If the energy distribution for fission neutrons from <sup>235</sup>U follows the functional approximation (for energy in MeV)

$$\chi(E) = 0.453e^{-1.036E} \sinh(\sqrt{2.29E}).$$

then the most probable energy of a neutron corresponds to the maximum of the function. A maximum value will be found at a critical point of the function, which can be found via differentiation (specifically when  $\frac{d\chi}{dE} = 0$ ):

$$\begin{split} \frac{d\chi}{dE} &= 0 = \frac{d}{dE} \left( 0.453 e^{-1.036E_{\text{max}}} \sinh(\sqrt{2.29E_{\text{max}}}) \right) \\ 0 &= 0.453 \frac{d}{dE} \left( e^{-1.036E_{\text{max}}} \sinh(\sqrt{2.29E_{\text{max}}}) \right) \\ 0 &= 0.453 \left[ e^{-1.036E} \frac{d}{dE} \sinh(\sqrt{2.29E_{\text{max}}}) + \frac{d}{dE} \left( e^{-1.036E_{\text{max}}} \right) \sinh(\sqrt{2.29E_{\text{max}}}) \right] \\ 0 &= 0.453 \left[ e^{-1.036E} \cosh(\sqrt{2.29E}) \frac{d}{dE} \left( \sqrt{2.29E_{\text{max}}} \right) - 1.036 e^{-1.036E} \sinh(\sqrt{2.29E}) \right] \\ 0 &= 0.453 \left[ e^{-1.036E} \cosh(\sqrt{2.29E}) \frac{\sqrt{2.29}}{2\sqrt{E_{\text{max}}}} - 1.036 e^{-1.036E} \sinh(\sqrt{2.29E}) \right] \\ 0 &= \frac{\sqrt{2.29} \cosh(\sqrt{2.29E})}{2\sqrt{E_{\text{max}}}} - 1.036 \sinh(\sqrt{2.29E}) \\ 0 &= 1 - 1.369 \sqrt{E_{\text{max}}} \tanh(\sqrt{2.29E_{\text{max}}}) \\ 1 &= 1.369 \sqrt{E_{\text{max}}} \tanh(\sqrt{2.29E_{\text{max}}}) \end{split}$$

$$E_{\rm max} = 0.724 \ {\rm MeV}$$

The average energy can be found by finding the expected value of the function on the domain  $[0,\infty)$ .

$$E_{\text{ave}} = \int_0^\infty E \, \chi(E) \, dE$$

$$= \int_0^\infty E \left( 0.453 e^{-1.036E} \sinh(\sqrt{2.29E}) \right) \, dE$$

$$= 0.453 \int_0^\infty E \, e^{-1.036E} \sinh(\sqrt{2.29E}) \, dE$$

This integral cannot be solved analytically. Solving numerically (with Wolram Alpha),

$$E_{\text{ave}} = 1.98 \text{ MeV}$$

A general reaction rate for process x as a function of energy can be defined as

$$R_x(E) = \Sigma_x(E)\phi(E)$$

where  $\Sigma_x(E)$  is the macroscopic cross section for reaction x and  $\phi$  is the neutron flux, both at energy E. The macroscopic cross section can be further decomposed, so that

$$R_x(E) = n_x \sigma_x(E) \phi(E)$$
.

Comparing neutron-neutron reactions with all neutron-nuclei reactions, the ratio of reaction rates is

$$\frac{R_{nn}(E)}{R_{tot}(E)} = \frac{n_n \sigma_{nn}(E)\phi(E)}{n_{\text{UO}_2} \sigma_{tot}(E)\phi(E)}.$$

We can know that the neutron flux is  $\phi(0.025 \text{ eV} = 10^{16} \text{ neutrons/(cm}^2 \cdot \text{s})$  and they are at thermal energies  $(E = 0.025 \text{ eV} \text{ and traveling at } v = \sqrt{\frac{2(0.025 \text{ eV})}{m_n}} = 2.190 \times 10^5 \text{ cm/s})$ . The neutrons that are then in a 1 cm<sup>3</sup> volume at any given second is

$$n_n = \frac{\phi(0.025 \text{ eV})}{v} = \frac{10^{16} \text{ neutrons/(cm}^2 \cdot \text{s})}{2.190 \times 10^5 \text{ cm/s}} = 4.566 \times 10^{10} \text{ neutrons/cm}^3$$

For UO<sub>2</sub>,  $\rho = 10.97 \text{ g/cm}^3$ ,  $m_{\rm O} = 16.0 \text{ g/mol}$ ,  $m_{\rm U8} = 238.05 \text{ g/mol}$ , and  $m_{\rm U5} = 235.04 \text{ g/mol}$ . If we use an enrichment of 5% (atom percent), then  $m_{\rm UO_2} = 0.95(238.05) + 0.05(235.04) + 2(16.0) = 269.90 \text{ g/mol}$ . For number density, we find

$$n_{\rm UO_2} = \frac{\rho N_A}{m_{\rm UO_2}} = \frac{(10.97~{\rm g/cm^3})(6.022\times10^{23})}{269.90~{\rm g/mol}} = 2.448\times10^{22}~{\rm molecules/cm^3}$$

Additionally, we're given that  $\sigma_{nn}=10$  b, and we can determine the microscopic cross section for UO<sub>2</sub> from tabulated data. (From ENDF/B-VII.1 at 0.025 eV:  $\sigma_{tot,U8}=11.962$  b,  $\sigma_{tot,U5}=698.856$  b, and  $\sigma_{tot,O}=3.852$  b)

$$\sigma_{tot} = 0.95\sigma_{tot,U8} + 0.05\sigma_{tot,U5} + 2\sigma_{tot,O}$$
  
$$\sigma_{tot} = 0.95(11.962 \text{ b}) + 0.05(698.856 \text{ b}) + 2(3.852 \text{ b})$$
  
$$\sigma_{tot} = 54.011 \text{ b}$$

We can now sove for the ratio of reaction rates (noting that  $\phi(E)$  cancels in the numerator and denominator)

$$\frac{R_{nn}(E)}{R_{tot}(E)} = \frac{n_n \sigma_{nn}(E)}{n_{\text{UO}_2} \sigma_{tot}(E)} = \frac{(4.566 \times 10^{10} \text{ neutrons/cm}^3)(10 \text{ b})}{(2.448 \times 10^{22} \text{ molecules/cm}^3)(54.011 \text{ b})}$$
$$\frac{R_{nn}(E)}{R_{tot}(E)} = 3.40 \times 10^{-13}$$

The rate of neutron-neutron collisions is 13 orders of magnitudes less than the rate of neutron-UO<sub>2</sub> collisions.

## Problem 4

First, we define the average scattering cosine  $\bar{\mu}_0$  as the average dot product,  $\langle \hat{\Omega} \cdot \hat{\Omega}' \rangle$ . When normalized by  $4\pi \Sigma_s$ , the total of cross sections for scattering from any angle  $\hat{\Omega}$  to any other angle  $\hat{\Omega}'$ , this is

$$\bar{\mu}_0 \equiv \langle \hat{\Omega} \cdot \hat{\Omega}' \rangle = \left(\frac{1}{4\pi \Sigma_s}\right) \int_{4\pi} d\hat{\Omega} \int_{4\pi} d\hat{\Omega}' \, \hat{\Omega} \cdot \hat{\Omega}' \Sigma_s (\hat{\Omega} \cdot \hat{\Omega}')$$

In the center of mass system, the probability that a particle scatters in any direction is roughly uniform,  $\Sigma_{\text{CM}}(\theta_C) = \frac{\Sigma_s}{4\pi}$ ,

$$\bar{\mu}_0 = \frac{1}{\Sigma_s} \int_{4\pi} d\hat{\Omega} \int_{4\pi} d\hat{\Omega}' \, \hat{\Omega} \cdot \hat{\Omega}' \Sigma_{\rm CM}(\theta_C).$$

A critical reactor has a multiplication factor of k=1. The multiplication factor (for an infinite reactor) can be defined as

$$k_{\infty} \equiv \frac{\# \text{ neutrons produced}}{\# \text{ neutrons absorbed}}$$

Mathematically, the number of neutrons produced is  $\int_0^\infty \nu \Sigma_f(E) \phi(E) dE$  and the number of neutrons absorbed is  $\int_0^\infty \Sigma_a(E)\phi(E) dE$ . Altogether, we can mathematically describe a critical reactor as

$$1 = \frac{\int_0^\infty \nu \Sigma_f(E) \phi(E) dE}{\int_0^\infty \Sigma_a(E) \phi(E) dE}$$

or equivalently

$$\int_0^\infty \nu \Sigma_f(E) \phi(E) dE = \int_0^\infty \Sigma_a(E) \phi(E) dE.$$

Since we are considering only thermal cross sections, we will let  $\Sigma_X(E) = \Sigma_X(0.025 \text{ eV}) = \Sigma_{X,T}$  and we find

$$\nu \Sigma_{f,T} \int_0^\infty \phi(E) dE = \Sigma_{a,T} \int_0^\infty \phi(E) dE.$$

The integrals over flux cancel, and so

$$\nu \Sigma_{f,T} = \Sigma_{a,T}.$$

The macroscopic cross sections can be rewritten as  $\Sigma_{f,T} = \Sigma_{f,T,\mathrm{f}}$  and  $\Sigma_{a,T} = \Sigma_{a,T,\mathrm{f}} + \Sigma_{a,T,\mathrm{m}}$  where subscripts f and m denote fuel and moderator, respectively. Furthermore, each macroscopic cross section for each material can be expressed in terms of the material's number density and microscopic cross section,  $\Sigma = n\sigma$ . In total

$$\nu n_{\rm f} \sigma_{f,T,\rm f} = n_{\rm f} \sigma_{a,T,\rm f} + n_{\rm m} \sigma_{a,T,\rm m}.$$

The fuel-to-moderator density at criticality can then be expressed as

$$\frac{n_{\rm f}}{n_{\rm m}} = \frac{\sigma_{a,T,{\rm m}}}{\nu \sigma_{f,T,{\rm f}} - \sigma_{a,T,{\rm f}}}. \label{eq:nf}$$

$$c.$$
) Water

$$\frac{n_{\rm f}}{n_{\rm res}} = 4.55 \times 10^{-6}$$

$$\frac{n_{\rm f}}{n_{\rm res}} = 1.36 \times 10^{-5}$$

$$\frac{n_{\rm f}}{n_{\rm m}} = 8.99 \times 10^{-4}$$

$$\frac{n_{\rm f}}{n_{\rm m}} = 4.55 \times 10^{-6} \qquad \qquad \frac{n_{\rm f}}{n_{\rm m}} = 1.36 \times 10^{-5} \qquad \qquad \frac{n_{\rm f}}{n_{\rm m}} = 8.99 \times 10^{-4} \qquad \qquad \frac{n_{\rm f}}{n_{\rm m}} = 1.64 \times 10^{-6}$$

(see attached Jupyter notebook for full calculations)

a.)

Like in Problem 5,

$$k_{\infty} = \frac{\int_0^{\infty} \nu \Sigma_f(E) \phi(E) dE}{\int_0^{\infty} \Sigma_a(E) \phi(E) dE}.$$

Again, since we are only considering cross sections averaged over the fast spectrum, we can state that  $\Sigma_X(E) = \Sigma_{X,F}$ , and then

$$k_{\infty} = \frac{\nu \Sigma_{f,F} \int_0^{\infty} \phi(E) \, dE}{\Sigma_{a,F} \int_0^{\infty} \phi(E) \, dE} = \frac{\nu \Sigma_{f,F}}{\Sigma_{a,F}}.$$

Next, we repeat our decomposition of the numerator and denominator, leaving

$$k_{\infty} = \frac{\nu_{\mathrm{PuO}_2} n_{\mathrm{PuO}_2} \sigma_{F,\mathrm{PuO}_2} + \nu_{\mathrm{UO}_2} n_{\mathrm{UO}_2} \sigma_{F,\mathrm{UO}_2}}{n_{\mathrm{PuO}_2} \sigma_{a,F,\mathrm{PuO}_2} + n_{\mathrm{UO}_2} \sigma_{a,F,\mathrm{UO}_2} + n_{\mathrm{Fe}} \sigma_{a,F,\mathrm{Fe}} + n_{\mathrm{Na}} \sigma_{a,F,\mathrm{Na}}}$$

After solving this equation with the given values, we find

$$k_{\infty} = 1.332$$

**b.**)

We can use the densities and volume fractions to determine the fraction of the core mass that is fuel.

c.)

If k=1 and the non-leakage probability is  $P_{NL}=0.90$ , then we can reexpress our earlier equation as

$$k = 1 = P_{NL} \frac{\nu \Sigma_{f,F} \int_0^\infty \phi(E) dE}{\Sigma_{a,F} \int_0^\infty \phi(E) dE} = P_{NL} \frac{\nu \Sigma_{f,F}}{\Sigma_{a,F}},$$

or

$$P_{NL}\nu\Sigma_{f,F}=\Sigma_{a,F}.$$

When separated into individual materials, this is

$$P_{NL} \sum_{i} \nu_{i} n_{i} \sigma_{f,i} = \sum_{j} n_{j} \sigma_{a,j}.$$

or explicitly

$$P_{NL}\left(\nu_{\mathrm{PuO}_2}n_{\mathrm{PuO}_2}\sigma_{F,\mathrm{PuO}_2} + \nu_{\mathrm{UO}_2}n_{\mathrm{UO}_2}\sigma_{F,\mathrm{UO}_2}\right) = n_{\mathrm{PuO}_2}\sigma_{a,F,\mathrm{PuO}_2} + n_{\mathrm{UO}_2}\sigma_{a,F,\mathrm{UO}_2} + n_{\mathrm{Fe}}\sigma_{a,F,\mathrm{Fe}} + n_{\mathrm{Na}}\sigma_{a,F,\mathrm{Na}}$$

We recall that  $n_{\text{PuO}_2} = f_{\text{f,PuO}_2}(0.3) \frac{\rho_{\text{PuO}_2} N_A}{m_{\text{PuO}_2}}$  and  $n_{\text{UO}_2} = (1 - f_{\text{f,PuO}_2})(0.3) \frac{\rho_{\text{UO}_2} N_A}{m_{\text{UO}_2}}$ , while  $n_{\text{Na}} = (0.5) \frac{\rho_{\text{Na}} N_A}{m_{\text{Na}}}$  and  $n_{\text{Fe}} = (0.2) \frac{\rho_{\text{Fe}} N_A}{m_{\text{Fe}}}$ . We can substitute these expressions into the equation above and solve for  $f_{\text{f,PuO}_2}$ .

The fuel must be 10.322%  $\mathrm{PuO}_2$  and 89.678%  $\mathrm{UO}_2$  for a critical reactor with  $P_{NL}=0.9$ 

(see attached Jupyter notebook for full calculations)

Answer the following questions as true or false, provide a one sentence justification for your answer

- 1. The integro-differential form of the transport equation expresses a local balance between neutron production and losses.
- 2. A vacuum boundary condition for the integro-differential transport equation implies a zero outgoing angular flux.
- 3. In the transport equation in curvilinear coordinates, the redistribution term allows neutrons to migrate between the directions as they move along a straight line.
- 4. A nuclear system is subcritical if its eigenvalues satisfy  $\max(\text{Re}(j))$ ; 1.
- 5. The energy spectrum of the fundamental eigenmode of the eigenvalue problem is skewed as though a 1 absorber is present.