

The magnetic structure and palaeomagnetic recording fidelity of sub-micron greigite (Fe₃S₄)

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Abstract

We present the results of a finite-element micromagnetic model of 30 nm to 300 nm greigite (Fe₃S₄) grains with a variety of equant morphologies. This grain size range covers the magnetic single-domain (SD) to pseudo single-domain (PSD) transition, and possibly also the PSD to multi-domain (MD) transition. The SD–PSD threshold d_0 is determined to be $50 \text{ nm} \leq d_0 \leq 56 \text{ nm}$ depending on grain shape. The nudged elastic-band method was used to determine the room temperature energy barriers between stable states and thus the blocking volumes. It is found that, in the absence of interparticle magnetostatic interactions, the magnetisation of equant SD greigite is not stable on a geological scale and only PSD grains $\geq 70 \text{ nm}$ can be expected to carry a stable magnetisation over billion-year timescales, i.e., all non-interacting SD particles are essentially superparamagnetic. We further identify a mechanism for the PSD to multi-domain (MD) transition, which is of a continuous nature from PSD nucleation up to 300 nm, when structures typical of MD behaviour like closure domains begin to form.

Keywords: Greigite, Pseudo Single-Domain, Micromagnetics

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1. Introduction

The ferrimagnetic mineral greigite (Fe_3S_4) is the iron sulphide analogue of the iron oxide magnetite (Fe_3O_4). It is commonly formed as a precursor to pyrite in early diagenetic anoxic sulphate-reducing sediments (Berner, 1984; Hunger and Benning, 2007) and as a product of biomineralisation by magnetotactic bacteria (Mann et al., 1990). Although thought to be thermodynamically unstable under most sedimentary regimes, it has been found to be stable on geological timescales (Roberts et al., 2011), making greigite a possible natural remanent magnetisation (NRM) carrier in many systems, e.g., lacustrine (Babinszki et al., 2007; Ron et al., 2007) and marine (Roberts and Turner, 1993; Roberts and Weaver, 2005; Rowan and Roberts, 2006; Rowan et al., 2009) sediments, oil fields' shallow overburdens (Abubakar et al., 2015; Donovan et al., 1984; Reynolds et al., 1993) and gas-hydrate-bearing sediments (Larrasoana et al., 2007). To further our understanding of the potential use of greigite as a proxy for environmental change, hydrocarbon exploration, magnetostratigraphy and in general the contribution of this iron sulphide to the magnetic properties of rocks, we have implemented micromagnetic numerical finite-element method (FEM) simulations of greigite. To characterise its basic properties we have modelled its magnetic domain state's shape and size dependence using the *DUNLOP* package (Nagy, 2016) and its stability on geological timescales at room temperature using the *MERRILL* package (Nagy et al., 2017).

The model solutions are dependent on a balance between various magnetic forces and thus it is important that the material's magnetic parameters be known as accurately as possible. Past difficulties in producing pure greigite samples on which to determine these parameters has resulted in a lack of accurate models in the literature. However, these difficulties have been overcome, and recent measurements on highly pure, high crystallinity synthetic greigite samples (Chang et al., 2008, 2009; Guowei et al., 2014; Winklhofer et al., 2014) now allow for numerical micromagnetic models to probe the magnetic microstructure of greigite.

31 The fundamental magnetic parameters of greigite used in this investigation
 32 are the saturation magnetisation $M_S = 3.51 \mu_B$ p.c.u. (Guowei et al., 2014) or
 33 $\sim 2.7 \times 10^5$ A/m which is $\sim 11\%$ higher than the value reported by Chang et al.
 34 (2009) of $3.25 \mu_B$ p.c.u. (and $\sim 57\%$ the value of M_S for magnetite) and the
 35 exchange stiffness constant $A = 2 \times 10^{-12}$ J/m (Chang et al., 2008) ($\sim 15\%$ the
 36 value of A for magnetite). Winklhofer et al. (2014) estimated a cubic magne-
 37 tocrystalline anisotropy (MCA) term $K_1 = -1.7 \times 10^4$ J/m³ ($\sim 42\%$ higher than
 38 the value for magnetite) and negligible second MCA constant K_2 to K_1 ratio,
 39 i.e., the easy axes are the $\langle 111 \rangle$.

40 Very few micromagnetic models of greigite have so far been attempted, with
 41 only the work of Muxworthy et al. (2013) having been published. They im-
 42 plemented a micromagnetic finite-difference (FD) method that used the earlier
 43 value of M_S from Chang et al. (2009) to model both individual grains and the
 44 effects of magnetostatic interactions between cuboidal particles of greigite ar-
 45 ranged in chains. For the single crystals they found good agreement with the
 46 analytical calculations of Diaz-Ricci and Kirschvink (1992) (for which crude esti-
 47 mates of the magnetic parameters were used). However, due to their structured
 48 spatial discretisation, FD methods are not as well suited for the simulation of
 49 the euhedral morphologies seen in natural samples (Snowball, 1997) as FEMs
 50 are. Elongation is a common feature of magnetosomal grains but not of parti-
 51 cles with non-biogenic origin. Although Muxworthy et al. (2013) have demon-
 52 strated that particle elongation and interparticle magnetostatic interactions are
 53 important, this study is limited to a variety of equant, isolated particles whose
 54 behaviour represent the limit as the particle concentration approaches zero. As
 55 such, this study provides a first step towards a more complete picture of the
 56 role of non-biogenic greigite in rock magnetism.

57 The grain size of a magnetic mineral strongly affects its magnetic behaviour
 58 and palaeomagnetic recording fidelity. As grain size distributions are usually
 59 inferred from bulk magnetic properties, a deep understanding of how a grain's
 60 magnetic properties depend on size is needed. We present here the results of
 61 a numerical FEM study of the stress-free, zero-field domain states of spheri-

cal and equant euhedral single grains of greigite in the single-domain (SD) to pseudo single-domain (PSD) range. To address this question we used the nudged elastic-band (NEB) method (Fabian and Shcherbakov, 2017) to calculate action-minimising paths (AMPs) between stable magnetic configurations which allow us to determine energy barriers and from these the blocking volumes of naturally occurring equant euhedral grains of greigite.

2. Methodology

2.1. The micromagnetic algorithm

For a ferromagnetic (in the broad sense, i.e., ferro- or ferri-) material—neglecting external fields, thermal and magnetostrictive effects—the Gibbs free-energy functional can be written as (Brown, 1963)

$$E_G = \int_{\Omega} (\phi_{\text{exchange}} + \phi_{\text{anisotropy}} + \phi_{\text{stray}}) d^3\mathbf{r}, \quad (1)$$

with Ω the ferromagnetic volume. Here,

$$\phi_{\text{exchange}} = A|\nabla\mathbf{m}|^2, \quad (2)$$

with the reduced (unitary) magnetisation vector \mathbf{m} , is the expression for the continuum approximation of the energy density due to the quantum-mechanical exchange forces (Landau and Lifshitz, 1935).

$$\phi_{\text{anisotropy}} = \frac{K_1}{2} \sum_{i \neq j} \gamma_i^2 \gamma_j^2 + K_2 \prod_i \gamma_i^2, \quad (3)$$

with γ_i the direction cosines, is the MCA energy density in the cubic anisotropy system. In the cartesian system this has the form:

$$\phi_{\text{anisotropy}} = K_1(m_x^2 m_y^2 + m_y^2 m_z^2 + m_z^2 m_x^2), \quad (4)$$

83 where K_2 has been neglected since in the cubic anisotropy system we are as-
 84 suming, K_1 is the dominant term at room temperature. Finally,

$$85 \quad \phi_{\text{stray}} = -\frac{\mu_0 M_S}{2} \mathbf{m} \cdot \mathbf{H}_{\text{stray}} \quad (5)$$

86 is the magnetostatic self-energy density, with $\mathbf{H}_{\text{stray}}$ the stray field produced by
 87 the ferromagnetic body. It is known from thermodynamics that under isother-
 88 mal and isobaric conditions a system will spontaneously evolve towards an equi-
 89 librium state with minimal Gibbs free-energy. It is the aim of a micromagnetic
 90 algorithm to find the equilibrium magnetisation \mathbf{m} .

91 The variational principle proposed by Brown (1963) $\delta E_G / \delta \mathbf{m} = 0$ leads to
 92 ‘Brown’s equations’

$$93 \quad \mathbf{m} \times (\mathbf{H}_{\text{exchange}} + \mathbf{H}_{\text{anisotropy}} + \mathbf{H}_{\text{stray}}) = 0, \quad (6)$$

94 with

$$95 \quad \mathbf{H}_{\text{exchange}} = \frac{2A}{\mu_0 M_S} \nabla^2 \mathbf{m} \quad (7)$$

96 and

$$97 \quad \mathbf{H}_{\text{anisotropy}} = \frac{2K_1}{\mu_0 M_S} [m_x(1 - m_x^2)\hat{\mathbf{i}} + m_y(1 - m_y^2)\hat{\mathbf{j}} + m_z(1 - m_z^2)\hat{\mathbf{k}}]. \quad (8)$$

98 Therefore, in equilibrium the magnetisation is parallel to an *effective* field
 99 $\mathbf{H}_{\text{eff.}} = \mathbf{H}_{\text{exchange}} + \mathbf{H}_{\text{anis}} + \mathbf{H}_{\text{stray}}$, while $\mu_0 M_S \mathbf{m} \times \mathbf{H}_{\text{eff.}}$ is the local torque
 100 produced on the magnetisation by the effective field at each magnetic site, so
 101 we have equilibrium via the vanishing of the torque (Brown, 1963).

102 We thus have the two main approaches towards finding a micromagnetic
 103 solution, i.e., the minimisation of the Gibbs free-energy functional, which is
 104 faster though prone to false minima, and the vanishing of the torque, which is
 105 slower but more robust and physically meaningful (Gilbert, 2004). The former is
 106 usually achieved with a variety of gradient descent methods (Fischbacher et al.,

2017) while the latter by numerical integration of the Landau-Lifshitz-Gilbert (LLG) equation (Gilbert, 2004)

109

$$\frac{\partial \mathbf{m}}{\partial t} = -\gamma' \mathbf{m} \times \mathbf{H}_{\text{eff}} - \alpha \gamma' \mathbf{m} \times (\mathbf{m} \times \mathbf{H}_{\text{eff}}), \quad (9)$$

where α is a phenomenological damping constant, and $\gamma' = \gamma/(1 + \alpha^2)$ with γ the gyromagnetic ratio.

FEMs allow for an unstructured discretisation of the spatial domain which in our case is decomposed into tetrahedral elements. In the treatment of the micromagnetic theory of Brown (1963) there are some linearisations, which means that there should not be large variations in the direction of \mathbf{m} between neighbouring nodes in the FE mesh. To model nonuniform structures it is sufficient that the spatial discretisation in the model be smaller than the exchange length $l_{\text{exch.}} = \sqrt{2A/\mu_0 M_S^2}$ (Rave et al., 1998), which for greigite is $l_{\text{exch.}} \approx 6.6$ nm; a maximum element size of 5 nm has been chosen for all the models. The non-local problem of calculating the stray field is handled via a transformation method (Imhoff et al., 1990) in *DUNLOP* and by a hybrid finite-element/boundary-element formulation (Fredkin and Koehler, 1990) in *MERRILL*.

2.2. Choice of morphologies

Based on scanning electron microscopy (SEM) and transmission electron microscopy (TEM) micrographs of synthetic greigite samples (Chang et al., 2008; Guowei et al., 2014) and of naturally occurring samples (Snowball, 1997; Vasiliev et al., 2008), five octahedral shapes have been chosen with increasing degrees of truncation of their corners: from no truncation at all (octahedron) to a ‘completely’ truncated shape (cuboctahedron) and three truncated octahedral shapes in-between to which we further refer to as minimally truncated octahedron, truncated octahedron (the regular case), and maximally truncated octahedron (Fig. 1). This series of shapes models the influence of the ratio of $\{001\}$ to $\{111\}$ faces, which increases with truncation. Furthermore, spherical shapes which, although unrealistic, serve as ‘control subjects’ that do not exhibit configurational anisotropy (Williams et al., 2006) and cubic shapes which

136 have been modelled before by Muxworthy et al. (2013) and represent the case
 137 of only $\{001\}$ faces. All the volumes are normalised to cubes, i.e., a particle
 138 sized 120 nm has a volume of $(120 \text{ nm})^3$.

139 2.3. Crystal growth model

140 The magnetic domain structure of a ferromagnetic nanoparticle obtained by
 141 a micromagnetic algorithm is not only a function of its mineralogy, size and
 142 shape, but also of its history. In particular, it is known that a SD particle can
 143 grow and remain SD up to a threshold size d_{max} after which it will become PSD
 144 (Enkin and Williams, 1994). (In this study we are concerned with the zero-field
 145 microstructure and properties; in this context, the onset of PSD behaviour
 146 is marked by the formation of a single-vortex structure). If the particle then
 147 decreases its volume it will remain PSD down to a threshold d_{min} ($< d_{\text{max}}$)
 148 below which it will become SD again. This defines a size range in which a
 149 ferromagnetic grain can be either SD or PSD dependent on its history (Mux-
 150 worthy and Williams, 2006). This phenomenon has been modelled for the seven
 151 morphologies (Fig. 1) in the 30 nm to 300 nm size range.

152 Starting from a 30 nm particle, we obtain the micromagnetic solution and
 153 extrapolate it to a larger grain. This becomes the new initial condition for
 154 which we solve and repeat, growing the particle in steps of 2 nm. Since a very
 155 fine incremental size step is used, much smaller than the exchange length, we
 156 can be certain that we are not missing any features from one step to the next.
 157 This process accurately models grain growth.

158 Once the particles have grown to 120 nm, the procedure is then reversed in
 159 decreasing steps of 2 nm until the initial size of 30 nm is reached. Since chemical
 160 alteration usually proceeds by alteration of the surfaces, the volume decreasing
 161 process can be thought of as a model for chemical alteration to a non-magnetic
 162 phase. This growth from 30 nm to 120 nm followed by the reverse process is
 163 referred to as the *main loop* (ML).

164 Solutions on the size-descending curve not found on the size-increasing curve
 165 were also subject to growth; the *secondary loop* (SL). The ML and SL allow us

to investigate the different domain states with size, shape and history. These micromagnetic solutions were performed by numerical integration of the LLG equation (Eqn. 9). Stable solutions at 120 nm, whether found on the ML or SL, were further grown up to 300 nm. Energy minimisation was used for calculations larger than 120 nm, as integration of the LLG equation is prohibitively slow at these sizes. The parallel *DUNLOP* package (Nagy, 2016) was used to model these grain size dependencies.

These models overlook the effect of thermal fluctuations. However, this limitation is addressed by calculation of the relaxation times. These allow the determination of the particles domain state at a given size beyond the capabilities of standard micromagnetics.

2.4. Relaxation times

Over a sufficiently long observation time, termed the *relaxation time*, a ferromagnetic particle will switch between different stable states due to thermal activation. The relaxation time is given by (Néel, 1955)

$$\tau = \tau_0 \exp(E_B/K_B T), \quad (10)$$

where K_B is Boltzmann's constant, T the temperature at which the transition occurs and τ_0 the *attempt time*, commonly with a value of $\sim 10^{-9}$ s (McNab et al., 1968). Any transition between stable states must occur along an AMP (Fabian and Shcherbakov, 2017); E_B is the energy barrier along such a path. When E_B is of the order of the thermal energy available on a timescale of interest, a particle is in a superparamagnetic (SP) state, spontaneously switching back and forth between different SD orientations. SD grains with an energy barrier larger than $\sim 60 K_B T$ have a relaxation time in the order of billions of years and are thus considered stable SD (SSD) reliable palaeomagnetic recorders (Dunlop and Özdemir, 1997). The SP-SSD threshold is therefore one of the key parameters to assess the palaeomagnetic significance of a NRM. A NEB method (Fabian and Shcherbakov, 2017) has been implemented in the micromagnetic package *MERRILL* (Nagy et al., 2017) (such methods are currently unavailable

195 in *DUNLOP*) to calculate the energy barriers between the stable configurations
 196 for the (truncated) octahedral particles, from 30 nm increasing in steps of 2 nm
 197 until the blocking volume is reached, which is here taken to be the volume for
 198 which the relaxation time surpasses four billion years.

199 Unlike early NEB methods (as applied to magnetic systems) which approxi-
 200 mated each particle as a single dipole (Berkov, 1998), here we are concerned with
 201 full micromagnetic solutions and thus, our results extend beyond SD configu-
 202 rations and coherent rotations. A similar method was implemented by Dittrich
 203 et al. (2002) to study the energy barriers in patterned granular media.

204 3. Results and Discussion

205 3.1. Magnetic structure in the SD-PSD regime

206 Starting at 30 nm from a randomised initial condition, all shapes relax to a
 207 SD state along an easy direction. Extrapolating onto larger grains, this state
 208 remains stable up to a threshold d_{\max} (Figs. 2a, c). When the SD state solutions
 209 are grown beyond d_{\max} , they are found to relax to an easy aligned vortex (EAV)
 210 (Figs. 2a, d). For all shapes, this vortex configuration is stable up to 120 nm.
 211 On reversal, the EAV is stable down to a threshold d_{EH} (Fig. 2a). Below d_{EH} ,
 212 the EAV goes to a hard-aligned vortex (HAV) preserving its chirality (Figs. 2a,
 213 e). Further decreasing the volume, the HAV is stable down to d_{\min} and below
 214 that, the HAV relaxes back to SD (Fig. 2a). This general behaviour on the
 215 ML is referred to as Type 1 (Figs. 2a, 3a, e, g, k), for which we introduce the
 216 notation:

$$217 \quad T_1 = \underbrace{\left(\text{SD}_{30}^{d_{\max}} \rightarrow \text{EAV}_{d_{\max}}^{120} \right)}_{\text{growth part}} \rightarrow \underbrace{\left(\text{EAV}_{120}^{d_{\text{EH}}} \rightarrow \text{HAV}_{d_{\text{EH}}}^{d_{\min}} \rightarrow \text{SD}_{d_{\min}}^{30} \right)}_{\text{volume-decreasing part}} \leftarrow, \quad (11)$$

218 with $30 \text{ nm} < d_{\min} < d_{\text{EH}} < d_{\max} < 120 \text{ nm}$.

219 Another sequence was observed. On the size-descending curve, the EAV
 220 is stable down to a threshold d_{EI} below which the EAV goes to a $\langle 011 \rangle$
 221 intermediate-aligned vortex (IAV) (Figs. 2f, g), stable down to d_{IH} . Below

that, the IAV goes to a HAV which remains down to d_{\min} . Finally, below d_{\min} the HAV relaxes back to SD. This ML behaviour is referred to as Type 2 (Figs. 3c, i), with the formula:

$$T_2 = \underbrace{\left(SD_{30}^{d_{\max}} \rightarrow EAV_{d_{\max}}^{120} \right)_{\rightarrow}}_{\text{growth part}} + \underbrace{\left(EAV_{120}^{d_{EI}} \rightarrow IAV_{d_{EI}}^{d_{IH}} \rightarrow HAV_{d_{IH}}^{d_{\min}} \rightarrow SD_{d_{\min}}^{30} \right)_{\leftarrow}}_{\text{volume-decreasing part}}, \quad (12)$$

with $30 \text{ nm} < d_{\min} < d_{IH} < d_{EI} < d_{\max} < 120 \text{ nm}$.

Growth of the vortex configurations found on the size-descending curve (HAVs, IAVs) forms the SL. When the ML is Type 1, the HAV is grown up to a threshold d_{HE} beyond which it realigns with an easy direction (Fig. 2a). When the ML is Type 2, the HAV goes to either an EAV (Fig. 3c) or to the IAV it nucleated from (Fig. 3i). When the IAV is grown beyond a threshold d_{IE} it realigns with an easy direction (Fig. 3c). This behaviour on the SL is referred to as Type A (Fig. 2a, 3c, e, g), with the general formula (in the case of growth of the HAV and realignment to an EAV):

$$T_A = \underbrace{\left(HAV_{d_{EH}}^{d_{HE}} \rightarrow EAV_{d_{HE}}^{120} \right)_{\rightarrow}}_{\text{secondary growth part}}. \quad (13)$$

There is also the possibility that the HAV/IAV is stable up to 120 nm. This SL is a Type B (Fig. 3a, i, k), with the general formula (in the case of growth of the HAV):

$$T_B = \underbrace{\left(HAV_{d_{EH}}^{120} \right)_{\rightarrow}}_{\text{secondary growth part}}. \quad (14)$$

3.1.1. Spheres

Fig. 2a shows the reduced magnetisation and Fig. 2b the energy density (against size) for the spherical shapes. Spheres showed a Type 1 ML (Eq. 11), with a specific formula:

$$ML_{\text{sph.}} = \left(SD_{30}^{92} \rightarrow EAV_{94}^{120} \right)_{\rightarrow} + \left(EAV_{120}^{54} \rightarrow HAV_{52}^{40} \rightarrow SD_{38}^{30} \right)_{\leftarrow}. \quad (15)$$

245 The SD state at $d_{\max} = 92$ nm and the EAV at 94 nm are shown in Figs. 2c–d.
 246 The SL is formed by growing the HAV found at 52 nm (Fig. 2e). This is a Type
 247 A SL (Eq. 13), with formula:

$$248 \quad \text{SL}_{\text{sph.}} = (\text{HAV}_{52}^{68} \rightarrow \text{EAV}_{70}^{120})_{\rightarrow}. \quad (16)$$

249 3.1.2. Octahedra and truncated octahedra

250 Fig. 3 shows the reduced magnetisation and energy density plots for the
 251 rest of the shapes. The octahedra (Figs. 3a–b) showed a Type 1 ML (Eq. 11),
 252 specifically:

$$253 \quad \text{ML}_{\text{oct.}} = (\text{SD}_{30}^{68} \rightarrow \text{EAV}_{70}^{120})_{\rightarrow} + (\text{EAV}_{120}^{52} \rightarrow \text{HAV}_{50}^{46} \rightarrow \text{SD}_{44}^{30})_{\leftarrow}. \quad (17)$$

254 The SD state at $d_{\max} = 68$ nm (Fig. 4a) shows significant flowering—deflection
 255 of the magnetisation onto edges and vertices. The EAV nucleated at 70 nm is
 256 shown in Fig. 4b. Growth of the HAV nucleated at 50 nm (Fig. 4c) showed a
 257 Type B SL (Eq. 14), *i.e.*, the HAV was found to be stable up to 120 nm,

$$258 \quad \text{SL}_{\text{oct.}} = (\text{HAV}_{50}^{120})_{\rightarrow}. \quad (18)$$

259 The minimally truncated octahedra (Figs. 3c–d) showed a Type 2 ML (Eq.
 260 12):

$$261 \quad \text{ML}_{\text{t.oct.}}^{\text{min.}} = (\text{SD}_{30}^{74} \rightarrow \text{EAV}_{76}^{120})_{\rightarrow} + (\text{EAV}_{120}^{50} \rightarrow \text{IAV}_{48}^{48} \rightarrow \text{HAV}_{46}^{44} \rightarrow \text{SD}_{42}^{30})_{\leftarrow}. \quad (19)$$

262 The SD state shows greater stability and less flowering at $d_{\max} = 74$ nm (Fig.
 263 4d) than for the octahedra, before relaxing to an EAV at 76 nm (Fig. 4e).
 264 Growth of both the IAV from 48 nm (Fig. 2f) and HAV from 46 nm (Fig. 4f)
 265 forms the composite SL. Both showed a Type A SL (Eq. 13), specifically:

$$266 \quad \text{SL}_{\text{t.oct.}}^{\text{min.}} = (\text{IAV}_{48}^{68} \rightarrow \text{EAV}_{70}^{120})_{\rightarrow} + (\text{HAV}_{56}^{70} \rightarrow \text{EAV}_{72}^{120})_{\rightarrow}. \quad (20)$$

267 The next two degrees of truncation, the regular truncated octahedra (Figs.
 268 3e–f) and the maximally truncated octahedra (Figs. 3g–h), both showed Type
 269 1 MLs (Eq. 11):

$$270 \quad \text{ML}_{\text{t.oct.}}^{\text{reg.}} = (\text{SD}_{30}^{78} \rightarrow \text{EAV}_{80}^{120})_{\rightarrow} + (\text{EAV}_{120}^{50} \rightarrow \text{HAV}_{48}^{42} \rightarrow \text{SD}_{40}^{30})_{\leftarrow}; \quad (21)$$

271 and

$$272 \quad \text{ML}_{\text{t.oct.}}^{\text{max.}} = (\text{SD}_{30}^{80} \rightarrow \text{EAV}_{82}^{120})_{\rightarrow} + (\text{EAV}_{120}^{48} \rightarrow \text{HAV}_{46}^{42} \rightarrow \text{SD}_{40}^{30})_{\leftarrow}. \quad (22)$$

273 The SD states show increasing stability and less flowering (Figs. 4g, 5a). The
 274 EAVs at 80 nm and 82 nm are shown in Figs. 4h, 5b. Growth of the HAVs
 275 nucleated at 48 nm (Fig. 4i) and at 46 nm (Fig. 5c) showed the SL was Type A
 276 (Eq. 13) for both shapes:

$$277 \quad \text{SL}_{\text{t.oct.}}^{\text{reg.}} = (\text{HAV}_{48}^{68} \rightarrow \text{EAV}_{70}^{120})_{\rightarrow}; \quad (23)$$

278 and

$$279 \quad \text{SL}_{\text{t.oct.}}^{\text{max.}} = (\text{HAV}_{46}^{70} \rightarrow \text{EAV}_{72}^{120})_{\rightarrow}. \quad (24)$$

280 3.1.3. Cuboctahedra

281 The cuboctahedra (Figs. 3i–j) showed a Type 2 ML (Eq. 12). The PSD
 282 state nucleated on the size-descending curve from the EAV is a distorted IAV
 283 (dIAV)—a sort of mixed state with the vortex mostly aligned with the $[0\bar{1}1]$
 284 direction and the ends of the vortex deflecting to a hard direction. The ML
 285 then has a formula:

$$286 \quad \text{ML}_{\text{oct.}}^{\text{cub.}} = (\text{SD}_{30}^{114} \rightarrow \text{EAV}_{116}^{120})_{\rightarrow} + (\text{EAV}_{120}^{86} \rightarrow \text{dIAV}_{84}^{62} \rightarrow \text{HAV}_{60}^{38} \rightarrow \text{SD}_{36}^{30})_{\leftarrow}. \quad (25)$$

287 The SD state is more stable, with d_{max} increasing to $d_{\text{max}} = 114$ nm (Fig. 5d)
 288 before relaxing to an EAV at 116 nm (Fig. 5e). Growth of the HAV from 60 nm
 289 (Fig. 5f) shows a Type A SL (Eq. 13), with the HAV going to the dIAV it

nucleated from. The dIAV (Fig. 2f) instead shows a Type B SL. The SL is then:

$$\text{SL}_{\text{oct.}}^{\text{cub}} = (\text{HAV}_{60}^{72} \rightarrow \text{dIAV}_{74}^{120})_{\rightarrow} + (\text{dIAV}_{84}^{120})_{\rightarrow}. \quad (26)$$

3.1.4. Cubes

The cubes (Figs. 3k–l) showed a Type 1 ML (Eq. 11). The EAV nucleated from the SD state is a distorted EAV (dEAV). The dEAV mostly aligns with an easy direction, but its ends deflect from the vertices to a hard direction, attaching to opposite faces. The ML is then:

$$\text{ML}_{\text{cube}} = (\text{SD}_{30}^{80} \rightarrow \text{dEAV}_{82}^{120})_{\rightarrow} + (\text{dEAV}_{120}^{64} \rightarrow \text{HAV}_{62}^{38} \rightarrow \text{SD}_{36}^{30})_{\leftarrow}. \quad (27)$$

The SD state at $d_{\text{max}} = 80$ nm is largely flowered (Fig. 5g). The dEAV initially nucleated at 82 nm (Fig. 5h) could also be interpreted as a distorted HAV. However, the alignment of such state with an easy direction becomes more obvious for larger cubes. Growth of the HAV from 62 nm (Fig. 5i) results in a Type B SL:

$$\text{SL}_{\text{cube}} = (\text{HAV}_{62}^{120})_{\rightarrow}. \quad (28)$$

3.1.5. Discussion

The values for d_{max} of the spheres are significantly larger than those found for the euhedral shapes (except for the cuboctahedra which have an anomalously large value), perhaps due to the absence of corners to act as nucleation points. Truncation of the octahedral particles increases the (numerical) stability of the SD solutions which is expressed as the increase in d_{max} from 68 nm for the octahedra (Fig. 3a) to 114 nm for the cuboctahedra (Fig. 3i) (Table 1). A less pronounced effect is the decrease of d_{min} with truncation from 46 nm for the octahedra to 38 nm for the cuboctahedra (Table 1). HAVs are more (numerically) stable at smaller sizes the more truncated the particle is. This is due to the large stray field energy created by a vortex pointing towards a vertex. It

is energetically favourable for a vortex to attach its ends to flat surfaces large enough to accomodate its core. This is seen in the distortion of the vortex structures shown in Figs. 2g, 5h. These avoid the production of large stray fields by attaching their ends to grain faces, at the cost of anisotropy and exchange energies needed to distort the otherwise straight structures of the vortices.

	main loop	d_{\max}	d_{EH}	d_{EI}	d_{IH}	d_{\min}	d_0
Spheres	T_1	92	54	N/A	N/A	40	51
Octahedra	T_1	68	52	N/A	N/A	46	56
Min. t. octahedra	T_2	74	N/A	50	48	44	54
T. octahedra	T_1	78	50	N/A	N/A	42	53
Max. t. octahedra	T_1	80	48	N/A	N/A	42	52
Cuboctahedra	T_2	114	N/A	86	62	38	50
Cubes	T_1	80	64	N/A	N/A	38	52

Table 1: Critical sizes for all shapes. All sizes in nm. d_0 , d_{\min} decrease with truncation while d_{\max} increases.

The energy plots (Figs. 2b, 3b, d, f, h, j, l) show the SD energy density is fairly constant with size for all shapes. For the octahedra, the intersection of the EAV and HAV energy curves occurs above the SD curve. This means that it is then the EAV energy curve which first intersects the SD curve at d_0 and thus this PSD state becomes the GEM thereon (Fig. 3b). With truncation, this intersection moves closer to the SD curve as can be seen in Fig. 3d in which all the different PSD states (EAV, IAV and HAV) and the SD energy curves meet at roughly the same point. Further truncation causes this intersection to eventually occur below the SD energy curve. This creates a narrow range of sizes for which the HAV is the lowest energy state (Figs. 3f, h). Completely truncated, the cuboctahedra show a *split* of this intersection into distinct crossings of the HAV/IAV and IAV/EAV energy curves (Fig. 3i, compare with Fig. 3d), creating a broad range of sizes, from ~ 50 nm to ~ 66 nm, for which the HAV has the lowest energy. A range for which the HAV has the lowest energy was also found for

spheres and cubes. The overall effect of truncation on d_0 is to decrease this threshold (Table 1).

For the cubes we found $d_{\max} = 80$ nm, much smaller than the value of 107 nm obtained by Muxworthy et al. (2013). We found $d_{\min} = 38$ nm, in agreement with the value by Muxworthy et al. (2013). We found the intersection of the SD and HAV energy curves $d_0 \approx 52$ nm (Fig. 3l), lower than the value of 58 nm by Muxworthy et al. (2013). Modelling the ML for cubes with the M_S value used by Muxworthy et al. (2013) we obtained $d_{\max} = 92$ nm, $d_{\min} = 42$ nm smaller and larger, respectively, than the values by Muxworthy et al. (2013). The differences in d_{\max} , d_{\min} can be due to Muxworthy et al. (2013) using a FD method as opposed to a FEM used here. However, excellent agreement of $d_0 \approx 60$ nm was found with the value of 58 nm by Muxworthy et al. (2013). The difference between d_0 obtained with the different M_S is significant as it is larger than the exchange length and thus, unlikely to be an effect of discretisation.

3.2. Identifying the PSD–MD transition

Unlike fine SD grains, bulk ferromagnetic materials can possess a (roughly) null net magnetisation. This is because bulk ferromagnets, in their lowest energy state, have a multi-domain (MD) magnetic structure (Dunlop and Özdemir, 1997). MD structure is characterised by the coexistence of multiple magnetic domains: small regions saturated in different easy directions and separated by narrow planar regions called domain walls where the magnetisation vector transforms continuously from the direction of one domain to that of its neighbour. Near the surface of a bulk sample the magnetic domains lie in directions parallel to the surfaces, which do not necessarily coincide with the magnetocrystalline easy axes; these are called *closure domains* as they enclose the magnetic flux. Then, MD grains minimise the stray field energy at the expense of exchange and MCA energies associated with the domain walls and closure domains. This balance becomes more energetically favourable as a particle grows. To identify a plausible PSD–MD transition, EAV states found to be stable at 120 nm were further grown up to 300 nm in steps of 2 nm.

3.2.1. Growth of EAVs

The octahedron EAV (Fig. 4b) was found to remain stable up to 300 nm—as was the case for the EAVs for all shapes. However, although the overall structure is preserved as the grain grows, there is a gradual, continuous transformation towards MD structure. As the crystal becomes larger, the vortex core (the cylindrical region encompassing the highest helicity $K = \mathbf{m} \cdot \nabla \times \mathbf{m}$) width remains the same. The regions radially far from the core become screened from its influence and more influenced by the effects of MCA, giving way to six ever larger expanses of the grain that become magnetised along easy axes. The sections with higher MCA energy become increasingly narrower, more planar and confined between the easy-aligned regions. With further growth, the regions close to the edges become magnetised along the edges thus completing a picture of MD structure: magnetic domains magnetised along easy directions (Fig. 6a); flat, narrow magnetic domain walls dividing these and closure domains formed along edges of the body (Fig. 6b). The spheres (not shown) show this same general behaviour of formation of magnetic domains and domain walls, but without the formation of closure domains as there are no edges to nucleate these type of domains.

At 300 nm, the regular truncated octahedron shows the same basic structure as the 300 nm octahedron except for the widening of the domain walls as they approach the square $\{001\}$ faces (Figs. 6c–d). Completely truncated, the cuboctahedra have a somewhat different structure. In the 300 nm cuboctahedron solution the domain walls have become so wide as they draw closer to the square faces that they engulf three of the magnetic domains (Fig. 6e), and hard-aligned Néel walls appear which are visible as blue streaks on the square faces (Fig. 6f).

3.2.2. Discussion

A tentative mechanism for a PSD–MD transition has been identified. This proceeds by a gradual formation of magnetic domains, domain walls and closure domains ‘seeded’ by the vortex structures nucleated in the SD–PSD transition

395 regime. However, this transition is more obvious for the octahedra than the
396 truncated octahedra and cuboctahedra.

397 Since our models do not account for the effects of thermal fluctuations,
398 the stability of the solutions is only numerical, e.g., the energy landscape of
399 a micromagnetic solution is somewhat flat in the vicinity of the solution, thus
400 small numerical perturbations can be insufficient for a micromagnetic algorithm
401 to drive the solution to a new local energy/torque minimum depending on the
402 sensitivity or control parameters of the algorithm. We find for the octahedra
403 and the cuboctahedra two solutions for each morphology up to 300 nm (Figs.
404 3a, e). Although the EAVs for both have the lowest energies, this leaves open
405 the question of whether we can expect to find metastable grains with the higher
406 energy structures. Likewise, in Section 3.1 we find that grains remain SD beyond
407 d_0 . In theory, a particle can remain in a metastable SD state beyond this
408 threshold if the energy barrier separating the SD from the PSD state is higher
409 than the thermal energy available. Knowledge of the energy barriers near the
410 SD-PSD transition is needed to answer these questions.

411 3.3. Energy barriers and blocking volumes

412 A nudged elastic-band (NEB) method (Fabian and Shcherbakov, 2017) was
413 implemented for the calculation of the energy barriers at room temperature as
414 a function of size and shape for the (truncated) octahedral morphologies. To
415 calculate the energy barrier for a given shape and size we obtain many solutions
416 from randomised initial conditions. From this set of solutions the state with the
417 lowest energy is identified. For the smaller particles we expect the GEM to be
418 SD and for larger grains PSD. Once the GEM has been identified from the set
419 of initial solutions, two appropriate solutions must be chosen to calculate the
420 action-minimising path (AMP) between them and thus the energy barrier.

421 When the GEM is a $\langle 111 \rangle$ -aligned SD state, the pair of appropriate solu-
422 tions are magnetised normal to contiguous $\{111\}$ faces. Above d_0 the GEM is
423 usually a HAV, then, the paths to test are between vortices at 90° from each
424 other, e.g., between $[100]$ and $[001]$ vortices. For some shapes the GEM is a

425 distorted vortex, for these, it is important to calculate several paths between
 426 different configurations to find the transition with the lowest energy barrier. At
 427 larger sizes, for all shapes the GEM is an EAV, this means that the paths to
 428 calculate are between vortices pointing towards contiguous $\{111\}$ faces, much
 429 like for SD particles. For all PSD to PSD transitions the vortices must have
 430 the same chirality as a change produces prohibitively large energy barriers one
 431 to two orders of magnitude larger than the AMP barrier and thus we can ne-
 432 glect the possibility of such transitions. It is not necessary to calculate the
 433 energy barriers for transitions other than the ones with the lowest energy as
 434 these dominate the behaviour and higher energy transitions usually proceed via
 435 lower energy ones (Nagy et al., 2017). For perfectly regular, equant grains as
 436 those modelled in this study, the lowest energy transition is degenerate and so,
 437 the relaxation time has to be divided by three (for the three distinct degenerate
 438 paths).

439 3.3.1. *Octahedra and truncated octahedra*

440 Fig. 7 shows three examples of transitions for the regular truncated octa-
 441 hedra which were found to be typical for all the (truncated) octahedra, but for
 442 the cuboctahedra. At the smaller sizes (Fig. 7, left column) the transition is
 443 from SD to SD via coherent rotation. The energy along such a path is plotted
 444 in Fig. 7a; the energy barrier is the difference between the easy (Figs. 7d, j)
 445 and intermediate-aligned SD (Fig. 7g) energies. The energy barrier for such
 446 transitions, for all shapes, is quite low, with relaxation times from a few mi-
 447 croseconds to a few days for the largest of these. A few nanometers before the
 448 SD-PSD transition the AMP is not a coherent rotation, but a transition via a
 449 curling mode (Fig. 7, center column). The energy barrier is also the difference
 450 between the easy and intermediate-aligned SD energies.

451 Once the GEM is an EAV, the transitions are between isochiral EAVs di-
 452 rected towards contiguous $\{111\}$ faces (Fig. 7, right column). The transition is
 453 a structured rotation of the vortex, through an IAV (Fig. 7i), which maintains
 454 its overall shape along the path. The energy along such paths is a double-bump

455 where the IAV sits in a shallow local energy minimum (LEM). This means that
 456 the energy barrier is not the difference between the EAV and IAV energies.
 457 Rather, it is given by the barrier separating the EAV from the IAV. However,
 458 with increasing size the IAV LEM becomes more shallow and thus the energy
 459 barrier is better approximated by the difference between the EAV and IAV
 460 energies—at 64 nm the IAV LEM is so shallow that the double-bump feature
 461 becomes flattened (Fig. 7c). In general, the PSD to PSD transitions are very
 462 similar to coherent rotations between SD states in that they are structured ro-
 463 tations of the vortex core. The left column of Video 1 (supplementary material)
 464 shows these typical transitions.

465 3.3.2. *Cuboctahedra*

466 The cuboctahedra show a very different behaviour. Their relaxation times
 467 increase exponentially for the SD to SD transitions, but do not drop for the
 468 first PSD to PSD transitions, at 48 nm, which are hard-aligned to hard-aligned
 469 (Fig. 8, left column) and pass through a dIAV (Fig. 8g). The energy barrier
 470 is an IAV (Fig. 8j). The relaxation times for these transitions then decrease
 471 as the HAV and dIAV energies get closer until, at 66 nm, the GEM is a dIAV.
 472 Then, since the straight IAV has a very high energy, transitions in which the
 473 dIAV ends reattach to contiguous faces are unfavorable. The transition with the
 474 lowest energy barrier is between two distorted vortices like the ones shown in
 475 Figs. 8e, k, which are (though distorted) $[0\bar{1}\bar{1}]$ - and $[1\bar{1}0]$ -aligned. The transition
 476 preserves the shape of the distorted vortex as it structuredly rotates keeping its
 477 ends attached to the square faces.

478 The relaxation times for this type of transitions plateau at a few microsec-
 479 onds up to 74 nm. In Fig. 3j the energies of the dIAV and the EAV intersect
 480 at roughly 90 nm and are very close up to 120 nm. Since the transition between
 481 two EAVs is likely to go through a dIAV, it is not expected that the relaxation
 482 time for EAV to EAV transitions will rapidly increase until sizes much larger
 483 than 90 nm. Indeed, we find that only for the particles larger than 110 nm the
 484 relaxation times start to grow exponentially—though at a rate slower than for

485 (truncated) octahedra transitions—surpassing 4 billion years at roughly 134 nm
 486 (Fig. 9c). A typical transition of this type at 132 nm is shown on the right
 487 column of Fig. 8 between $[111]$ - and $[11\bar{1}]$ -aligned vortices via vortex distortion
 488 (not shown). The highest energy state along the path is a straight IAV (Fig. 8i).
 489 The right column of Video 1 (supplementary material) shows these transitions.

490 3.3.3. Discussion

491 Fig. 9 summarizes the results of the relaxation times obtained for the five
 492 (truncated) octahedral shapes in the 30 nm to 74 nm range and the relaxation
 493 times for the cuboctahedra in the 120 nm to 134 nm range (Fig. 9, inset). For
 494 all shapes, except the cuboctahedral, the general behaviour is an exponential
 495 increase of the relaxation times for the SD to SD transitions from a few tens of
 496 microseconds to a few seconds for the regular truncated octahedra and up to
 497 12 days for the octahedra, followed by a sharp drop at 46 nm (for the regular
 498 truncated octahedra) to 50 nm (for the octahedra). This drop in the relaxation
 499 times marks the SD–PSD threshold. Then, the PSD to PSD transitions have
 500 increasingly lower relaxation times until reaching a global minimum at ~50 nm to
 501 ~54 nm. Once the GEM is the EAV the relaxation time shoots up exponentially
 502 with crystal growth, surpassing 4 billion years at ~68 nm to ~72 nm.

503 These calculations show that equant, non-interacting, room-temperature SD
 504 grains of greigite are on a geological timescale, SP given that the largest relax-
 505 ation time found for SD particles (octahedra) of 12 days is only stable on lab-
 506 oratory timescales. Therefore, the expected SP–SSD threshold does not exist
 507 but rather a SP–PSD threshold since it is only the PSD grains which can hold a
 508 remanence for extended periods of time. Muxworthy et al. (2013) estimated the
 509 energy barriers for non-interacting cubes of greigite from the anisotropy energy
 510 surface and found a SP–SSD threshold of 56 nm, for a relaxation time of four
 511 billion years. In this study the blocking volumes for cubes and spheres were not
 512 calculated as these are not naturally occurring morphologies for greigite. How-
 513 ever, given the results for the five (truncated) octahedral shapes, it is unlikely
 514 that a SP–SSD threshold should exist for either cubes or spheres.

515 Since the relaxation times for the larger SD particles are so small and since
 516 the transitions from SD to SD just before the GEM becomes PSD are through
 517 a curling mode (vortex nucleation), we conclude that the SD–PSD threshold is
 518 defined only by the first intersection of a PSD energy curve with the SD energy
 519 curve, d_0 (Figs. 2b, 3b, d, f, h, j, l). That SD–SD transitions just below the
 520 SD–PSD threshold occur via a curling mode was also observed for magnetite
 521 by Enkin and Williams (1994) and shows that greigite particles are likely to
 522 show PSD aspects even below the SD–PSD threshold. It is highly unlikely that
 523 metastable SD grains exist beyond d_0 in the absence of grain–grain interactions.
 524 The relaxation times between PSD states just above d_0 are the shortest of all
 525 and only start to grow exponentially for the EAV to EAV transitions. We thus
 526 conclude that the PSD–MD transition through a crystal growth mechanism,
 527 must occur through an EAV path.

528 The splitting of the energy curves’ intersections exhibited by the cuboc-
 529 tahedra not only has repercussions on the SD–PSD threshold, but also on its
 530 stability against thermal fluctuations. All the (truncated) octahedral particles
 531 have blocking volumes of ~ 72 nm while for the cuboctahedra this volume is
 532 ~ 134 nm. This could be explained by the fact that the first PSD to PSD tran-
 533 sitions, between hard-aligned PSD states, traverse through IAVs (Fig. 8, left
 534 column): as the cuboctahedral particle grows, the energy difference between the
 535 IAV and HAV becomes smaller (Fig. 3j) and this is reflected in a diminishing of
 536 the relaxation times for this type of transition (Fig. 9). Since for the rest of the
 537 shapes the PSD and SD energy curves all meet in a relatively short span, this
 538 effect of diminished stability with increasing size only lasts for a much smaller
 539 range of sizes, and once the EAV is the GEM, the relaxation times for tran-
 540 sitions between this type of domain state grow very rapidly. However, for the
 541 cuboctahedra the EAV becomes the GEM at a larger size than for the rest of
 542 the shapes and the energy difference with the dIAV, through which the AMP
 543 goes through, is very small.

544 4. Conclusion

545 We have presented calculations of the SD-PSD threshold for realistic shapes
546 of equant greigite grains. We have found that octahedral shapes have the largest
547 SD-PSD threshold $d_0 \approx 56$ nm, a value lower than the threshold obtained for
548 magnetite octahedra by Witt et al. (2005) of ~ 73 nm. This value decreases by
549 the effects of truncation to $d_0 \approx 53$ nm for the regular truncated octahedra and
550 further down to $d_0 \approx 50$ nm for the cuboctahedra. The SD-PSD threshold for
551 greigite cubes was lower ($d_0 \approx 52$ nm) than that of Muxworthy et al. (2013)
552 ($=58$ nm); however, we used a different value for M_S . When we modelled using
553 the same parameters as Muxworthy et al. (2013) we found excellent agreement,
554 obtaining a $d_0 \approx 60$ nm.

555 NEB method calculations of the room-temperature blocking volumes of
556 greigite show the importance of thermal fluctuations in determining the mag-
557 netic structure of a ferromagnetic grain beyond simpler micromagnetic models.
558 We found that equant SD greigite grains cannot be expected to be reliable
559 palaeomagnetic recorders: only PSD grains larger than ~ 72 nm are able to hold
560 a remanence in geological timescales, though there is a strong grain shape de-
561 pendence.

562 We found that cuboctahedral particles have a blocking volume of ~ 134 nm, a
563 volume almost seven times larger than for the (truncated) octahedral particles.
564 This highlights the importance of addressing not only the size distribution in
565 sedimentary rock samples, but also the shapes of the greigite grains present.

566 We found that the transitions between PSD states occur via well-defined
567 states by a structured rotation of the vortex cores. This could be useful for
568 future analytical models beyond the SD coherent rotation theory.

569 Although in this study we have disregarded the effects of particle elonga-
570 tions, which are sure to effect the blocking volumes (Muxworthy et al., 2013),
571 our results are representative of widespread authigenic greigite grains of abiotic
572 origin. The effects of grain-grain magnetostatic interactions have also been ex-
573 cluded from this study. Isolated greigite grains are not very common, rather,

574 natural samples show that greigite is more commonly aggregated in tight clus-
575 ters. However, this study provides a stepping-stone towards understanding the
576 more complex greigite clusters.

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