

Radiative-convective equilibrium in a grey atmosphere

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Complex systems in climate physics, 3 October 2023

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Introduction

Radiative equilibrium

Radiative-convective equilibrium

Conclusion

- Average vertical temperature profile $T(t, z)$ of atmosphere.

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- Average vertical temperature profile $T(t, z)$ of atmosphere.
- Radiative Transfer Equation (RTE).

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- Average vertical temperature profile $T(t, z)$ of atmosphere.
- Radiative Transfer Equation (RTE).
- Fluid dynamics equations.

- Thermodynamic energy equation in Local Thermodynamic Equilibrium (LTE):

$$\frac{\partial T}{\partial t} = -\frac{1}{\rho c_P} \frac{\partial q}{\partial z} \quad . \quad (1)$$

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- Radiative-convective equilibrium.
- Grey atmosphere.

- Hypotheses on the planet.

Additional hypotheses

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- Hypotheses on the composition of atmosphere.

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- Hypotheses on total heat flux.
- Resulting thermodynamic energy equation:

$$\frac{\partial T}{\partial t} = -\frac{1}{\rho c_P} \frac{\partial}{\partial z} (E_U - E_D) \quad . \quad (2)$$

- Relation between pressure and altitude:

$$P(z) = P_g \exp \left(- \frac{z - z_g}{z_0} \right) . \quad (3)$$

- Relation between optical depth and pressure:

$$\delta(P) = \frac{\mu_m}{g} (P - P_{\text{TOA}}) . \quad (4)$$

- Relation between optical depth and altitude:

$$\delta(z) = \frac{\mu_m}{g} \left(P_g \exp \left(- \frac{z - z_g}{z_0} \right) - P_{\text{TOA}} \right) . \quad (5)$$

Equations in radiative equilibrium

- RTE for non-scattering medium in LTE:

$$\frac{1}{\mu} \frac{\partial L}{\partial z} = B_{\nu} - L \quad . \quad (6)$$

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- Equations for irradiances:

$$-\frac{\partial}{\partial \delta'} E_U(t, \delta') = \sigma T(t, \delta')^4 - E_U(t, \delta') \quad , \quad (7)$$

$$\frac{\partial}{\partial \delta'} E_D(t, \delta') = \sigma T(t, \delta')^4 - E_D(t, \delta') \quad . \quad (8)$$

Initial Value Problem

- Steady state.

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- Initial conditions on irradiances:

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- Additional relations:

$$\frac{d}{d\delta'} (E_U(\delta') - E_D(\delta')) = 0 \quad , \quad (11)$$

$$E_U(\delta') - E_D(\delta') = S_t \quad . \quad (12)$$

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- Initial condition on temperature:

$$T(0) = \left(\frac{S_t}{2\sigma} \right)^{\frac{1}{4}} \quad . \quad (13)$$

- Temperature:

$$T(\delta) = \left(\frac{S_t}{2\sigma} (1 + D\delta) \right)^{\frac{1}{4}} . \quad (14)$$

- Irradiances:

$$E_U(\delta) = \frac{S_t}{2} (2 + D\delta) , \quad (15)$$

$$E_D(\delta) = \frac{S_t}{2} D\delta . \quad (16)$$

- Normalisation:

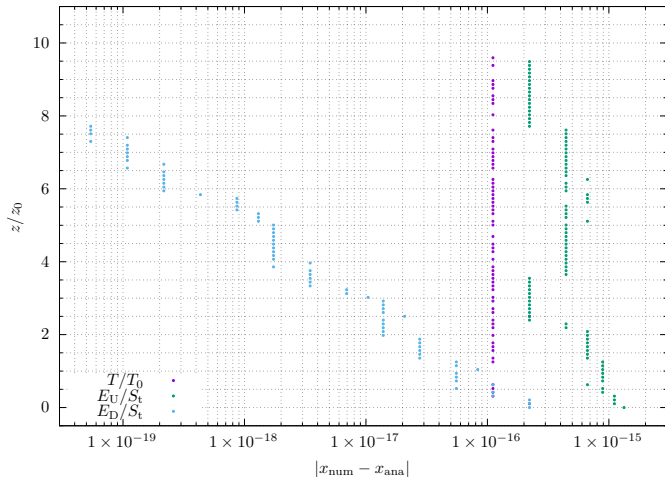
$$Y_0 = \frac{T^4}{T_0^4} \quad , \quad Y_1 = \frac{E_U}{S_t} \quad , \quad Y_2 = \frac{E_U}{S_t} \quad (17)$$

- Resulting system of ODEs:

$$\begin{cases} \frac{dY_0}{d\delta} = \frac{D}{2} \\ \frac{dY_1}{d\delta} = D(Y_1 - Y_0) \\ \frac{dY_2}{d\delta} = D(Y_0 - Y_2) \end{cases} \quad (18)$$

- Runge-Kutta method of order 4.
- Non-uniform step size.

Errors of numerical solutions



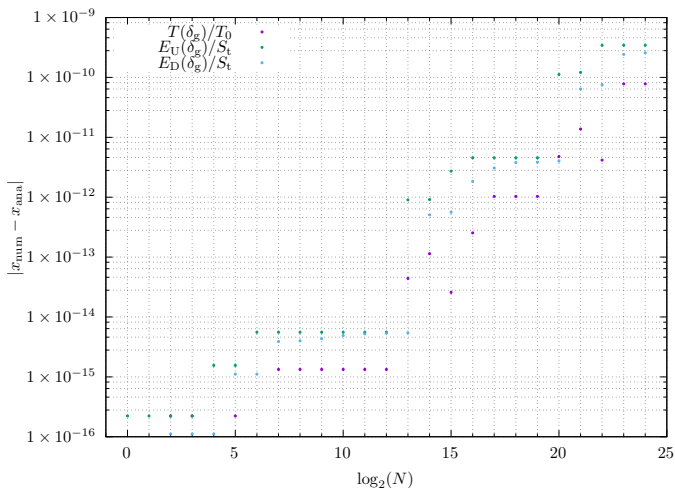
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- Iterative procedure.

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- Additional normalisation:

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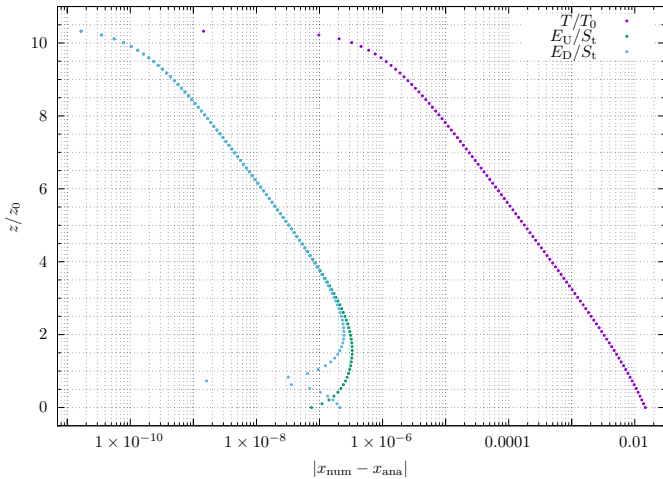
$$Y_3 = \frac{T}{T_0} \quad . \quad (19)$$

- Resulting system of PDEs:

$$\begin{cases} \frac{\partial}{\partial t} Y_3(t, \delta) = \frac{\mu_m S_t}{c_P T_0} \frac{\partial}{\partial \delta} (Y_1(t, \delta) - Y_2(t, \delta)) \\ \frac{\partial}{\partial \delta} Y_1(t, \delta) = D(Y_1(t, \delta) - Y_3(t, \delta)^4) \\ \frac{\partial}{\partial \delta} Y_2(t, \delta) = D(Y_3(t, \delta)^4 - Y_2(t, \delta)) \end{cases} \quad . \quad (20)$$

- Euler method for temporal integration.
- Constant time step.

Errors of time integration at steady state



Radiative-convective equilibrium

- Lapse rate at steady state.

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- Lapse rate at steady state.
- Convection faster than radiative processes.

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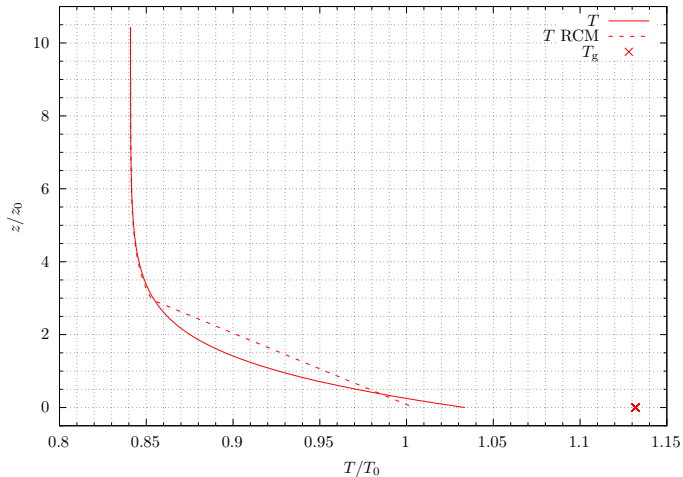
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- Lapse rate at steady state.
- Convection faster than radiative processes.
- Convective adjustment if $-\frac{\partial T}{\partial z} > \Gamma_0$.

Temperature plot



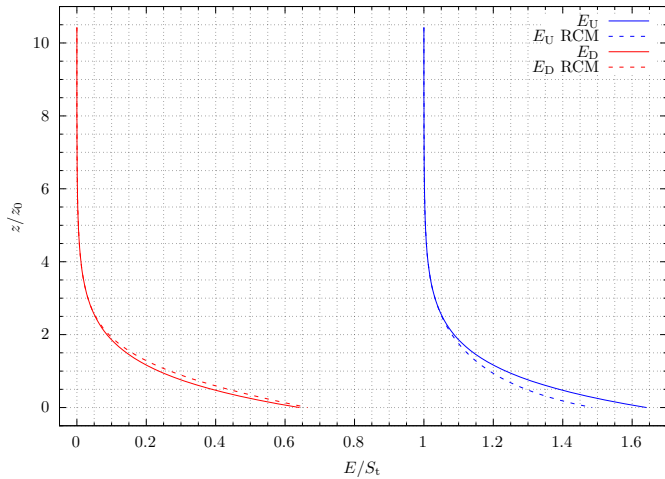
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- Vertical temperature profile of a grey atmosphere described in general by a system of PDEs and at the steady state by a system of ODEs.
- Negligible errors for ODEs numerical solutions.
- Possible improvements on description of physical phenomena.
- Possible improvements on precision of PDEs numerical solutions.