_toNEWPRESSURE, where:

**OLDNAME** is the name of the file with the structure at the initial pressure

**NEWPRESSURE** is the final pressure

Example: python change\_pressure.py 50 58

#### 8 relax\_new\_pressure.py

It allows an accurate translation of the structures contained in many CIF files to a different pressure, by relaxing them with VASP. The script works in a folder with many CIF files and it is meant to be run on a cluster. Before running it, add the following lines to your .bashrc file:

- 1. export VASP\_SCRIPT=PATH/TO/relax\_new\_pressure/run\_vasp.py
- 2. export VASP\_PP\_PATH=PATH/TO/YOUR/pp/FOLDER

Then tune the <code>relax\_new\_pressure.py</code> file with the desired new pressure and the appropriate VASP parameters for your calculation and run the python script on a computing node. By default, VASP relaxations are performed sequentially and every relaxation takes 4 computing cores. For more details about the implementation, have a look at the ASE VASP calculator.

# Appendices

## A Spectra comparison

We developed a code which computes, from the experimental spectrum and the USPEX output, the degree of agreement (fitness) of each relaxed structure with the experimental data. The USPEX calculation and the experimental spectrum do not need to be exactly at the same pressure, but the two pressures need to be *close*, that is, no more than 20 GPa apart.

Since we have a pressure difference, we first translate each calculated structure to the experimental pressure by using the Birch-Murnaghan equation

$$\Delta P = \frac{3B_0}{2} \left[ \left( \frac{V_0}{V} \right)^{\frac{7}{3}} - \left( \frac{V_0}{V} \right)^{\frac{5}{3}} \right] \left\{ 1 + \frac{3}{4} \left( B_0' - 4 \right) \left[ \left( \frac{V_0}{V} \right)^{\frac{2}{3}} - 1 \right] \right\}$$

where  $\Delta P$  is the pressure difference,  $V_0$  is the volume of the unit cell at the calculated pressure, V is the volume at the experimental pressure,  $B_0$  is the bulk modulus and  $B'_0$  is the derivative of the bulk modulus with respect to pressure. For  $B_0$  and  $B'_0$  we take the average values  $B_0 = 300$  and  $B'_0 = 3$ .

Assuming that, for small pressure variations, also  $V_0/V$  will be small, we approximate the Birch-Murnaghan equation to a second order Taylor expansion in  $V_0/V$  and we get the volume V at the experimental pressure:

$$V = \frac{300}{150 + \sqrt{22500 + 300\Delta P}} V_0$$

Then we define the following scaling factor

$$k = \sqrt[3]{\frac{300}{150 + \sqrt{22500 + 300\Delta P}}}$$

that will be used to rescale the lattice parameters of all the calculated structures.

After this rescaling, the relaxed structures are symmetrized with a tolerance of 0.2 and the theoretical XRD spectra are computed. Both the theoretical and the experimental spectra are in the form of a series of peaks. For each peak we know the diffraction angle and the relative intensity (while the intensity of the highest peak in each spectrum has been conventionally given the value 100). We define a *match* between a theoretical and an experimental peak if the two peaks are less than *match\_tol* degrees apart, regardless of their intensities.

The fitness between a calculated XRD spectrum and the experimental spectrum is defined by the following fitness function

$$F = \sum_{i,j}^{match} \frac{(x_i^{exp} - x_j^{th})^2}{\Delta \alpha^2} f(h_i^{exp}) + \sum_i^{rest} \frac{(x_i^{exp})^2}{\Delta \alpha^2} f(h_i^{exp}) + \sum_i^{rest} \frac{(x_i^{th})^2}{\Delta \alpha^2} f(h_i^{th})$$

$$\sum_{i,j}^{match} \frac{(h_i^{exp} - h_j^{th})^2}{100^2} f(h_i^{exp}) + \sum_i^{rest} \frac{(h_i^{exp})^2}{100^2} f(h_i^{exp}) + \sum_i^{rest} \frac{(h_i^{th})^2}{100^2} f(h_i^{th})$$

where the  $x_i$  are diffraction angles, the  $h_i$  are intensities,  $\Delta \alpha$  is the total width of the spectra and  $f(h_i)$  is an importance coefficient given by the following piecewise-defined function

$$f(t) = \begin{cases} 5 & \text{if } t > 90\\ 1 & \text{if } 50 < t \le 90\\ 0.25 & \text{if } 10 < t \le 50\\ 0.02 & \text{if } 1 < t \le 10\\ 0 & \text{if } t \le 1 \end{cases}$$

The first term of the fitness is a sum over the matched peaks, and each addend of this sum will be smaller the more the two peaks are close to each other. At the denominator we find the total width  $\Delta \alpha$  of the spectra, which is the

maximum value for an angle and it allows us to get dimensionless addends. The importance coefficient just tells us that higher peaks will give a more significant contribution to the fitness than smaller peaks, and that peaks with intensity less than 1 will give no contribution. The second term is a sum over the experimental peaks that are left after our matching, we call these peaks *experimental rest*. It has a similar fashion to the first term, but its addends will have a much higher value since there is no subtraction at the numerator. The third term, analogously, is a sum over the *theoretical rest*. The remaining three terms are very similar to the first three, but angles are substituted by intensities.

It is clear that a low value of F gives a good agreement between calculated and experimental spectra, allowing a quick identification of promising candidates in the USPEX output.