

Quick guide to running python code for computing relic abundance for the Quintuplet

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1. Requirements: **python3**, **scipy**, **numpy**, **mpmath**, **matplotlib**.
2. To compute the abundance for a fixed mass (say 13 TeV) execute the following

```
python boltzmann_eqs_solve.py 13.e3
```

The output is as follows:

mdm	Omega_eff	Omega_eff(gs.const)	Omega_network	Omega_network (with I→J)
time(s)				
13000.0	0.12941000226553		0.12404274053778493	0.11032009304201815
0.11104055336618876		67.53104209899902		

The average execution time is about 60 seconds to compute all the four cases above.

3. To perform a scan over the masses (default $m_\chi \in [2, 15]$ TeV)

```
python mass_scan.py
```

The output table is saved in file **mx_omega_huelthen_network.dat**, in the same column order as above.

4. Plot results with

```
python plot_omega.py
```

5. To extract the yield curve for a fixed mass (say 13 TeV)

```
python yield_compare.py 13.e3
```

0.1 Basic structure of the code:

1. Main file **boltzmann_eqs_solve.py**: imports degrees of freedom from folder **dof**. Thermally averaged cross section tables imported from folder **bsf_sigmas**. Summary in table [1](#)
2. Folder **dof**: degrees of freedom as a function of T, including the derivative term.
3. Folder **bsf_sigmas**: All thermally averaged cross sections are computed here.

Function name	purpose	some details
dydzbseff dydzbseffgsc	Effective dY/dz Effective dY/dz	$t=x=z$, $y_i = Y_\chi$ (const. gs)
dydz9	Eqs. for dY/dz	with DM and 9 bound states $t=x=z$, $y_i[0] = Y_\chi$, $y_i[1]-y_i[9]$ = all bound states dydx0 for DM, and dydxi[0]-dydxi[8] eqs for bound states
dydz9ibs	Eqs. for dY/dz	same as above + inter bound-state transitions
solve_ivp	Solves the above differential eqs.	Method='Radau' (an implicit method) $y0ia$ = initial values, $teval$ = integration range , $z = 3 - 10^6$

Table 1: Summary of the main code boltzmann_eqs.solve.py